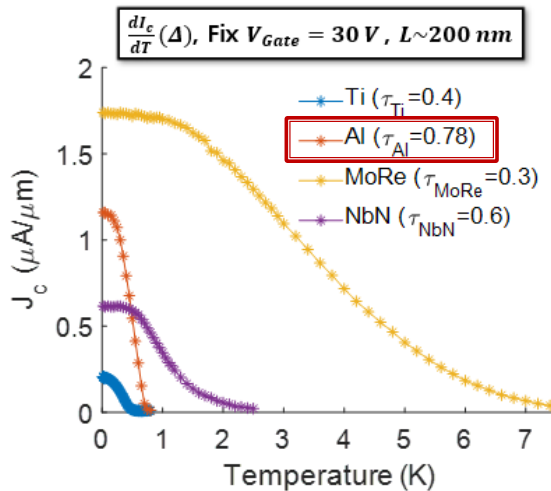
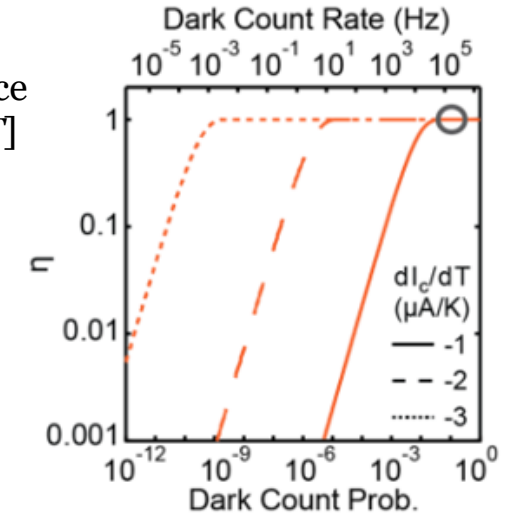


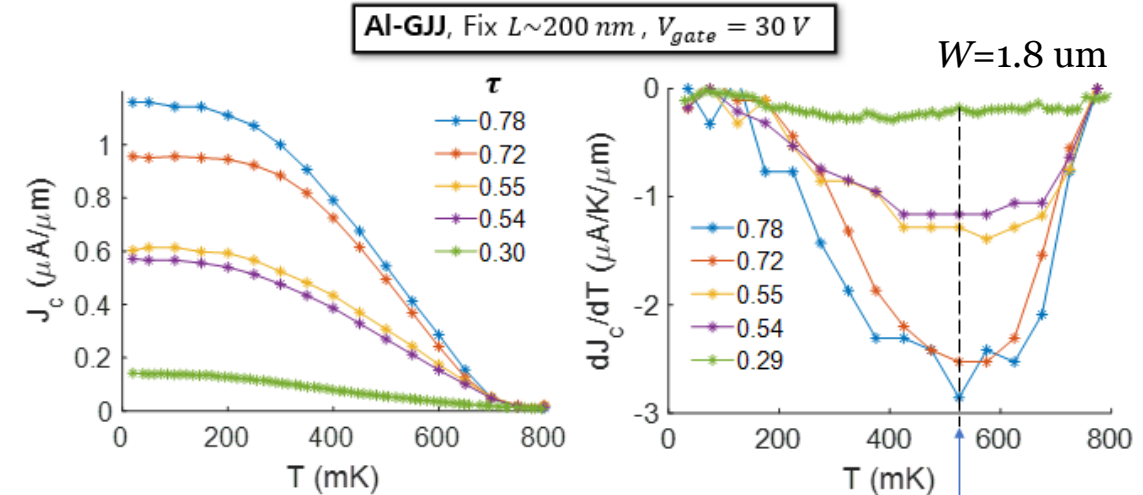
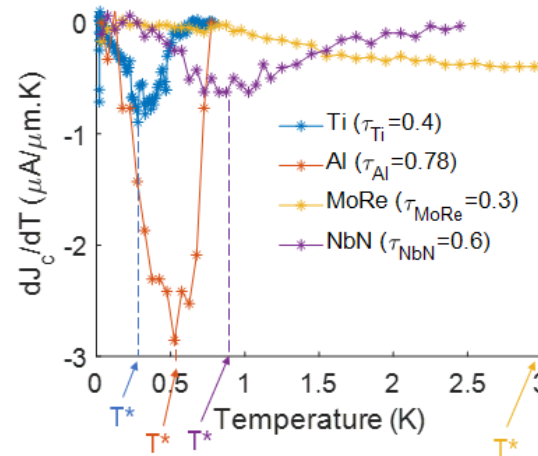
Optimization of $|dI_c/dT|$ for GJJ

- ✓ $|dI_c/dT| \uparrow \rightarrow$ Detection efficiency $\uparrow$ & Dark count rate $\downarrow$
- ✓ Lower T_c of $J_c(=I_c/W)$ & larger J_c at low temperature $\rightarrow$ Aluminum
- ✓ Contact transparency (τ) $\uparrow \rightarrow |dJ_c/dT| \uparrow$
- ✓ For Al & $\tau=0.78$, $|dI_c/dT| \sim 5 \mu\text{A/K}$ (@ $T=500 \text{ mK}$) is realized.

