

Almost
One-hundred
year's QUEST :

Are (active sub-eV)
neutrinos
Dirac or Majorana ?



1930 ~



$$\nu \neq \nu^c$$

Dirac particle

ν

Majorana particle



$$\nu = \nu^c = C\bar{\nu}^T$$

First Part

practical Dirac-Majorana Confusion Theorem for sub-eV active neutrino, and How to overcome it?

NCTS-NTU Taiwan 03/05/2025
Academia Sinica 02/26/2025

KMI, Nagoya Univ. 07/16/2024
UW Wisc-Pheno, 08/23/2023
DPF, Mexico Physics Society, 06/12/2023
SNU-Physics, 03/30/2023
MPI-Heidelberg, Germany, 01/23/2023
High Energy Physics in LHC Era, Chile, 01/10/2023
Apres-LHC Workshop, 08/09/2022
4th Mini WS on CUBES, 11/05/2022
WS on Physics of Dark Cosmos, 10/23/2022
KEK-IPNS, Japan, 08/22/2022
IHEP, China, 07/15/2022

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Yonsei University: ChonNam National University-SRC



CSK,MM,DS, PRD105 (2022)
CSK,JR,DS, EPJC83 (2023)
CSK, EPJC83 (2023)
CSK,DS,KV, EPJC84(2024)

(Is the sub-eV active neutrino Dirac or Majorana ?)

Alternative to $0\nu\beta\beta$:

Second Part
Work in progress

Quantum Statics + $2\nu\beta\beta$!!

KyungPook National Univ. 06/08/2026
IHEP, China 06/24/2026
Hebei U 07/06/2026
nEXO@Yemilab 08/31/2026
YITP@Stonybrook 09/09/2026

Pittsburgh U 03/03/2026, 03/05/2026
Maryland U 02/09/2026
Seoul Tech 10/30/2025
IBS-CUP 10/15/2025
UW-Madison 06/27/2025
FNAL 06/20/2025
CNU-SRC 05/16/2025
CUBES-7 04/26/2025

Choong Sun Kim

Yonsei University: CNU-SRC



CSK,MM,DS, PRD105 (2022); 2106.11785
CSK,JR,DS, EPJC83 (2023); 2209.10110
CSK, EPJC83 (2023); 2307.05654
CSK, DS, KV, EPJC84 (2024); 2405.17341



First Part

Second Part

CONTENTS

- Introduction (nu Mass, Seesaw & 0nuBB)
- practical Dirac-Majorana Confusion Theorem (pDMCT)
- How to overcome pDMCT?
- -----
- Alternatives to $0\nu\beta\beta$
- Back-to-back $\nu - \bar{\nu}$ in $2\nu\beta^+\beta^-$: $[\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)]$
- Summary (& Comments)

INTRODUCTION

Sub-eV active neutrino mass

Seesaw mechanism

$\Delta L = 2$ processes & 0-nu-Beta-Beta

(sub-eV active) neutrinos have mass

- ❖ Neutrinos are massless in SM, $m_\nu = 0$.
All neutrinos are only left-handed (ν_L).

$$\mathcal{L}_{\text{mass}}^D = -m_\nu (\bar{\nu}_R \nu_L + \bar{\nu}_L \nu_R), \quad m_\nu = \frac{Y_\nu v}{\sqrt{2}},$$

where Y_ν = Higgs-neutrino Yukawa coupling constant,
and v = Higgs VEV.

No way to generate mass without right-handed neutrinos (ν_R).

- ❖ But observations of **neutrino oscillation** imply that **neutrinos have mass**, $m_\nu \neq 0$.

The Nobel Prize in Physics 2015 was awarded jointly to Takaaki Kajita (Super-Kamiokande) and Arthur B. McDonald (Sudbury Neutrino Observatory) “for the discovery of neutrino oscillations, which shows that neutrinos have mass”.



How to give neutrinos mass?

There are various suggestions as to how neutrinos can get mass.

❖ Dirac mass:

- Assumption: ν_R exists.
- Lagrangian:

$$\mathcal{L}_{\text{mass}}^D = -m_\nu^D (\bar{\nu}_R \nu_L + \bar{\nu}_L \nu_R).$$

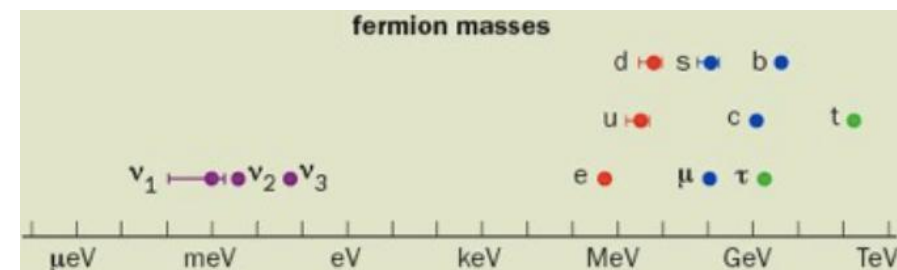
- Disadvantage: No reason for m_ν^D to be small.

❖ Majorana mass:

- Assumption: neutrino $\equiv$ anti-neutrino.
- Lagrangian:

$$\mathcal{L}_{\text{mass}}^M = \frac{1}{2} m_\nu^M (\bar{\nu}_L^C \nu_L + \bar{\nu}_L \nu_L^C).$$

- Disadvantage: $\mathcal{L}_{\text{mass}}^M$ is not invariant under $SU(2)_L \times U(1)_Y$ gauge group, so not allowed by SM.



How to give neutrinos (very small) mass ?

- ❖ **See-saw mechanism:** A simpler version of Dirac-Majorana mass, with a nice twist.

[PM,PLB67(1977)421]

- *Assumptions:* $m_\nu^L = 0$ and $m_\nu^D \ll m_\nu^R$.

- *Lagrangian:*

$$\mathcal{L}_{\text{mass}}^{D+M} = \frac{1}{2} m_\nu^R (\overline{\nu_R^C} \nu_R) - m_\nu^D (\overline{\nu_R} \nu_L) + \text{H.c.} = \frac{1}{2} \overline{N_L^C} M N_L + \text{H.c.}, \text{ where}$$

$$N_L = \begin{pmatrix} \nu_L \\ \nu_R^C \end{pmatrix} \text{ and } M = \begin{pmatrix} 0 & m_\nu^D \\ m_\nu^D & m_\nu^R \end{pmatrix} \text{ is the mass matrix.}$$

- *Mass eigenvalues:*

$$m_{2,1} = \frac{1}{2} \left(m_\nu^R \pm \sqrt{(m_\nu^R)^2 + 4(m_\nu^D)^2} \right)$$

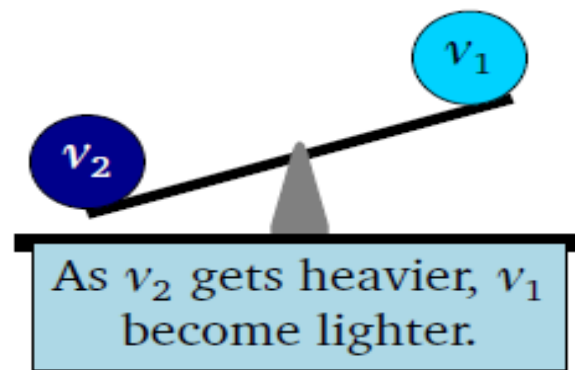
$$\approx \frac{1}{2} m_\nu^R \left(1 \pm 1 \pm 2 \left(\frac{m_\nu^D}{m_\nu^R} \right)^2 \right).$$

$$\Rightarrow m_1 \approx -\frac{(m_\nu^D)^2}{m_\nu^R} \text{ and } m_2 \approx m_\nu^R. \quad \sin \theta \approx m_\nu^D / m_\nu^R$$

- *Advantage:* $m_1 \ll m_2$, so light neutrinos are possible.

- *Challenges:*

- To find the heavy ν_2 experimentally.
- To prove that both the light ν_1 and heavy ν_2 are Majorana neutrinos.

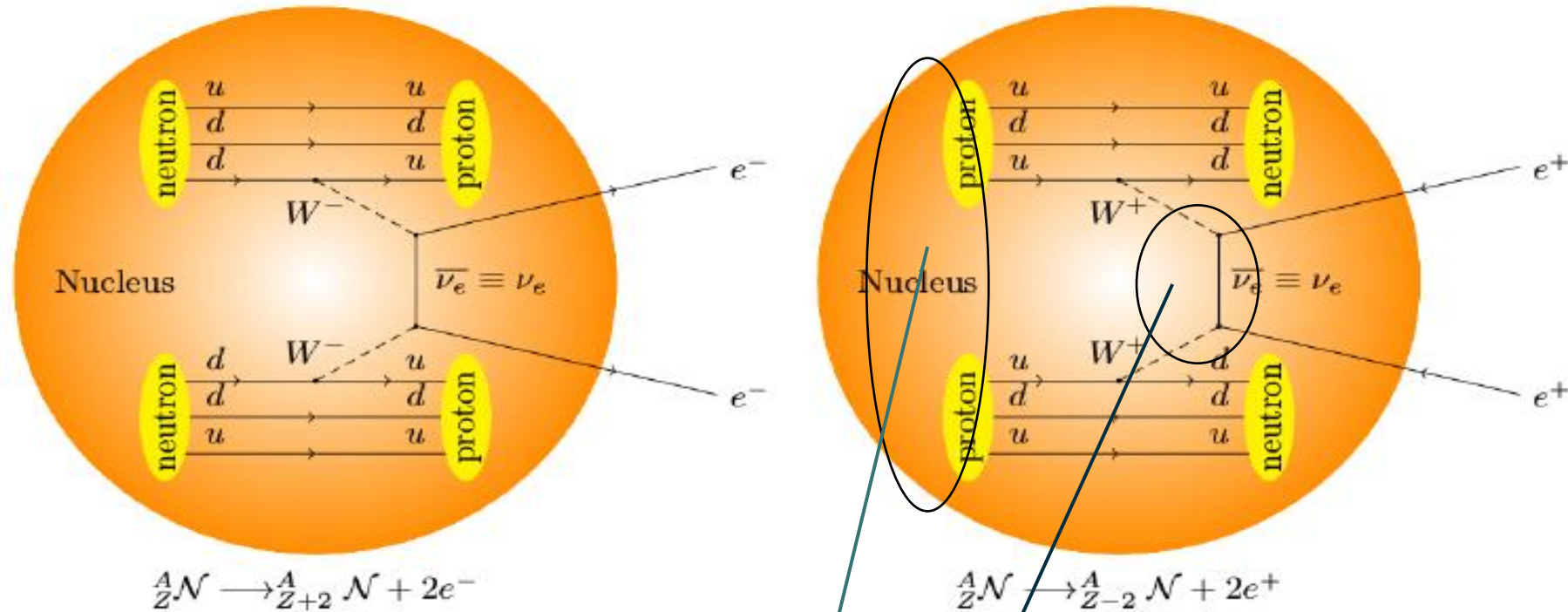


Neutrino-less double-beta decay ($0\nu\beta\beta$) (1)

❖ Process:

Lepton Number Violation (LNV)

