

BSM ν_τ Appearance at DUNE-PRISM

Min-Gwa Park

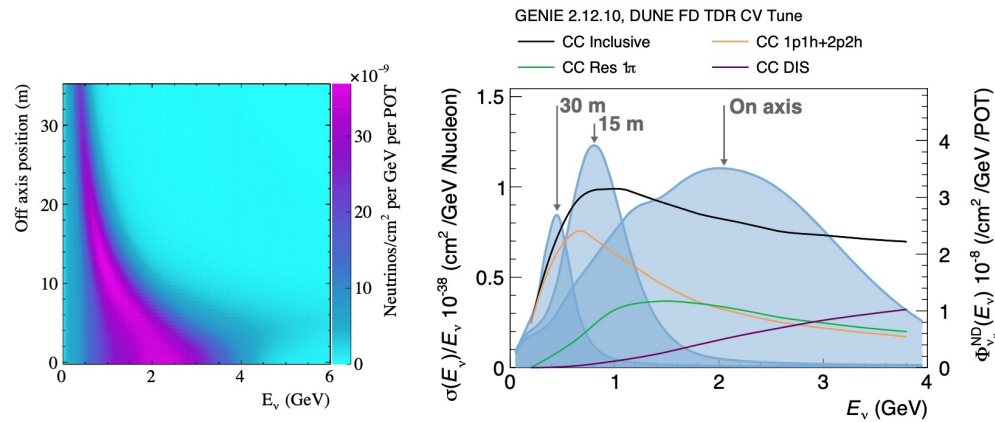
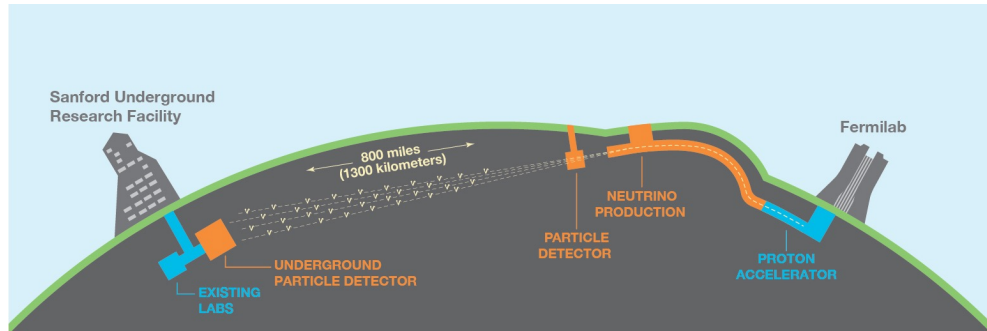
Northwestern University

Work with Doojin Kim (South Dakota U.), Seodong Shin (JBNU) [in progress]

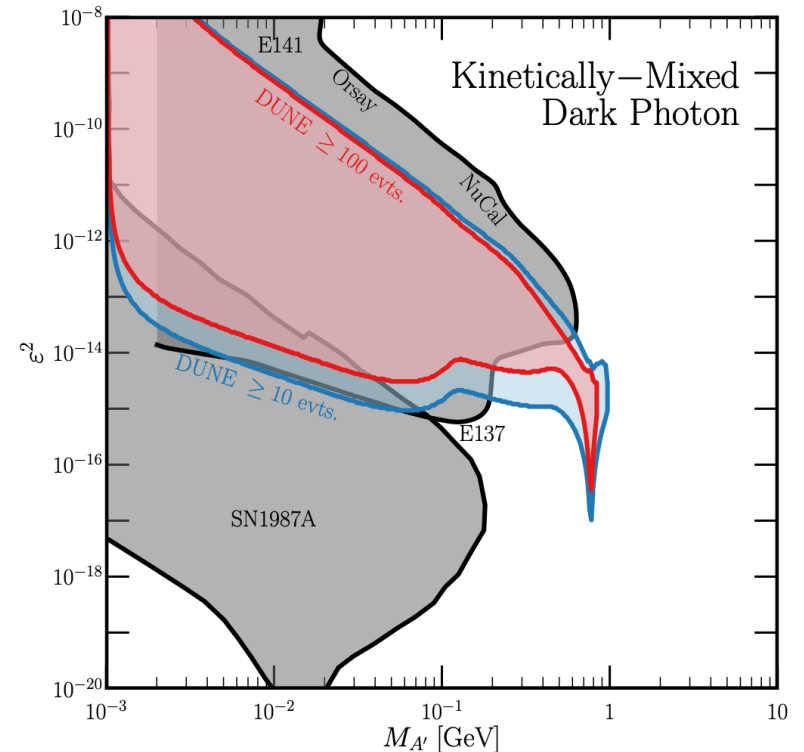
SSU 2026, Busan

Aug 26, 2026

DUNE: Accelerator Neutrino Experiment



DUNE Near Detector Conceptual Design Report

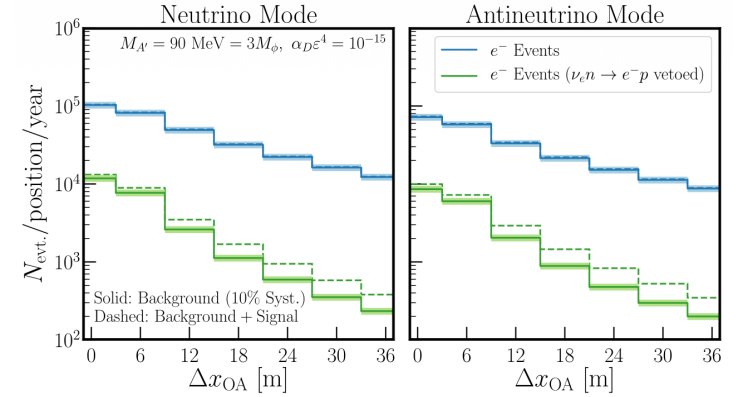
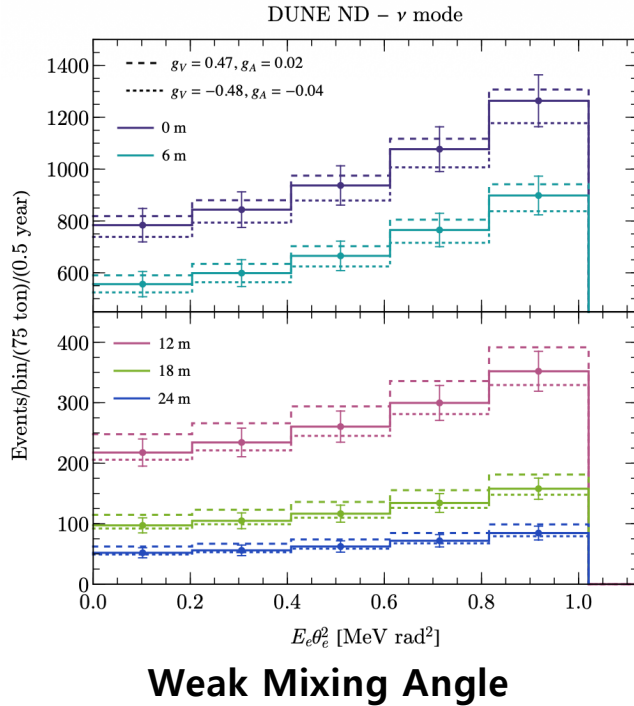


New Physics at DUNE Near Detector

Berryman, de Gouvea, Fox, Kayser, Kelly, Raaf (2020)

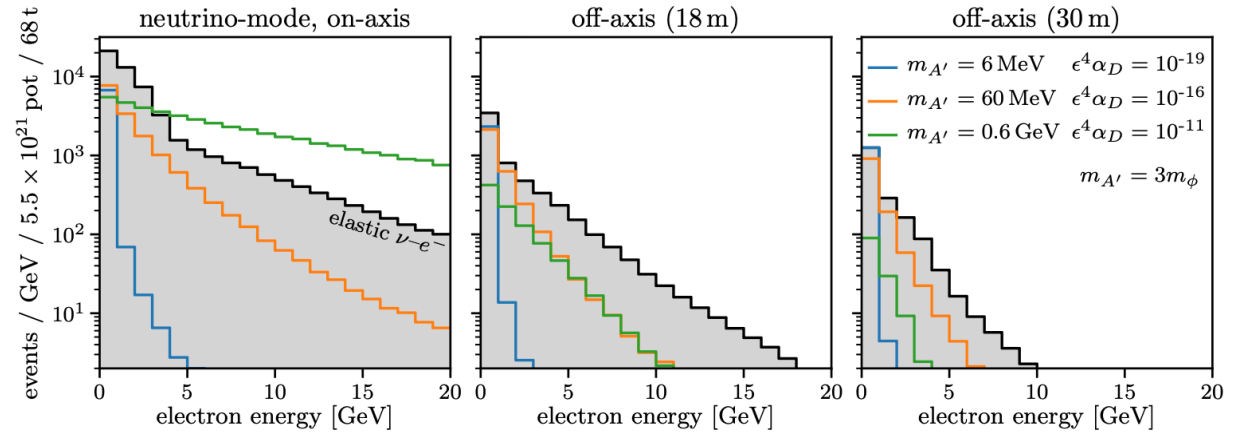
Physics Search with DUNE-PRISM

de Gouvea, Machado, Perez-Gonzales, Tabrizi (2022)



New Physics

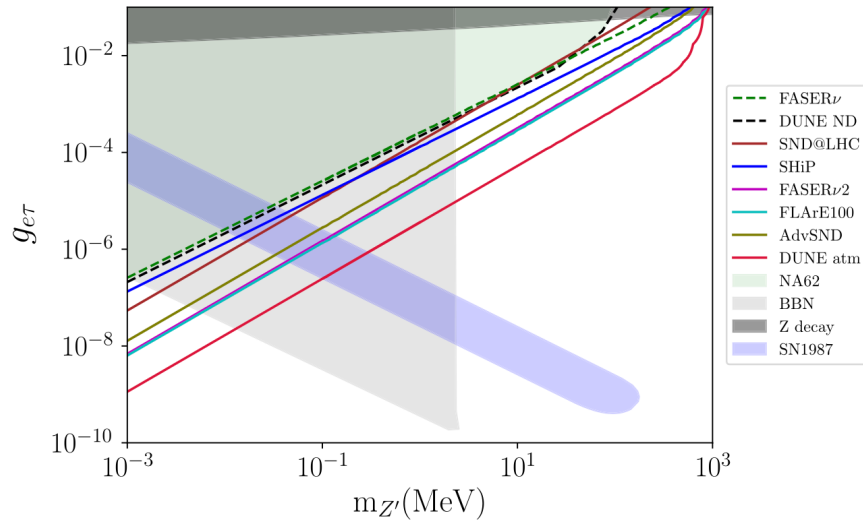
De Romeri, Kelly, Machado (2019)