[Detector performance with different dI_c/dT]



[dJ_c/dT of GJJ with various superconductors]

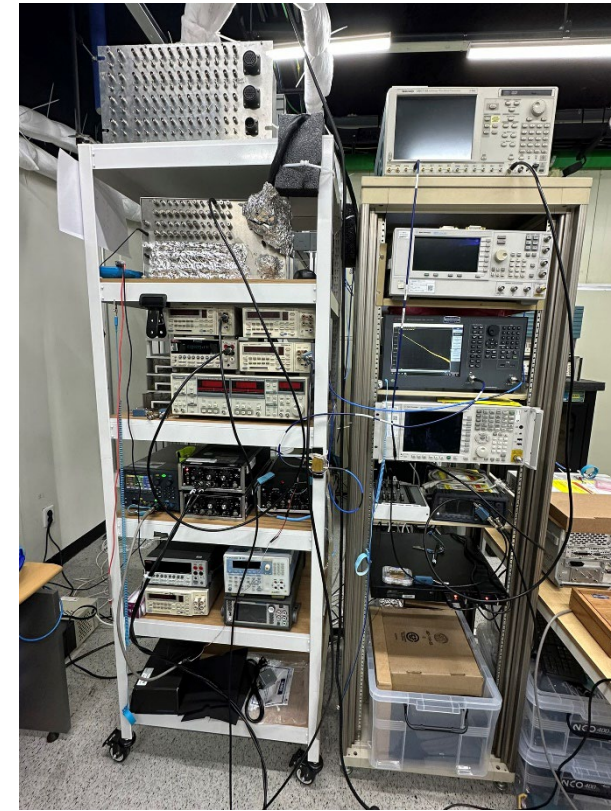


[dJ_c/dT with different transparency]

