

# On the Origins of Varying Gauge Couplings

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# Motivation

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- \* Several constructions utilize the following effective Lagrangian:

$$\mathcal{L} \supset -\frac{1}{2} \left( \frac{1}{g^2} - \frac{\phi}{\Lambda} \right) \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$$

- \* **Varying fine-structure** ([Lee, et. al., '22](#), [Chluba & Hart, '23](#), [Ghosh, et. al., '25](#)), **electroweak confinement** ([Berger, et. al., '19](#), [Howard, et. al., '21](#), [Bhalla-Ladd, et. al., '25](#)), **ultralight scalars** ([Brzeminski, et. al., '20](#), [Banerjee, et. al., '22](#), [Becker, et. al., '25](#)), **early QCD confinement** ([Ipek & Tait, '18](#), [Lu, et. al., '22](#))...
- \* Before BBN/EWPT little is known, dynamics may have differed drastically from standard cosmological evolution
- \* Gauge couplings drive early-universe dynamics and since  $g \rightarrow g(\mu)$ , **different cosmological epochs map to different energy scales**
- \* RGE-driven evolution depends only on heavy thresholds, only **slow (logarithmic!)** variations in energy are possible

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- \* Before BBN,  $\phi$  is constant,  $g$  is constant, standard cosmological evolution
- \* Gauge couplings change at low energies?  $g \rightarrow g(\mu)$ , **different cosmological epochs map to different energy scales**
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- \* Drastic changes of couplings implies  $\langle \phi \rangle \sim \Lambda$ ... e.g. confined EW requires  $\frac{1}{g^2} \approx \frac{\langle \phi \rangle}{\Lambda} \sim 2.37$

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# Varying Gauge Coupling Induced by Singlet Scalar...

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- \* The EFT description breaks down in this regime!

$$\mathcal{L} \supset -\frac{1}{2} \left[ \frac{1}{g^2} - \sum_n c_n \left( \frac{\phi}{\Lambda} \right)^n \right] \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$$

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- \* Simplest mechanism: matter **charged under the gauge group** whose mass tracks a gauge-singlet scalar  $\phi$

$$\mathcal{L} \supset - (M + y\phi) \bar{\psi} \psi \implies m(\phi) = |M + y\phi|$$

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- \* As  $\phi$  shifts,  $m(\phi)$  moves the RG threshold where  $\psi$  enters the  $\beta(\mu)$ , inducing a field-dependent low-energy coupling

# Outline

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1. The Higgs-like (4D) charged matter UV-completion
2. The restriction of gauge invariance
3. A caveat in 4D...
4. A “natural” alternative
5. Phenomenological implications
6. The bottom line

# The Charged Matter UV-Completion

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\* Leading term in the **one-loop** effective action (Shifman, et. al., '79):  $\mathcal{L} \supset -\frac{1}{2g^2(\mu)} \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$

\*  $g(\mu)$  determined by RGE, well known solution:  $\frac{1}{g^2(\mu)} = \frac{1}{g^2(\Lambda)} + 2\tilde{b} \ln \left( \frac{\mu}{\Lambda} \right)$

\* The slope is fixed by **all light charged degrees of freedom** present at that scale:

$$\tilde{b} = \frac{1}{4\pi^2} \left[ \frac{11}{3} C_2(G) - \frac{2}{3} \sum_f S_2(r_f) - \frac{1}{3} \sum_s S_2(r_s) \right]$$

\* Distinct coefficients  $\tilde{b}_>$  ( $m(\phi) < \mu$ ) and  $\tilde{b}_<$  ( $m(\phi) > \mu$ ),  $g(\mu)$  **continuous** across thresholds at leading order

\*  **$\phi$  dependence** of the gauge coupling arises when  $\mu < m(\phi)$ ... **Running and matching gives**

$$\frac{1}{g^2(\mu, \phi)} = \frac{1}{g^2(\Lambda)} + 2\tilde{b}_> \ln \left( \frac{\mu}{\Lambda} \right) + 2\tilde{b}_\Delta \ln \left[ \frac{\mu}{m(\phi)} \right] \Theta [m(\phi) - \mu]$$