→ not allowed within SM



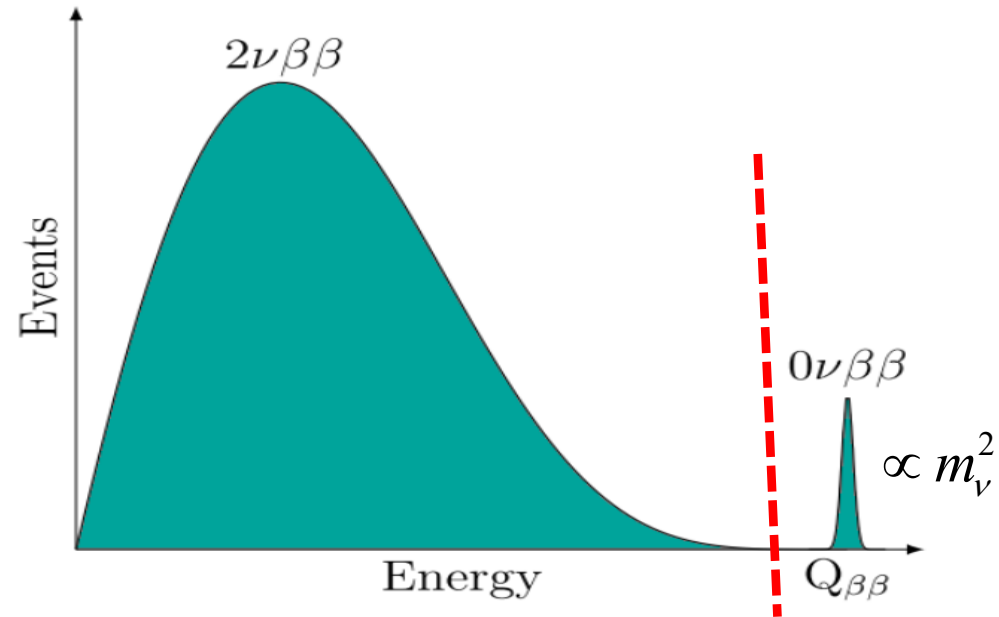
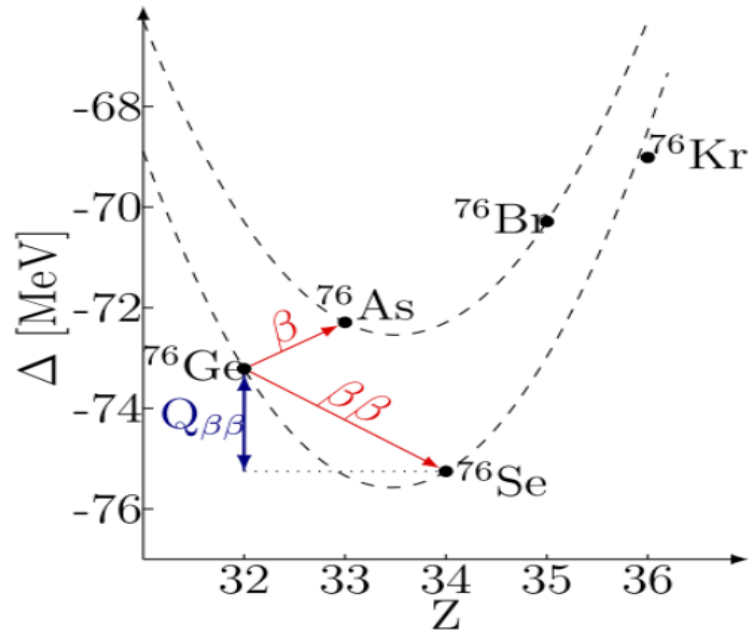
Doubly weak charged current process

❖ The half-life of a nucleus decaying via $0\nu\beta\beta$ is,

$$\left[T_{1/2}^{0\nu}\right]^{-1} = G_{0\nu} |M_{0\nu}| |m_{\beta\beta}|^2,$$

Neutrino-less double-beta decay ($0\nu\beta\beta$) (2)

- ❖ Double-beta ($2\nu\beta\beta$) decay has been observed in 10 isotopes, ^{48}Ca , ^{76}Ge , ^{82}Se , ^{96}Zr , ^{100}Mo , ^{116}Cd , ^{128}Te , ^{130}Te , ^{150}Nd , ^{238}U , with half-life $T_{1/2} \approx 10^{18} - 10^{24}$ years.



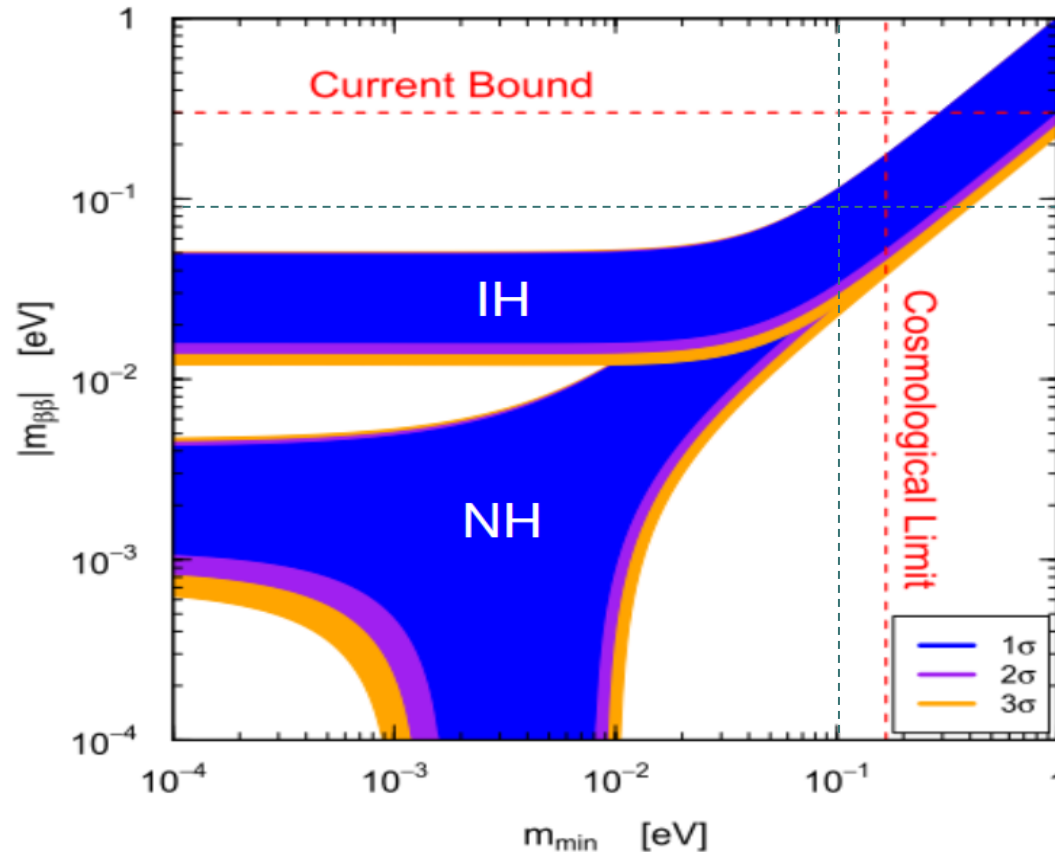
Giovanni Benato (for the GERDA collaboration), arXiv:1509.07792

- ❖ $0\nu\beta\beta$ (forbidden in SM) is yet to be observed in any experiment.

$$T_{1/2}^{0\nu} [^{76}\text{Ge}] > 2.1 \times 10^{25} \text{ years (90\% C.L.)}.$$

M. Agostini *et al.* (GERDA Collaboration) Phys. Rev. Lett. **111**, 122503 (2013).

Neutrino-less double-beta decay ($0\nu\beta\beta$) (3)



NH: Normal hierarchy
IH: Inverted hierarchy

S. M. Bilenky and C. Giunti

Mod. Phys. Lett. A **27**, 1230015 (2012),
arXiv:1203.5250

- ❖ If $m_{\beta\beta} < 10^{-2}$, only NH is viable and the $T_{1/2}^{0\nu}$ will be much larger than the current experimental lower bound $\left[T_{1/2}^{0\nu}\right]^{-1} = G_{0\nu} |M_{0\nu}| |m_{\beta\beta}|^2$.

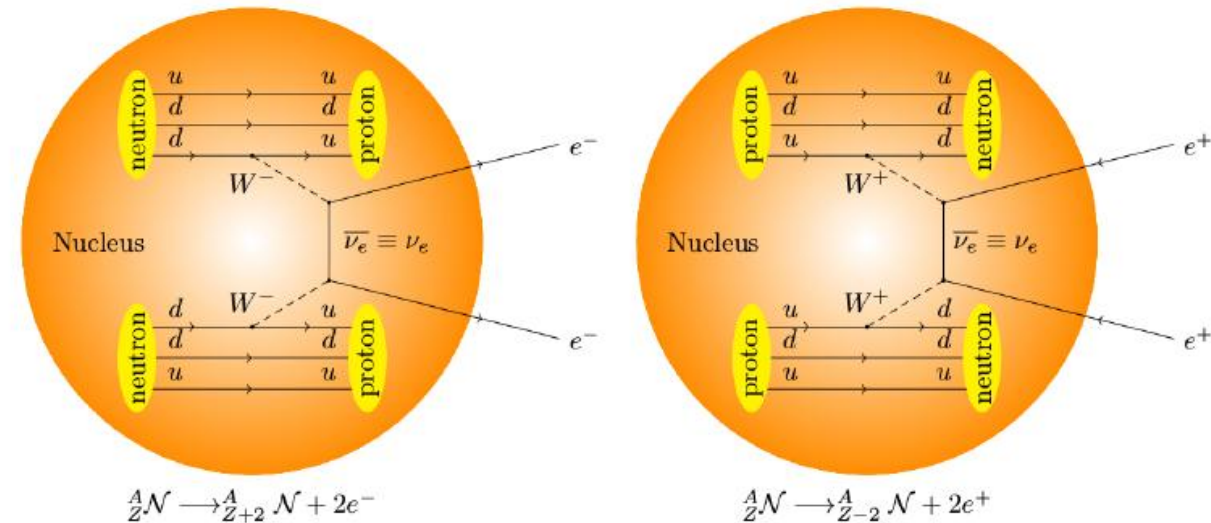
Neutrino-less double-beta decay ($0\nu\beta\beta$) (4)

Is $0\nu\beta\beta$ the evidence for (sub-eV) active neutrino being Majorana fermion ?

→ Evidence of **Lepton Number Violation**

→ **NOT unequivocal evidence** of active (sub-eV) neutrino being Majorana neutrino !!