Overview of Our Measurement Set-up



❖ Dedicated Dilution refrigerator



Arbitrary Waveform Generator

Analog Signal Generator

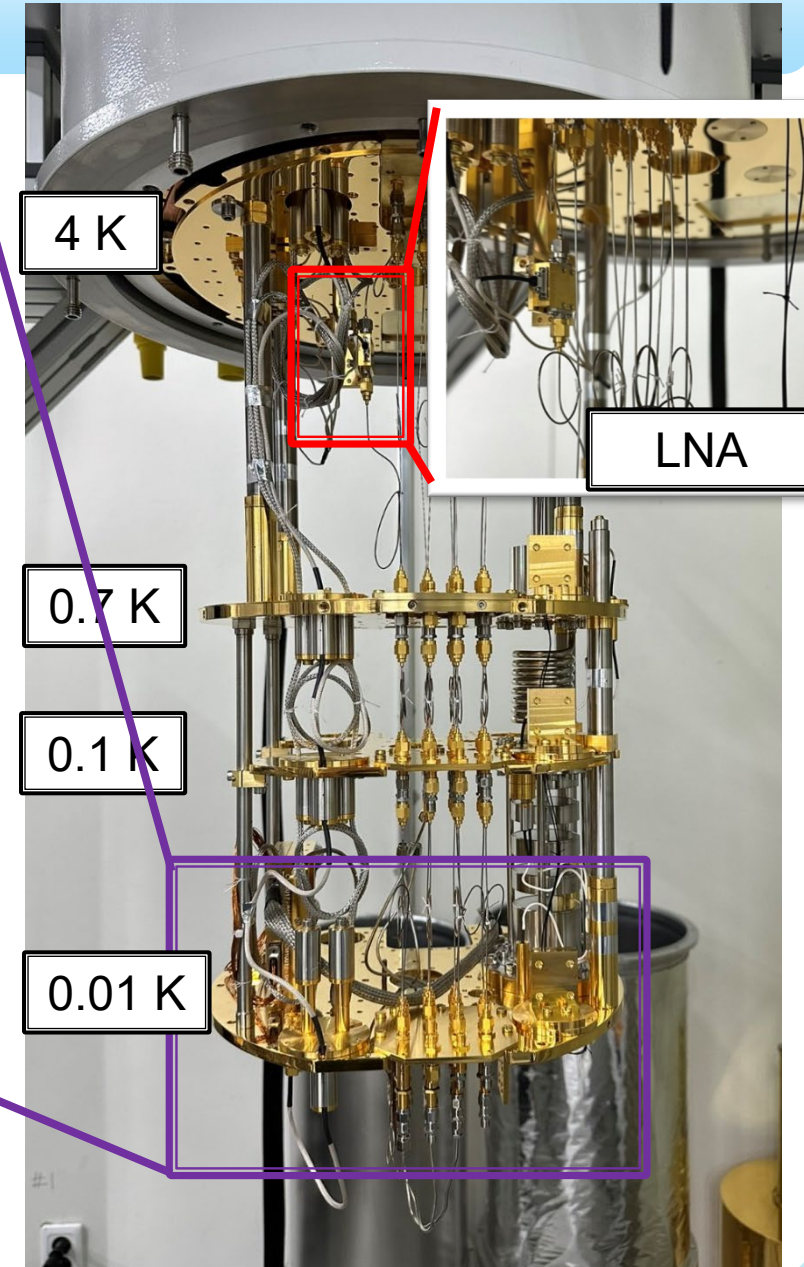