# Logarithmic, Not Linear

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\* Re-expressed via the coupling measured today,  $g(\mu_0, \phi_0)$ :

$$\frac{1}{g^2(\mu, \phi)} = \frac{1}{g^2(\mu_0, \phi_0)} + 2\tilde{b}_> \ln\left(\frac{\mu}{\mu_0}\right) + 2\tilde{b}_\Delta \left\{ \ln\left[\frac{\mu}{m(\phi)}\right] \Theta[m(\phi) - \mu] - \ln\left[\frac{\mu_0}{m(\phi_0)}\right] \Theta[m(\phi_0) - \mu_0] \right\}$$

$$\text{assuming } \mu_0 < m(\phi_0) \implies \mathcal{L}_{\text{eff}} = -\frac{1}{2} \left[ \frac{1}{g^2(\mu_0, \phi_0)} + \tilde{b}_< \ln\left(\frac{\mu}{\mu_0}\right) - 2\tilde{b}_\Delta \ln\left(\frac{|M + y\phi|}{m(\phi_0)}\right) \right] \text{Tr}\left(F_{\mu\nu}F^{\mu\nu}\right)$$

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- \* The dependence on the scalar is **logarithmic, not linear**, re-summing the EFT tower keeps it valid for  $|\Delta m|/m \gtrsim 1$
- \* Applies for any **charged matter** whose mass depends on  $\phi$ ; **multiple fields summed!**
- \* Small excursions recover the dimension-5 form, but the variation is **far slower than linear**

$$\mathcal{L} \supset -\frac{1}{2} \left[ \frac{1}{g^2(\mu, \phi_0)} - 2\tilde{b}_\Delta \frac{1}{m} \frac{dm}{d\phi} \right]_0 (\phi - \phi_0) + \dots \text{Tr}\left(F_{\mu\nu}F^{\mu\nu}\right)$$

# Couplings in the Early Universe: Higgs-Like

- \* Toy abelian model:  $N_f$  vector-like fermions +  $\phi$ ; a **second-order transition** profile

$$V(\phi, T) = \frac{\lambda}{2} (\phi^2 - \langle \phi \rangle^2)^2 + f(\lambda, y, \dots) \phi^2 T^2$$

$$\Rightarrow \langle \phi(T) \rangle = \Theta(T_c - T) \langle \phi \rangle \sqrt{1 - \left( \frac{T}{T_c} \right)^2} \quad m(\phi) = |M \pm y \langle \phi(T) \rangle|$$

- \* The probed energy scale **is the temperature** ( $\mu \rightarrow T$ ) of the bath; as the universe cools, thresholds shift
- \* With  $N_f = 100$ , many distinct UV theories flow to the **same IR coupling**
- \* The naïve linear EFT drastically **overestimates** the achievable variation.

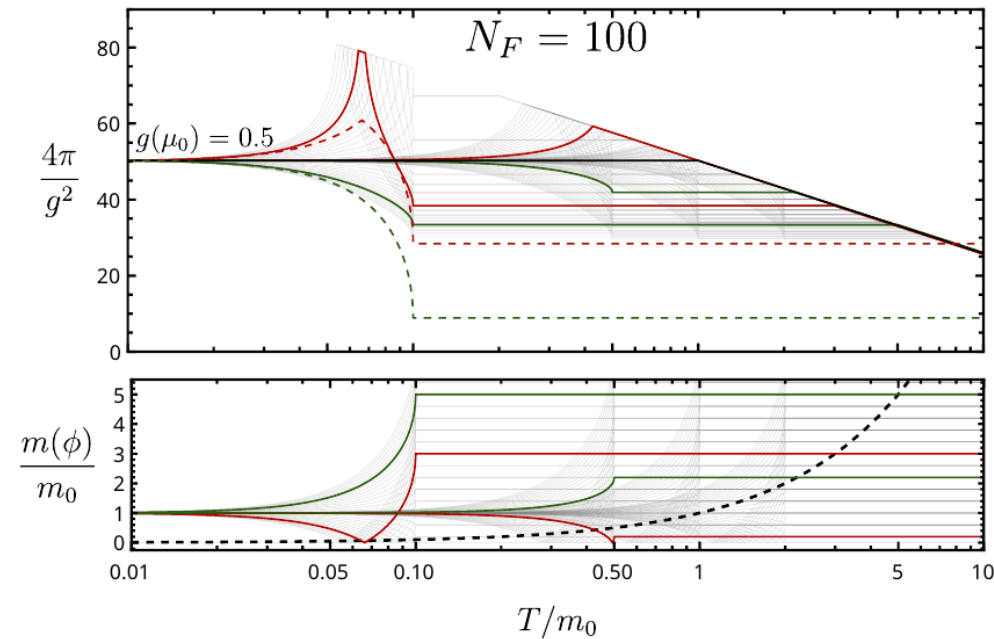


Fig. 1 Coupling fixed in the IR; dashed = naïve linear

# Symmetries on the Scalar

- \* When  $\phi$  transforms non-trivially under a symmetry, dependence of the effective coupling can be modified
- \* Consider theory with  $\phi$  and  $N_f$  fermions  $\psi_n$ ,  $n = 0, 1, 2, \dots, N_f - 1$  and invariant under  $\mathbb{Z}_{N_f}$  symmetry  
 $\phi \rightarrow e^{2i\pi/N_f} \phi$ ,  $\psi_n \rightarrow e^{-2i\pi n/N_f} \psi_n$