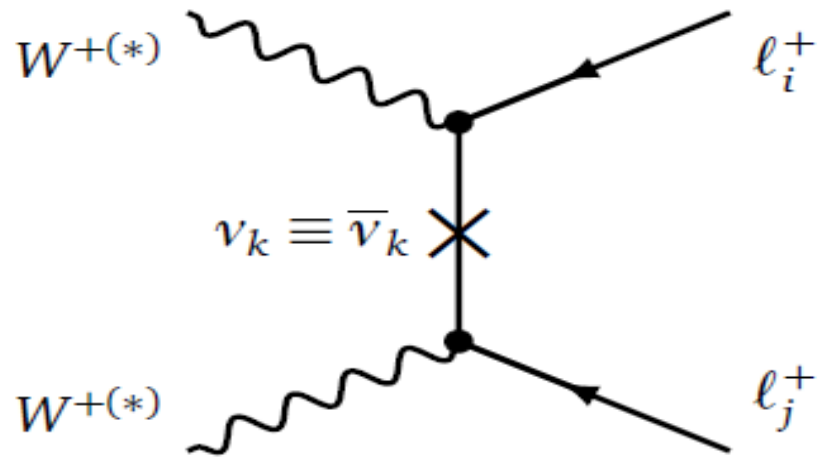
❖ Process:



e.g. heavy sterile neutrinos can give the same effects

Looking for Majorana neutrinos via $\Delta L = 2$ processes (1)

- ❖ Neutrinos are the only *elementary fermions* known to us that *can* have Majorana nature.
- ❖ Majorana neutrinos: $\nu \equiv \bar{\nu}$.
- ❖ Majorana neutrinos violate lepton flavor number (L), they mediate $\Delta L = 2$ processes.



$$\propto \int \frac{d^4 p}{(2\pi)^4} \sum_k U_{\ell_i k} U_{\ell_j k} \frac{m_k + \not{p}}{p^2 - m_k^2}$$

- ❖ $\Delta L = 2$ processes play crucial role to probe Majorana nature of ν 's.
- | | |
|---------------------|---|
| $m_k^2 \ll p^2$ | ○ neutrinoless double-beta ($0\nu\beta\beta$) decay |
| $m_k^2 \approx p^2$ | ○ Rare meson decays with $\Delta L = 2$ |
| $m_k^2 \gg p^2$ | ○ Collider searches at LHC |

Looking for Majorana neutrinos via $\Delta L = 2$ processes (2)

❖ Decay rate of any $\Delta L = 2$ process with final leptons $\ell_1^+ \ell_2^+$:

$$\Gamma_{\Delta L=2} \propto \left| \sum_k U_{\ell_1 k} U_{\ell_2 k} \frac{m_k}{p^2 - m_k^2 + im_k \Gamma_k} \right|^2,$$

where we have used the fact that $(1 - \gamma^5) \not{p}(1 - \gamma^5) = 0$.

○ Light ν :

$$m_k^2 \ll p^2$$

$$\Gamma_{\Delta L=2} \propto \left| \sum_k U_{\ell_1 k} U_{\ell_2 k} m_k \right|^2 = |m_{\ell_1 \ell_2}|^2.$$

$$|m_{\beta\beta}|^2 \sim (0.01 \text{eV})^2 = (10^{-22}) \text{GeV}^2$$

○ Heavy ν :

$$m_k^2 \gg p^2$$

$$\Gamma_{\Delta L=2} \propto \left| \sum_k \frac{U_{\ell_1 k} U_{\ell_2 k}}{m_k} \right|^2.$$

$$m_\nu^R \sim 10^5 \text{GeV} \quad U \approx \frac{m_\nu^D}{m_\nu^R} \sim 10^{-3}$$

○ Resonant ν :

$$m_k^2 \approx p^2$$

$$\Gamma_{\Delta L=2} \propto \frac{\Gamma(N \rightarrow i) \Gamma(N \rightarrow f)}{m_N \Gamma_N}.$$

$$B^+ \rightarrow K^- \mu^+ \mu^+, \dots$$

[CS Kim et al, PRD82(2010);1005.4282]



[C.S. Kim, M. Murthy, D. Sahoo, arXiv:2106.11785 [PRD105(2022)]; C.S. Kim, arXiv:2307.05654 [EPJC83(2023)]]

practical Dirac-Majorana Confusion Theorem

practical Dirac-Majorana Confusion Theorem (pDMCT)

Discussions on pDMCT

practical Dirac-Majorana Confusion Theorem

The difference between Dirac and Majorana neutrinos via **any kinematical observable** would be practically impossible to determine due to fact that **the observable difference between Dirac and Majorana neutrinos is proportional to the tiny neutrino mass.**

B. Kayser, PRD26, 1662 (1982)

(hidden/overlooked) Issues :

1. Historically, only SM allowed **neutral current interaction** mediated processes were analyzed.
2. No way to gain any information regarding individual neutrino antineutrino energies or 3-momenta.
This invariably leads to **integration over neutrino antineutrino related kinematic variables.**

No model-independent, process-independent and observable-independent proof of this so-called “theorem”. All processes where it was shown to hold involved **full integration over the 4-momenta of missing neutrinos and/or only for $Z^{(*)} \rightarrow \nu \bar{\nu}$ case.**

History trying to overcome DMCT, but only confirming

All for weak neutral current process in SM

$$\gamma^* \rightarrow \nu \bar{\nu}$$

$$Z \rightarrow \nu \bar{\nu}$$

$$e^+ e^- \rightarrow \nu \bar{\nu}$$

$$K^+ \rightarrow \pi^+ \nu \bar{\nu}$$

$$e^+ e^- \rightarrow \nu \bar{\nu} \gamma$$

$$|es\rangle \rightarrow |gs\rangle + \gamma \nu \bar{\nu}$$

$$e^- \gamma \rightarrow e^- \nu \bar{\nu}$$

[B Kayser, PRD26(1982)] **_[0]

[1] S.P.Rosen,PRL48(1982);
W.Rodejohann etal,JHEP05(2017).

[RE Shrock, eConf(1982)]

[E Ma, JT Pantaleone, PRD40(1989)]

[JF Nieves, PB Pal, PRD32(1985)]

[T Chabra, PR Babu, PRD46(1992)] **_[2]

[Y Yoshimura, PRD75(2007)],

[JM Berryman etal, PRD98(2018)]

** All practically impossible to measure momenta of nu-nubar → Need integrate out → pDMCT

practical Dirac-Majorana Confusion Theorem (1)

Consider the SM allowed decay, e.g. $B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2),$

Amplitude for Dirac case

$$\mathcal{M}^D = \mathcal{M}(p_1, p_2),$$

$$M^D = M[\mu^-(p_{\mu-}, s_{\mu-}), \mu^+(p_{\mu+}, s_{\mu+}), \bar{\nu}(p_{\bar{\nu}}, s_{\bar{\nu}}), \nu(p_\nu, s_\nu)]$$

For Majorana case

$$\mathcal{M}^M = \frac{1}{\sqrt{2}}(\mathcal{M}(p_1, p_2) - \mathcal{M}(p_2, p_1)).$$

not helicity flip, but momentum exchange

Difference between D and M

Initial spin averaged, final spin summed

$$\begin{aligned} |\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 &= \frac{1}{2} \left(\underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} - \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}} \right) \\ &\quad + \underbrace{\text{Re}(\mathcal{M}(p_1, p_2)^* \mathcal{M}(p_2, p_1))}_{\text{Interference term}}. \end{aligned}$$

Dirac-Majorana Confusion Theorem (2)

Interference term
(within SM)

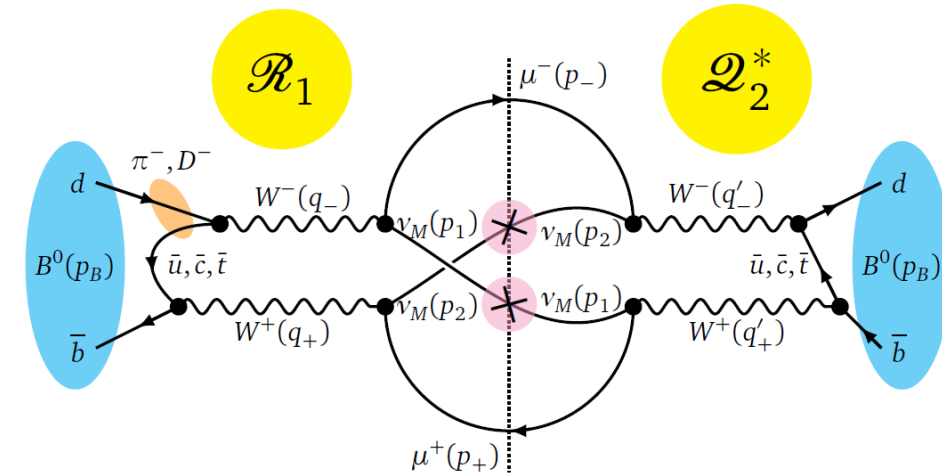
$$\text{Re}(\mathcal{M}(p_1, p_2)^* \mathcal{M}(p_2, p_1)) \propto m_\nu^2.$$

In general

$$\underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} \neq \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}}.$$

However, after integration
(required if momenta p_1 and p_2 are unobservable and missing)

$$\iint \underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} d^4 p_1 d^4 p_2 = \iint \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}} d^4 p_1 d^4 p_2,$$



← Possibly only source of D-M difference w/ SM

Dirac-Majorana Confusion Theorem (2)

Interference term

In general

(useful if p_1 and/or p_2 are known)

However, after integration

(required if momenta p_1 and p_2 are unobservable)

**** Special case :** $|M(p_1, p_2)|^2 = |M(p_2, p_1)|^2$

$$|\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 = \frac{1}{2} \left(\underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} - \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}} \right) + \underbrace{\text{Re}(\mathcal{M}(p_1, p_2)^* \mathcal{M}(p_2, p_1))}_{\text{Interference term}} \longrightarrow m_\nu^2$$

pDMCT holds at the Amplitude-Squared level

→ Difference between Dirac and Majorana neutrinos becomes 0 at every observable

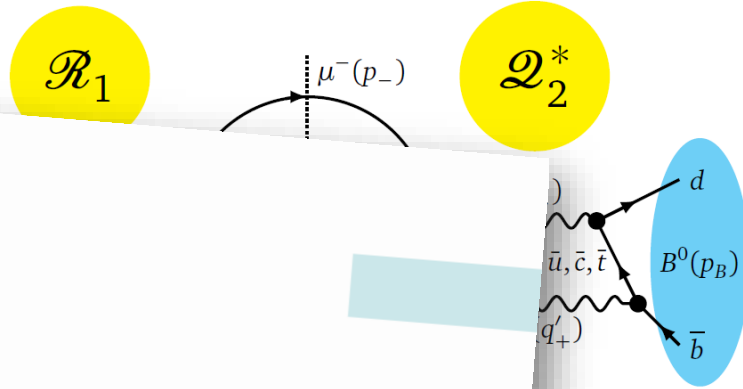
→ e.g. Within the SM, weak neutral current decays: $Z^{(*)} \rightarrow \nu \bar{\nu}$