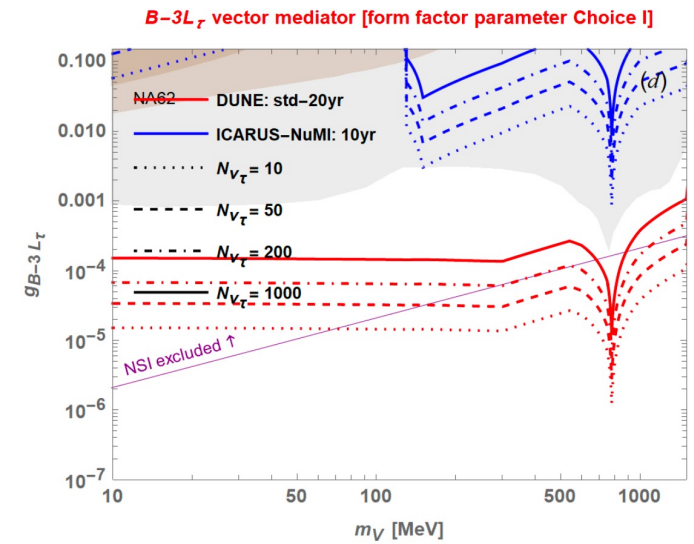
Breitbach, Buonocore, Frugieuele, Kopp, Mitnacht (2022)

Angular, Spectral information increases sensitivity

ν_τ Appearance at DUNE Near Detector



Bakhti, Rajaei, Shin (2024)

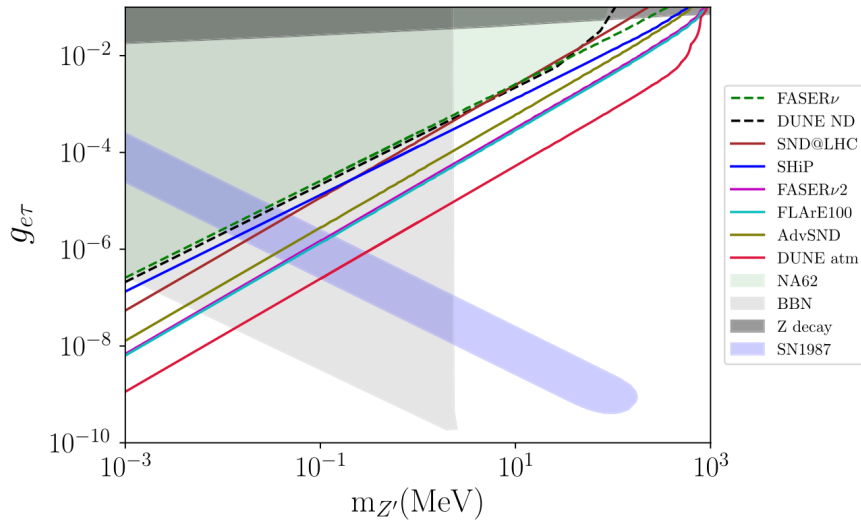


Dev, Dutta, Han, Kim (2024)

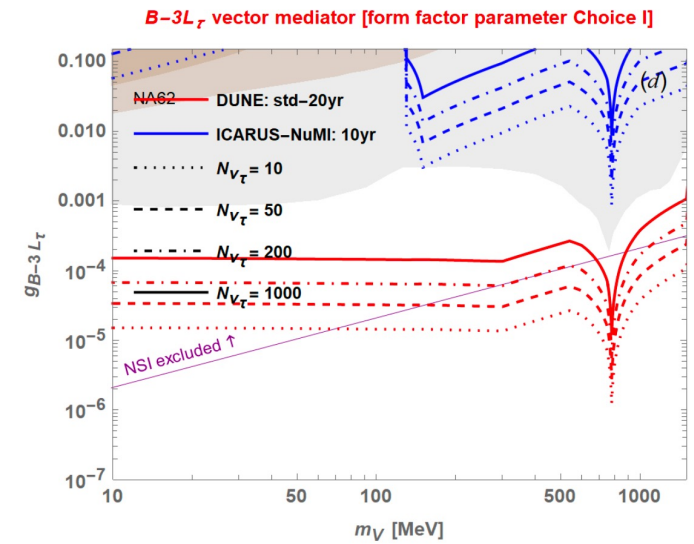
Anomalous meson decay or Proton bremsstrahlung

→ BSM ν_τ flux at DUNE Near Detector

ν_τ Appearance at DUNE Near Detector



Bakhti, Rajaei, Shin (2024)

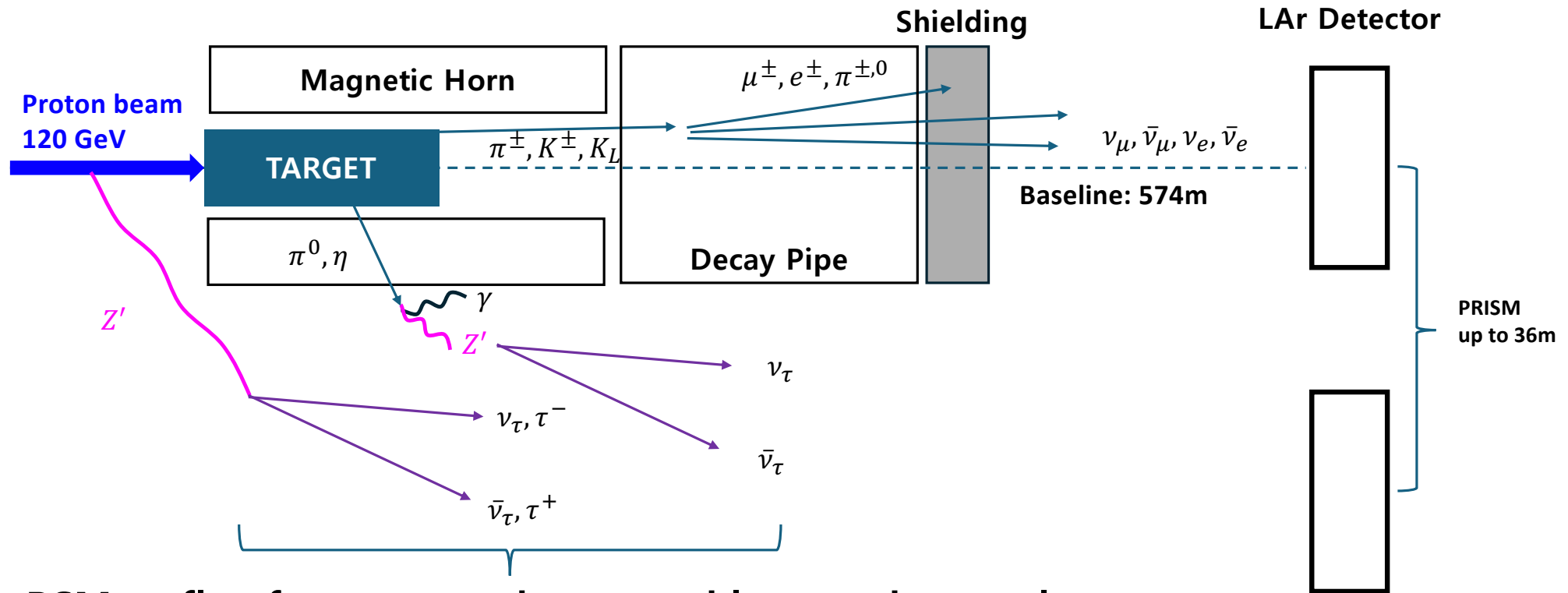


Dev, Dutta, Han, Kim (2024)

Hard & Wide neutrino spectrum from heavy Z'

→ What can we learn about BSM ν_τ flux at DUNE-PRISM?

$B - 3L_\tau$ model Search at DUNE-PRISM



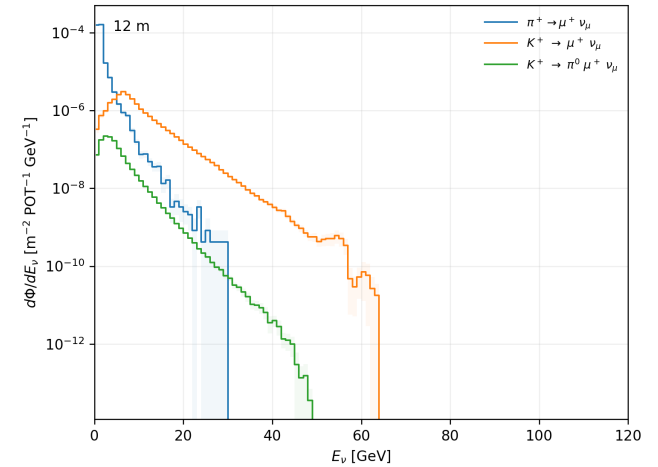
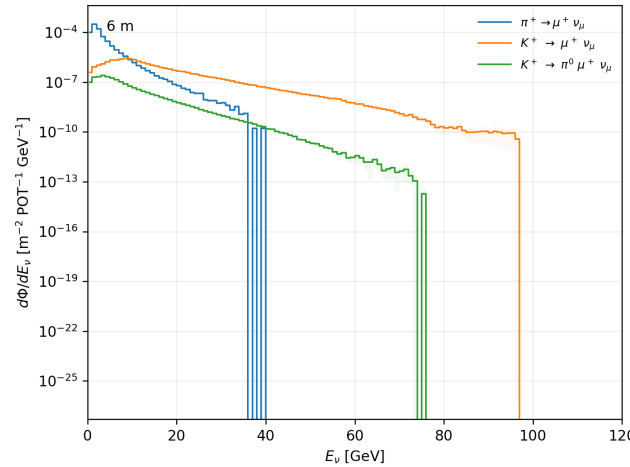
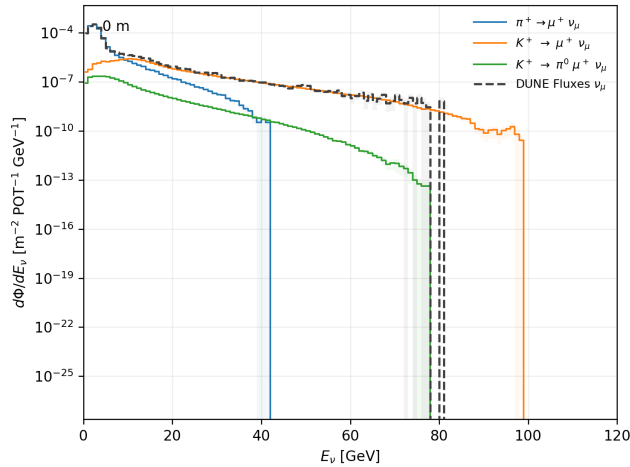
BSM ν_τ flux from proton bremsstrahlung and neutral meson