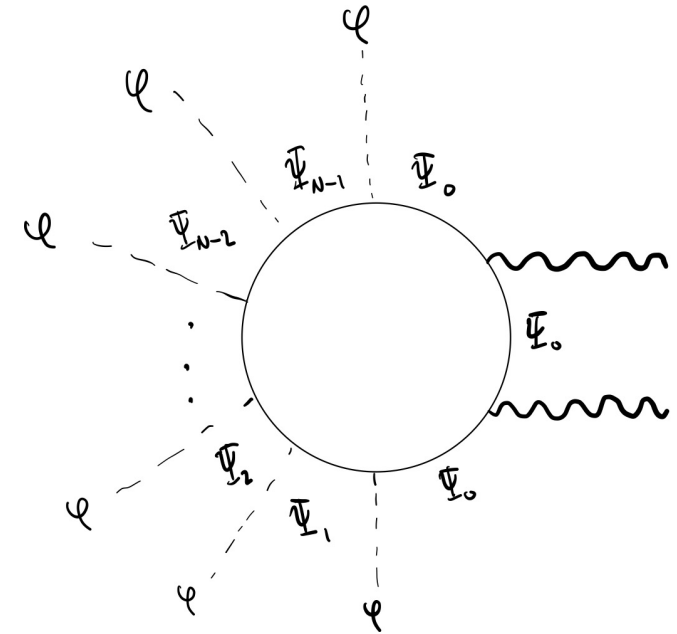
$$\mathcal{L} = - \sum_{k=0}^{N_f-1} \left[ M_k \bar{\psi}_k \psi_k + y_k \phi \bar{\psi}_k \psi_{k+1} + Y \bar{\psi}_0 \phi \psi_{N_f} + \text{h.c.} \right]$$

$$\Rightarrow \frac{1}{g^2(\mu, \phi)} = \frac{1}{g^2(\Lambda)} + 2\tilde{b}_> \ln \left( \frac{\mu}{\Lambda} \right) + 2\tilde{b}_\Delta \ln \left[ \frac{\mu^{N_f}}{|\det m(\phi)|} \right]$$

- \* The symmetry is encoded in  $|\det m(\phi)| \dots$  e.g.  $N_f = 3$ :

$$\det m(\phi) = M_0 M_1 M_2 - M_0 |y_1 \phi|^2 - M_1 |y_2 \phi|^2 - M_2 |y_0 \phi|^2 + (y_0 y_1 Y \phi^3 + \text{h.c.})$$

- \*  $\phi^3$  term  $\propto Y \dots$   $Y \rightarrow 0$  restores a larger  $U(1)$  symmetry.



# Enhanced Symmetry $\rightarrow$ Higher Operators

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\* Generalizing...  $\det m(\phi) \supset \prod_i M_i |y_{i+1}\phi|^2 + Y\phi^{N_f}$

\* Expand  $\phi = \phi_0 + \Delta\phi$ ...

$$\ln |\det m(\phi)| = \ln |\det m(\phi_0)| + \text{Re} \left[ \text{Tr} \left( m^{-1} \frac{dm}{d\phi} \right)_{\phi_0} \right] \Delta\phi + \frac{1}{2} \text{Re} \left[ \text{Tr} \left( m^{-1} \frac{d^2 m}{d\phi^2} \right)_{\phi_0} - \text{Tr} \left( m^{-1} \frac{dm}{d\phi} \right)_{\phi_0}^2 \right] (\Delta\phi)^2$$

\* Generic point: the trace is non-zero  $\rightarrow$  leading dependence is **linear** (a dimension-5 operator).

\* **Enhanced-symmetry point** (e.g.  $\phi_0 = 0$ ): leading dependence is **quadratic or higher** (dimension-6+).