3/1/2023

practical Dirac-Majorana Confusion Theorem

Direct term

32

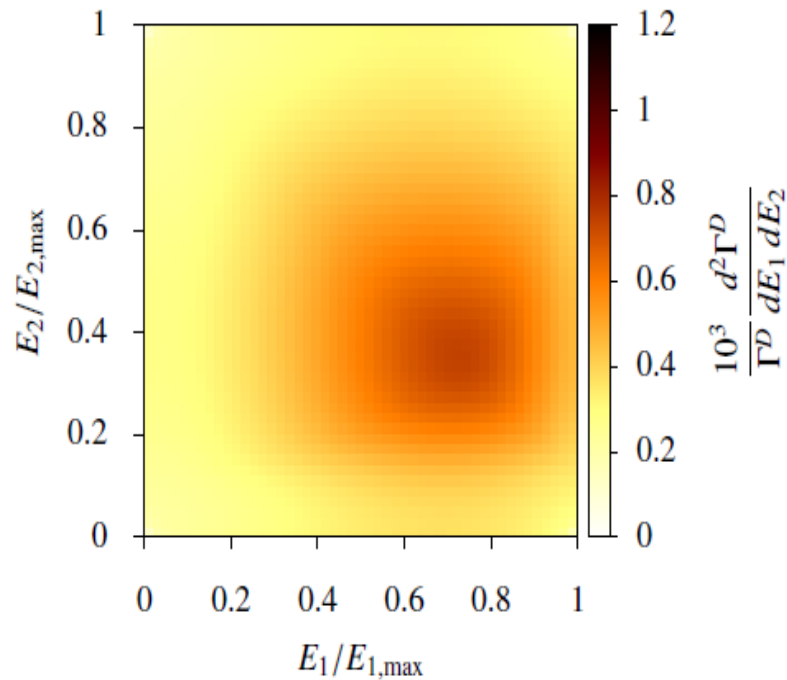


nu nubar

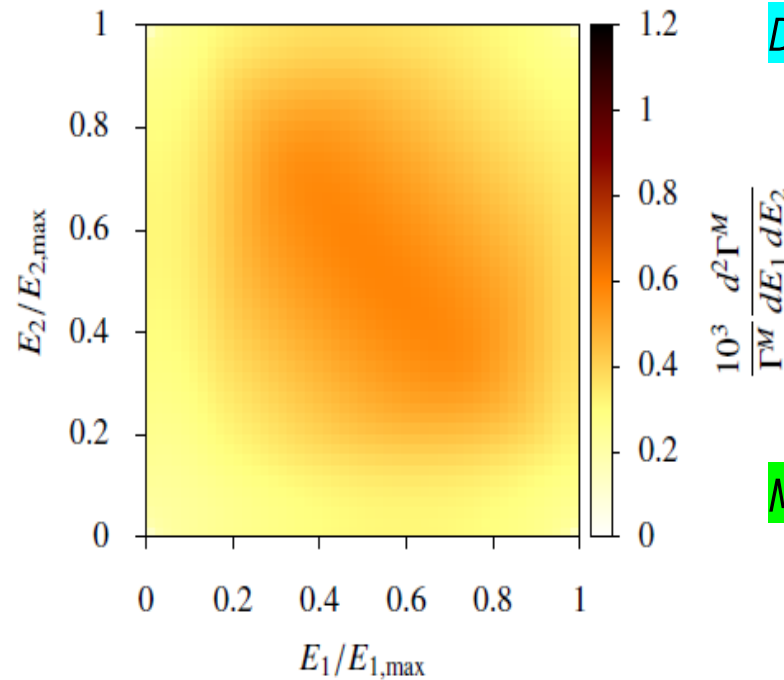
Different distribution, but the same total rate (DMCT)

$$\mathcal{M}^D = \mathcal{M}(p_1, p_2), \quad \mathcal{M}^M = \frac{1}{\sqrt{2}}(\mathcal{M}(p_1, p_2) - \mathcal{M}(p_2, p_1)).$$

Dirac case



Majorana case

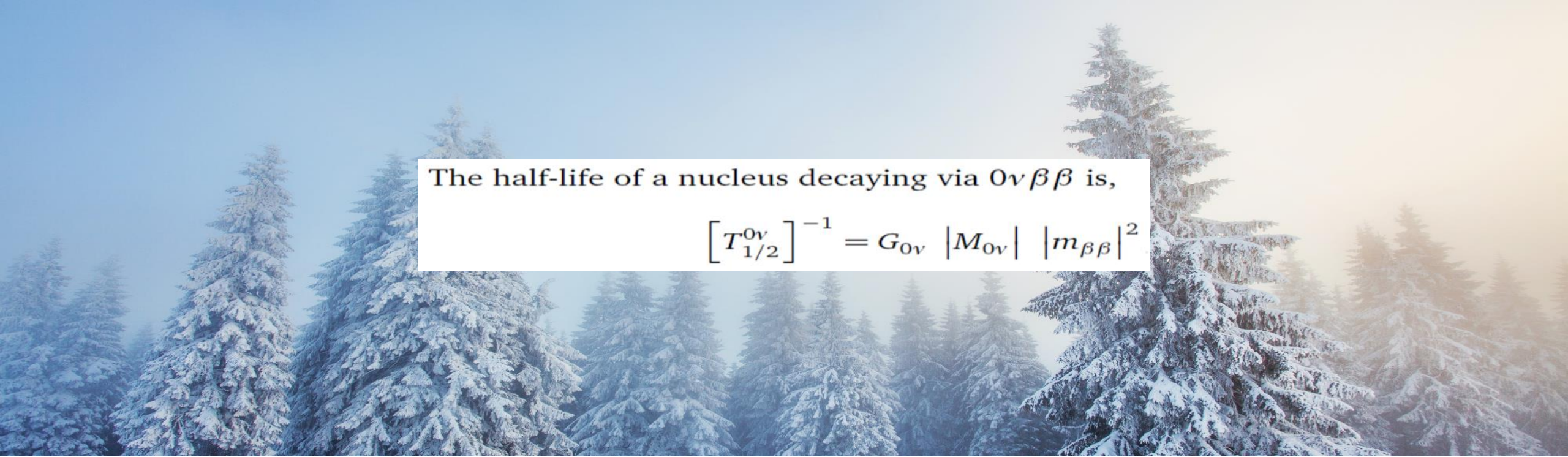


D =in general, no reason to be symmetric

M =always, must be symmetric

$$\underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} \neq \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}}.$$

$$\iint \left(|\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 \right) d^4 p_1 d^4 p_2 \propto m_v^2.$$



The half-life of a nucleus decaying via $0\nu\beta\beta$ is,

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{0\nu} \left| M_{0\nu} \right| \left| m_{\beta\beta} \right|^2$$

How to overcome pDMCT ?

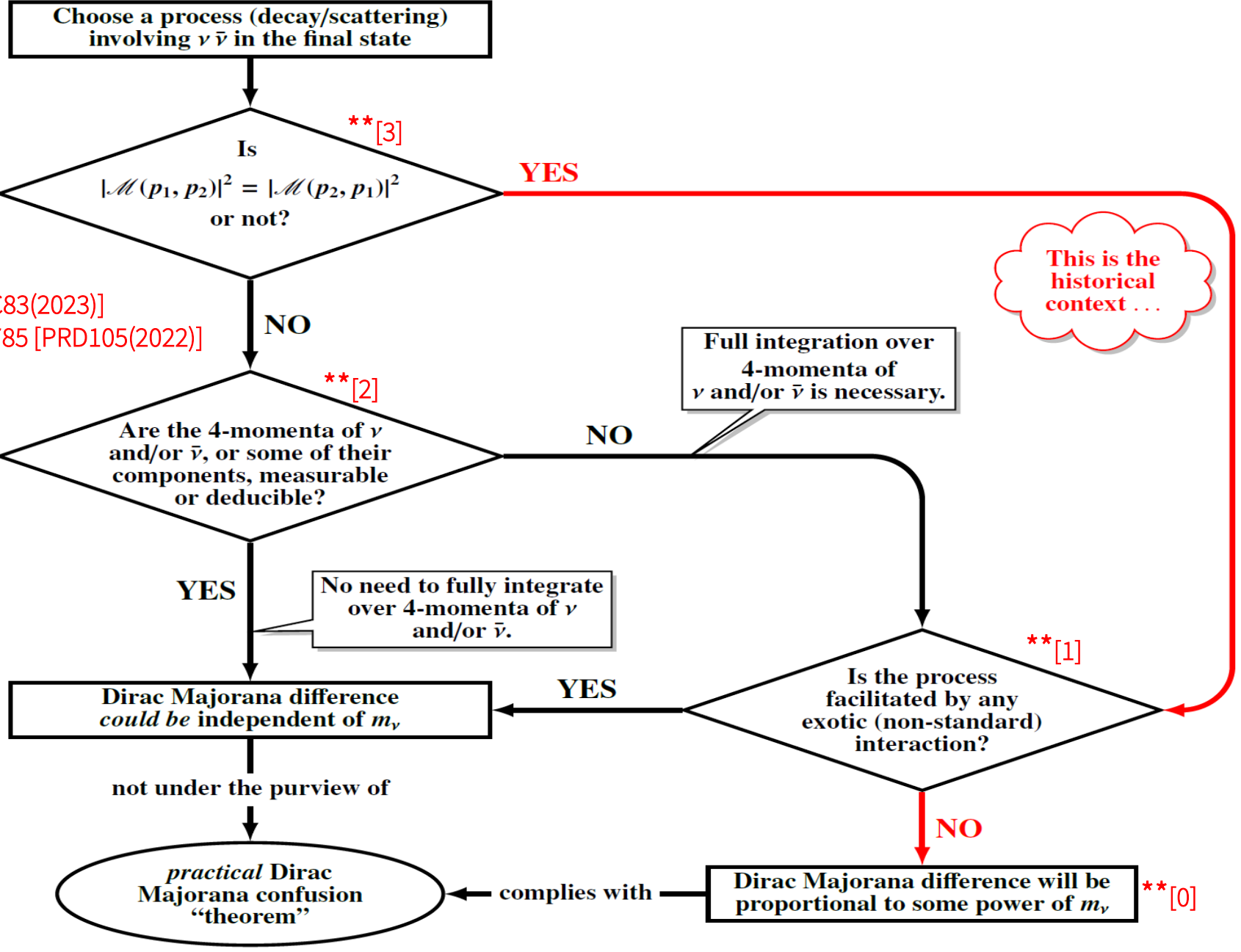
New physics effects in neutrino interaction (beyond SM)

[C. S.Kim, J.Rosiek, D.Sahoo, arXiv:2209.10110 [EPJC83 (2023)]; C. S.Kim, D.Sahoo, K.Vishnudath, arXiv:2405.17341 [EPJC84 (2024)]]

$B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2)$, and deduction of neutrino energy-momenta (within SM)

[C.S. Kim, M. Murthy, D. Sahoo, arXiv:2106.11785 [PRD105(2022)]; C.S. Kim, arXiv:2307.05654 [EPJC83(2023)]]

Issues in pDMCT (Analytic Continuity & F-D Statistics) [C.S. Kim, arXiv:2307.05654 [EPJC83(2023)]]



This is the historical context ...