→ Not focused by the magnetic horn

DUNE Fluxes

L. Fields (<https://glaucus.crc.nd.edu/DUNEFluxes/>)

: Optimized 3-horn Design with 1.5 m target and fully engineered Horns (Sept 2021)

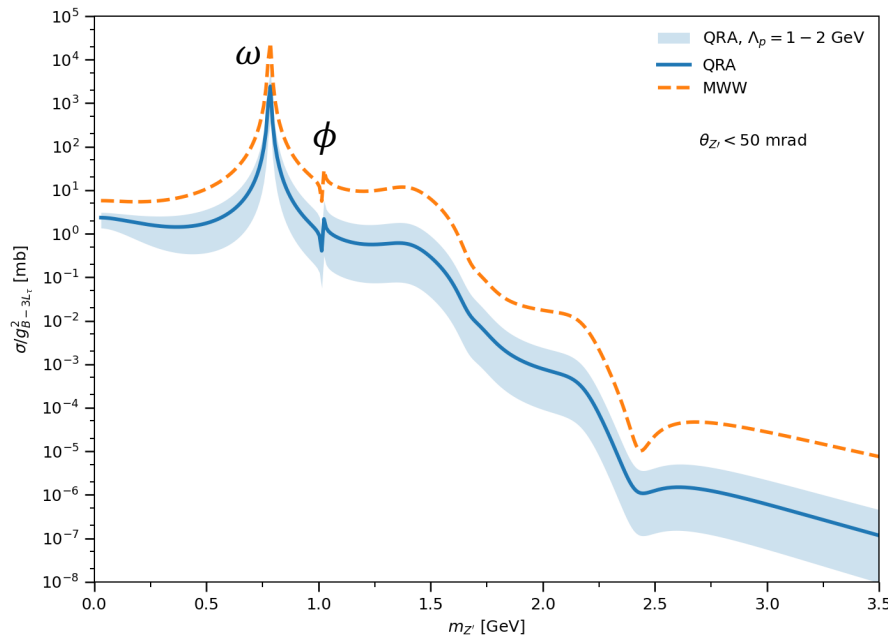


dk2nu file contains $\pi^\pm, K^\pm, K_L$ and their daughter particles (total 10⁷ POT)

Re-decayed parent mesons to get fine MC spectra
& Compared to 'DUNE Fluxes' channel by channel

D_s, π^0, η spectra separately generated with Pythia 8 (SoftQCD:all)

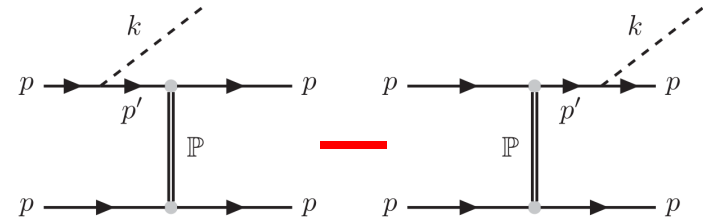
Proton Bremsstrahlung for $B - 3L_\tau$ model



Simulated following

Foroughi-Abari, Ritz (2022),
Foroughi-Abari, Reimitz, Ritz (2025),
Kling, Reimitz, Ritz (2026)

Foroughi-Abari, Ritz (2022)

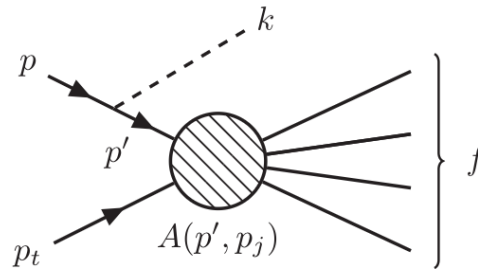


Strong **destructive** interference between QE ISR & FSR

Previously widely used

$$d\sigma_{MWW} = dP_{ISR}^{split} \times (\sigma^{EL} + \sigma^{SD} + \sigma^{NSD})$$

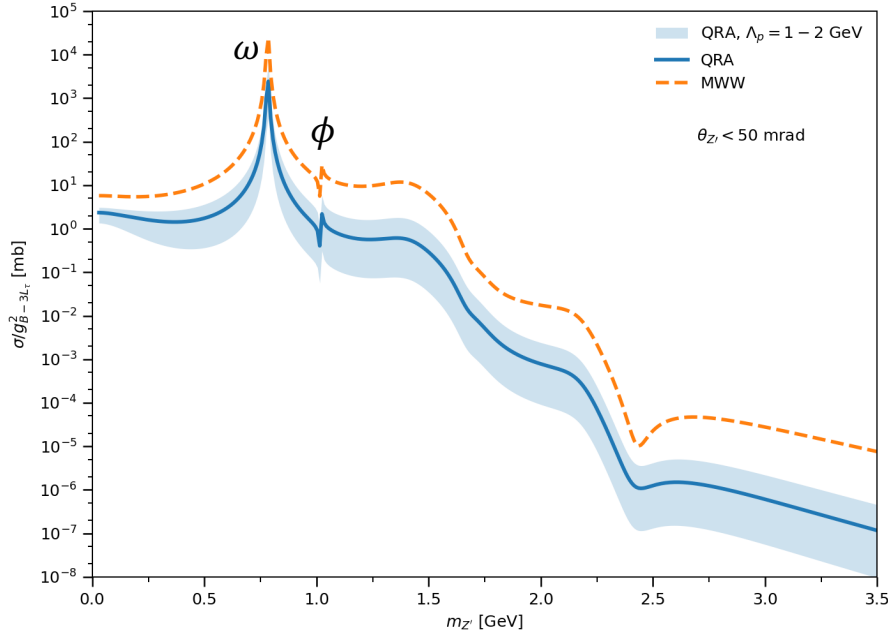
$$>> d\sigma_{QRA} = dP_{ISR}^{split} \times (\cancel{\sigma^{EL}} + \cancel{\sigma^{SD}} + \sigma^{NSD})$$



Non-Single Diffraction:

$$\mathbf{p} + \mathbf{p} \rightarrow \mathbf{X} + \mathbf{Y}$$

Proton Bremsstrahlung for $B - 3L_\tau$ model



Simulated following

Foroughi-Abari, Ritz (2022),
Foroughi-Abari, Reimitz, Ritz (2025),
Kling, Reimitz, Ritz (2026)

$$\mathcal{L} \supset g \sum_f x_f \bar{f} \gamma^\mu f Z'_\mu$$

$$x_u = x_d = x_s = \frac{1}{3}, \quad x_\tau = x_{\nu_\tau} = -3$$

$$x_\omega = 3(x_u + x_d) = 2$$

$$x_\phi = -3x_s = -1$$

$$x_\rho = x_u - x_d = 0$$

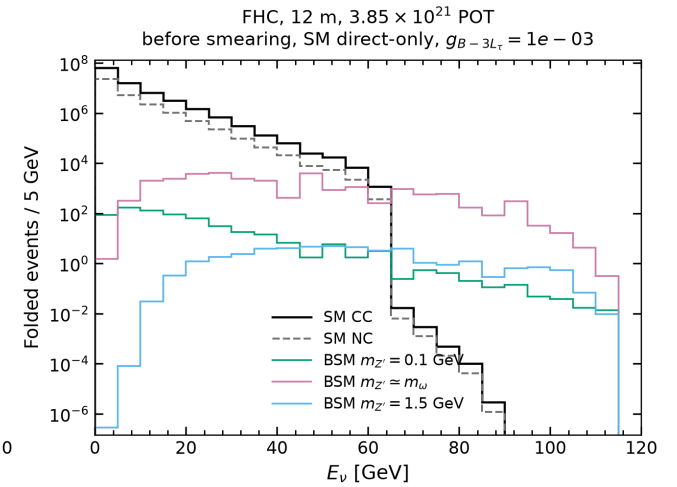
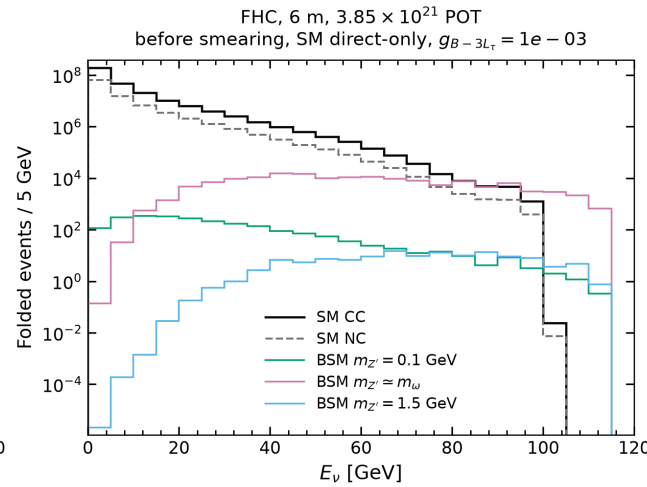
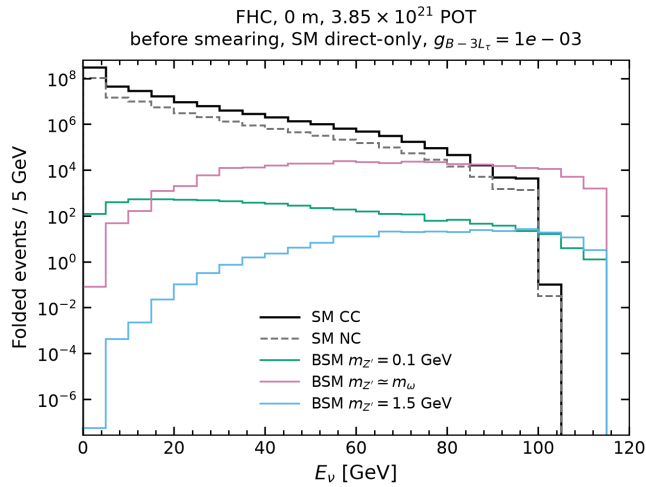
Proton form factors:

$$F_i^p(t) = x_\omega F_{i,\omega}(t) + x_\phi F_{i,\phi}(t) + x_\rho F_{i,\rho}(t)$$

