

Search for Invisible Particles at GRANDE

On behalf of the KAPAE & GRANDE Collaboration

Speaker: Hong Joo Kim

Center for High Energy Physics (CHEP)

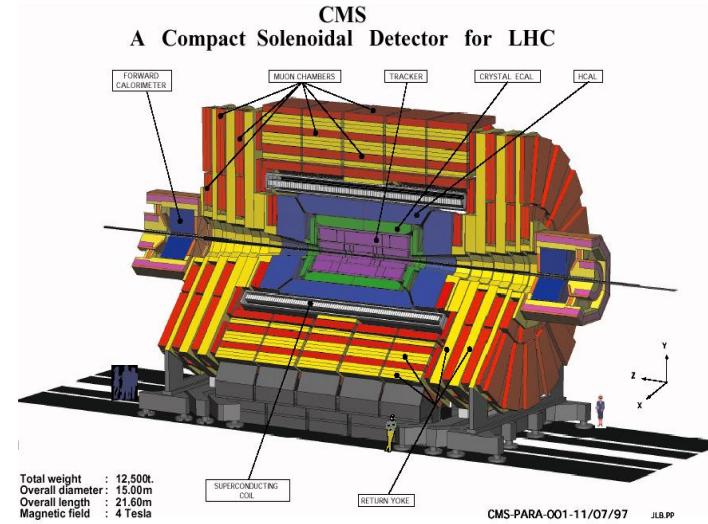
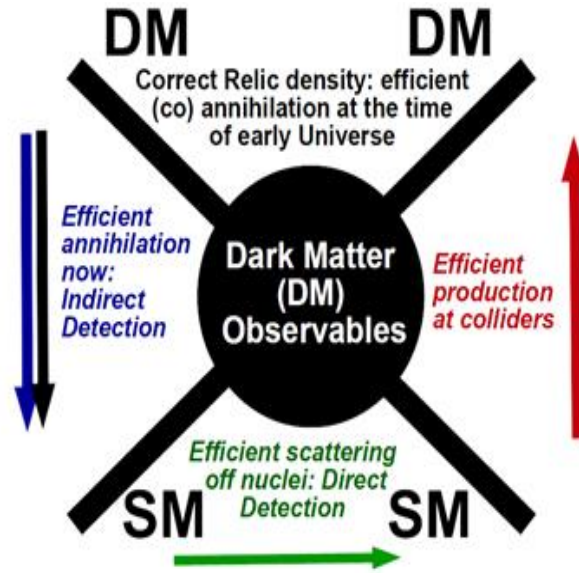
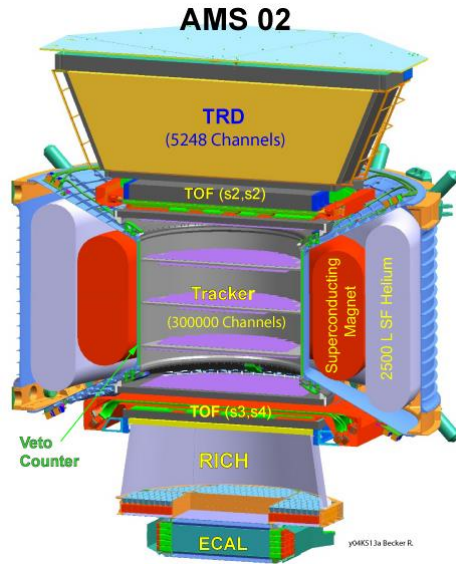
Department of Physics, Kyungpook National University



CUBES 2026

Apr 24 – 26, 2026
The-K Jirisan Family Hotel, Sandong, Gurye
Asia/Seoul timezone

Why universe is dark?



Dark Sector Search

- WIMP
- SUSY
- Axion
- Dark photon

.....
 ??????

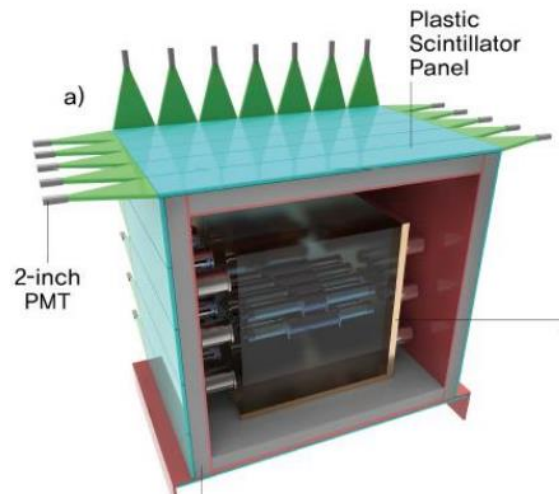
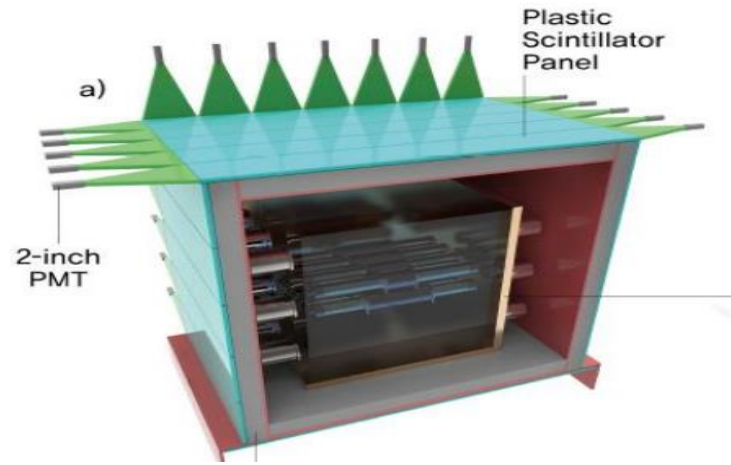
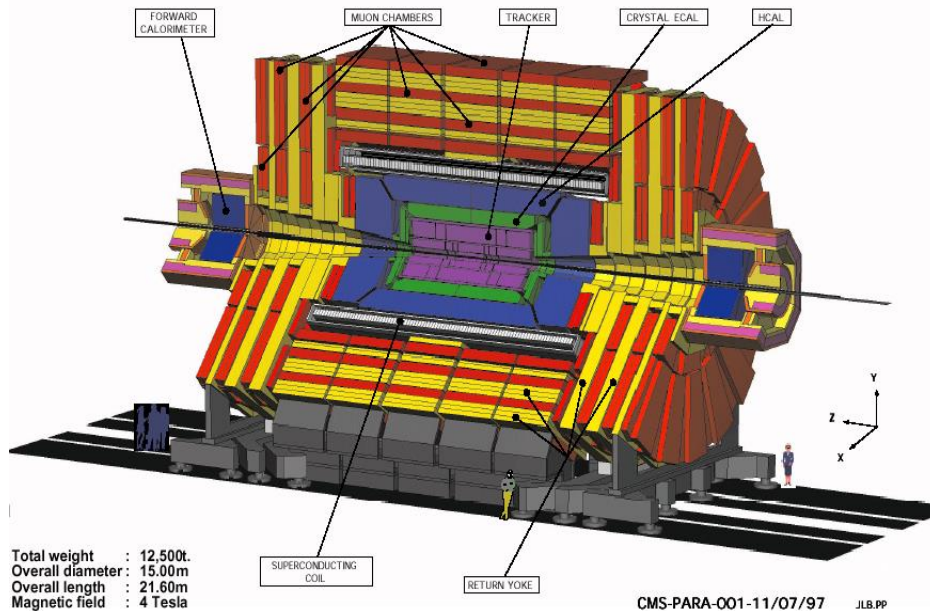


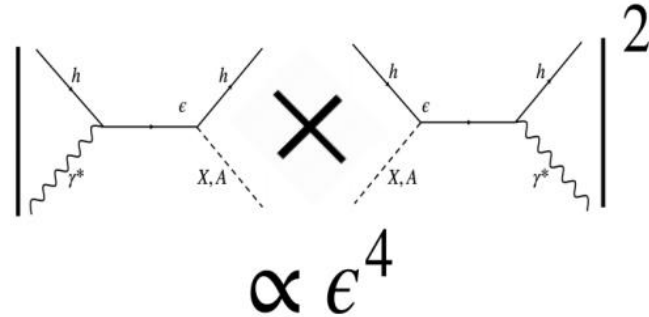
Table top scale for dark sector search?



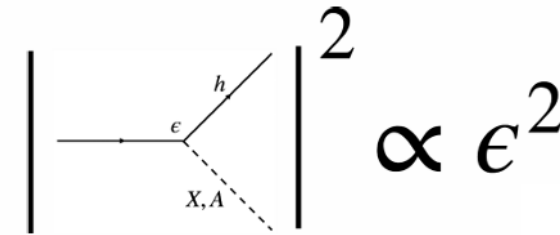
CMS
A Compact Solenoidal Detector for LHC



Radioactive sources for Dark Sector Search : $E < 1$ MeV



Appearance Exp.
Accelerator, Reactor
 $\sim 10^{24}$ POT
 $\rightarrow \epsilon \sim 10^{-6}$



Disappearance Exp.
Radioactive source
 $\sim 10^{12-14}$ Decay
 $\rightarrow \epsilon \sim 10^{-6 \sim 7}$

Advantage

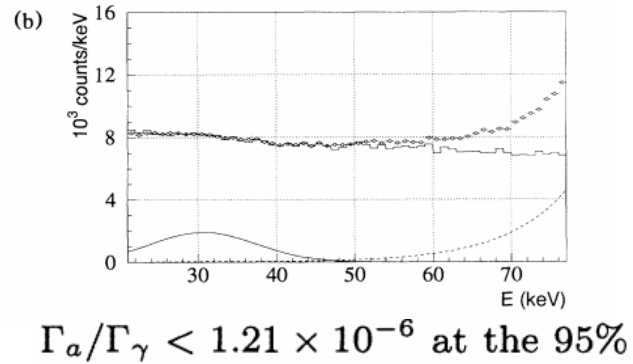
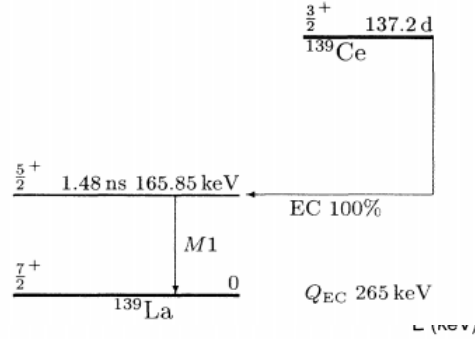
- Decay process: Coupling is **proportional to ϵ^2** not ϵ^4 (Accelerator, Reactor)
- Tabletop-scale experiment (**much lower cost**)
- Source-detector technique (radioactive doping in fast scintillator)
- Time-delayed coincidence method to eliminate backgrounds in the case of isomeric states
- Underground experiment with Cu, Pb shielding

Limitation

- Low mass only (typically ≤ 1 MeV due to detector costs)
- Activity limitations

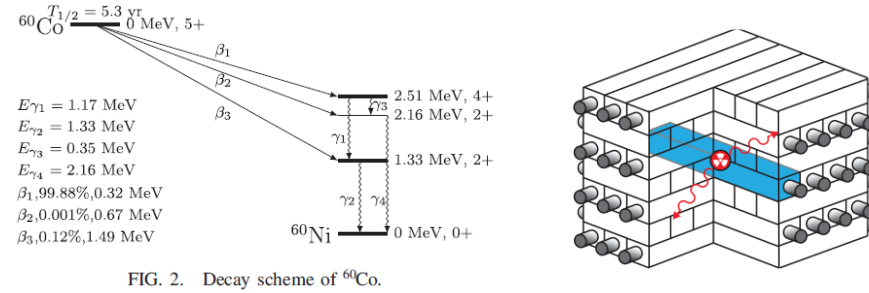
Tabletop experiments with radioactive source?