[0] B.Kayser,PRD26(1982)
[1] S.P.Rosen,PRL48(1982); CSK etal. 2209.10110 [EPJC83(2023)]
[2] T.Chabra,PR.Babu,PRD46(1992); CSK etal. 2106.11785 [PRD105(2022)]
[3] CS Kim, 2307.05654 (2023) [EPJC83(2023)]

Extra 1 [C. S.Kim, J.Rosiek, D.Sahoo, arXiv:2209.10110 [EPJC83 (2023)];
C. S.Kim, D.Sahoo, K.Vishnudath, arXiv:2405.17341 [EPJC84 (2024)]]

New Physics Effects to distinguish Majorana from Dirac

General Comments on New Physics Scenario

[S.P.Rosen,PRL48(1982): W.Rodejohann et al,JHEP05(2017)]

Detailed study on $Z \rightarrow \nu_\ell \bar{\nu}_\ell$

Detailed study of $B \rightarrow K \nu \bar{\nu}, K \rightarrow \pi \nu \bar{\nu}$

Detailed study of $Z \rightarrow \nu_\ell \bar{\nu}_\ell$ (1)

Most general amplitude for Dirac neutrino

$$\mathcal{M}^D = \mathcal{M}(p_1, p_2) = -\frac{i g_Z}{2} \epsilon_\alpha \left[\bar{u}(p_1) \gamma^\alpha (C_V^\ell - C_A^\ell \gamma^5) v(p_2) \right],$$

$$C_{V,A}^\ell = \frac{1}{2} + \varepsilon_{V,A}^\ell, \quad \varepsilon_{V,A}^\ell = 0. \quad (\text{SM})$$

Then,

$$\mathcal{M}^M = \frac{1}{\sqrt{2}} (\mathcal{M}(p_1, p_2) - \mathcal{M}(p_2, p_1)) = \frac{i g_Z C_A^\ell}{\sqrt{2}} \epsilon_\alpha \left[\bar{u}(p_1) \gamma^\alpha \gamma^5 v(p_2) \right].$$

$$|\mathcal{M}^D|^2 = \frac{g_Z^2}{3} \left(((C_V^\ell)^2 + (C_A^\ell)^2) (m_Z^2 - m_\nu^2) + 3 ((C_V^\ell)^2 - (C_A^\ell)^2) m_\nu^2 \right),$$

$$|\mathcal{M}^M|^2 = \frac{2 g_Z^2 (C_A^\ell)^2}{3} (m_Z^2 - 4 m_\nu^2), \quad |\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 = \underbrace{\frac{1}{2} (|\mathcal{M}(p_1, p_2)|^2 - |\mathcal{M}(p_2, p_1)|^2)}_{\text{Direct term}} + \underbrace{\text{Re}(\mathcal{M}(p_1, p_2)^* \mathcal{M}(p_2, p_1))}_{\text{Exchange term}} + \underbrace{\text{Re}(\mathcal{M}(p_1, p_2)^* \mathcal{M}(p_2, p_1))}_{\text{Interference term}}.$$

such that

$$\begin{aligned} |\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 &= \frac{g_Z^2}{3} \left(((C_V^\ell)^2 - (C_A^\ell)^2) (m_Z^2 + 2 m_\nu^2) + 6 (C_A^\ell)^2 m_\nu^2 \right) \\ &= \begin{cases} \frac{g_Z^2}{2} m_\nu^2, & (\text{for SM alone}) \\ \frac{g_Z^2}{3} (\varepsilon_V^\ell - \varepsilon_A^\ell) m_Z^2, & (\text{with NP but neglecting } m_\nu) \end{cases} \end{aligned}$$

[DMCT at amplitude²]

[DMCT x with NP]

Detailed study of $Z \rightarrow \nu_\ell \bar{\nu}_\ell$ (2)

Therefore, neglecting neutrino mass,

$$\Gamma^D(Z \rightarrow \nu_\ell \bar{\nu}_\ell) = \Gamma_Z^0 (1 + 2\varepsilon_V^\ell + 2\varepsilon_A^\ell),$$

$$\Gamma^M(Z \rightarrow \nu_\ell \bar{\nu}_\ell) = \Gamma_Z^0 (1 + 4\varepsilon_A^\ell),$$

$$\Gamma_Z^0 = \frac{G_F m_Z^3}{12 \sqrt{2} \pi},$$

Then,

$$\Gamma_{Z,\text{inv}} = \begin{cases} \Gamma_Z^0 \left(3 + 2 \sum_{\ell=e,\mu,\tau} (\varepsilon_V^\ell + \varepsilon_A^\ell) \right), & \text{(for Dirac neutrinos)} \\ \Gamma_Z^0 \left(3 + 4 \sum_{\ell=e,\mu,\tau} \varepsilon_A^\ell \right). & \text{(for Majorana neutrinos)} \end{cases}$$

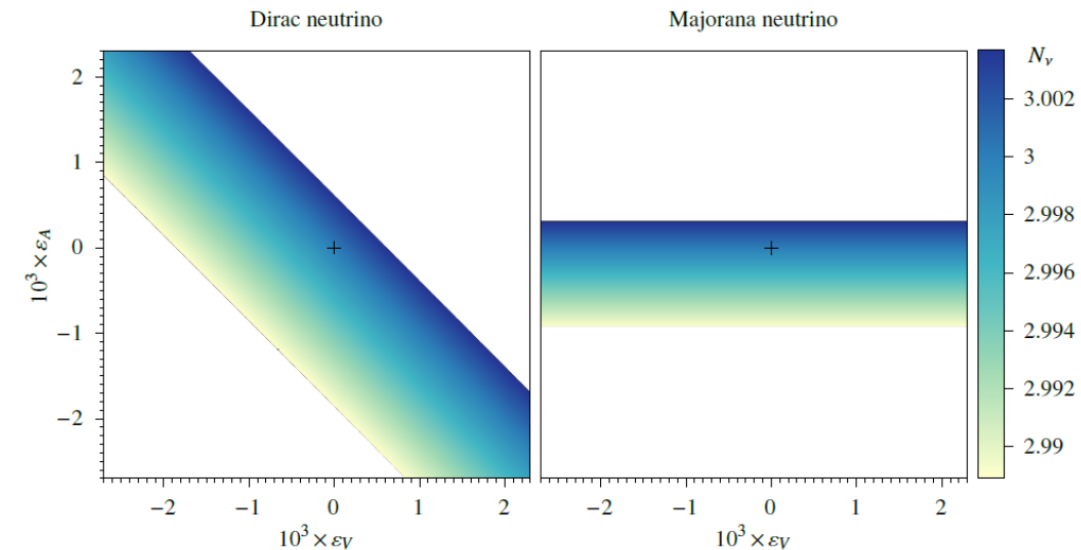
$$\varepsilon_{V,A}^e = \varepsilon_{V,A}^\mu = \varepsilon_{V,A}^\tau \equiv \varepsilon_{V,A}$$

$$N_\nu = \Gamma_{Z,\text{inv}} / \Gamma_Z^0 = 2.9963 \pm 0.0074.$$

[pdg(2021), PLB803(2020)]

$$\sum_{\ell=e,\mu,\tau} (\varepsilon_V^\ell + \varepsilon_A^\ell) = -0.0018 \pm 0.0037, \quad \text{(for Dirac neutrinos)}$$

$$\sum_{\ell=e,\mu,\tau} \varepsilon_A^\ell = -0.0009 \pm 0.0018, \quad \text{(for Majorana neutrinos)}$$





Extra 2 [C.S. Kim, M. Murthy, D. Sahoo, arXiv:2106.11785 [PRD105(2022)];
C.S. Kim, arXiv:2307.05654 [EPJC83(2023)]]

BACK-TO-BACK muons (ie. B2B $\nu - \bar{\nu}$)
in $B^0(p_B) \rightarrow \mu^-(p_-)\mu^+(p_+)\bar{\nu}_\mu(p_1)\nu_\mu(p_2)$, decay.

$$\underbrace{|\mathcal{M}(p_1, p_2)|^2}_{\text{Direct term}} \neq \underbrace{|\mathcal{M}(p_2, p_1)|^2}_{\text{Exchange term}} .$$

Doublyweak charged current process in SM

$B^0(B_s) \rightarrow \mu\mu(Z^* \rightarrow \nu\nu)$
not good, too small BR (FCNC)

- Exception to pDMCT within the SM

Back-to-back muons, (easily measurable exception to pDMCT)

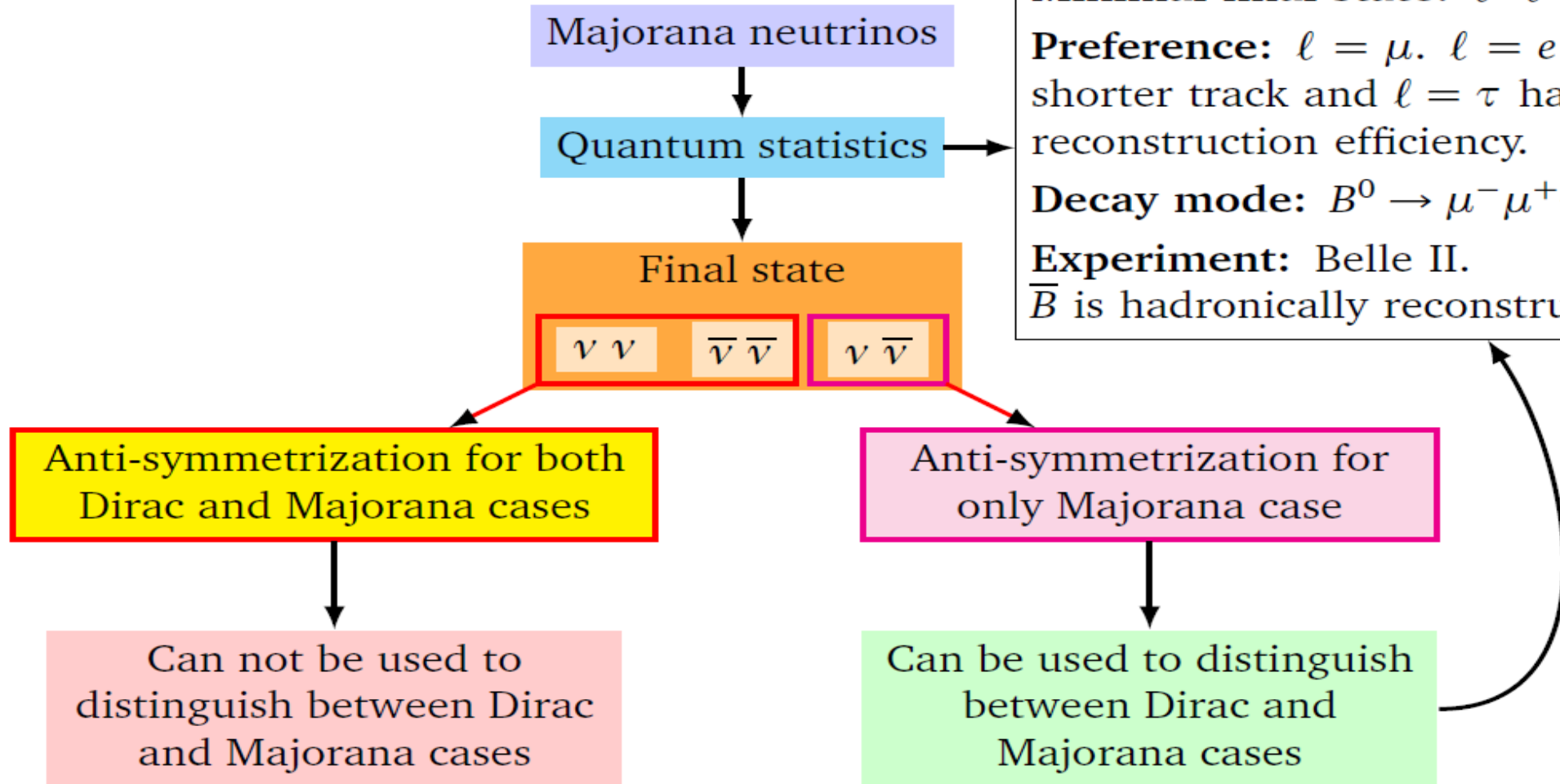
[CSK,MM,DS, 2106.11785, PRD105,113005(2022)]

[CS Kim, 2307.05654, EPJC83(2023)]

Allowed within SM

Minimal final state: $\ell^- \ell^+ \nu_\ell \bar{\nu}_\ell$
Preference: $\ell = \mu$. $\ell = e$ has shorter track and $\ell = \tau$ has low reconstruction efficiency.
Decay mode: $B^0 \rightarrow \mu^- \mu^+ \nu_\mu \bar{\nu}_\mu$
Experiment: Belle II.
 $\bar{B}$ is hadronically reconstructed.