$\phi = s\bar{s}$: OZI suppressed

DUNE Flux + BSM ν_τ Flux events

Folded GENIE cross section



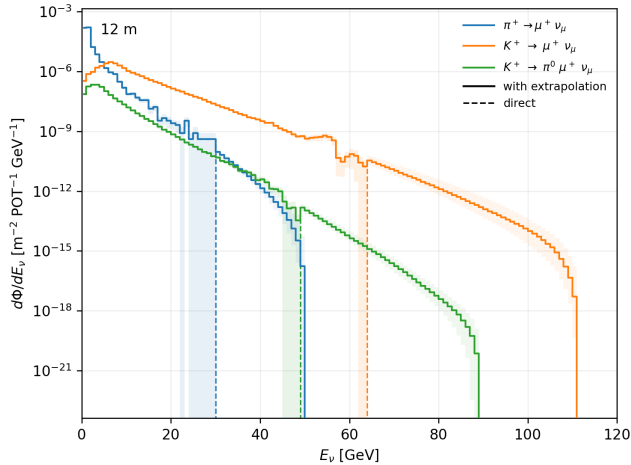
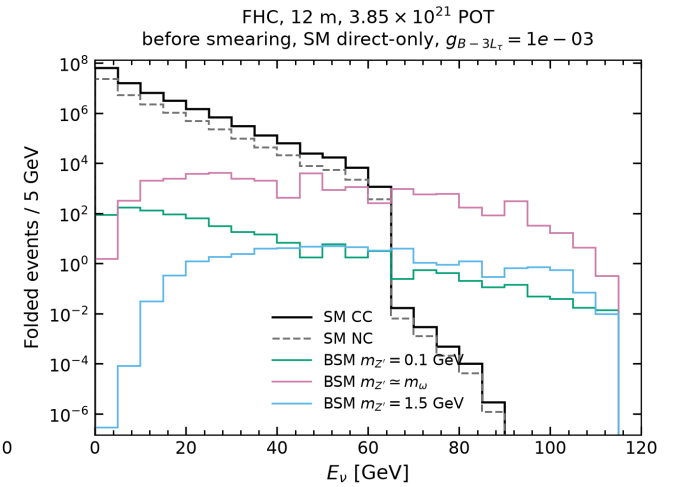
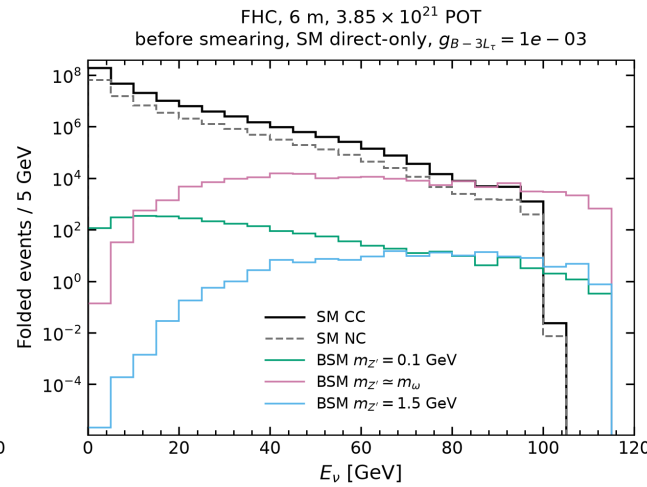
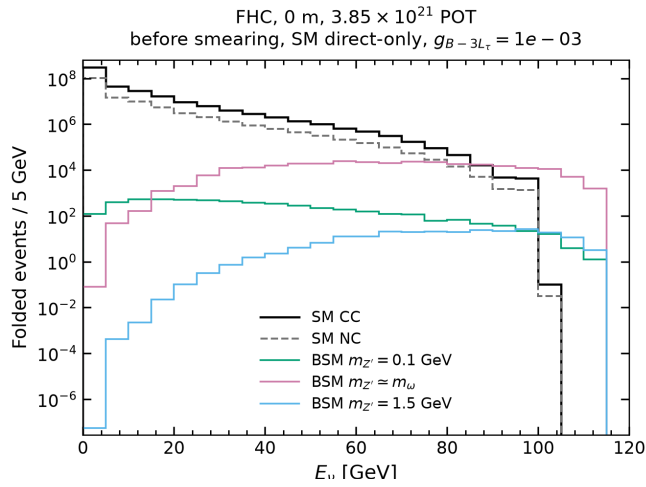
Our 10^7 POT dk2nu files **don't include enough high energy mesons,**

so even many re-decays can't provide high energy neutrino flux.

NOT a fair SM background comparison with BSM ν_τ flux!

DUNE Flux + BSM ν_τ Flux events

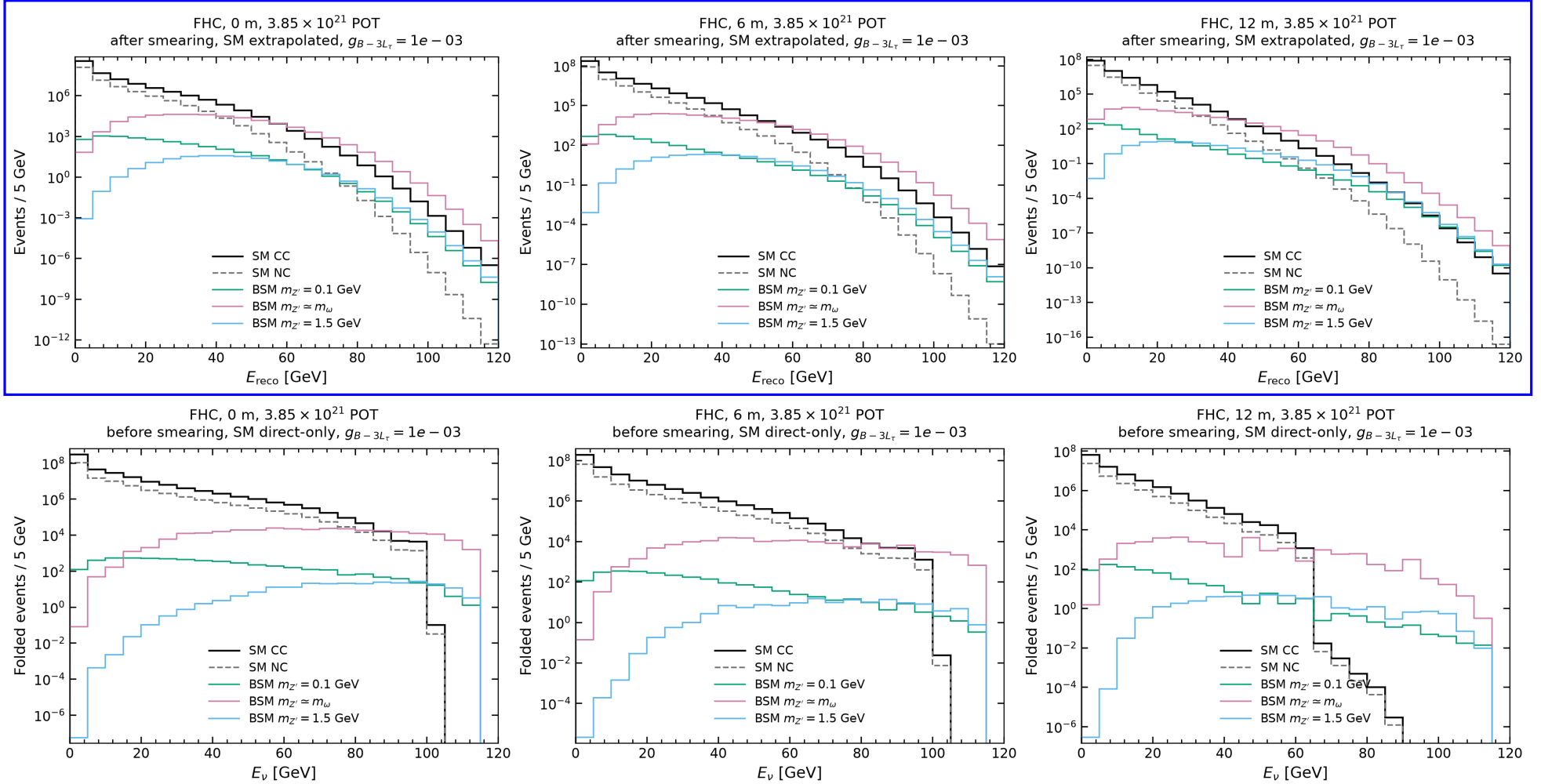
Folded GENIE cross section



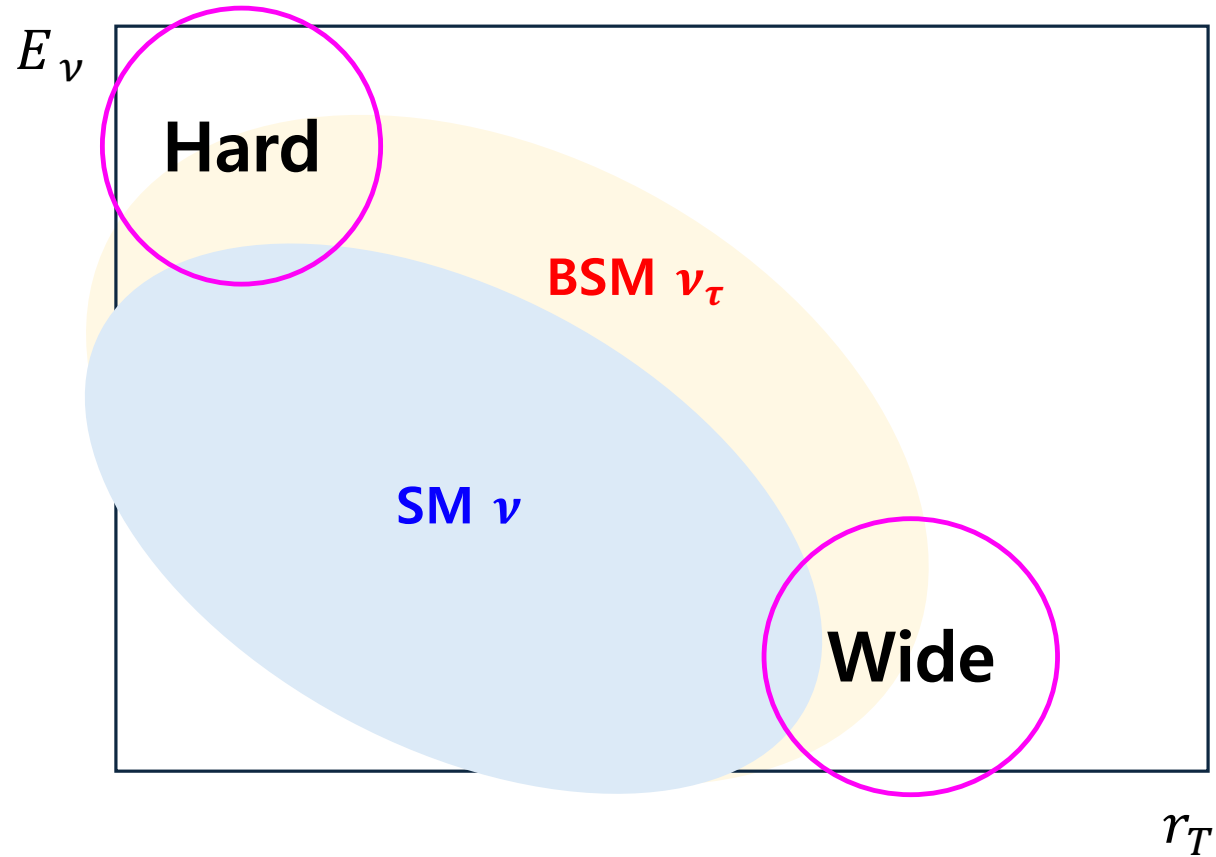
We extrapolated SM neutrino flux at each x_{OA} until kinematically allowed endpoints,

which is more fair comparison!