- Background is too high
No shielding, over ground

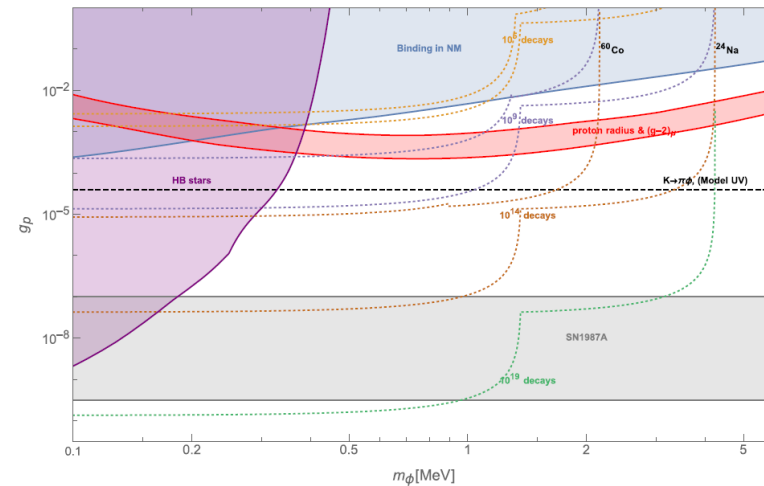


Mass of hadronic axion > 26.7 keV by 95%CL
M.Minowa et al, Phys. Rev. Lett. 71, 4120

- Background is the serious issue
- Detector is too thin for \sim MeV gamma



INVISIBLE DECAY MODES IN NUCLEAR GAMMA CASCADES PHYS. REV. D **99**, 035025 (2019)



When experimentalist meets theorist

PHYSICAL REVIEW LETTERS **133**, 232501 (2024)

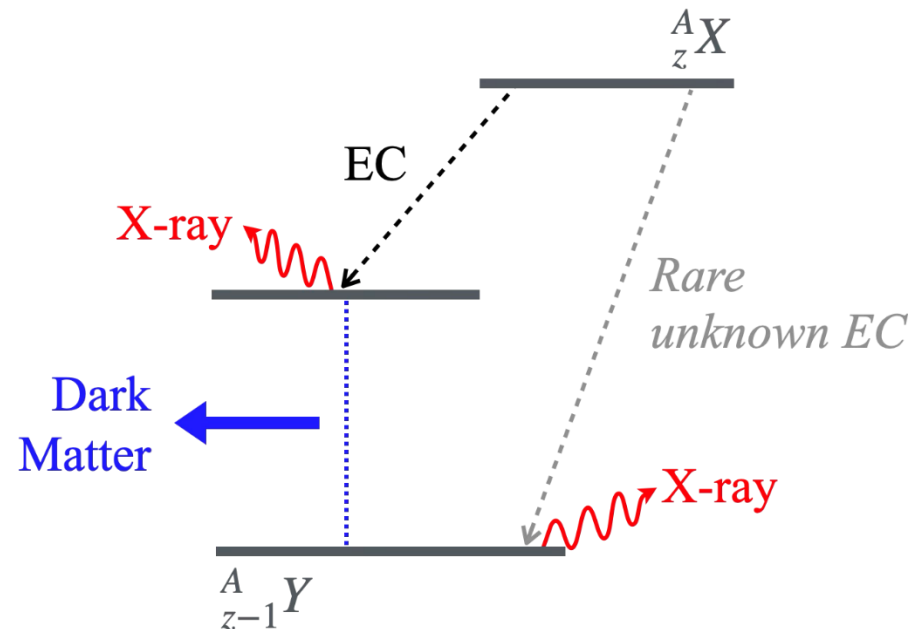
Constraints for Rare Electron-Capture Decays Mimicking Detection of Dark-Matter Particles in Nuclear Transitions

Aagrah Agnihotri^{*} and Jouni Suhonen[†]

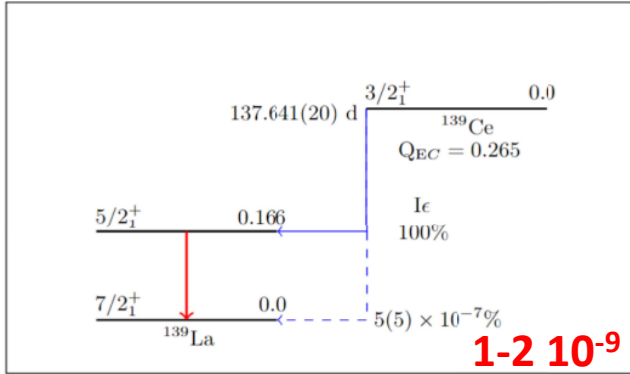
University of Jyväskylä, Department of Physics, P.O. Box 35, FI-40014 Jyväskylä, Finland

Hong Joo Kim[‡]

Department of Physics, Kyungpook National University, Daegu 41566, Republic of Korea



Major possibilities for dark matter particle searches through EC transitions



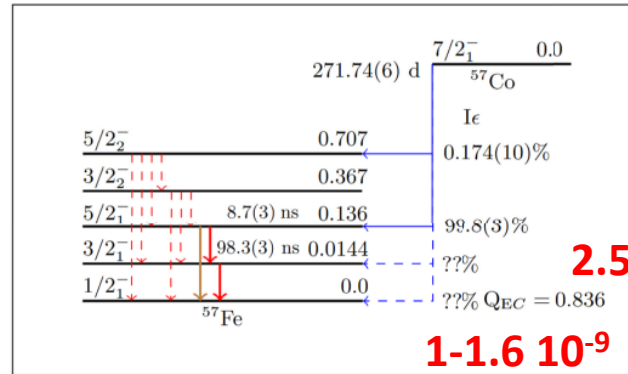
1-2 10⁻⁹

#1
M1 5/2₁⁺ → 7/2₁⁺: 165.86 keV
 → K_α X-ray: ~33 keV



Easiest case
Search for 33 keV X-ray only

CeBr₃:Tl⁴⁴
 CeBr₃:Co⁵⁷
 CeBr₃:Ce¹³⁹

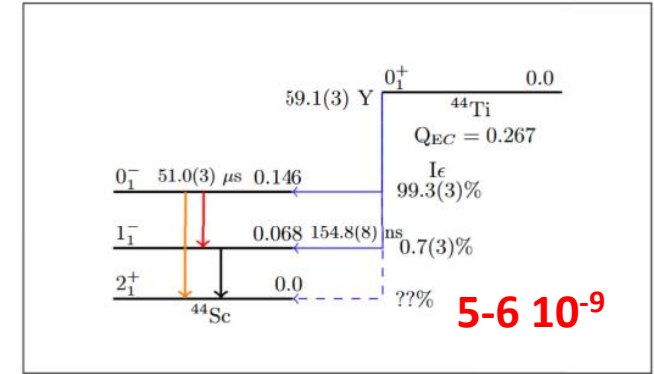
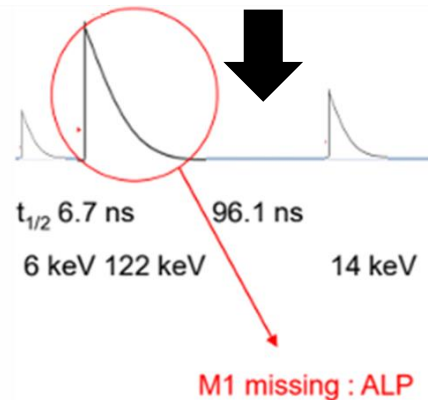


1-1.6 10⁻⁹

2.5-3.3 10⁻⁸

#1
M1 5/2₁⁻ → 3/2₁⁻: 122.06 keV
 → M1 3/2₁⁻ → 1/2₁⁻: 14.4 keV
 → K X-ray: ~6 keV

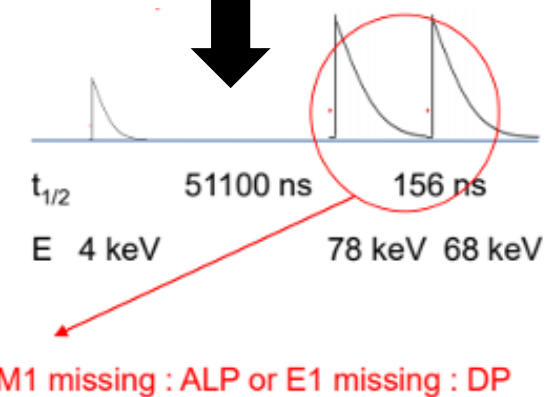
#2
E2 5/2₁⁻ → 1/2₁⁻: 136.47 keV
 → K X-ray: ~6 keV



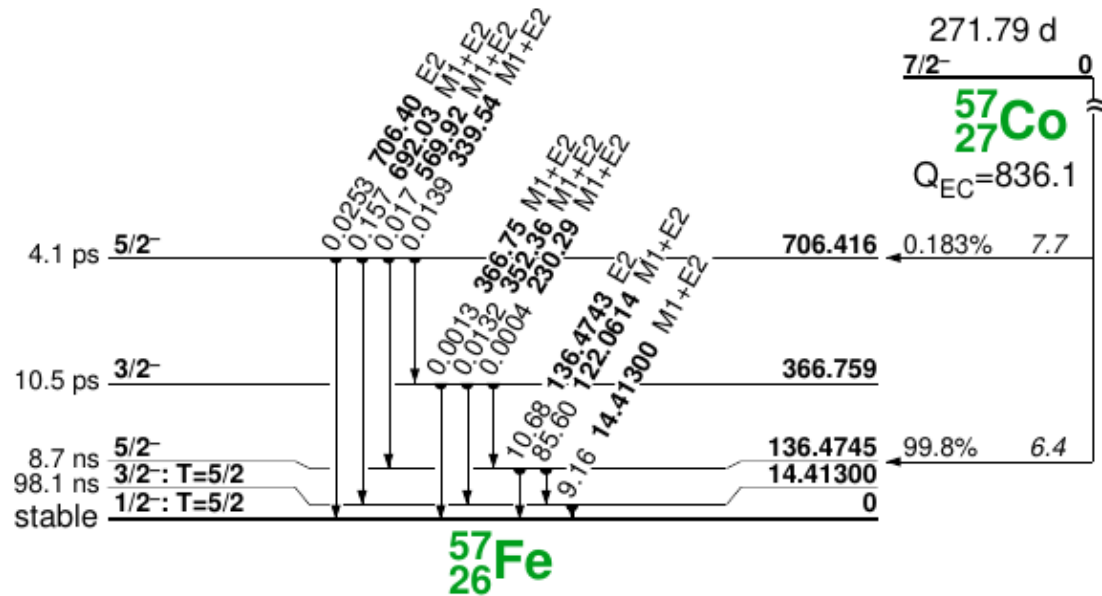
5-6 10⁻⁹

#1
M1 0₁⁻ → 1₁⁻: 78.33 keV
 → E1 1₁⁻ → 2₁⁺: 67.87 keV
 → K_α X-ray: 4 keV

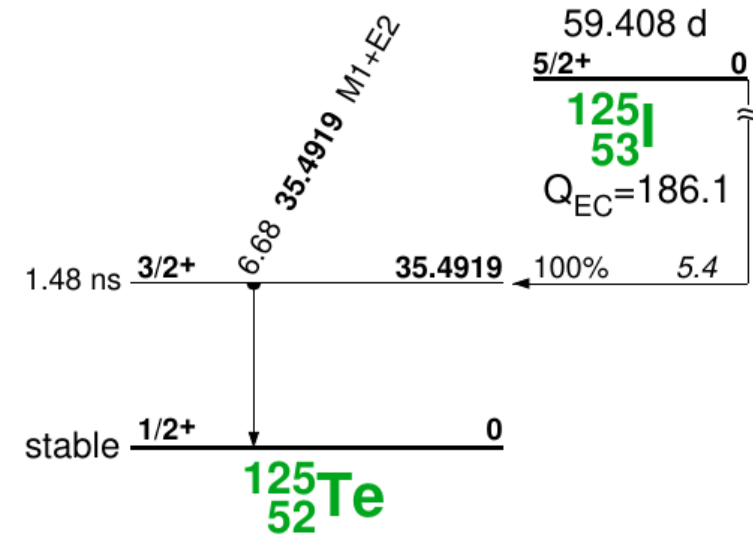
#2
M2 0₁⁻ → 2₁⁺: 146.212 keV
 → K_α X-ray: 4 keV



Limitation of Co-57 and new candidate



Even if branching ratio is 0.15%, 692 keV gamma will escape if BGO thickness is not enough and it create unwanted background signals.