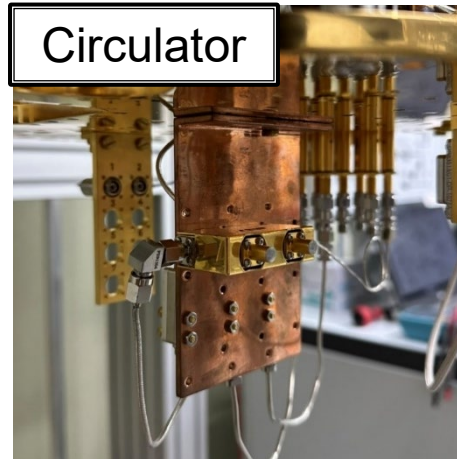
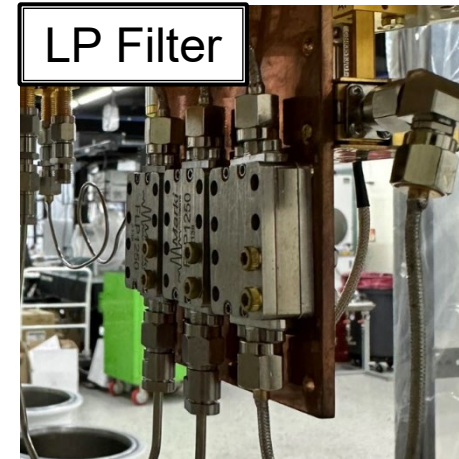
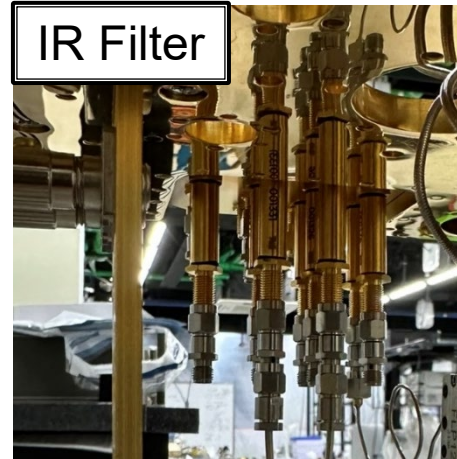
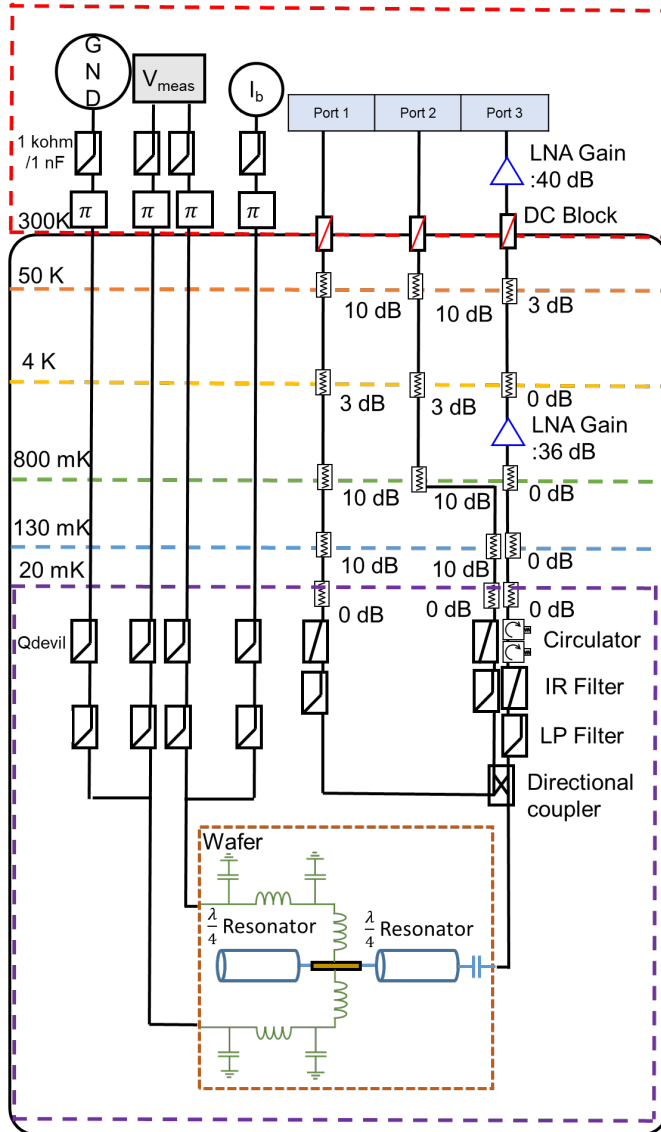
Vector Network Analyzer