DUNE Flux **extrapolated** + BSM ν_τ Flux events

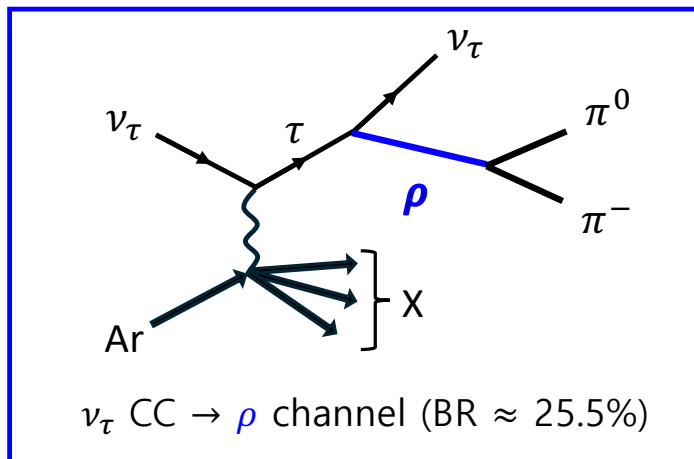


Focusing on Hard & Wide BSM ν_τ Spectrum

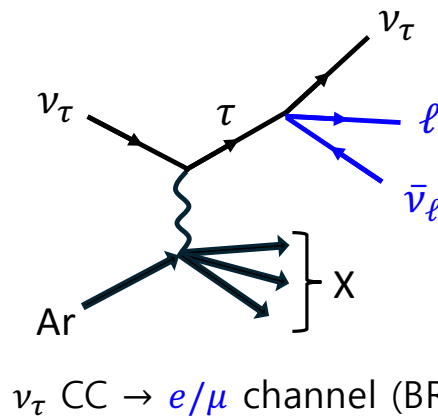


Signal / Background for ν_τ at LAr detector

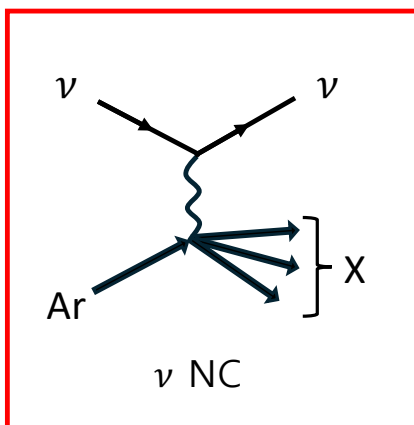
Signal



Low efficiency due to ν_e/ν_μ CC

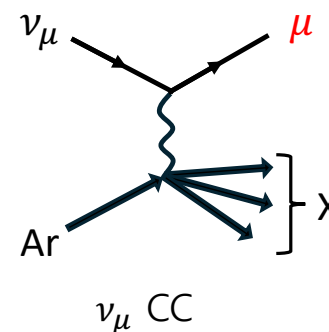
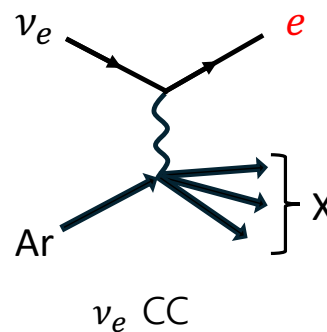


Background



$\sim 10\%$ leakage

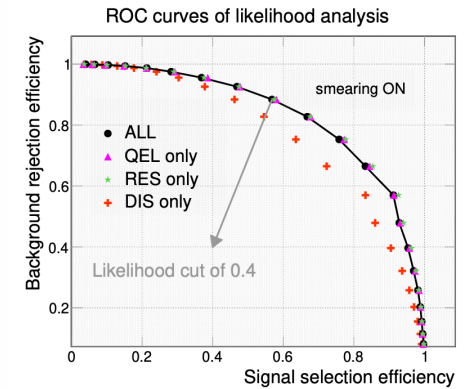
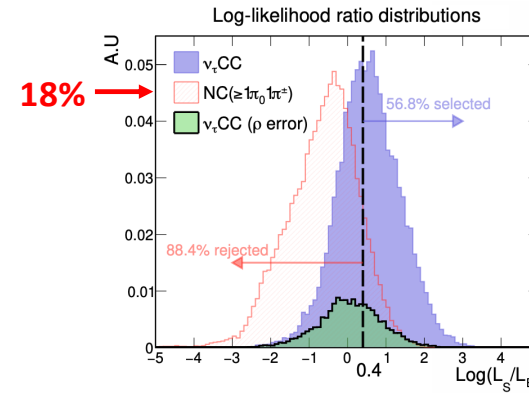
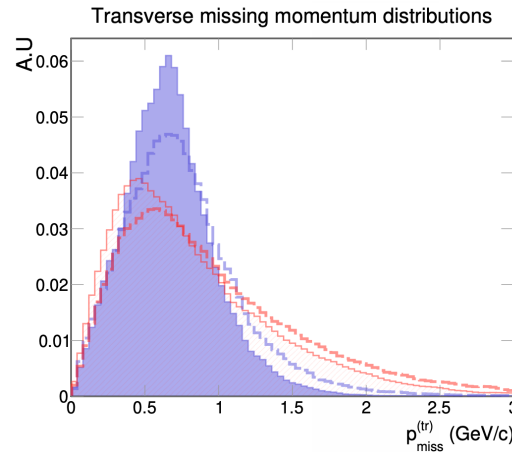
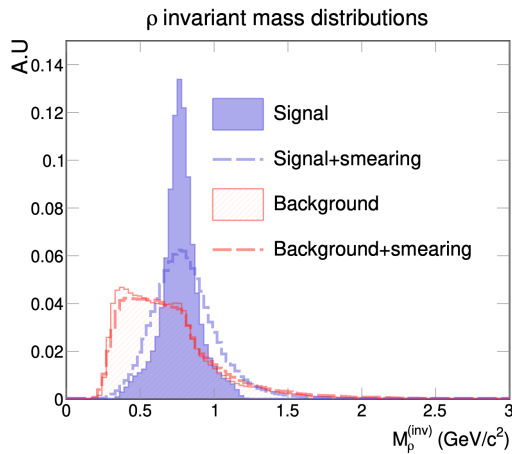
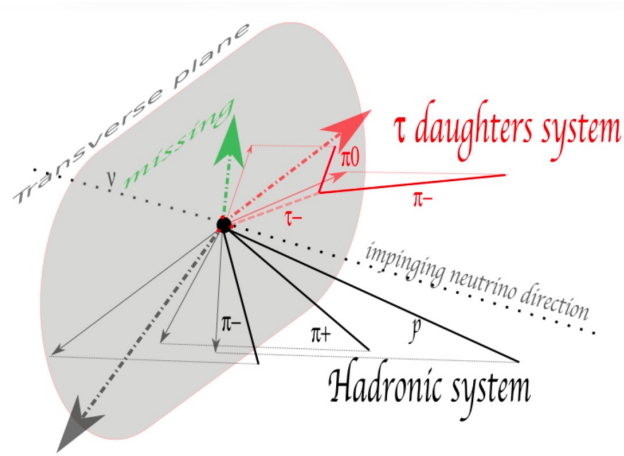
Going out of active volume, $\mu^\pm/\pi^\pm$ misidentification, ...



Transverse Plane Kinematics

Albright & Schrock, 1979

Thomas Kosc (PhD Thesis), 2021



(a)	(b)						
Likelihood cut	0	0.2	0.4	0.6	0.8	1.0	1.2
fraction of signal kept (%)	75.8	66.8	56.8	46.9	36.9	28.2	21.4
Cases of fake ρ propagated	8.8	7.1	5.6	4.2	2.9	2.0	1.3
fraction of bck contamination	24.8	17.3	11.6	7.4	4.5	2.5	1.3

**18% of NC events survives topology criteria
+ Transverse missing momentum
+ ρ invariant mass distribution**

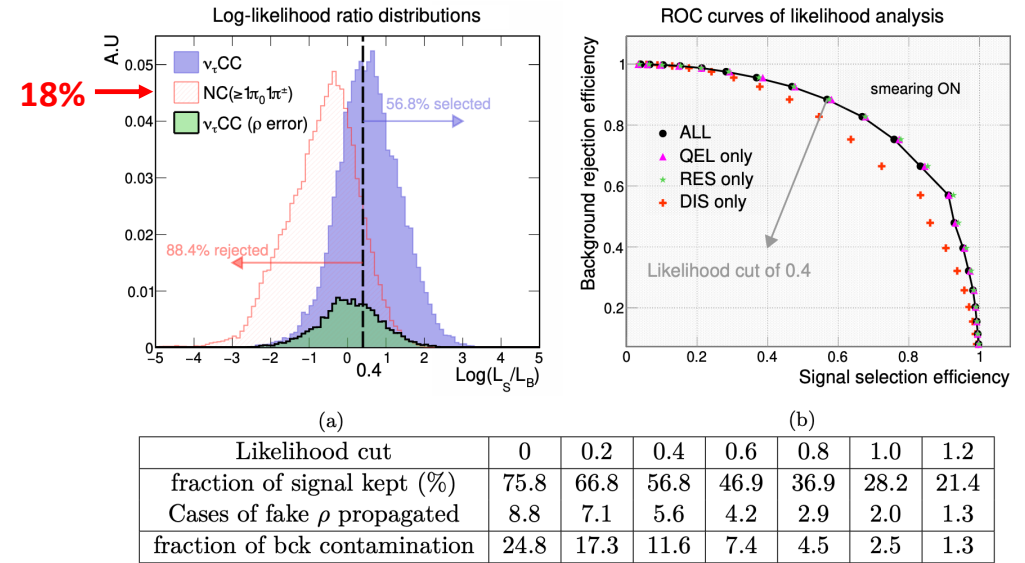
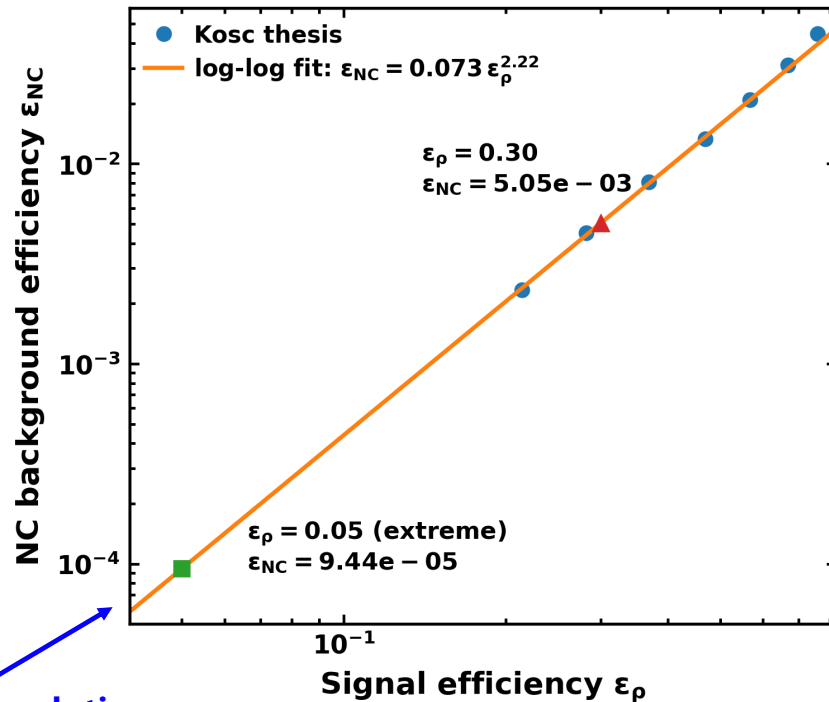
→ Gives NC-like background efficiency