(1 to 4 body process)
Many observables to deduce $E(\nu)$!!



$B^0 (B_s) \rightarrow \mu \mu (Z^* \rightarrow \nu \nu)$
not good, too small BR (FCNC)

Back-to-back muons, (easily measurable exception to pDMCT)

$$B^0(p_B) \rightarrow \mu^-(p_-)\mu^+(p_+)\bar{\nu}_\mu(p_1)\nu_\mu(p_2),$$

In the rest frame
 IF muon- and m
 → nu and nu-b

$$E_1 = E_2$$

$$m_{\mu\mu}^2$$

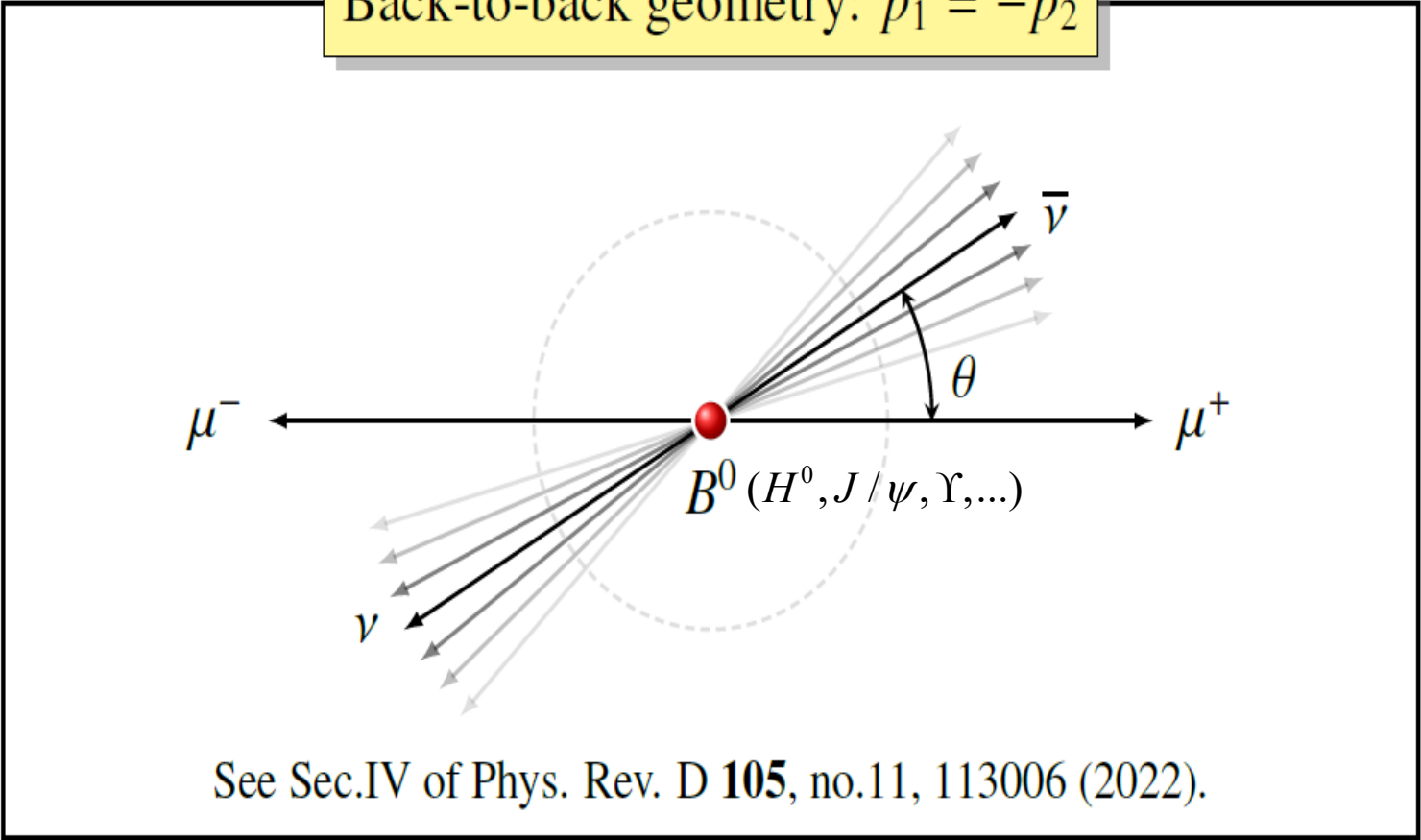
$$Y_m = \sqrt{($$

$$Y_n$$

Need not

only unmeasurable angle (θ) integrate out

Back-to-back geometry: $\vec{p}_1 = -\vec{p}_2$



t opposite direction

ble or measurable,

$\bar{\nu}$ and $\mu_- - \mu_+$)

$$\frac{d\Gamma}{dE_\nu^2 d\sin\theta}$$

Overcoming DMCT constraint

Back-to-back muons, (easily measurable exception to DMCT)

$$B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2),$$

In the rest frame of parent B meson,

IF muon- and muon+ are back-to-back, ie. flying with 3 momenta of equal magnitude but opposite direction

→ ν and $\bar{\nu}$ also back-to-back

$$E_1 = E_2 = E_\nu = m_B/2 - E_\mu$$

$$m_{\nu\nu}^2 = 4E_\nu^2$$

$$m_{\mu\mu}^2 = (m_B - 2E_\nu)^2$$

$$Y_m = \sqrt{(m_B/2 - E_\nu)^2 - m_\mu^2}$$

$$Y_n = \sqrt{E_\nu^2 - m_\nu^2}$$

→ All kinematic variables are calculable or measurable,

Only the angle (between $\nu - \bar{\nu}$ and $\mu_- - \mu_+$)
UNKNOWN

$$\frac{d\Gamma}{dE_\mu^2 d\sin\theta}$$



$$\frac{d\Gamma}{dE_\nu^2 d\sin\theta}$$

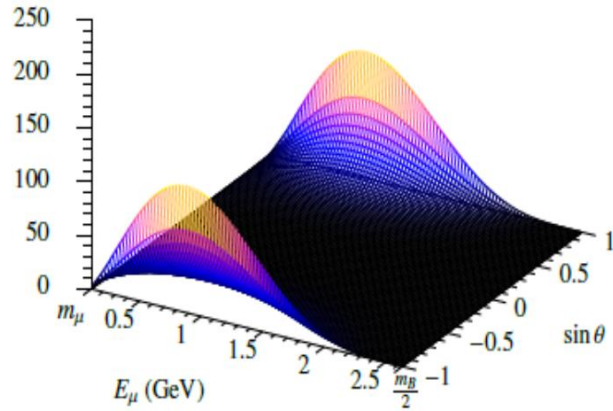
Need not integrate out ν - $\bar{\nu}$ full phase space,
only unmeasurable angle (θ) integrate out

→ Overcoming DMCT constraint

Back-to-back muons in

$$B^0(p_B) \rightarrow \mu^-(p_-)\mu^+(p_+)\bar{\nu}_\mu(p_1)\nu_\mu(p_2),$$

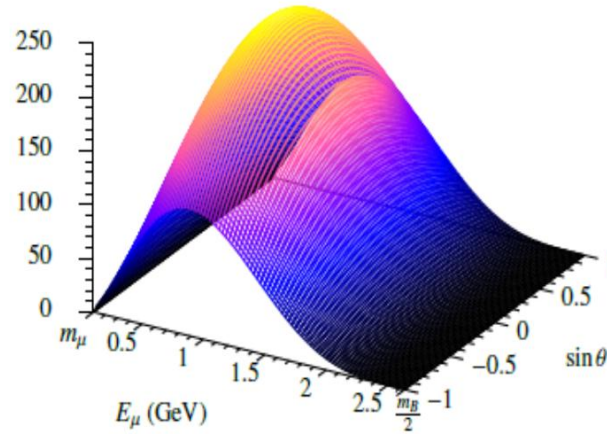
$$\frac{512\pi^6 m_B}{G_F^4 |\mathbb{F}_d|^2} \frac{d^3\Gamma_{\leftrightarrow}^D}{dE_\mu^2 d\sin\theta} (\text{GeV}^6)$$



(a) Three dimensional view of the differential decay rate for Dirac case with an appropriate normalization as mentioned.

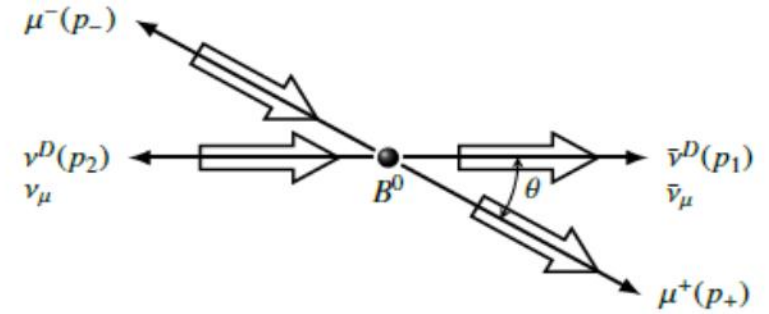
$$|\mathcal{M}_{\leftrightarrow}^D|^2 \propto \underbrace{(1 - \cos\theta)^2}_{\text{Direct term}}.$$

$$\frac{512\pi^6 m_B}{G_F^4 |\mathbb{F}_d|^2} \frac{d^3\Gamma_{\leftrightarrow}^M}{dE_\mu^2 d\sin\theta} (\text{GeV}^6)$$

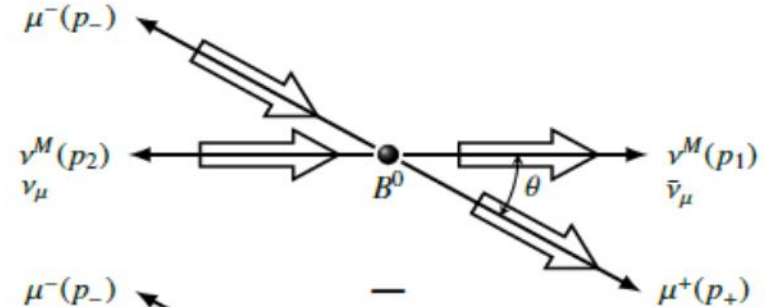


(b) Three dimensional view of the differential decay rate for Majorana case with an appropriate normalization as mentioned.

$$|\mathcal{M}_{\leftrightarrow}^M|^2 \propto \frac{1}{2} \left[\underbrace{(1 - \cos\theta)^2}_{\text{Direct term}} + \underbrace{(1 - \cos(\pi - \theta))^2}_{\text{Exchange term}} - \underbrace{\mathcal{O}(m_\nu^2)}_{\text{Interference term}} \right] \simeq 1 + \cos^2\theta.$$



(a) Helicity configuration involving Dirac neutrinos, $\nu_\mu \equiv \nu^D$, $\bar{\nu}_\mu \equiv \bar{\nu}^D$.



(b) Helicity configuration involving Majorana neutrinos, $\nu_\mu = \bar{\nu}_\mu \equiv \nu^M$.