\* **Qualitatively different** low-energy EFT... significant **phenomenological implications** ([Brzeminski, et. al., '20](#), [Banerjee, et. al., '22](#), [Becker, et. al., '25](#), [Ghosh, et. al., '25](#))

# The Restriction of Gauge Invariance

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# Gauge Invariance in 4D ensures logarithmic variation

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- \* Gauge invariance forbids mass/kinetic mixing of gauge bosons with neutral fields (exception: Abelian kinetic mixing)
- \* Up to dimension four, gauge bosons couple **only to charged matter**, which feeds back at one loop
- \* To change the coupling we must **change charged matter...** only masses and non-gauge couplings may vary
- \* At one loop, **only masses** reach the gauge running
- \* **How does a moving mass threshold enter the coupling?**

# The Logarithm Is Generic

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- \*  $G_{>} \supseteq G_{<}$  matched at massive threshold  $m$
- \* Spontaneous breaking of scale invariance by  $\phi \rightarrow e^\lambda \phi$
- \* Above  $m$  have renormalizable theory, **no  $\phi$  dependence**; scale anomaly coefficient at one-loop:  $\partial_\mu D^\mu = \tilde{b}_{>} \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$
- \* Below  $m(\phi)$  expand in powers of  $g$ :

$$\mathcal{L}_{\text{eff}} \supset -\frac{1}{2} \left[ \frac{1}{g^2(\mu)} + f(\phi) + \mathcal{O}(g^2) \right] \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$$

- \* Scale anomaly coefficient **below threshold**:

$$\partial_\mu D^\mu = \left( \tilde{b}_{<} - \frac{1}{2} \frac{df}{d \ln \phi} \right) \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$$

- \* **Matching** the light-field anomaly across the threshold **fixes  $f(\phi)$** :

$$\frac{df}{d(\ln \phi)} = -2(\tilde{b}_{<} - \tilde{b}_{>}) = -2\tilde{b}_\Delta \implies f(\phi) = \text{constant} - 2\tilde{b}_\Delta \ln \phi$$

# The Logarithm Is Generic

- \*  $G_{>} \supseteq G_{<}$  matched at massive threshold  $m$ ... arises from scale-invariant interactions
- \* Spontaneous breaking of scale invariance by  $\phi \rightarrow e^\lambda \phi$
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- \* Below  $m(\phi)$  expand in powers of  $g$ :

$$\mathcal{L}_{\text{eff}} \supset -\frac{1}{2} \left[ \frac{1}{\Lambda^2} \text{Tr} (F_{\mu\nu} F^{\mu\nu}) + \dots \right]$$

There are some caveats!

- \* Scale anomaly coefficient **below threshold**:

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# A Caveat in 4D

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# Yukawas Vary at Tree Level

- \* Yukawas are **not protected by gauge invariance**

- \* Froggatt–Nielsen-style UV completion ([Froggatt & Nielsen, '79](#), [Leurer, et. al., '92](#), [Leurer, et. al., '93](#)):

$$-\mathcal{L} \supset \xi^c \Phi (y\eta + Y\psi) + (m + f\phi)\psi^c \eta + (M + F\phi)\psi^c \psi$$

|         | $\eta$       | $\xi^c$            | $\psi$       | $\psi^c$     | $\Phi$       | $\phi$       |
|---------|--------------|--------------------|--------------|--------------|--------------|--------------|
| $SU(N)$ | $\mathbb{1}$ | $\bar{\mathbf{N}}$ | $\mathbb{1}$ | $\mathbb{1}$ | $\mathbf{N}$ | $\mathbb{1}$ |
| $U(1)$  | 1            | -1                 | 1            | -1           | 0            | 0            |

- \* Integrate out  $\psi$  and **re-sum  $\phi$  insertions**:

$$-\mathcal{L}_{\text{eff}} \supset \xi^c \Phi \eta \left( y + \frac{Yf\phi}{M + F\phi} \right) \equiv y(\phi) \xi^c \Phi \eta \quad y(\phi) = y + \frac{fY\phi}{M + F\phi}$$

- \* A genuine **linear** dependence (EFT restructuring once  $F\phi \gtrsim M$ )

# Faster-than-Linear Yukawas

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- \* Add a second heavy pair protected by a  $\mathbb{Z}_2$ ; the leading insertion is now  $\phi^2$ :

$$-\mathcal{L}_{\text{eff}} \supset \xi^c \Phi \eta \left( y + \frac{Y f F_{12} \phi^2}{M_1 M_2 - F_{12} F_{21} \phi^2} \right) \equiv y(\phi) \xi^c \Phi \eta \quad y(\phi) = y + \frac{Y f F_{12} \phi^2}{M_1 M_2 - F_{12} F_{21} \phi^2}$$

- \* Larger discrete symmetries push the leading power higher, giving

$$y(\phi) \simeq Y \left[ \frac{\phi}{\Lambda(\phi)} \right]^p$$

- \* Exponent **controlled by symmetry...** power-law sensitivity, not from any logarithm!