Transverse Plane Kinematics

Albright & Schrock, 1979

Thomas Kosc (PhD Thesis), 2021

Our signal efficiency choices



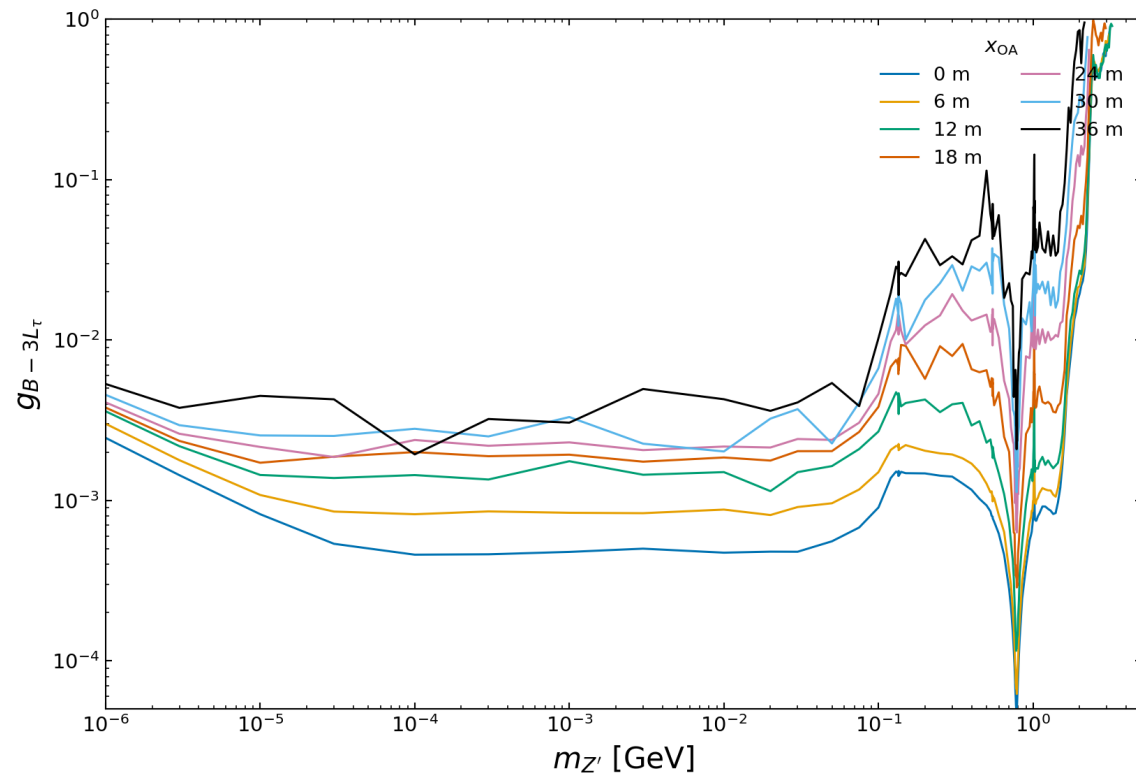
Δ - de Gouvea et al (2019): $\epsilon_{hadron} = 30\%$, $\epsilon_{NC} = 0.5\%$

■ - Dev et al (2024): $\epsilon_{\rho} = 5\%$, $\epsilon_{NC} \approx 0$ (from Aurisano)

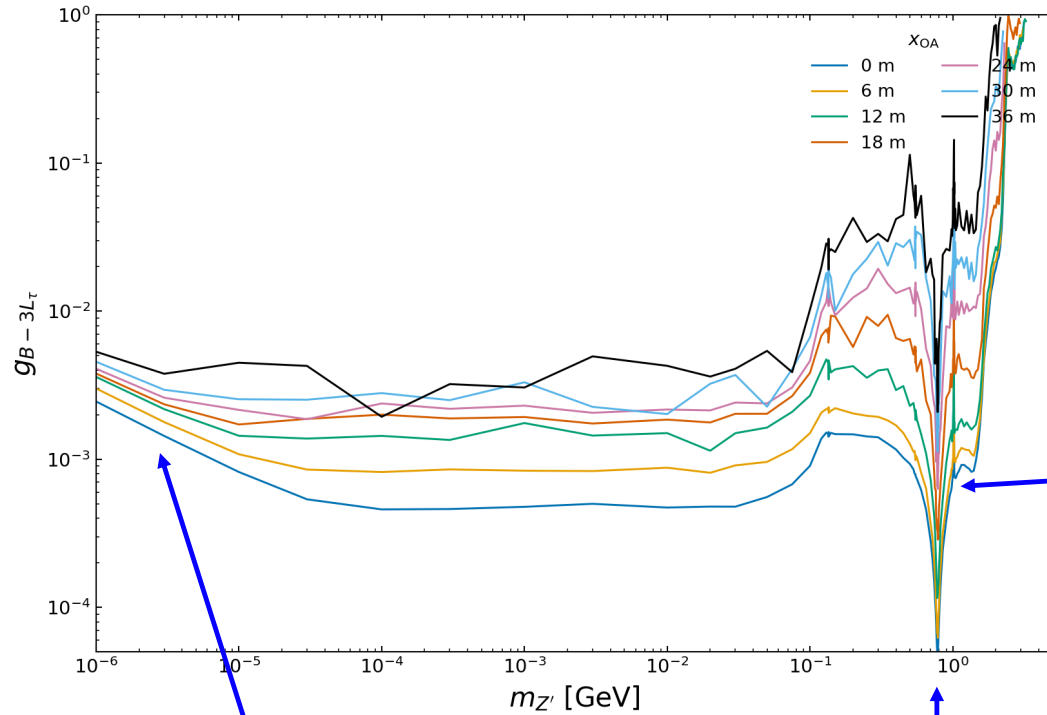
x0.18

Slide along ROC curve fit → Conservative efficiency assumption compared to previous works

Constraints – off-axis overlayed



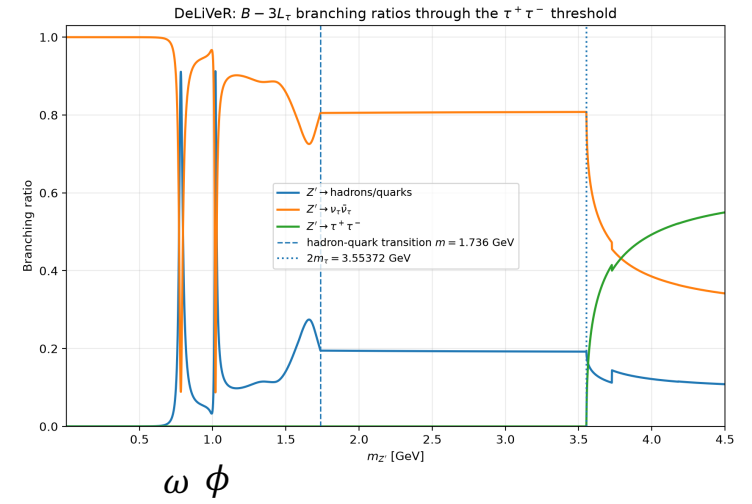
Constraints – off-axis overlayed



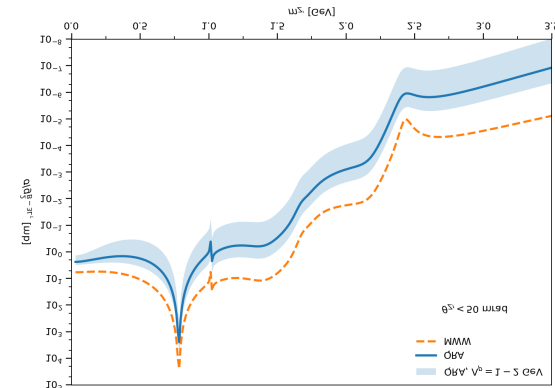
Z' propagation length become larger, so do not decay before the detector