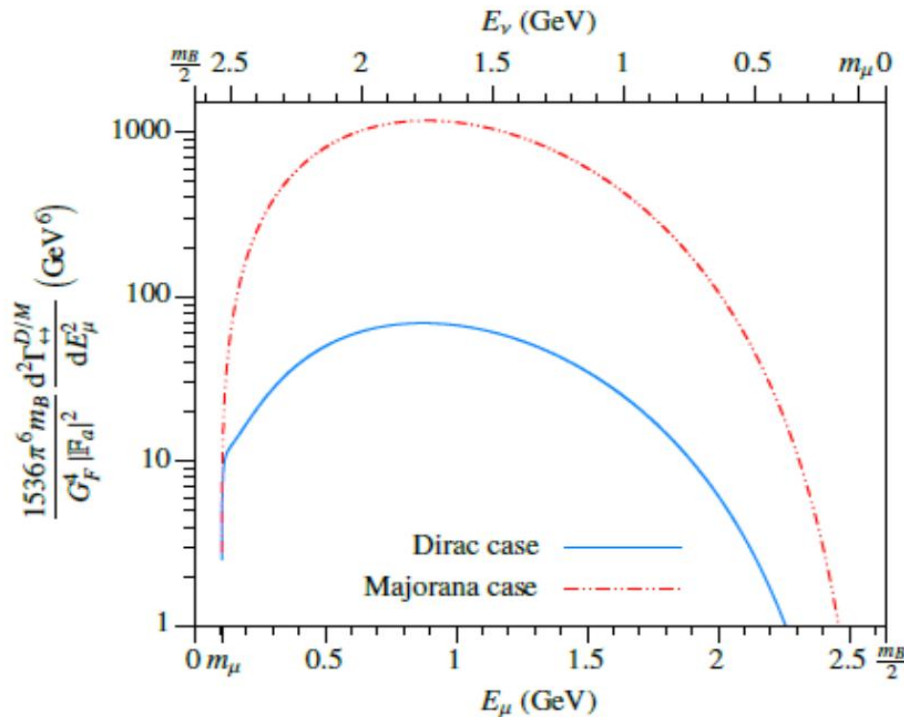
Back-to-back muons in

$$B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2),$$

$$\frac{d^3\Gamma_{\leftrightarrow}^D}{dE_\mu^2 d\sin\theta} = \frac{G_F^4 |\mathbb{F}_a|^2 (m_B - 2E_\mu)^4 K_\mu}{512 \pi^6 m_B E_\mu} (E_\mu - K_\mu \cos\theta)^2,$$

$$\frac{d^3\Gamma_{\leftrightarrow}^M}{dE_\mu^2 d\sin\theta} = \frac{G_F^4 |\mathbb{F}_a|^2 (m_B - 2E_\mu)^4 K_\mu}{512 \pi^6 m_B E_\mu} (E_\mu^2 + K_\mu^2 \cos^2\theta),$$

$$K_\mu = \sqrt{E_\mu^2 - m_\mu^2}.$$



(c) Comparison of muon energy distributions between Dirac and Majorana cases in the back-to-back scenario.

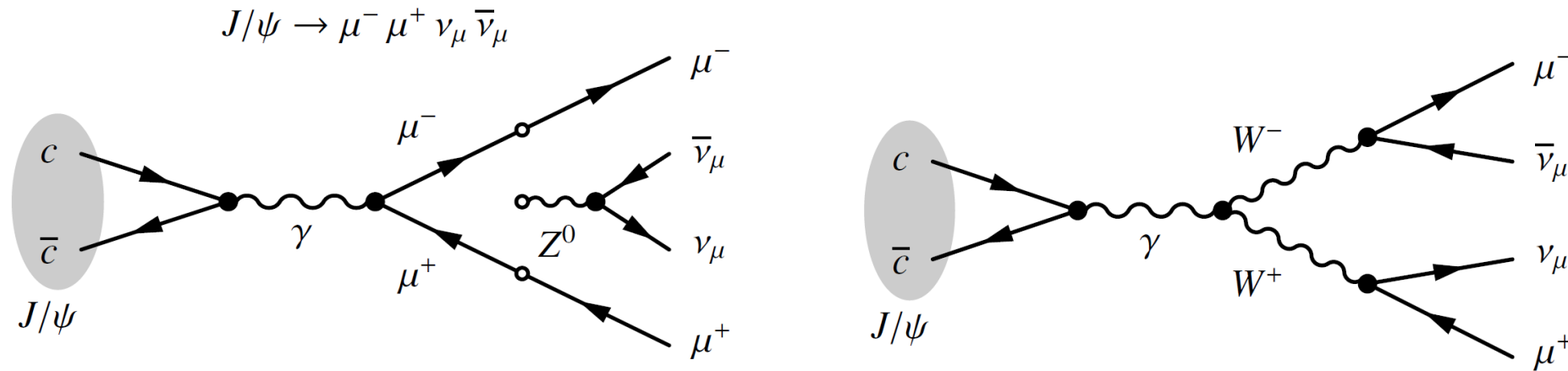
$$BR(D - M) = \int (D_{\leftrightarrow} - M_{\leftrightarrow}) \neq m_\nu^2$$

$$\mathcal{B}_{\leftrightarrow}^D = \Gamma_{\leftrightarrow}^D / \Gamma_B \approx 1.1 \times 10^{-12} \text{ GeV}^{-2} \times |\mathbb{F}_a|^2,$$

$$\mathcal{B}_{\leftrightarrow}^M = \Gamma_{\leftrightarrow}^M / \Gamma_B \approx 1.8 \times 10^{-11} \text{ GeV}^{-2} \times |\mathbb{F}_a|^2,$$

Discussions on $J/\psi \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ -- B2B muons (1)

Cases for $J/\psi \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ at BES III, $\Upsilon(1s) \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ at Belle2



** Present measurements:

$$Br(B^0 \rightarrow \mu^+ \mu^-) = (1.1 \pm 1.4) * 10^{-10}$$

$$B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2),$$

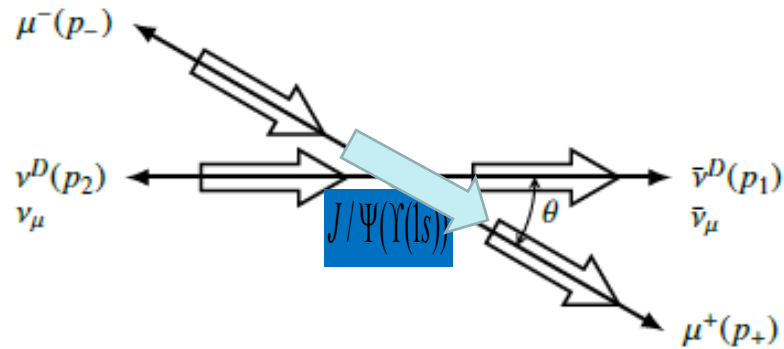
FCNC decay via photonic Penguin
Doubly weak charged current decay

$$Br(J/\Psi \rightarrow \mu^+ \mu^-) = (5.961 \pm 0.033)\%$$

$$Br(\Upsilon(1s) \rightarrow \mu^+ \mu^-) = (2.48 \pm 0.05)\%$$

Discussions on $J/\psi \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ -- B2B muons (2)

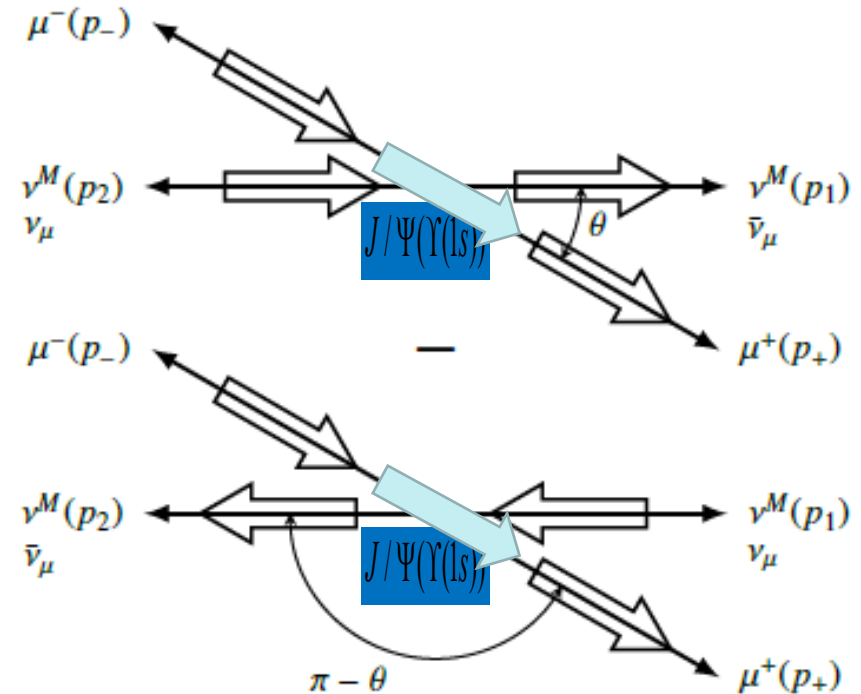
Cases for $J/\psi \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ at BES III, $\Upsilon(1s) \rightarrow \mu^+ \mu^- \nu_\mu \bar{\nu}_\mu$ at Belle2
 ** (simple rough) Helicity Consideration



(a) Helicity configuration involving Dirac neutrinos, $\nu_\mu \equiv \nu_\mu^D$, $\bar{\nu}_\mu \equiv \bar{\nu}_\mu^D$.

$$|\mathfrak{M}_{\leftrightarrow}^D|^2 \propto (1 + \cos \theta)^2$$

$$|\mathfrak{M}_{\leftrightarrow}^M|^2 \propto 1 + \cos^2 \theta$$



(b) Helicity configuration involving Majorana neutrinos, $\nu_\mu = \bar{\nu}_\mu \equiv \nu_\mu^M$.

Extra 3

[C.S. Kim, arXiv:2307.05654 [EPJC83(2023)]]



- Issues in pDMCT (Analytic Continuity & F-D statistics)

1. Analytic Continuity

Is there any one-to-one correspondence between Dirac and Majorana neutrinos in the massless limit, $m_\nu \rightarrow 0$?

2. Anti-Symmetrization of Amplitude

Should the amplitude be anti-symmetrized for pair of Majorana neutrinos of the same flavor ?

Analytic Continuity when $m \rightarrow 0$ limit ? (1)

****** Is there smooth transition between Majorana to Dirac neutrinos under $m \rightarrow 0$ limit ??

1) In the context of a *specific process and observable*,
there is indeed *no one-to-one correspondence between the Dirac and Majorana neutrino scenarios*.