Spectrum Analyzer

DC
Instrument

RF
Instrument

Our Measurement Set-up (RF)



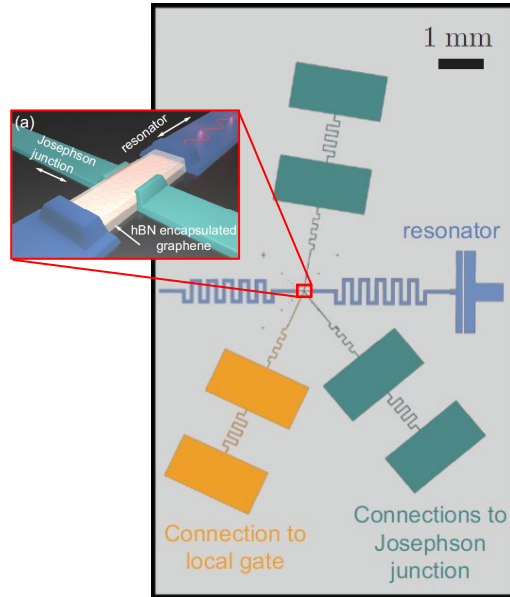
Calibration*

*No $O(\text{THz} \sim \text{meV})$ single-photon sources

➔ How can we check the response of our detector to DM scattering?

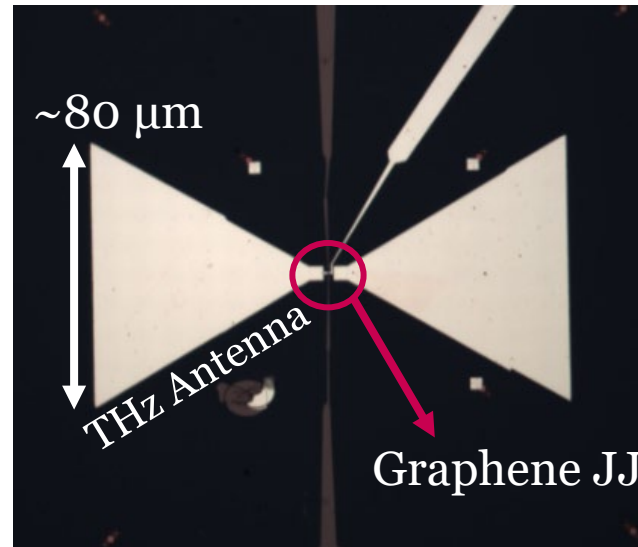
How to Calibrate Detector

Microwave (μW)



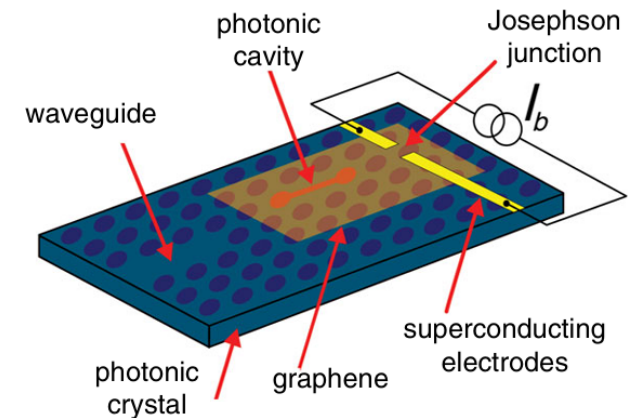
Photon energy $E = 0.1 \text{ meV}$
Wavelength $\lambda = 1 \text{ cm}$