It will have a 35.5 keV gamma + 30 keV X-ray that 30 keV X-ray only will be ALP or anapole signal.

Theoretical calculation of EC to ground state is necessary.

GRANDE

(Global Rare Anomalous Nuclear Decay Experiment)

GRANDE Collaboration (5 countries, 9 institution, 17 members)

- H.J. Kim, N.T. Luan, D.W. Jeong, J.Y. Cho, J.Y. Lee, L. Truc (Kyungpook National University, KOREA)
- M.H. Lee (Center for Underground Physics, IBS, KOREA)
- J.H. So (Yemi operation center, IBS, KOREA)
- J. Suhonen, A. Agnihotri (University of Jyväskylä, Finland)
- J. Kaewkhao (Nakhon Pathom Rajabhat University, Thailand)
- S. Kothan (Chiangmai University, Thailand)
- M. Haroon, S. Khan (Kohat University of Science and Technology, Pakistan)
- F. Qayyum, M. A. Program, G. Rooh (Abdul Wali Khan University Mardan, Pakistan)
- I. Gkialas (University of the Aegean, Greece)

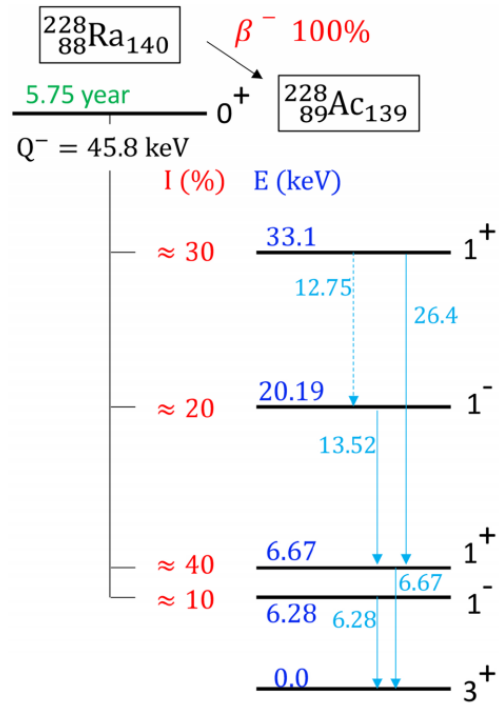
Scope of the GRANDE

- Experiment setup at the Yemi Underground Lab with low-background Cu+Pb shielding
- Radioactive source embedded in a crystal scintillator
(CeBr₃ : fast, high light yield, low background)
- 4 π VETO with BGO
- Aiming for extremely low background condition
- Measurement of Rare EC process, rare beta, and alpha decay and searching for new particles.

Radioactive source doped crystal scintillators?

Measurement of New Ac-228 Isomers Using a Ra-228 doped CeBr3

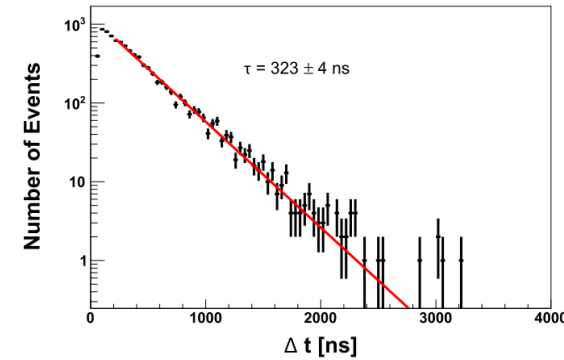
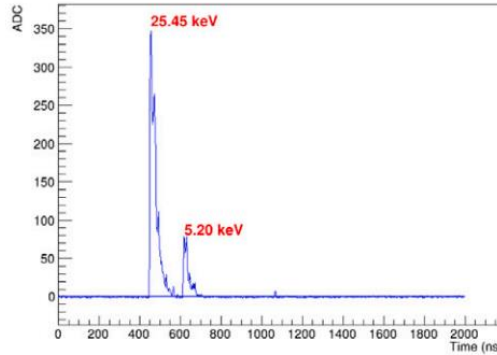
D.H. Lee et al, Radiation Physics and Chemistry 238 (2026) 113148



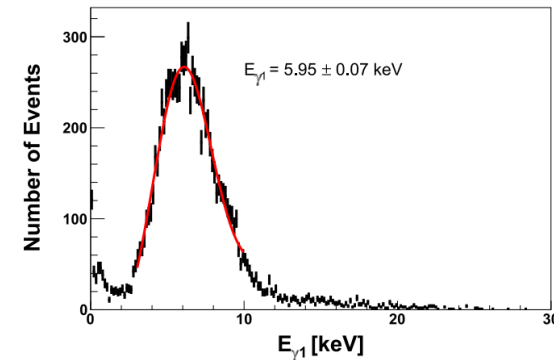
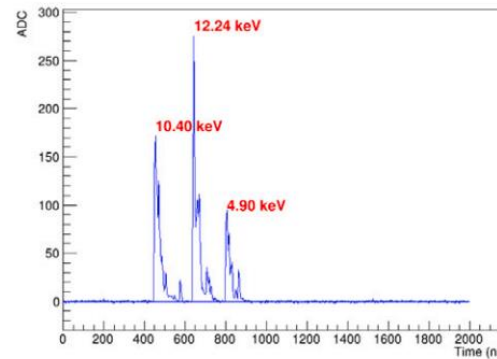
(5 mm \times Φ 7 mm)



Encapsulated $\text{CeBr}_3:^{228}\text{Ra}$



(a)



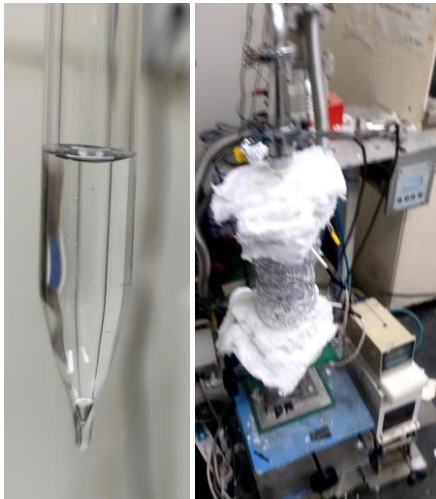
Low energy gamma (X-ray), beta and internal conversion can be measured efficiently.

Radioactive source doped $CeBr_3$ crystal growth

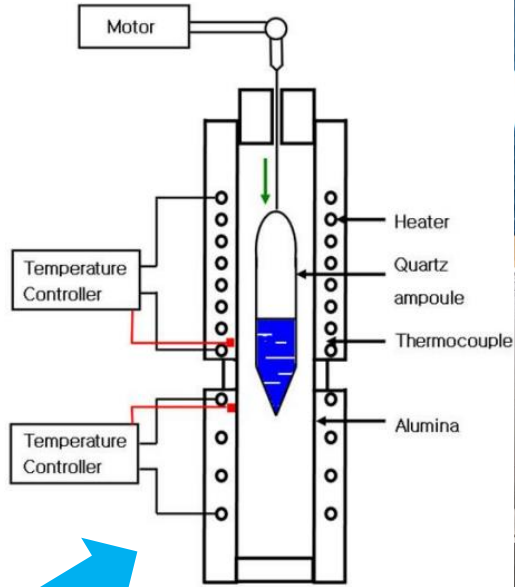
Rad. Source in Solution



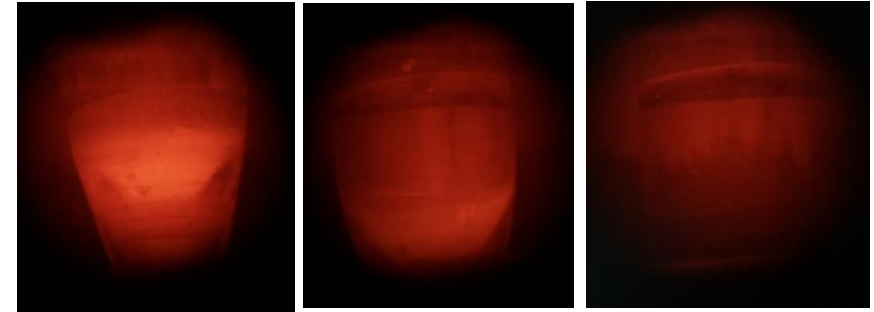
Rad. ampoule coating
(under vacuum)



"home-made" Bridgman furnace



Realtime camera assistant interface monitor
monitoring and changing growth rate

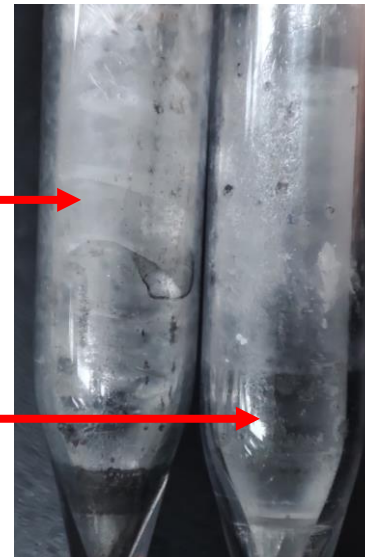


Fail process (dehydrate)
(solved)



#Early
Dirty ampoule
Bad cracks

#NOW
#Optimized
Clean
No cracks
Long



Cutting



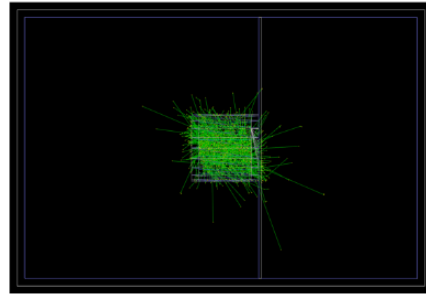
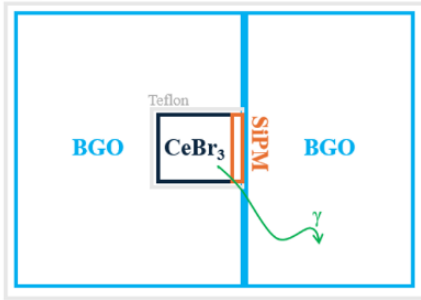
Crystal