# Two-Loop Caveat: Yukawa Couplings Augment the Running

- \* At two loops the Yukawa enters the gauge  $\beta$ -function:

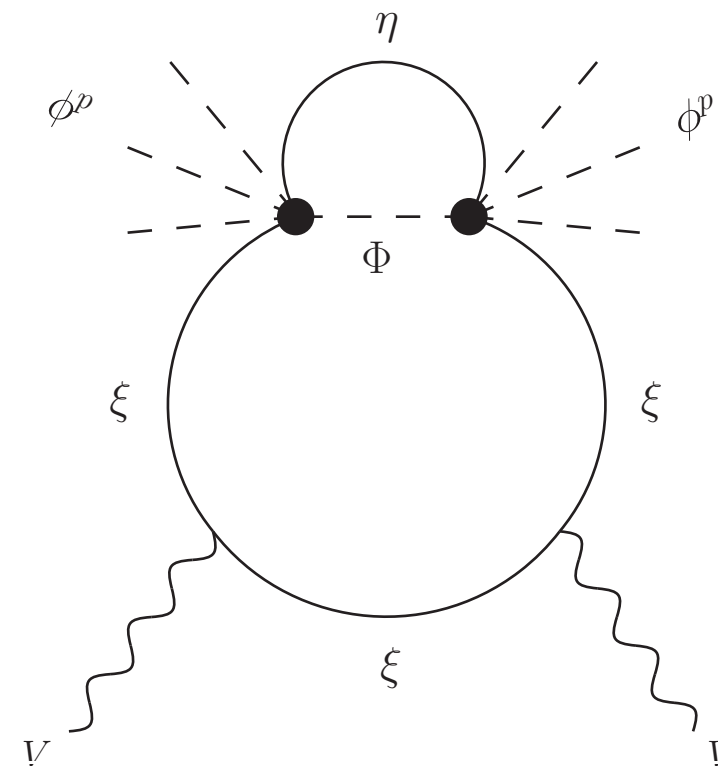
$$\frac{dg}{dt} \simeq -\frac{b_1}{(4\pi)^2} \left[ 1 + \frac{b_2^{(y)}}{b_1(4\pi)^2} y^2 \right] g^3$$

- \* Integrating (with Yukawa as effective  $\beta$  coefficient):

$$\frac{1}{g^2(\mu)} = \frac{1}{g^2(\Lambda)} + \frac{2b_1}{(4\pi)^2} \left[ 1 + \frac{b_2^{(y)}}{b_1(4\pi)^2} y^2 \right] \ln \left( \frac{\mu}{\Lambda} \right)$$

- \* With  $y \rightarrow y(\phi) \simeq Y \left[ \frac{\phi}{\Lambda(\phi)} \right]^p$ , the gauge coupling inherits a **power-law** piece:

$$\left( \frac{1}{g^2} \right)_{2\text{-loop}} \propto y^2(\phi) \ln \left( \frac{\mu}{\Lambda} \right) \sim Y^2 \left[ \frac{\phi}{\Lambda(\phi)} \right]^{2p} \ln \left( \frac{\mu}{\Lambda} \right)$$



# A “Natural” Alternative

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\* In 5D  $[g] = M^{-1}$ : possible to have stronger variation

\* Kaluza-Klein (KK) decomposition provides natural setting for large multiplicity of charged states

\* Consider the 5D theory with  $z \in \{0, R\}$ :  $\mathcal{L}_5 = -\frac{1}{4}F_{MN}F^{MN} + \frac{1}{2}\bar{\Psi}i\Gamma^M D_M \Psi - M\bar{\Psi}\Psi - \frac{y}{\sqrt{R}}\Phi\bar{\Psi}\Psi + \frac{1}{2}|\partial\Phi|^2 - V(\Phi)$

\* Potential induces  $\Phi = \sqrt{R}\phi$ ; take  $\Psi_L|_0 = 0$ ,  $\Psi_R|_R = 0$

\* Zero-mode forbidden:  $m_n^2(\phi) = (M + y\phi)^2 + \left(\frac{\omega_n}{R}\right)^2$ ,  $\tan(\omega_n) = -\frac{\omega_n}{|M + y\phi|R}$

\* For  $\mu_0 < m_1$ :  $\frac{1}{g^2(\mu, \phi)} = \frac{1}{g^2(\mu_0, \phi_0)} + \tilde{b}_\Psi \sum_n^\infty \ln\left(\frac{\mu^2}{m_n^2}\right) \Theta(\mu - m_n)$ ,  $\tilde{b}_\Psi = -\frac{4Q^2}{3(4\pi)^2}$