Terahertz (THz)



$E = 10 \text{ meV}$
 $\lambda = 100 \mu\text{m}$

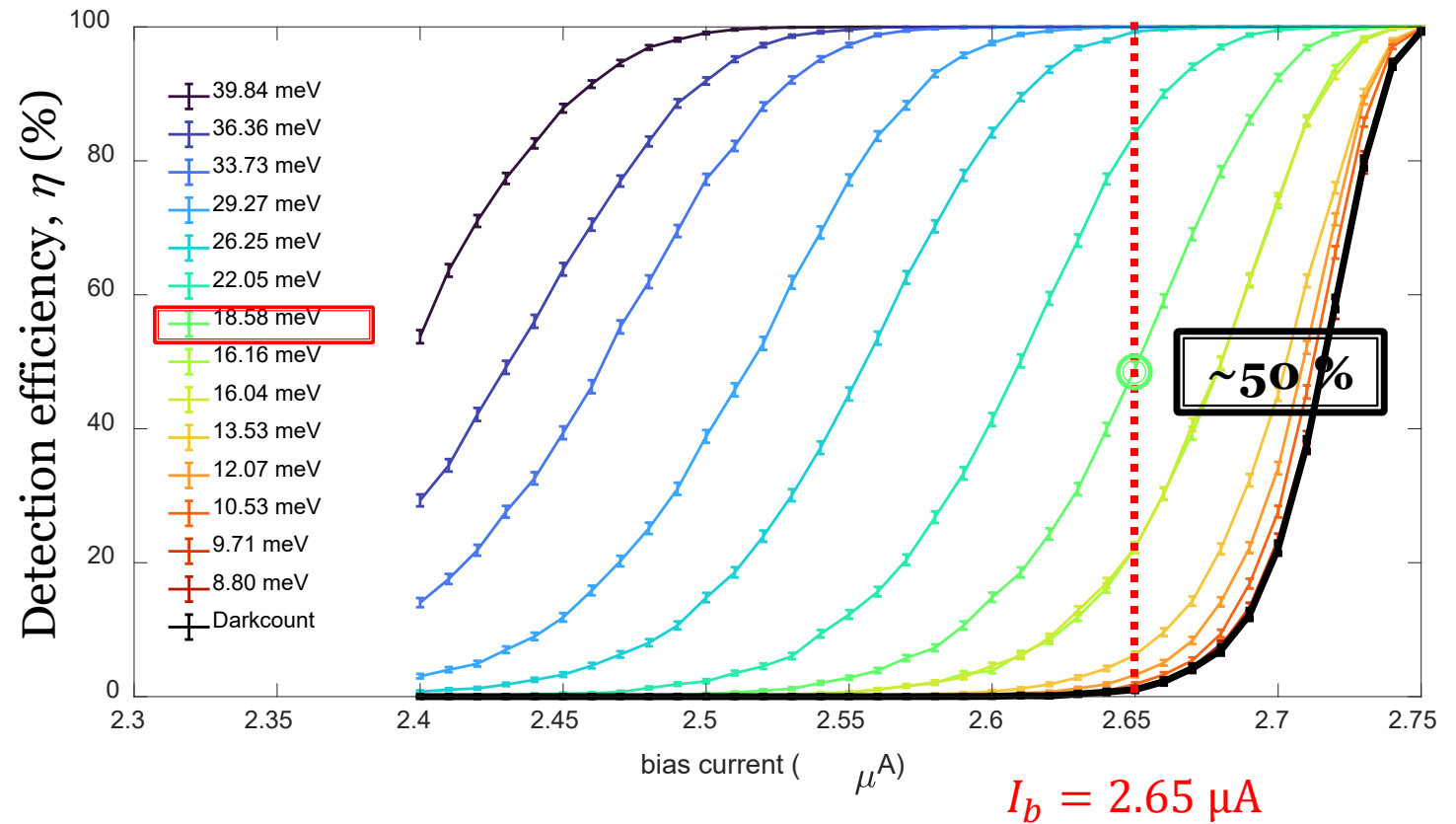
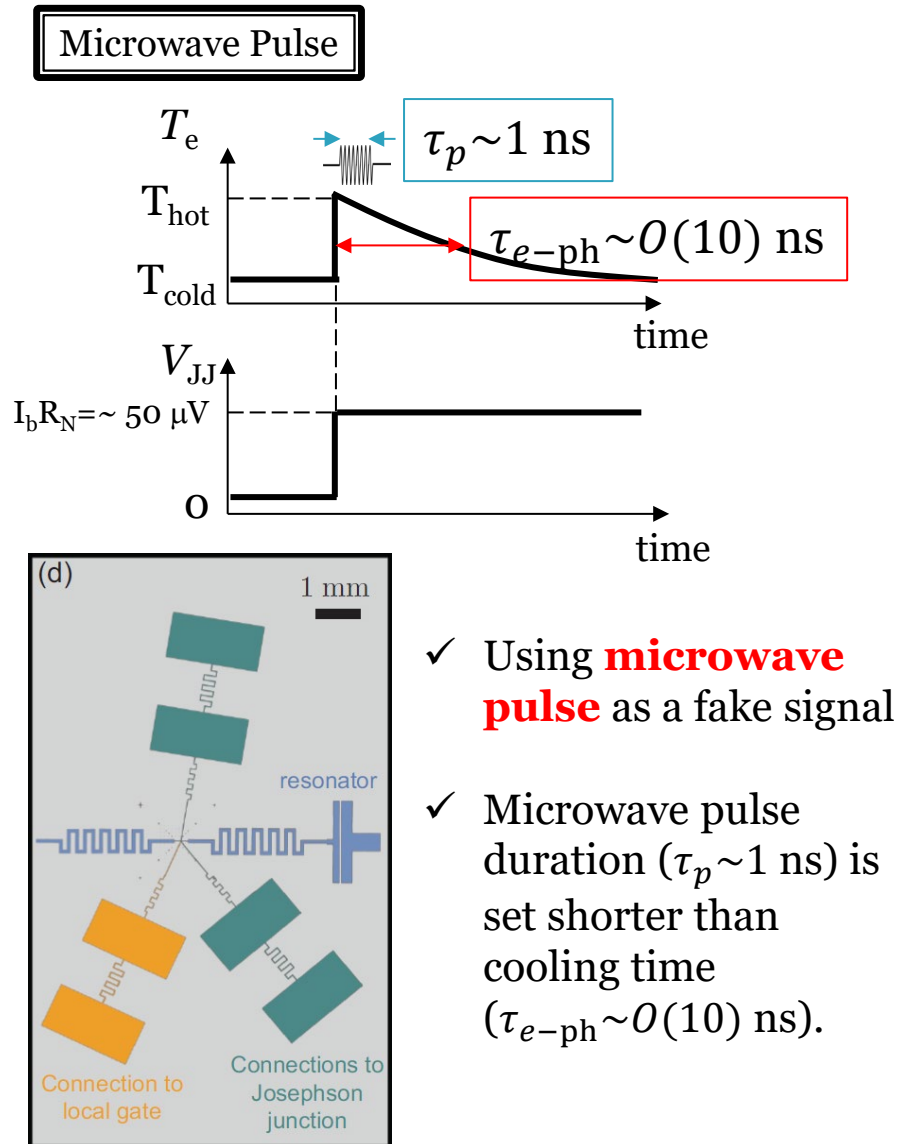
Infrared (IR)