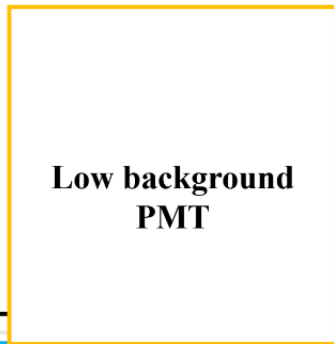
Source-as-Detector Experiment

CeBr₃ advantages : Fast decay time, high light yield, good energy resolution

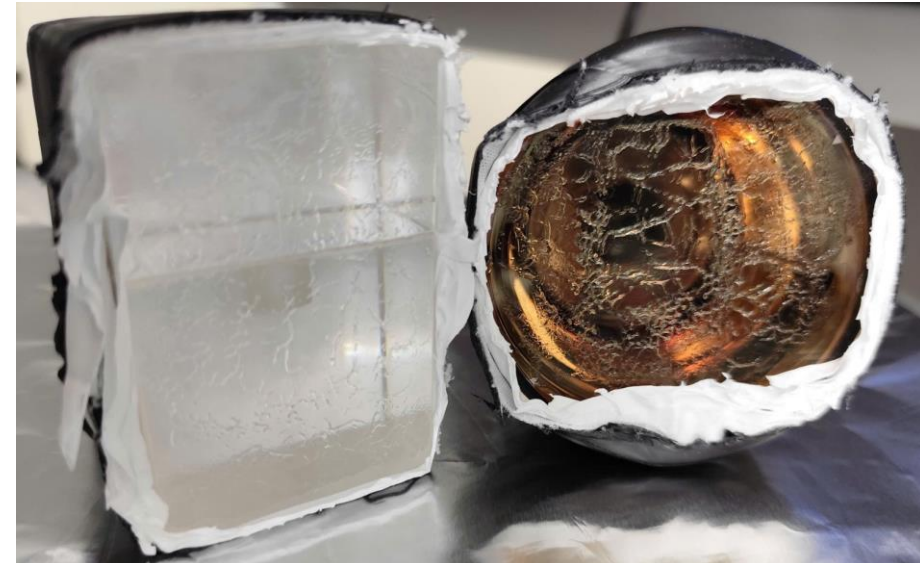
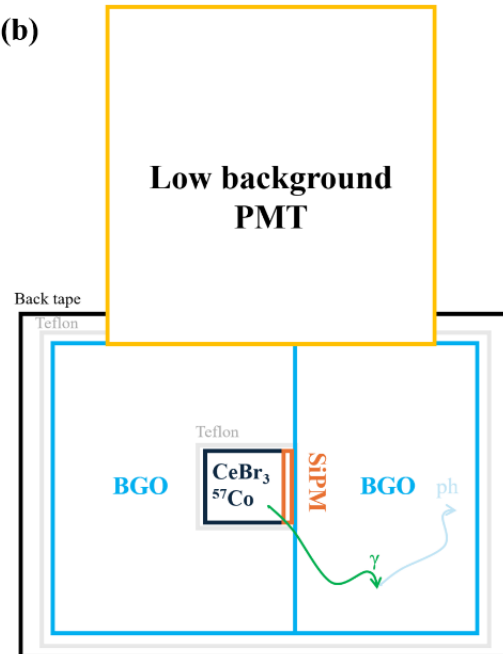
Disadvantages : very hygroscopic, internal background?



(a)

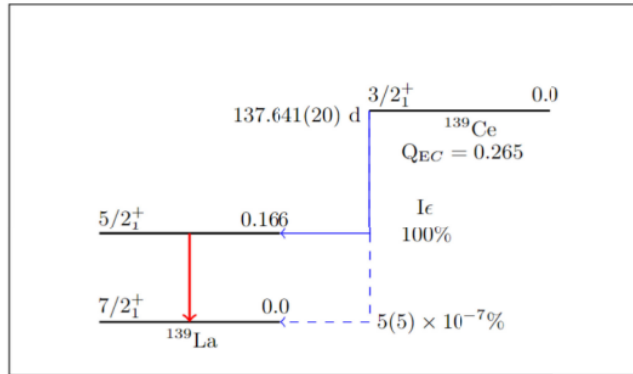


(b)



Phoswich detector

1st experiment using ¹³⁹Ce Source



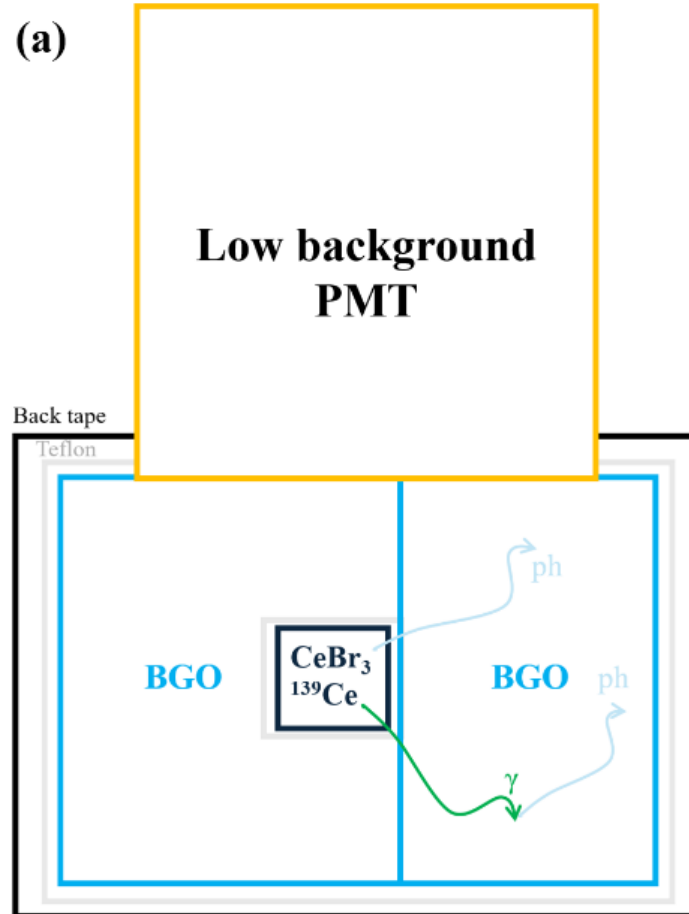
#1

M1 $5/2^+_1 \rightarrow 7/2^+_1$: 165.86 keV

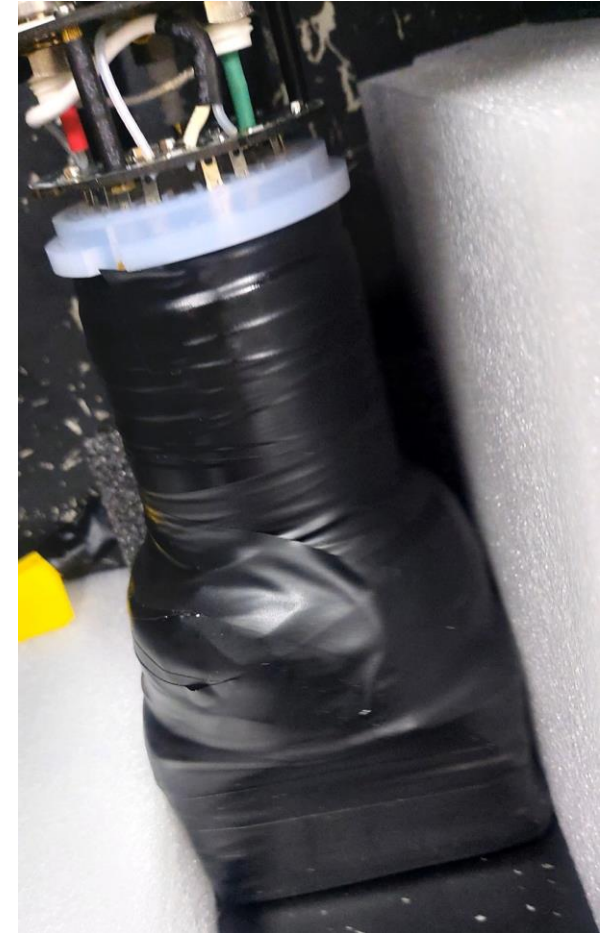
→ K_α X-ray: ~33 keV



Easiest case
Search for single X-ray



Phoswich detector

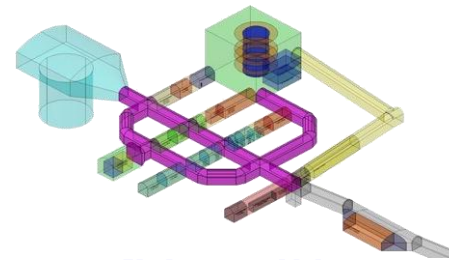
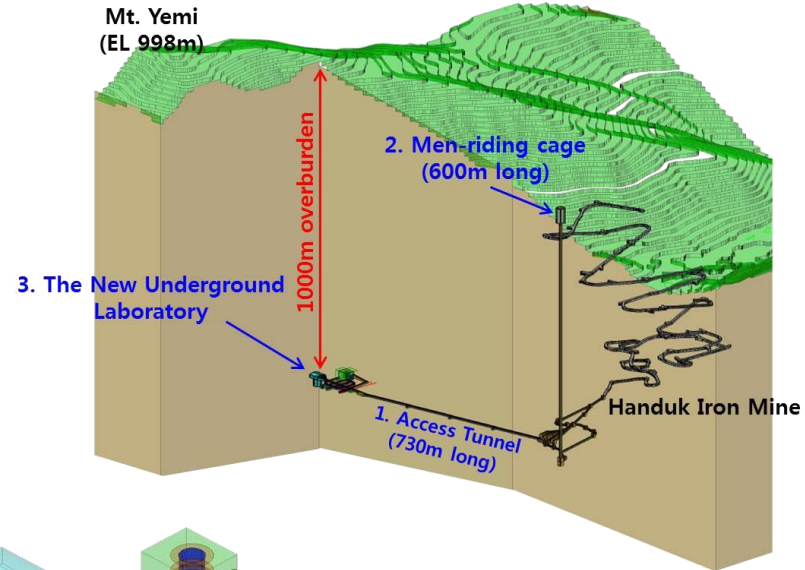
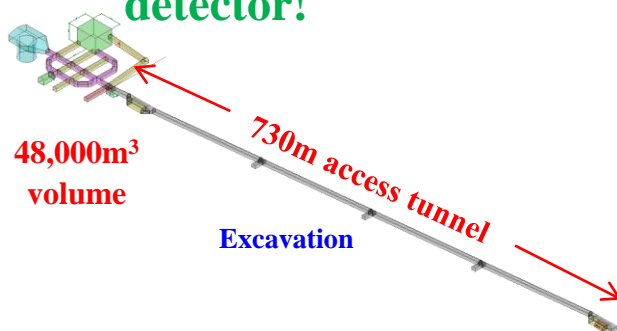


Yemi Underground Laboratory (Center for Underground Physics)

- **COSINE** Dark matter search
- **AMoRE** $0\nu\beta\beta$ experiment
- **Rare decay experiment (KAPAE-II, GRANDE)**