ω meson resonance in production

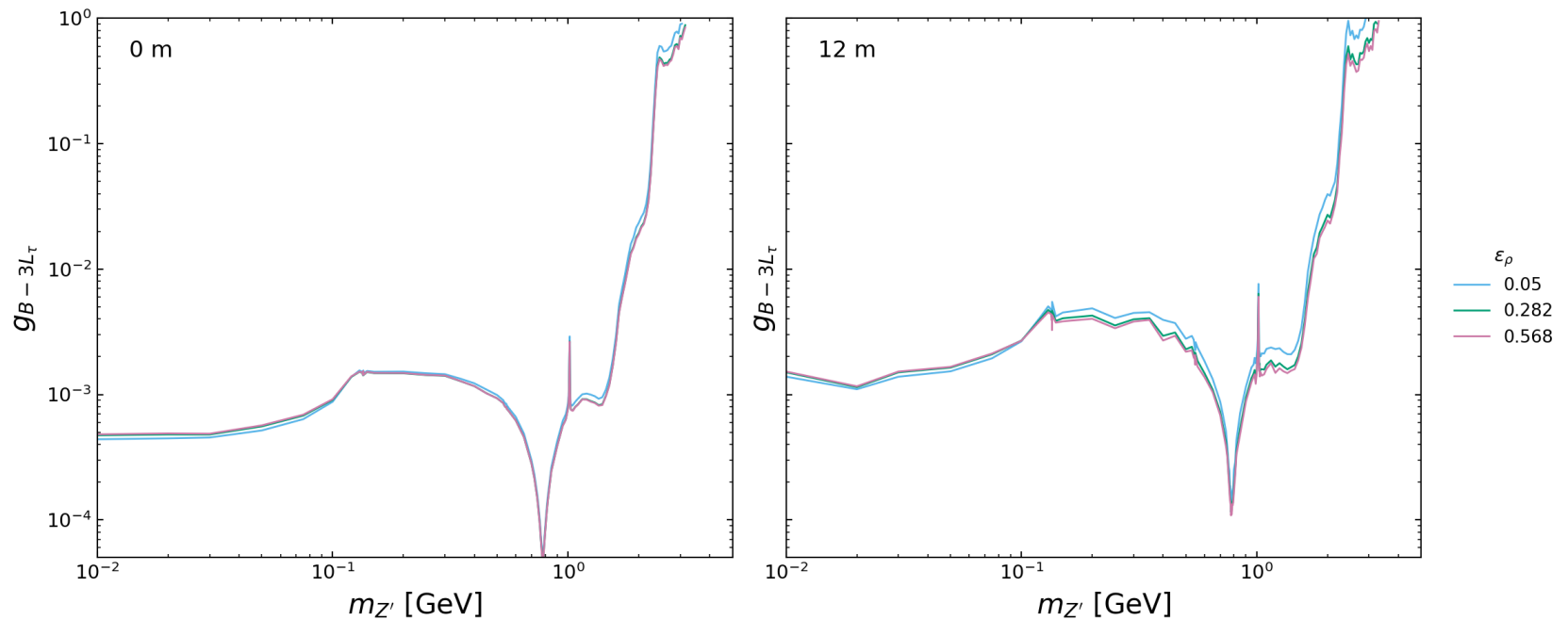
Foguel, Reimitz, Funchal - DeLiVeR



ϕ meson resonance in decay

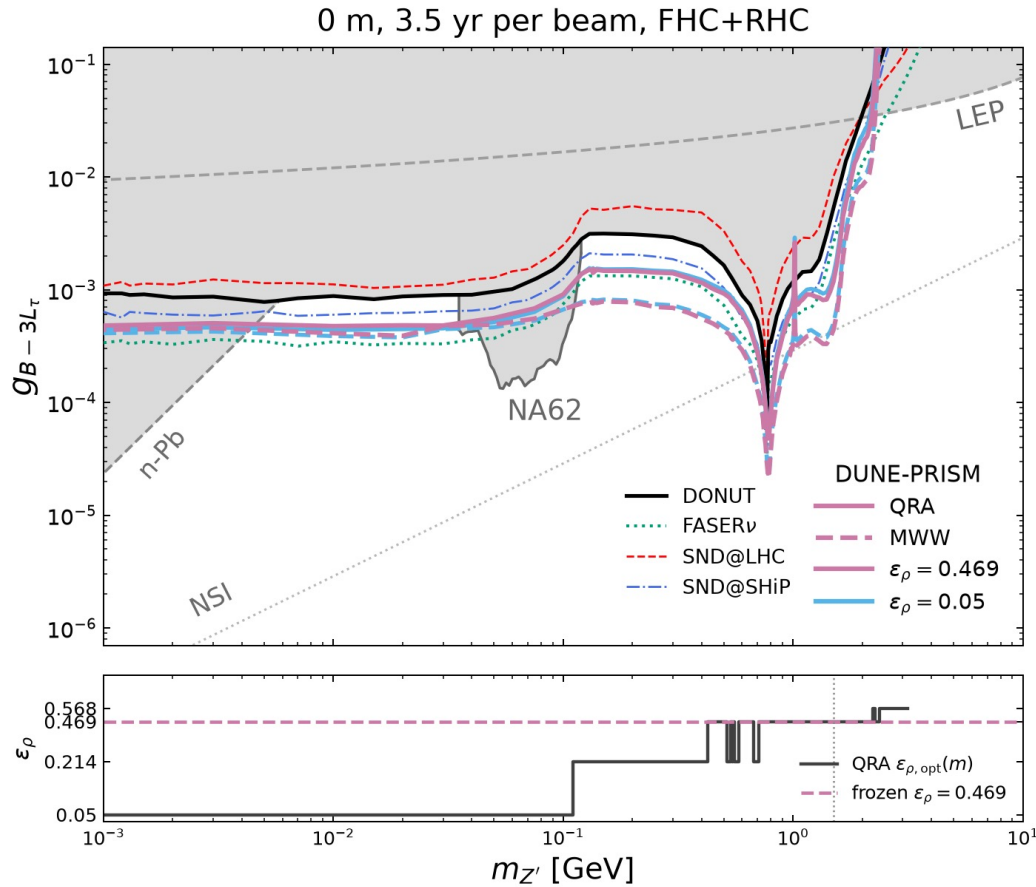


Constraints – off-axis overlayed

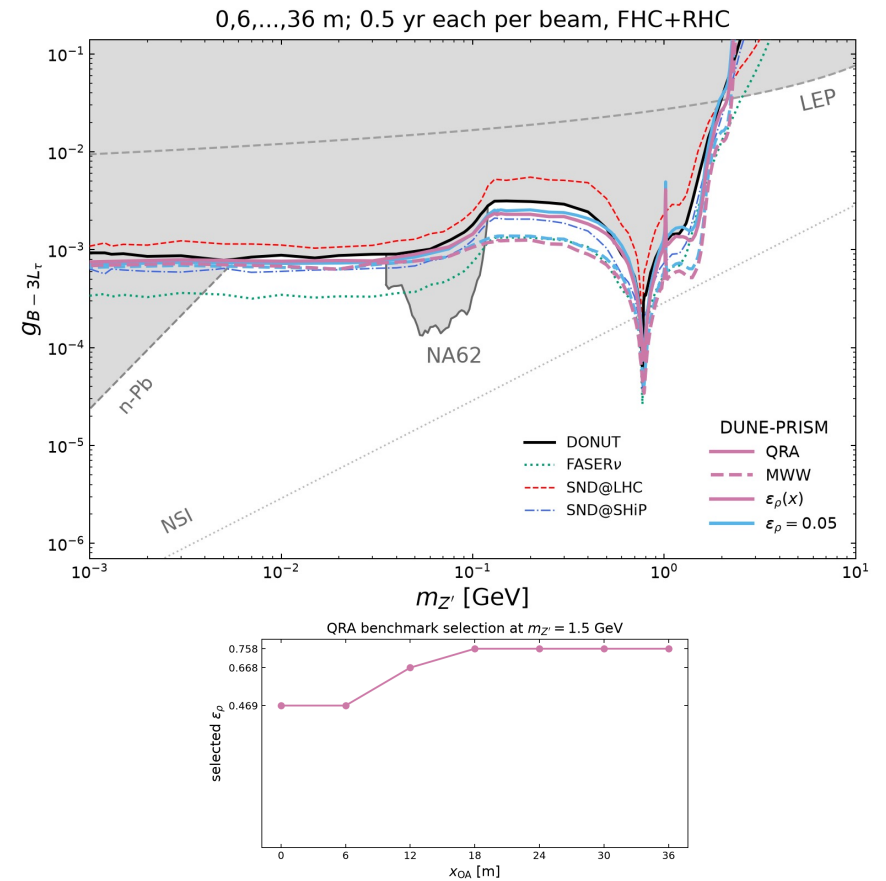


Slide along ROC curve fit → Give better sensitivity, especially at 1.5 GeV shoulder region

Constraints



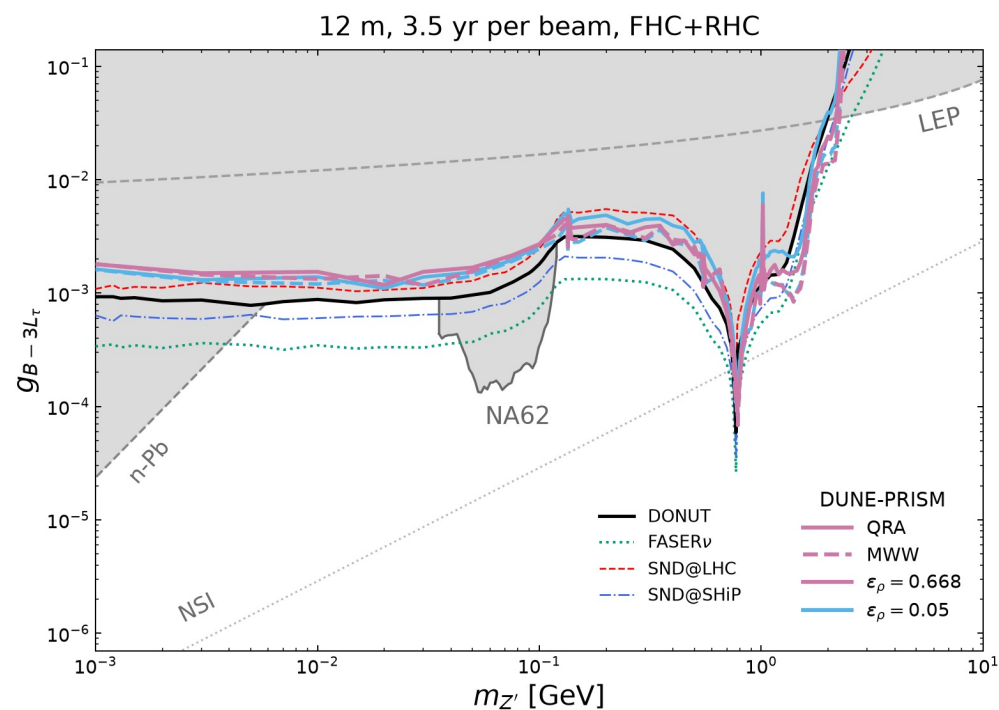
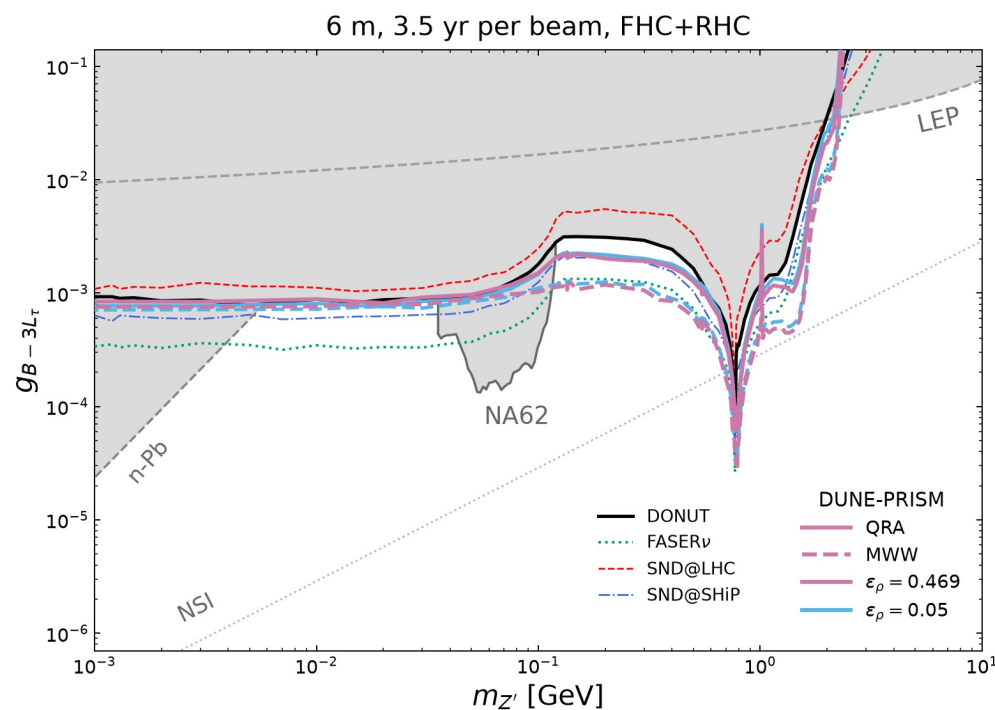
Comparison with Kling (2020)