\* Running scale  $\mu < \Lambda \simeq \frac{4\pi^2}{g^2 R} \implies n \lesssim \frac{4\pi}{g^2}$

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\*  $\mu \gg \frac{\pi}{R}, |M + y\phi| \implies N \simeq \frac{\mu R}{\pi}$ : 
$$\frac{1}{g^2(\mu, \phi)} = \frac{1}{g^2(\mu_0, \phi)} + 2b_\Psi \left[ \ln\left(\frac{\mu R}{N\pi}\right) + \ln\left(\frac{N^N}{N!}\right) - \sum_{n=1}^N \ln\left(\frac{m_n R}{n\pi}\right) \right]$$
$$\simeq \frac{1}{g^2(\mu_0, \phi)} - 2b_\Psi \left[ \left(\frac{\mu R}{\pi}\right) + \sum_{n=1}^N \ln\left(\frac{m_n R}{n\pi}\right) \right]$$

\*  $\phi$  VEV does not affect UV gauge coupling, **fractional changes in KK masses decrease for increasing KK levels**

\* Matching gives IR relation: 
$$\frac{1}{g^2(\mu, \phi)} \simeq \frac{1}{g^2(\mu_0, \phi_0)} - b_\Psi \sum_{n=1}^{\infty} \ln\left(\frac{m_n^2(\phi)}{m_n^2(\phi_0)}\right)$$

# Couplings in the Early Universe: Flat Extra Dimension

- \* Toy bulk abelian model generates  $\Phi \rightarrow \sqrt{R}\phi$  with same **second-order transition** profile:

$$\langle \phi(T) \rangle = \Theta(T_c - T) \langle \phi \rangle \sqrt{1 - \left( \frac{T}{T_c} \right)^2}$$

$$m_n^2(\phi) = (M \pm y \langle \phi(T) \rangle)^2 + \left( \frac{\omega_n}{R} \right)^2$$

- \* Again take **temperature** ( $\mu \rightarrow T$ ) of the bath; as the universe cools, thresholds shift
- \* Sensitive to number of states present at low-energy, **key difference: if more modes stack together and allow for faster changes**
- \* Increasing  $R$  allows higher multiplicity of modes... **edge of reliability of KK description**

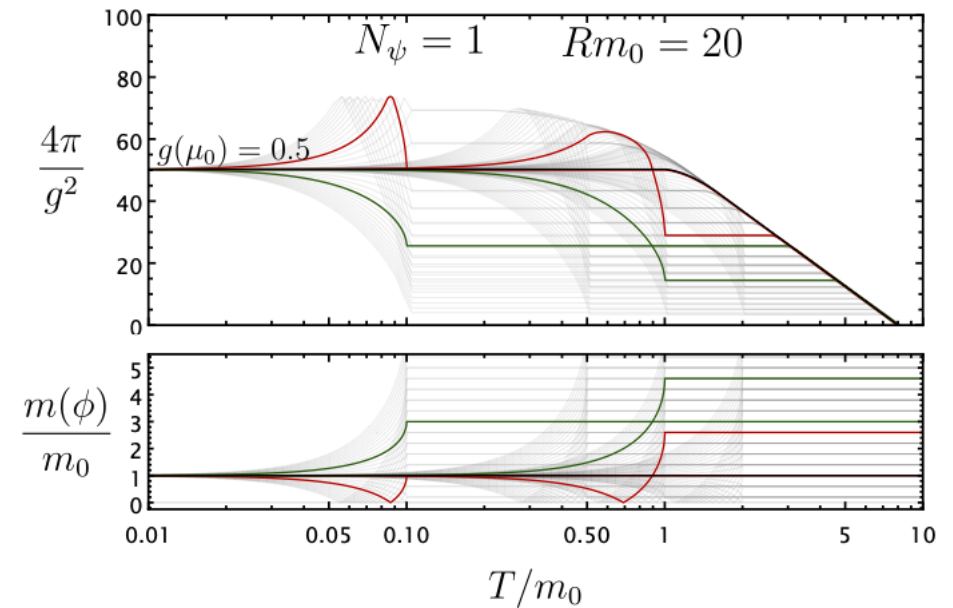


Fig. 1 Coupling fixed in the IR

# Phenomenological Implications (mostly) in 4D

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# Phenomenology of New Charged Matter