$E = 1 \text{ eV}$
 $\lambda = 1 \mu\text{m}$

Different energy of photons requires different coupling scheme.

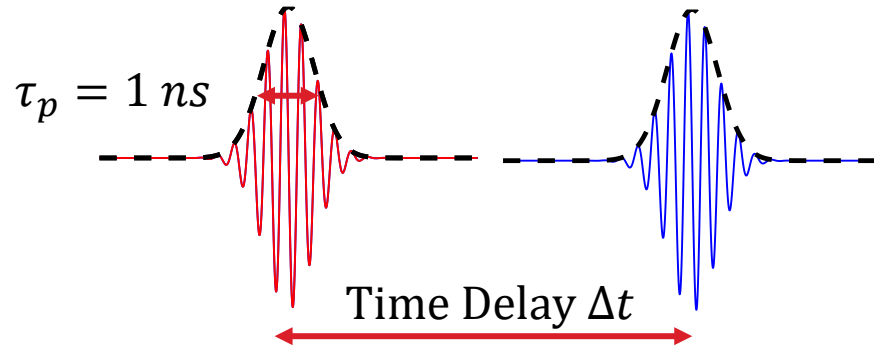
Detector Calibration Scheme



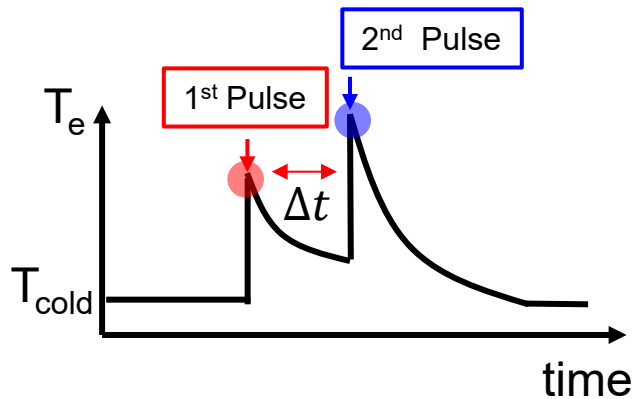
Threshold $E \sim 20 \text{ meV}$

Thermal Relaxation Time Measurement Scheme

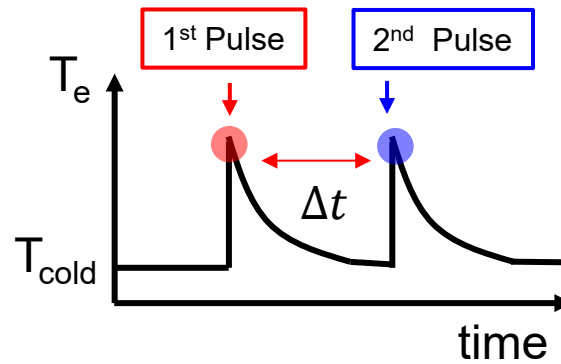
Two-pulse sequence



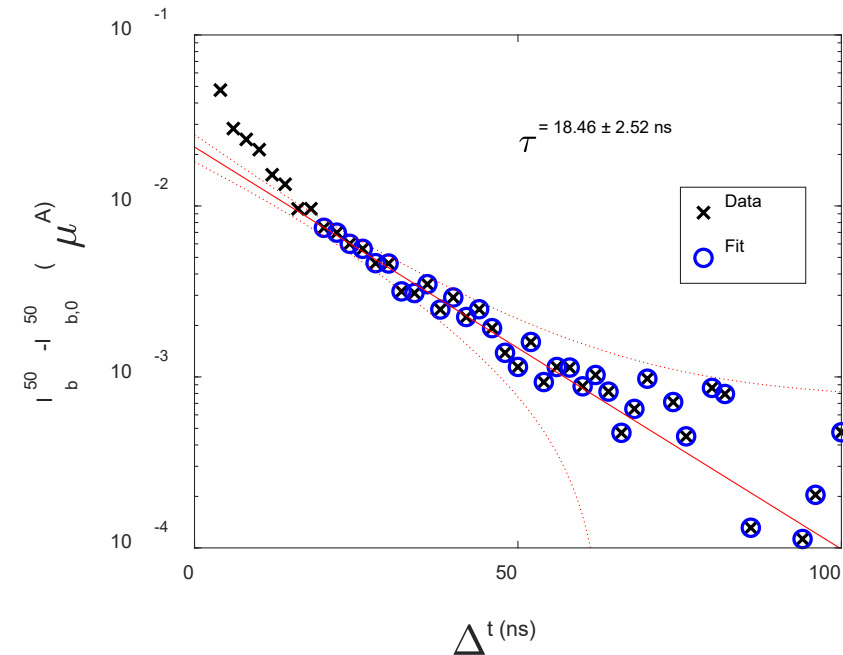
Case 1, $\Delta t \ll \tau_{th}$



Case 2, $\Delta t \gg \tau_{th}$



Measure the switching probability varying Δt
 → calculate the **thermal relaxation time**