Future possibility

- **k-Ton LSC** for DB, solar ν
- Rare decay search with HPGe
- **Ton scale DB exp.**
- **Dark photon exp.**
- **New dark matter search**
- **New idea with new detector!**



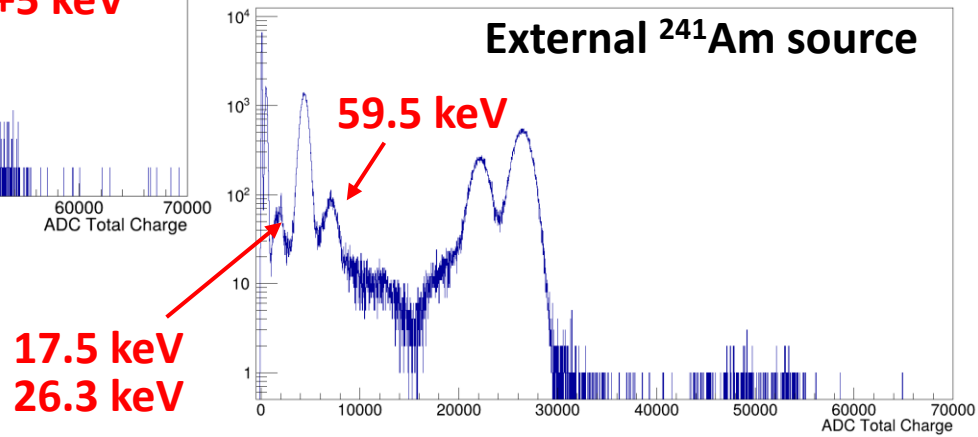
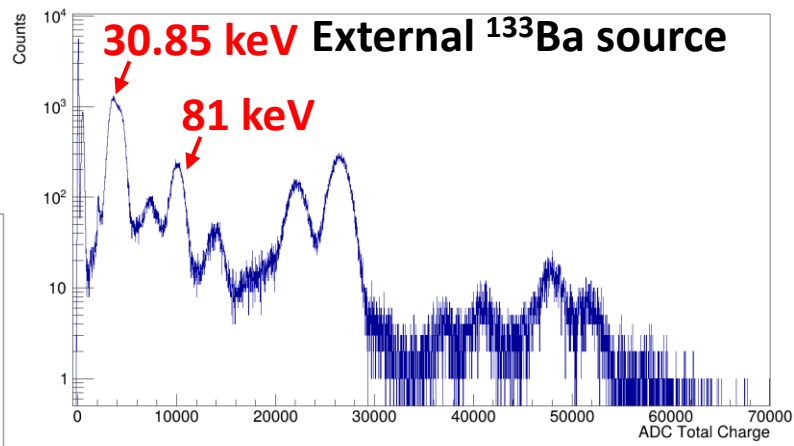
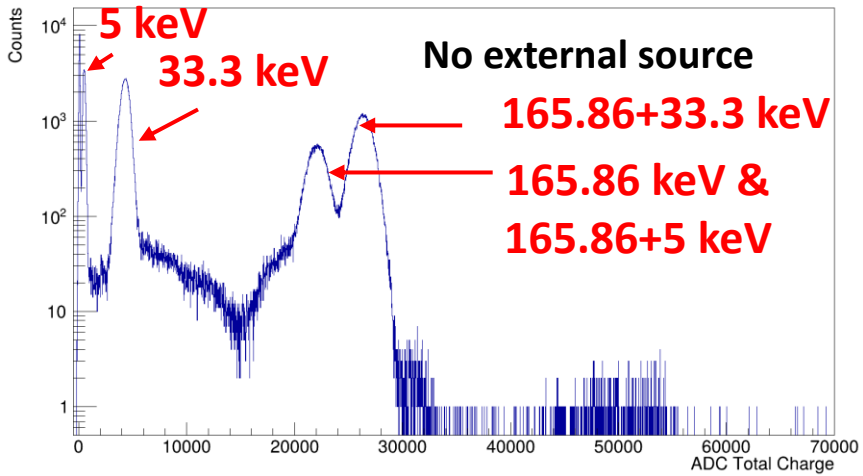
Experiment using ^{139}Ce Source



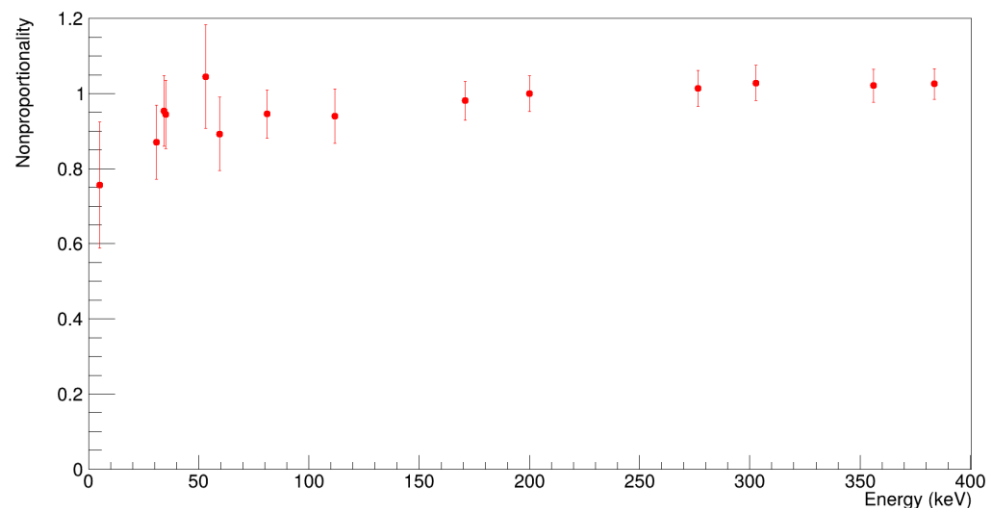
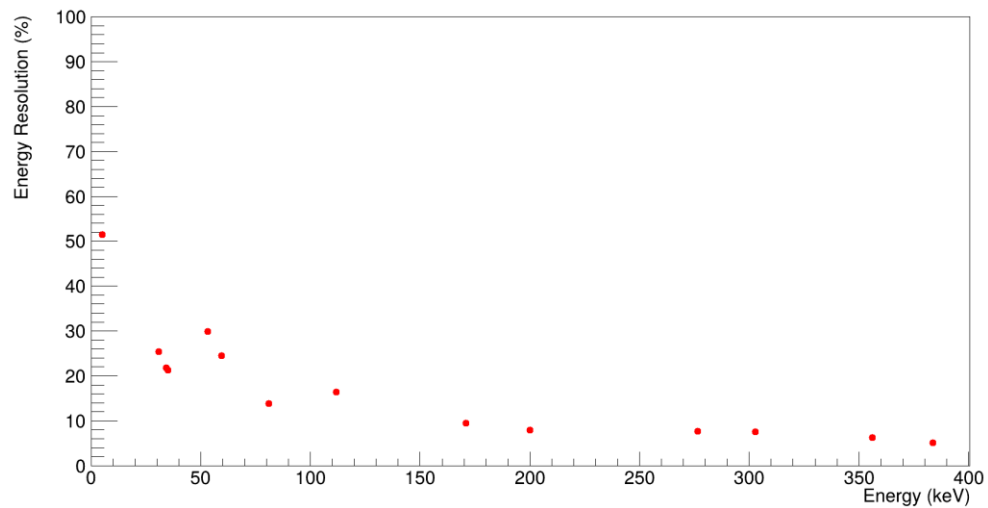
- Yemi Underground lab. with 15cm Pb shielding.
- Super bi-alkali PMT 12669SEL (High quantum efficiency, low background)
- Notice 500-MHz flash Analog to Digital Converter
- Caen High Voltage Supplier
- PC with 18Tb storage
- Data rate ~ 20 kHz