Slide along ROC curve fit → Give better sensitivity, especially at 1.5 GeV shoulder region

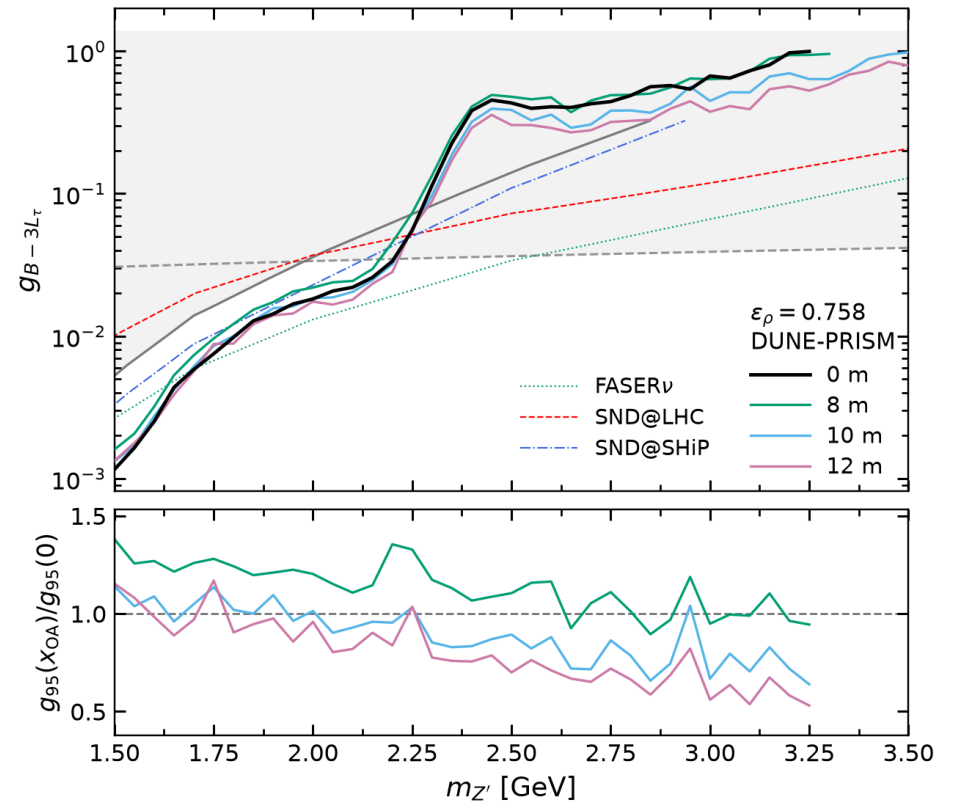
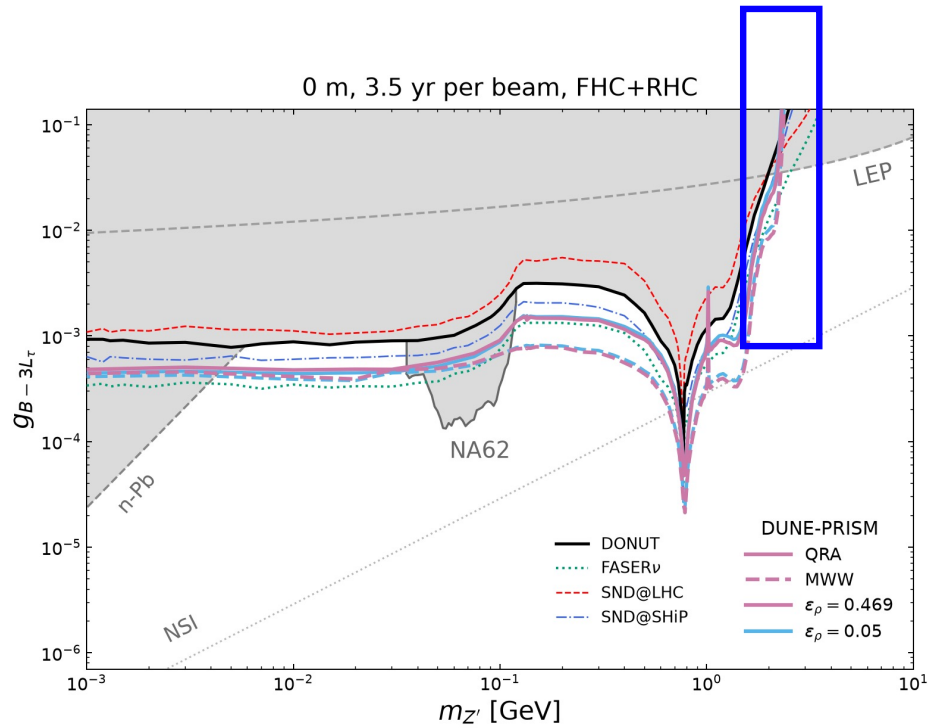
Constraints – off axis effect

Comparison with Kling (2020)



Efficiency choice becomes more important at **more off axis** position.

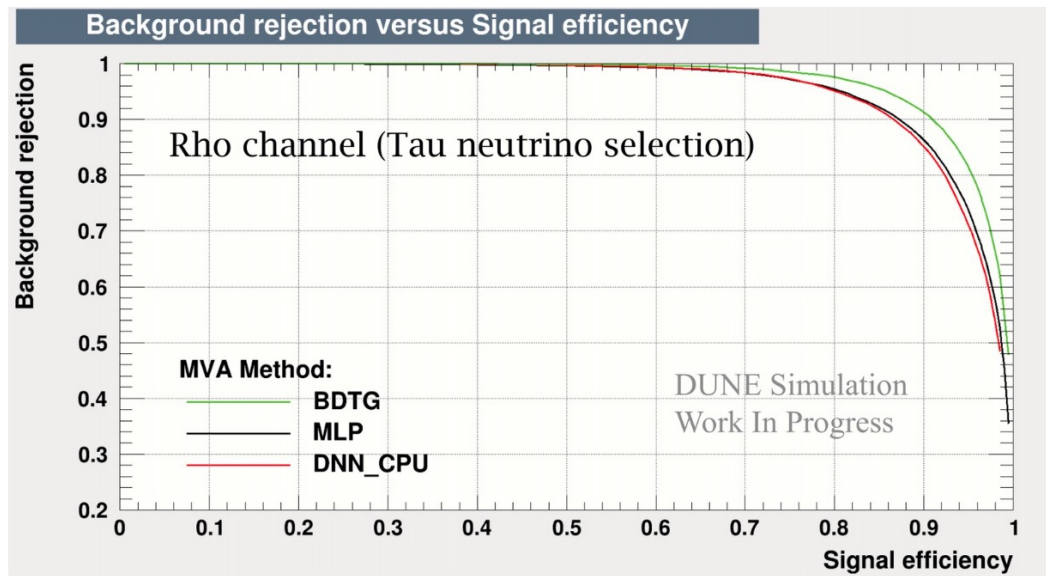
Constraints – off axis effect



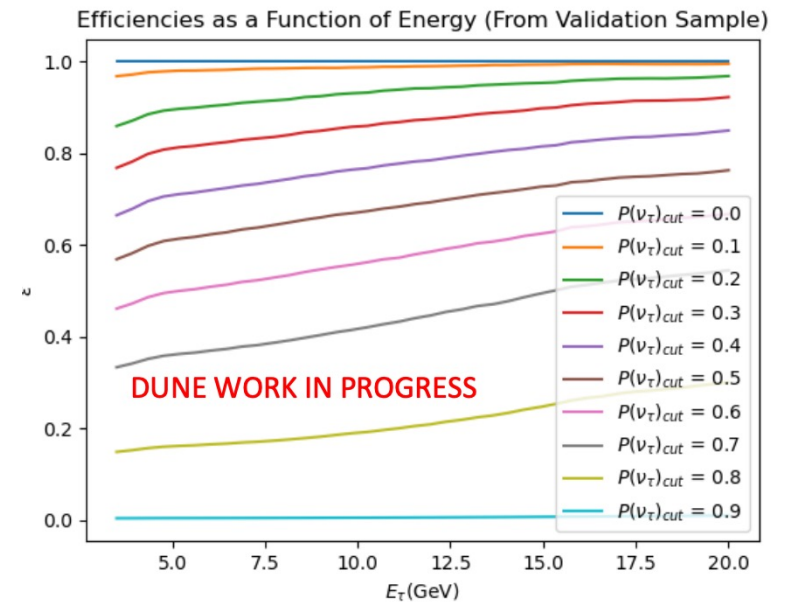
Small advantage at off-axis over on-axis for **heavier Z'** , but mostly already constrained.

Optimistic Background Scenarios?

A. Aurisano (NuFact 2021)



W. Dallaway (NuFact 2025)



Conclusion

Heavy BSM mediator

→ Hard and Wide neutrino spectrum

→ DUNE-PRISM would be nice place to study

Hard spectrum

→ Benefited by loose efficiency choice

Wide spectrum

→ Small effect, but still gives lessen