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- \* New, stable charged matter very constrained, cosmologically
- \* Can decay **directly** or **indirectly** to SM
  - \* Requires non-trivial electroweak charge  $\implies$  changing one gauge coupling changes others
  - \* **Cascade decays: e.g. electroweak multiplet dark matter** ([Cirelli et. al. Nucl. Phys. B 753 \(2006\) 178](#))
- \* Additional collider constraints on electroweak and color charged matter
  - \*  $\tau \gtrsim 1\mu\text{s}$  requires  $m \gtrsim 2\text{ TeV}$  (color) and  $m \gtrsim 400 - 1200\text{ GeV}$  (electrically charged) ([CMS Collaboration, JHEP 04 \(2025\) 109](#), [ATLAS Collaboration, JHEP 7 \(2025\) 140](#))
  - \* More freedom with cascade decays,  $m \gtrsim 140\text{ GeV}$  (Higgsino-like) and  $m \gtrsim 1000\text{ GeV}$  (color) ([ATLAS Collaboration JHEP 06 \(2026\) 094](#), [CMS Collaboration, arXiv:hep-ph/2604.25604](#))

# Dynamical Gauge Variations via Singlet (Thermalized)

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- \* 4D UV completion of EFT with  $\phi \rightarrow \phi_0$  coupling to new matter charged under  $SU(3)_C$  or  $SU(2)_L$
- \* For  $\mu < m(\phi)$ , change in the coupling relative to the running SM value  $\alpha_i(\mu, \phi_0) = g_i^2/4\pi$

$$\Delta\alpha_i(\mu, \phi) \simeq \frac{b_\Delta}{2\pi} \alpha_i^2(\mu, \phi_0) \ln \left( \frac{m(\phi)}{m(\phi_0)} \right), \quad b_\Delta = (4\pi)^2 \tilde{b}_\Delta = \frac{2N_\psi}{3}$$

- \* Baryogenesis example ([S.A.R Ellis, et. al., JHEP8 \(2019\) 002](#)):  $\Delta\alpha_2 > 0$  or  $\Delta\alpha_3 < 0$  at  $T \sim 100$  GeV could enable EWBG
- \* Generating  $Y_B$  with  $\Delta\alpha_3 = 0$  requires  $\Delta\alpha_2 \gtrsim 0.01 \implies b_\Delta \ln(m(\phi)/m(\phi_0)) \gtrsim 55$  or  $N_\psi \gtrsim 83$  Dirac fundamentals for  $\ln(m/m_0) = 1$
- \* Generating  $Y_B$  with  $\Delta\alpha_2 = 0$  requires  $\Delta\alpha_3 \lesssim -0.03 \implies b_\Delta \ln(m(\phi)/m(\phi_0)) \lesssim -16$  or  $N_\psi \gtrsim 24$  Dirac fundamentals for  $\ln(m/m_0) = -1$

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- \* Early QCD phase transition ([D. Berger, et. al. JHEP 07 \(2020\) 192](#), [S. Ipek, T.M.P Tait PRL 22 \(2019\) 112001](#),...):

$$\Lambda(\phi) \simeq \Lambda_{\text{QCD}} \left( \frac{m(\phi)}{m(\phi_0)} \right)^{b_\Delta/b_{\text{QCD}}}, \quad \Lambda_{\text{QCD}} \simeq 50 \text{ MeV}, \quad b_{\text{QCD}} = 7, \quad b_\Delta = \frac{2}{3}N_\psi$$

- \* Envision  $\Lambda(\phi) \gtrsim T_{\text{EWPT}} \sim 100 \text{ GeV}$ , requires  $\mathcal{O}(10^3)$  enhancement

$$\implies \frac{m(\phi)}{m(\phi_0)} \gtrsim 1400, \text{ for } N_\psi = 10, \quad \frac{m(\phi)}{m(\phi_0)} \gtrsim 10, \text{ for } N_\psi = 30, \quad \frac{m(\phi)}{m(\phi_0)} \gtrsim 2, \text{ for } N_\psi = 100$$

# Dynamical Gauge Variations via Singlet (Non-Thermalized)

- \* Dynamical variations can also be generated by an non-thermalized singlet scalar
- \* After an initial displacement, oscillate freely until  $H < m_\phi$
- \* Parameterizing the physical excitation as  $\eta(x) = \phi(x) - \phi$ , the coupling to photons, gluons and light quarks is parameterized as ([D. Antypas, et. al. 2203.14915](#), [D.B. Kaplan and M.B. Wise, JHEP 08 \(2000\) 037](#),...)

$$\mathcal{L} \supset \sum_{n \geq 1} \frac{1}{n} \left( \frac{\eta}{\sqrt{2} M_{\text{Pl}}} \right)^n \left[ \frac{d_e^{(n)}}{4e^2} F_{\mu\nu} F^{\mu\nu} - \frac{\beta_3 d_g^{(n)}}{2g_3^2} G_{\mu\nu}^a G^{a\mu\nu} - \sum_q \left( d_{m_q}^{(n)} + \gamma_m d_g^{(n)} \right) m_q \bar{q} q \right]$$