**Running 5 month without problem!
No degradation of CeBr3 crystal**

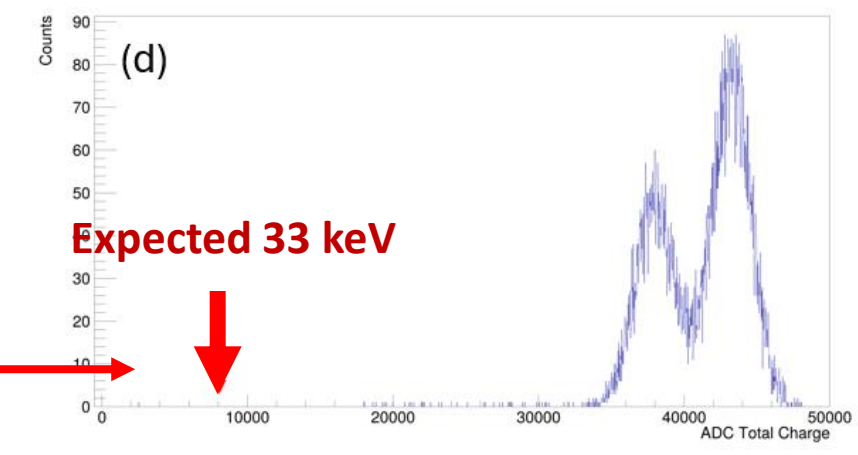
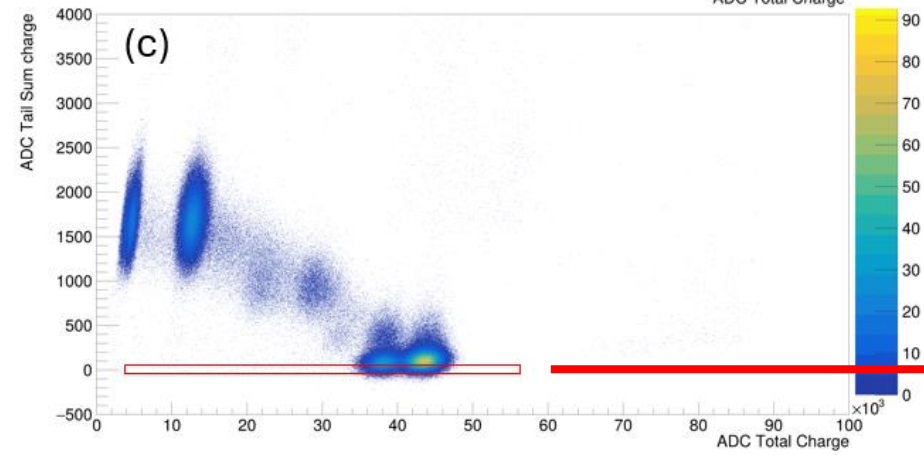
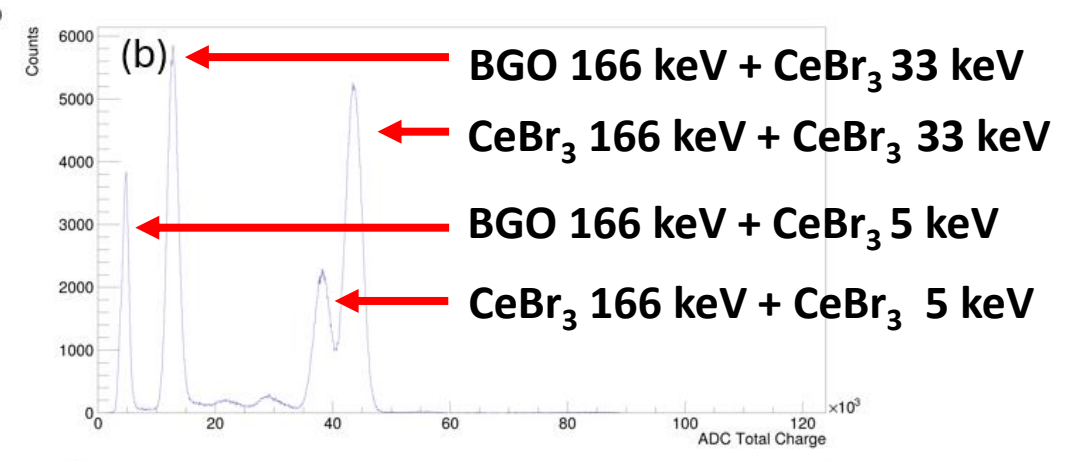
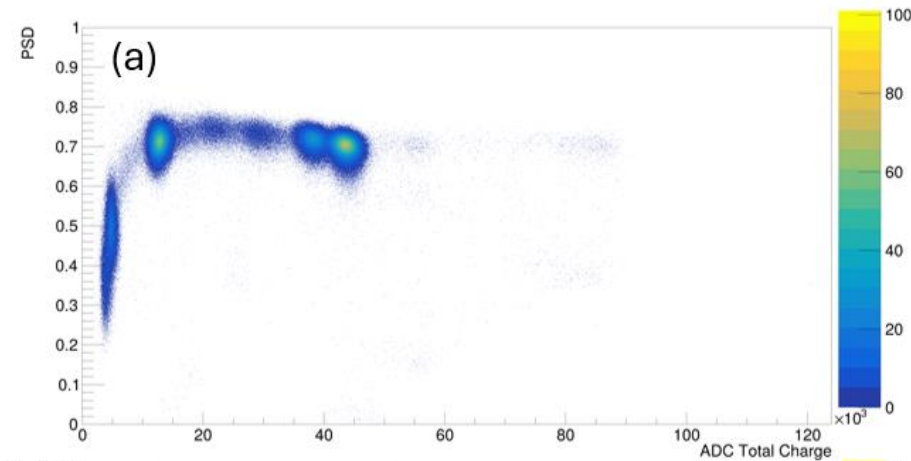
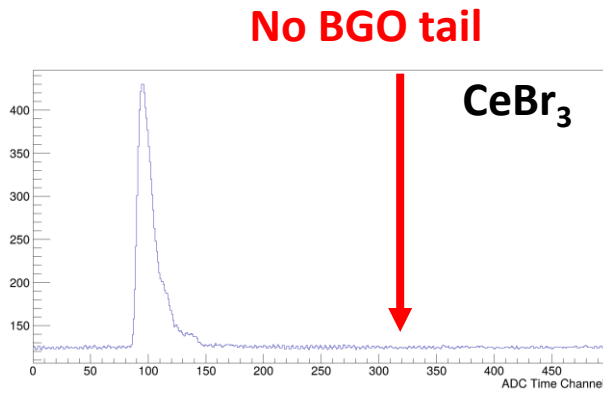
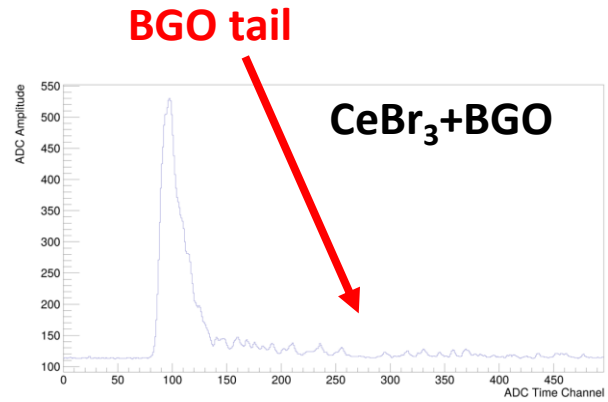
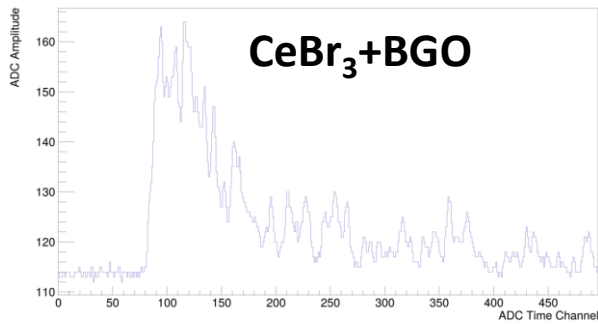
CeBr₃:¹³⁹Ce



Isotope	Energy (keV)
¹³⁹ Ce	5
	33.
	170
	200
²⁴¹ Am	59.5
¹³³ Ba	30.85
	35.1
	53.15
	81
	111.85
	276.4
	302.85
	356.02
383.85	

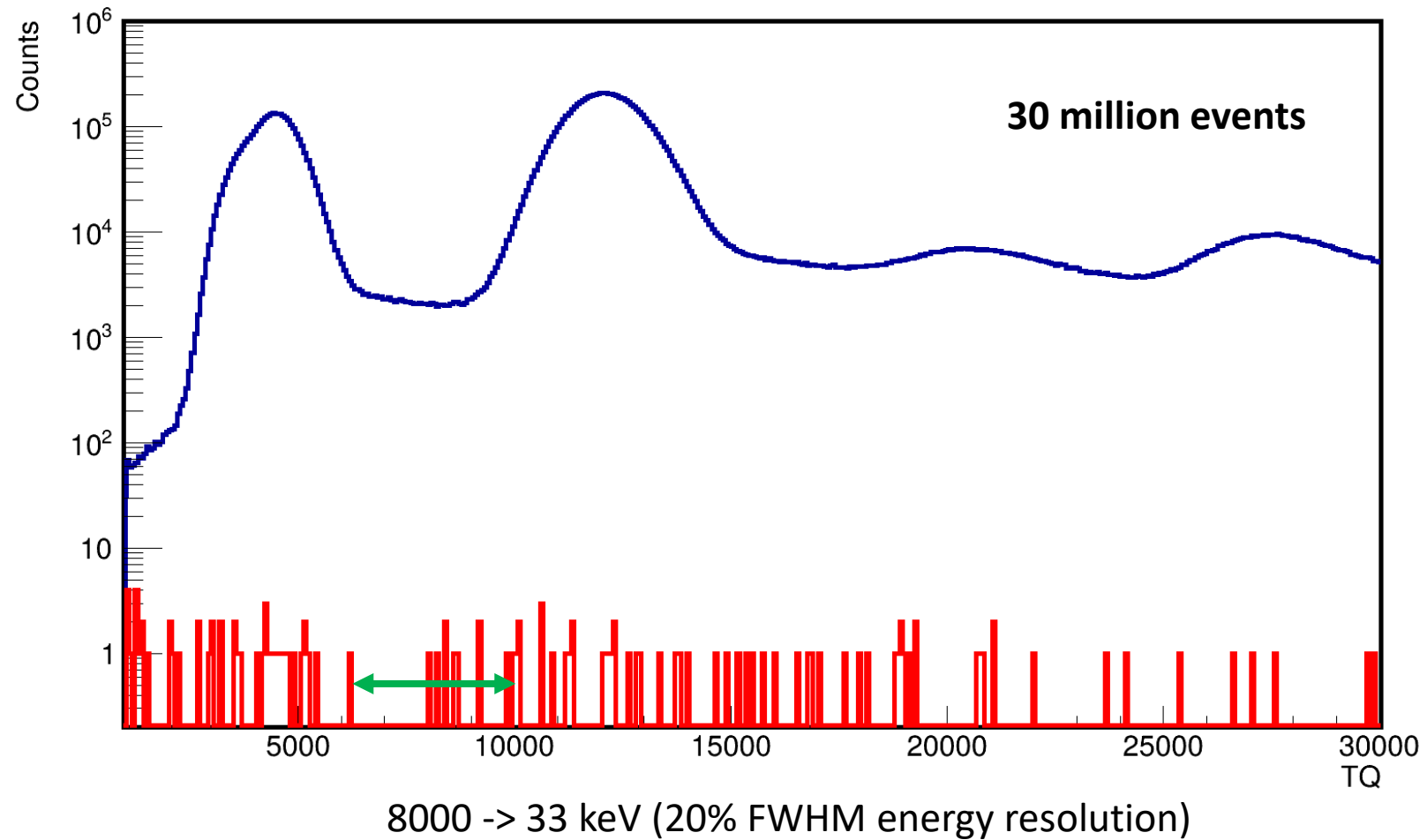


CeBr₃:¹³⁹Ce + BGO veto



CeBr₃:¹³⁹Ce + BGO veto

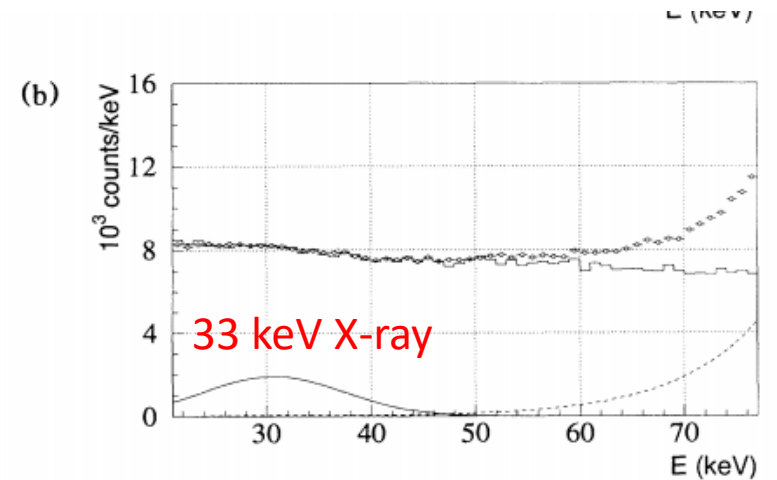
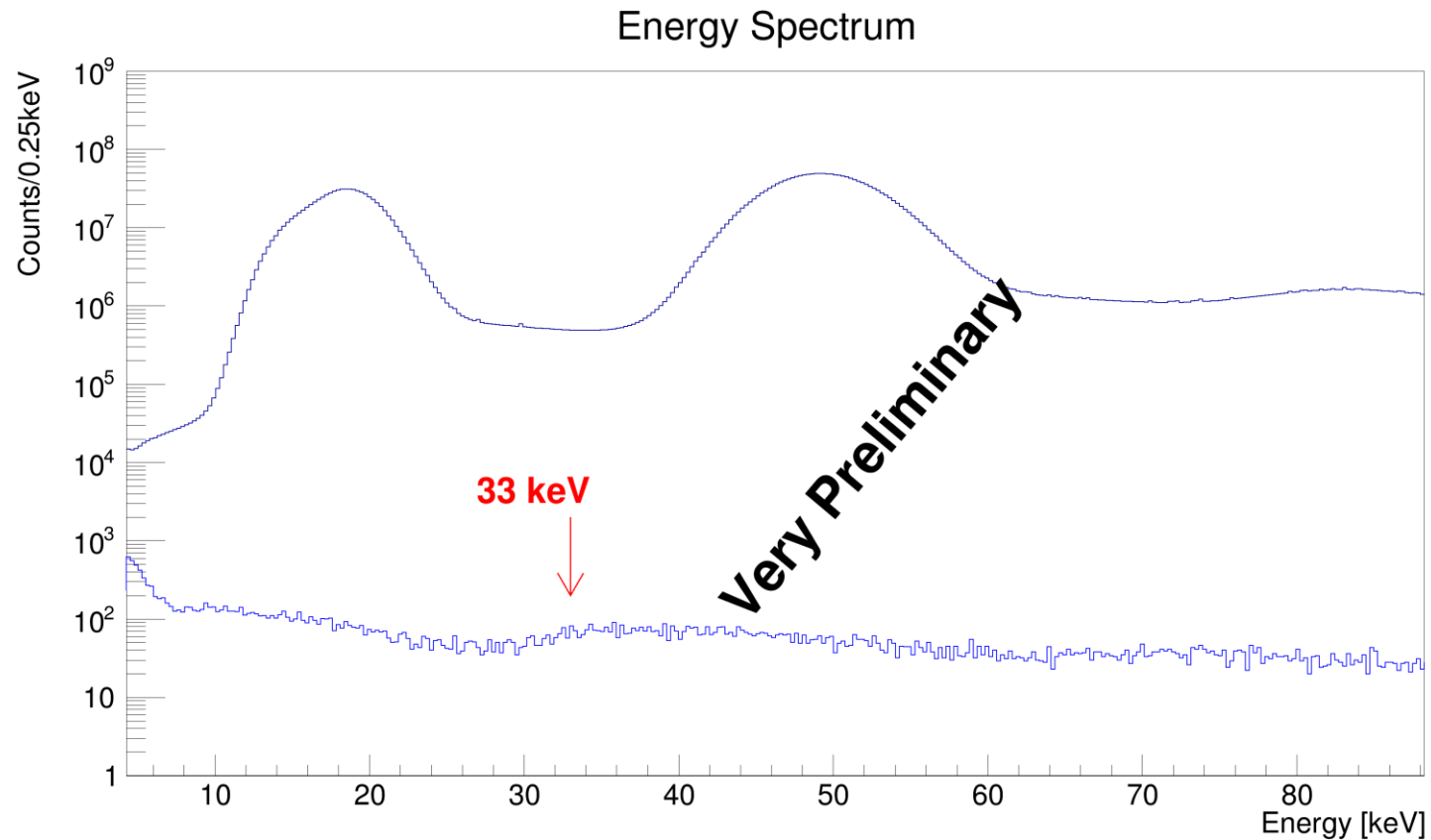
- 2 million events → reach 1×10^{-6}
- 30 million events → reach 4×10^{-7}