$$\tau_{th}^* = 18 \text{ ns} \gg \tau_p = 1 \text{ ns}$$



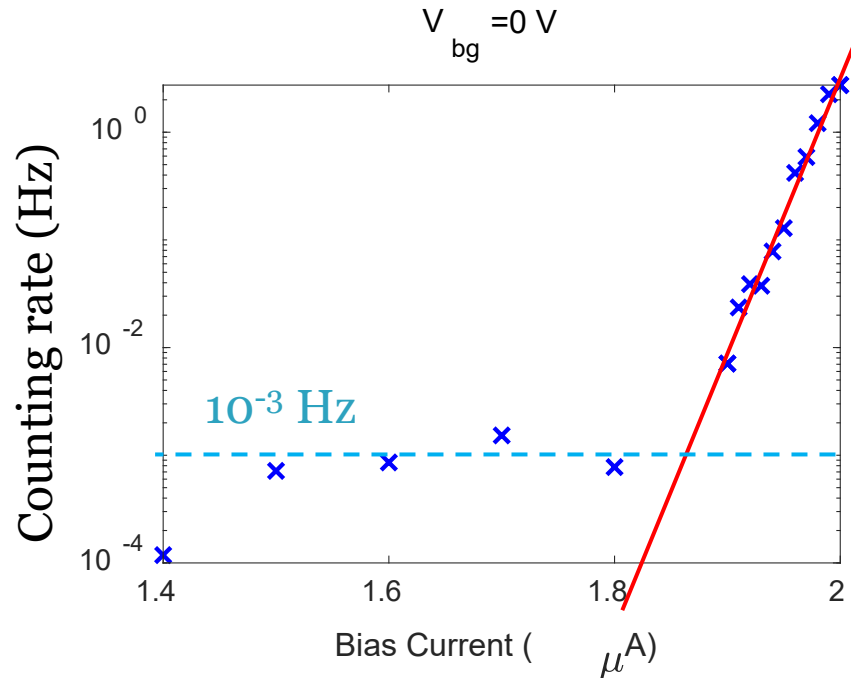
LEE*

***Low energy excess**

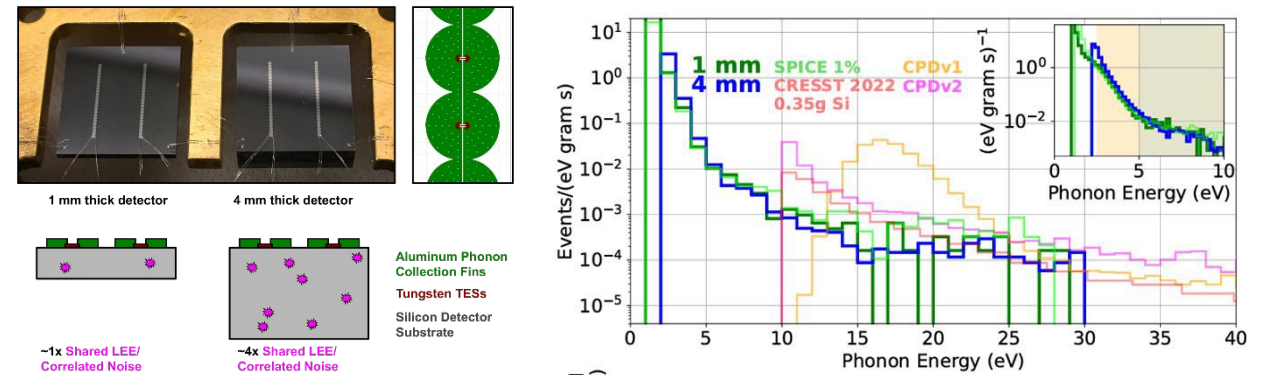
Low Energy Excess (LEE) in GJJ Detector

- Dark matter search was performed at various bias currents.
- Counting rate saturates around 10^{-3} Hz, which is too high to be due to dark matter.
- ➔ Low Energy Excess (LEE)

- **Cosmic ray** may cause such dark counts.

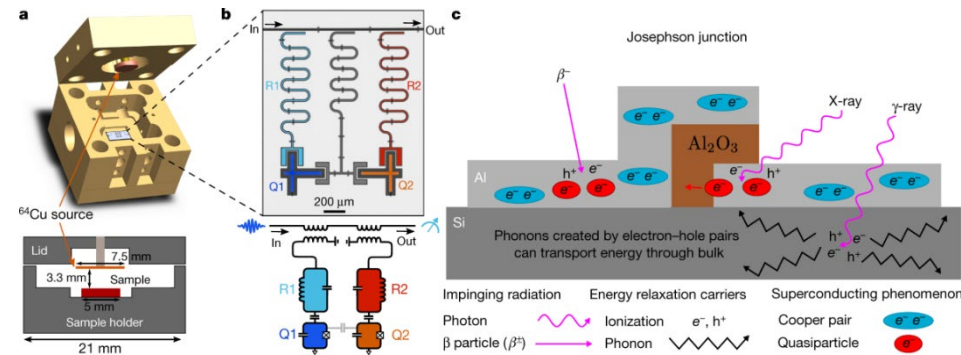


Counting rate of GJJ vs. Bias current



Substrate volume $\uparrow$, DCR $\uparrow$

[arXiv:2505.16092]



[Nature 584, 551–556 (2020)]

Investigating LEE in GJJ Detector

- To avoid LEE and search for dark matter, we need to investigate the origin of LEE.
- Currently, we are comparing devices with different areas of superconducting electrodes.
- We plan to compare devices with different volume of substrates.