- \* Constrained by equivalence principle violation ([D. Antypas, et. al. 2203.14915](#)), nuclear mass changes ([A. Hees, et. al. PRL 117\(2016\) 061301](#)), from BBN/CMB ([S. Ghosh, et. al. PRD 114 \(2026\) 023525](#))

- \* Especially strong constraints on variations arising on  $d^{(1)} = \frac{\alpha b_\Delta}{2\pi} \frac{\sqrt{2} M_{\text{Pl}}}{m} \frac{dm(\phi)}{d\phi} \Big|_{\phi=\phi_0}$

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- \* With  $\phi \rightarrow -\phi$  and  $\phi_0 = 0$  can forbid  $n = 1$  with  $m^2(\phi) = M^2 \mp y^2 \phi^2$

- \* For  $n = 2$  the variation is  $\frac{\Delta\alpha}{\alpha} \simeq \mp \frac{d^{(2)} M^2}{4y^2 M_{\text{Pl}}^2} \ln \left( 1 \mp \frac{y^2 \phi^2}{M^2} \right)$

- \* Greatly relaxes constraints in this case ([T. Bouley, et. al. JHEP 03 \(2023\) 104](#), [S. Ghosh, et. al. PRD 114 \(2026\) 023525](#))

# Gauge Couplings vs. Yukawa Couplings in 4D

## GAUGE COUPLINGS

$$\mathcal{L}_{\text{eff}} \supset -\frac{1}{2} \left[ \frac{1}{g^2(\mu)} - c \ln \left( 1 + \frac{\Delta m(\phi)}{M} \right) \right] \text{Tr} \left( F_{\mu\nu} F^{\mu\nu} \right)$$

- \* Protected by gauge invariance
- \* Enter only at **one loop**, only through **mass thresholds**
- \* Sensitivity is **logarithmic** in  $\phi$

## YUKAWA COUPLINGS

$$y(\phi) \simeq Y \left[ \frac{\phi}{\Lambda(\phi)} \right]^p$$

- \* Unprotected can be modified at tree-level
- \* **Symmetry** and **mass hierarchy** set power  $p$
- \* Sensitivity can be **power-law** in  $\phi$

**Bottom line:** in UV-complete 4D theories, gauge couplings must vary logarithmically; reaching **power-law** and/or drastic variation costs an extra loop (via power-law Yukawas) or a high multiplicity of charged matter (via extra dimensions).

Thank You!



# Back-up Stuff on Dynamics of Singlet Scalar

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# Dynamics of the Singlet Scalar

---

\*  $\psi + \bar{\psi} \rightarrow \phi + A_\mu$  populates the scalar drives it toward equilibrium

\* Charged matter coupling  $\Rightarrow$  corrections to  $V_{\text{eff}}(\phi)$

\* Quantum corrections at  $T = 0$ :  $\Delta m_\phi^2 \sim \frac{d(r_\psi)N_\psi y^2}{(4\pi)^2} M^2, \Delta\lambda \sim \frac{d(r_\psi)N_\psi y^4}{(4\pi)^2}$

\* Quantum corrections at  $T \neq 0$ :  $\Delta V(\phi) \propto \begin{cases} m_\psi^2(\phi)T^2, & T \gg m(\phi) \\ g^2(\mu \sim T, \phi)T^4, & T \lesssim m(\phi) \end{cases}$

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\* When thermalized, finite  $T$  corrections  $\Rightarrow$  phase transition, changing  $\phi$  and  $g(\mu, \phi)$

\* Significant changes need **large  $\Delta m(\phi)$**  driven by  $\phi$  couplings to *lighter* fields

\* Thermalized  $\phi$  produced abundantly, **decays to  $\psi + \bar{\psi}$**  ( $m_\phi > 2m_\psi$ ) **or to vector bosons** ( $m_\phi < 2m_\psi$ )

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\* Non-thermalized scalar coherently oscillates at late times

\* In  $y \rightarrow 0$  limit with initial displacement  $\phi_i$  and  $\phi = 0$  today, trapped until  $H \sim m_\phi$

\*  $\phi$  rolls to minimum, oscillates, redshifts as matter

$$f_{\text{DM}} \simeq \left( \frac{m_\phi}{4.5 \times 10^{-17} \text{ eV}} \right)^2 \left( \frac{\bar{\phi}(t_0)}{100 \text{ GeV}} \right)^2$$