CeBr₃:¹³⁹Ce + BGO veto

☐ Very preliminary results with 10 billion events (5 month data taking)

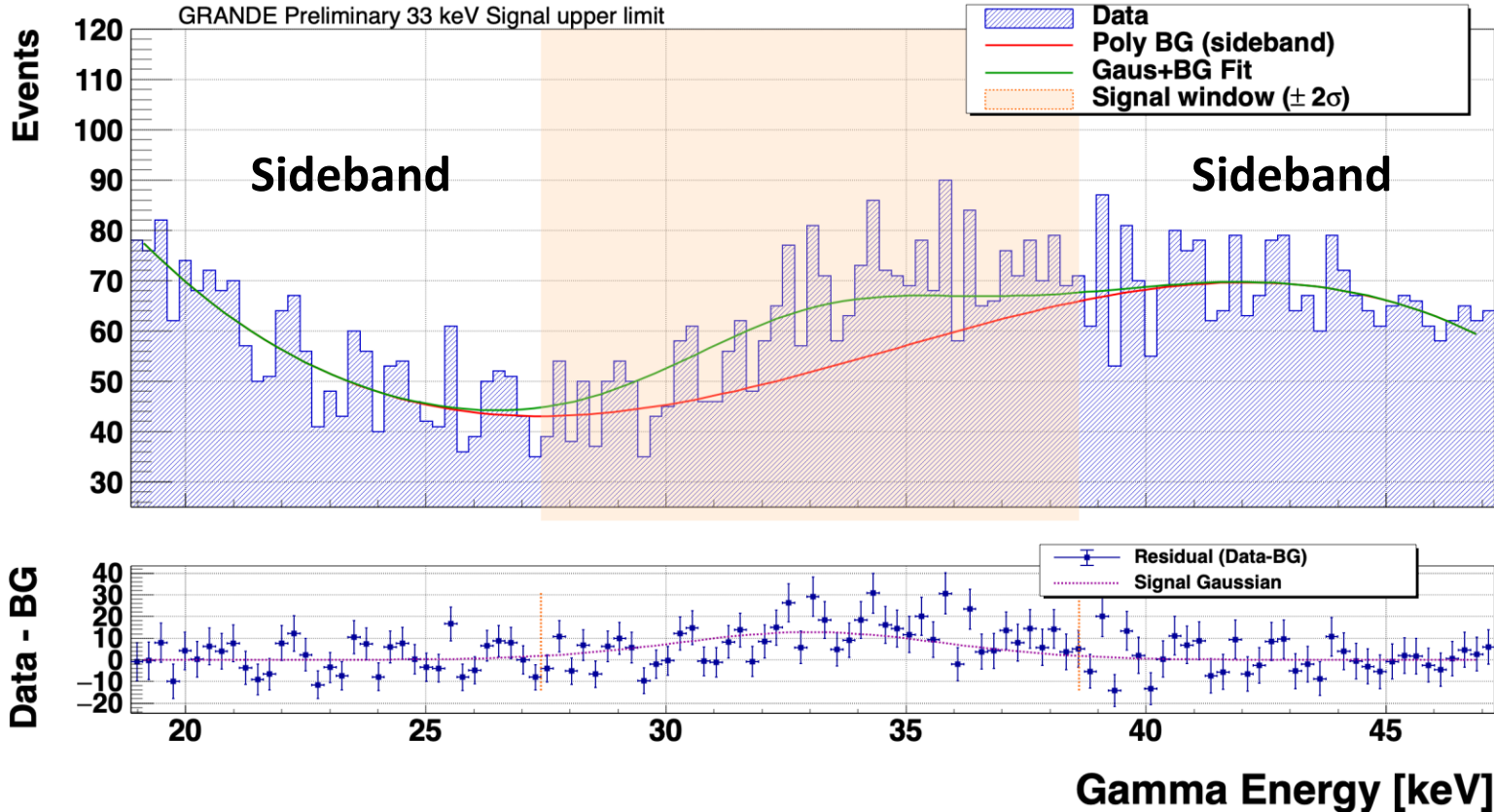
☐ 200 times lower background than M. Minowa's experiment



$$\Gamma_a/\Gamma_\gamma < 1.21 \times 10^{-6} \text{ at the 95\%}$$

Mass of hadronic axion > 26.7 keV by 95%CL
M.Minowa et al, Phys. Rev. Lett. 71, 4120

Preliminary Branching ratio upper limit of ^{139}Ce



- **Signal window:** 33 keV, $\pm 2\sigma$ region

- **Sideband:** 33 keV, $\pm 5\sigma$ region, excluding the signal window ($\pm 2\sigma$) with 3rd-order polynomial

- **Branching ratio upper limit:** 1.1×10^{-8} (90% C.L.)

Preliminary

- **Previous experiment**

$$\Gamma_a/\Gamma_\gamma < 1.21 \times 10^{-6} \text{ at the 95\%}$$

Background signals due to 166 keV gamma escape from BGO since it is only 2 cm thickness.

Upgrade of $\text{CeBr}_3:^{139}\text{Ce}$ experiment with KAPAE-II

Phoswitch to KAPAE-II

1. BGO thickness is only 2.0cm (166 keV γ leak) \rightarrow 6cm for KAPAE-II setup
2. Phoswitch does not give a perfect separation between CeBr_3 and BGO signal

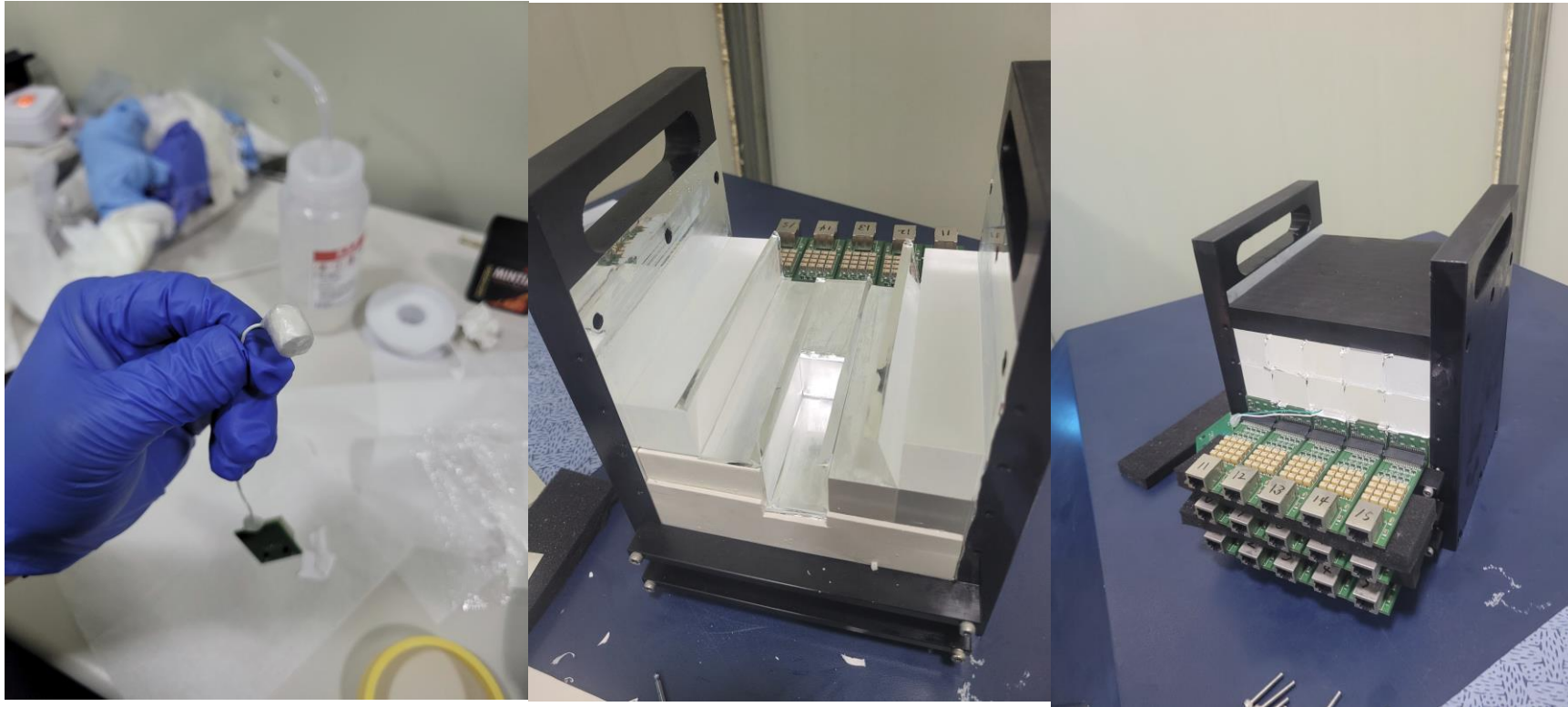
Phoswitch setup



KAPAE-II Setup



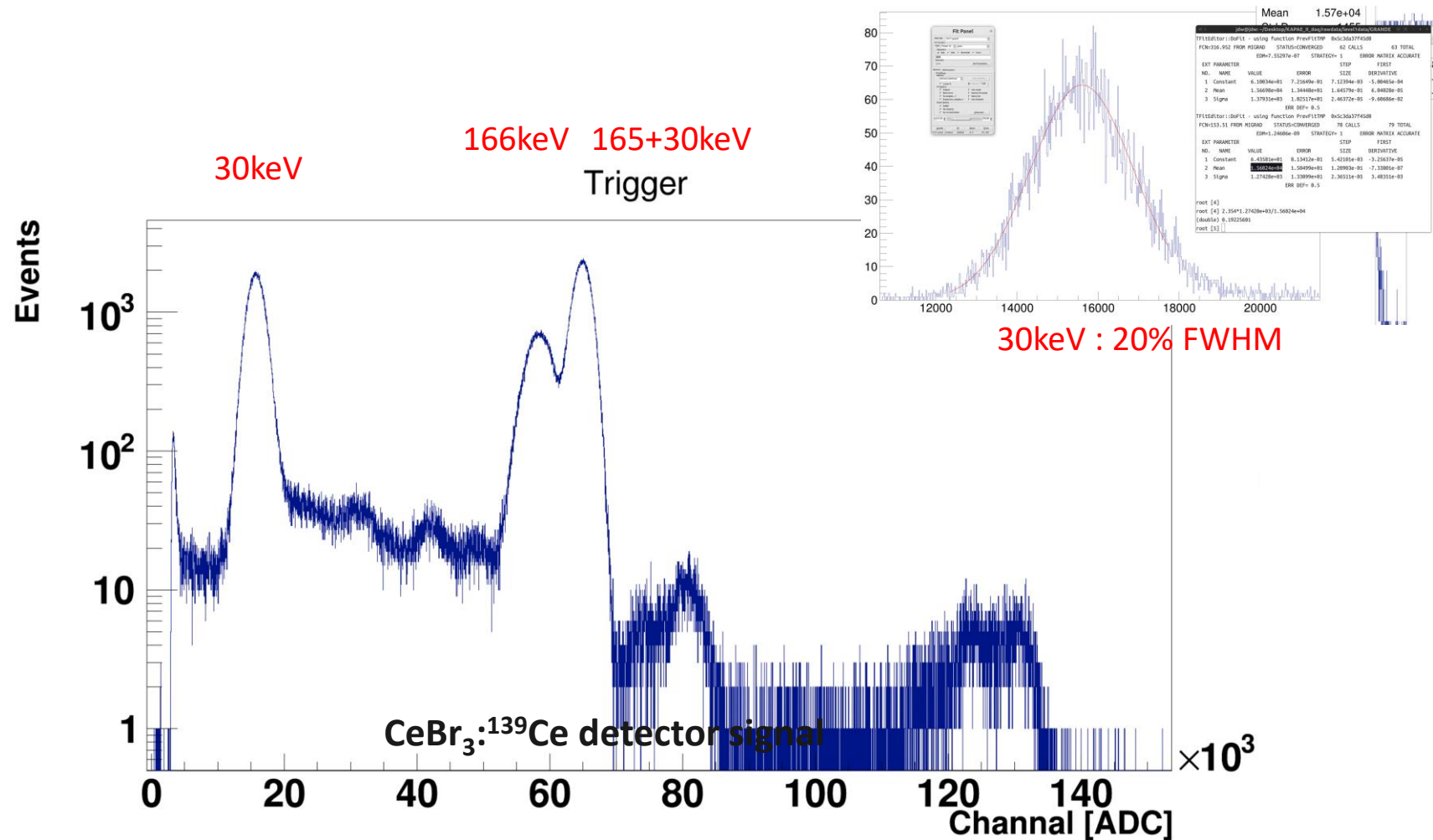
^{139}Ce experiment with KAPAE-II Setup



$\text{CeBr}_3:^{139}\text{Ce}$ Crystal with
SiPM connection

^{139}Ce experiment with KAPAE-II Setup : CeBr_3 signal

Data just started Oct 17, 2025 and will be continue for 6 month

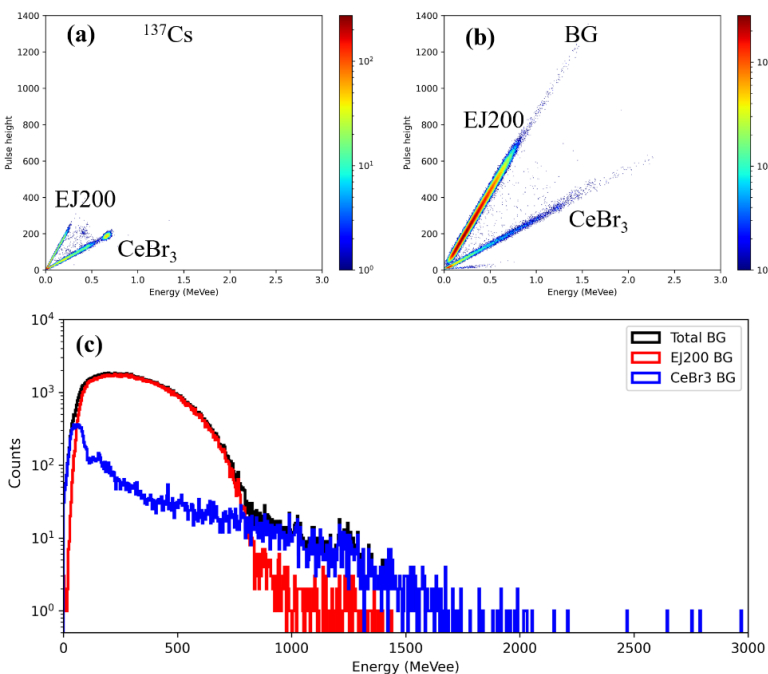
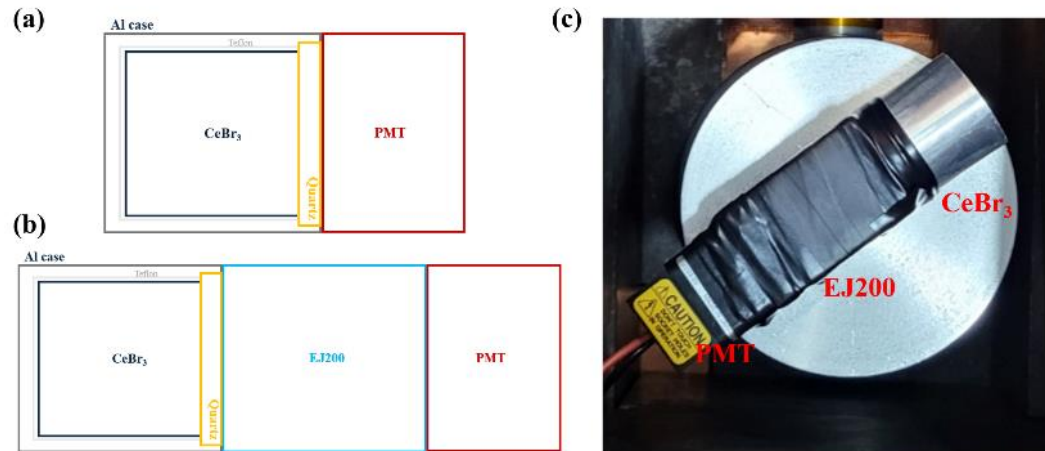


Summary and Prospect

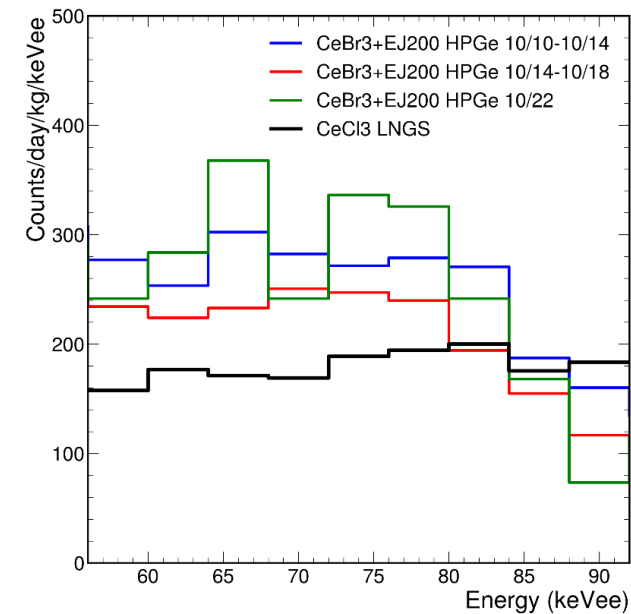
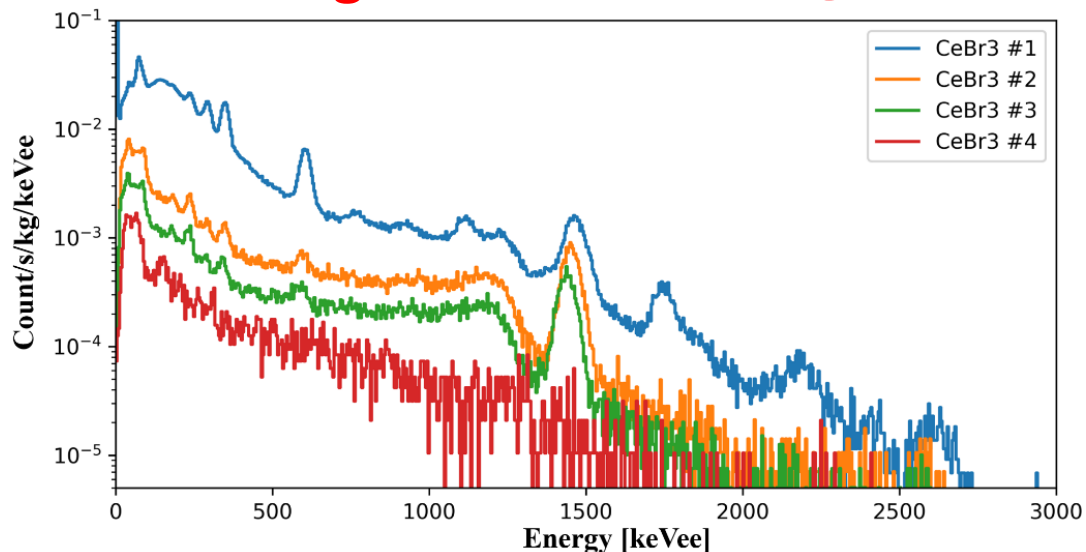
- ❑ Radioactive sources are good for the Tabletop experiments with low budget.
- ❑ Search for the exotic dark matter search and extremely rare nuclear EC transition measurement will be performed with GRANDE.
- ❑ Source-Detector technique is used with CeBr_3 doped with radio-active source.
- ❑ To reduced external radio-active background, 4p BGO veto is used and experiment is on going at Yemi underground laboratory in Cu and Pb shielding.
- ❑ Missing gamma study with ^{139}Ce , ^{125}I and ^{44}Ti of EC M1/E1/E2 transition, exotic dark matter search of ALP, dark photon and anapole will be performed in GRANDE.
- ❑ Are there any new theoretical motivation at $E < 1\text{MeV}$ for new particle searches related with invisible particles such as new bosonic particle?

Thank You !

Pure 1" CeBr₃ internal background @ Yemi



Good background condition @ Yemi



PLAN

Current