E.g. For $Z \rightarrow \nu_\ell \bar{\nu}_\ell$

$$\begin{aligned} |\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 &= \frac{g_Z^2}{3} \left(((C_V^\ell)^2 - (C_A^\ell)^2) (m_Z^2 + 2m_\nu^2) + 6(C_A^\ell)^2 m_\nu^2 \right) \\ &= \begin{cases} \frac{g_Z^2}{2} m_\nu^2, & \text{(for SM alone)} \\ \frac{g_Z^2}{3} (\varepsilon_V^\ell - \varepsilon_A^\ell) m_Z^2, & \text{(with NP but neglecting } m_\nu) \end{cases} \end{aligned}$$

$$C_{V,A}^\ell = \frac{1}{2} + \varepsilon_{V,A}^\ell, \quad \varepsilon_{V,A}^\ell = 0. \text{ (SM)}$$

Even within SM, doubly weak charged current decay processes, such as $B^0 \rightarrow \mu^- \mu^+ \nu_\mu \bar{\nu}_\mu$, give

$$|\mathcal{M}^D|^2 - |\mathcal{M}^M|^2 \not\propto m_\nu^2.$$

[C.S. Kim et al, arXiv:2106.11785 [PRD105(2022)]

Analytic Continuity when $m \rightarrow 0$ limit ? (2)

**** Is there smooth transition** between Majorana to Dirac neutrinos **under $m \rightarrow 0$ limit ??**

2) Now assuming Standard Model Lagrangian, due to the (V-A) structure of weak interaction, always **only** $\nu_L, \bar{\nu}_R$ are produced (and detected).

Assuming $mass(\nu) = 0$, chirality of neutrinos is conserved and invariant.

And $\nu_L, \bar{\nu}_R$ can be distinguished by weak charged current interaction.

→ **massless** $\nu_L, \bar{\nu}_R$ are **Dirac neutrino** (No mass-less Majorana neutrino exists).

3) Assuming $mass(\nu) \neq 0$, chirality of neutrinos is not conserved (still invariant), as $m \rightarrow 0$.

space-time evolution

$$\begin{aligned} i \gamma^\mu \partial_\mu \psi_R &= m \psi_L, \\ i \gamma^\mu \partial_\mu \psi_L &= m \psi_R. \end{aligned} \quad \psi_L \rightarrow \begin{pmatrix} \frac{m}{E} \psi_R \\ \psi_L \end{pmatrix} \quad \begin{cases} m \rightarrow 0 \\ E \neq 0 \end{cases} \quad \begin{cases} m \rightarrow 0 \\ E \rightarrow 0 \end{cases}$$

(relevant parameter, not m , but $m/E \rightarrow 0$)

$$\frac{m}{E} \rightarrow 0$$

$$\frac{m}{E} \sim \text{finite}$$

Analytic Continuity when $m \rightarrow 0$ limit ? (3)

****** Is there smooth transition between Majorana to Dirac neutrinos under $m \rightarrow 0$ limit ??

4) Most importantly, **Quantum Statistics** **DEPENDS** on **Spin & Identicalness of neutrino**,
not DEPENDS on **Mass of neutrino** !!

E.g. For $B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2)$,

if ν = Dirac $\mathcal{M}^D = \mathcal{M}(p_1, p_2)$,

if ν = Majorana $\mathcal{M}^M = \frac{1}{\sqrt{2}}(\mathcal{M}(p_1, p_2) - \mathcal{M}(p_2, p_1))$. (for $m_\nu = 0$ or $m_\nu \neq 0$)

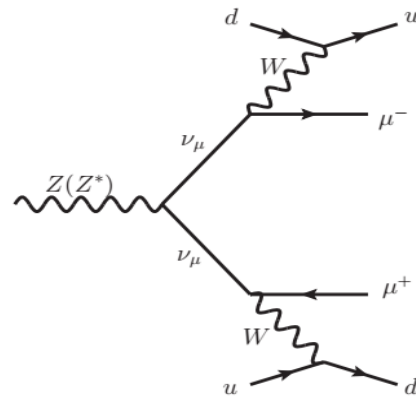
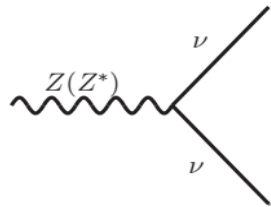
Therefore, even if $\Gamma_D = \Gamma_M(m_\nu \rightarrow 0)$, possibly $d\Gamma_D \neq d\Gamma_M(m_\nu \rightarrow 0)$

→ If we can deduce e.g. E_ν , we can find $\frac{d\Gamma_D}{dE_\nu} \neq \frac{d\Gamma_M}{dE_\nu}$ independent of m_ν

Anti-symmetrization of Amplitude : F-D Statistics

****** Should the amplitude be anti-symmetrized for pair of Majorana neutrinos of the same flavor ??

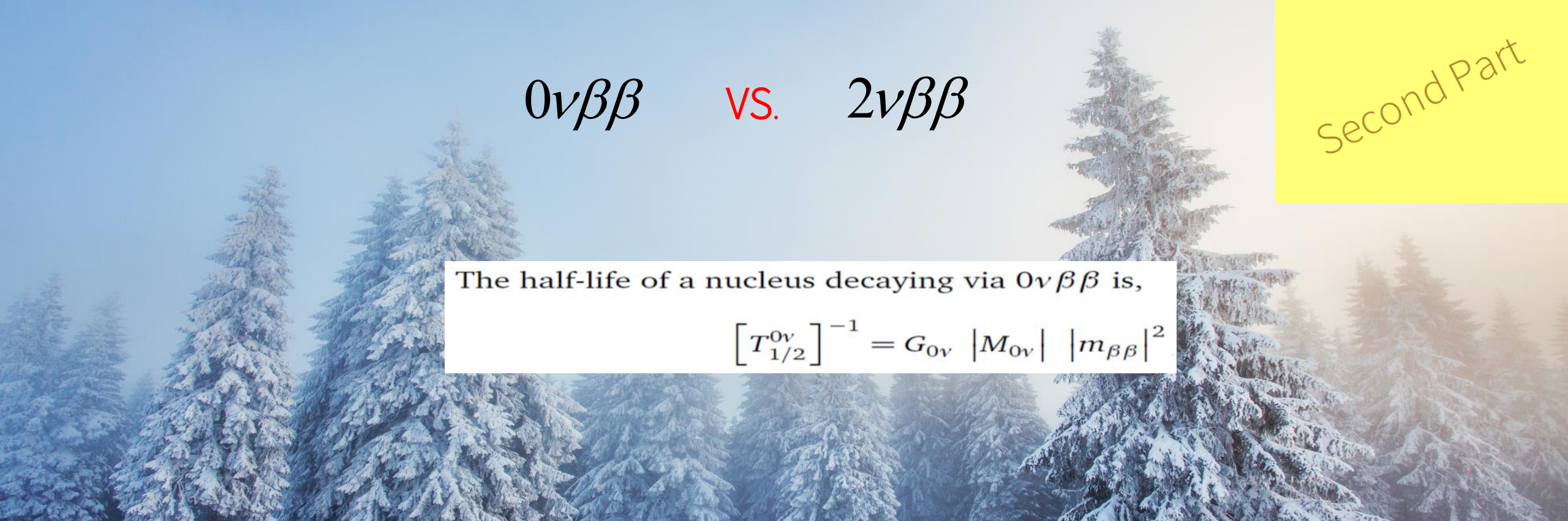
- 1) Quantum statistics requires **absolutely identical indistinguishable particles**.
- 2) **Quantum measurement effect** is analogous to the well-known fact that the interference pattern in a double slit experiment is destroyed if one could somehow know through which slit the photon has passed.



→ requires a method deducing E (or p) of neutrino without destroying quantum effects.

Need **Anti-symmetrization**

No **Anti-symmetrization** (→ similar to the double slit experiment)
once we know $\nu_1 \neq \nu_2$


$$0\nu\beta\beta \quad \text{vs.} \quad 2\nu\beta\beta$$

Second Part

The half-life of a nucleus decaying via $0\nu\beta\beta$ is,

$$\left[T_{1/2}^{0\nu} \right]^{-1} = G_{0\nu} \left| M_{0\nu} \right| \left| m_{\beta\beta} \right|^2$$

Alternatives to $0\nu\beta\beta$ ($\Delta L = 2$ process)

Neutrino Casimir Force

EDM & MDM of neutrinos

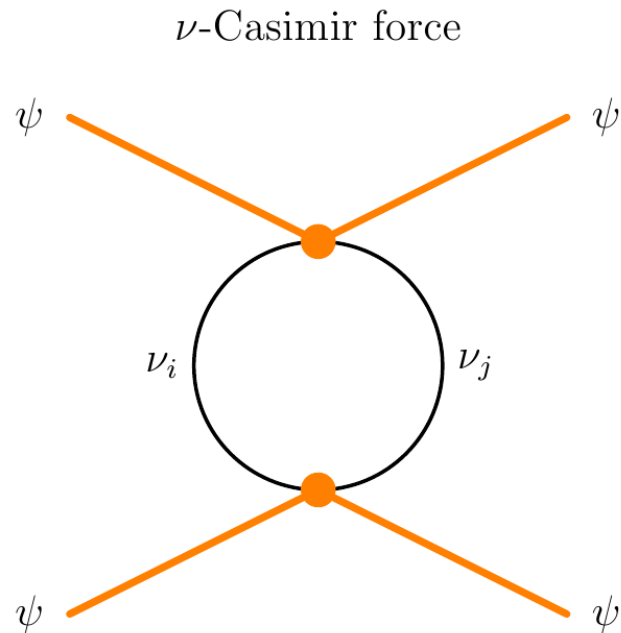
Fermi-Dirac Statistics for Fermion (ν) – Quantum Statistics

Quantum Statistics + ($2\nu\beta\beta$)

Alternative to 0nuBB (1) – Neutrino Casimir force

Principle: Exchange of pair of neutrinos can give rise to long-range quantum force (aka **neutrino Casimir force or the neutrino exchange force**) between macroscopic objects, and **the effective potential can differentiate Dirac and Majorana neutrinos.**

G Feinberg, J Sucher, PRD166(1968)
Phys. Rev. D 101, no.11, 116006 (2020); JHEP 09, 122 (2020)
arXiv:2209.07082 [hep-ph]



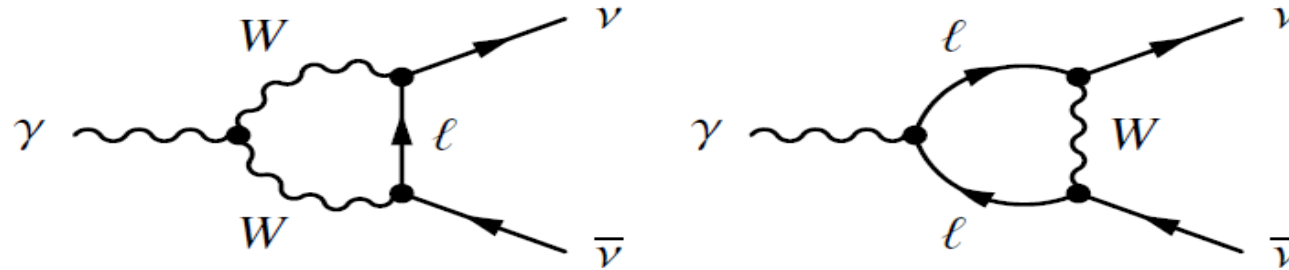
Issue: The potential (and hence the force) is proportional to product of the tiny neutrino masses in the loop.

- ** Thermal fluctuation, van der Waals force
- ** Very weak force. For $r > 1$ nm, Gravitational force between two protons is bigger than this force.

Status: Experimental study is still awaited.

Alternative to 0nuBB (2) – e.d.m. & m.d.m. of Neutrino

- ❖ Neutrinos: electrically neutral ($Q_\nu = 0$).
But might carry non-zero electric and magnetic dipole moments.
- ❖ Neutrino has no interaction with photon at tree-level.
At one loop-level involving charged lepton and W boson in loop such interaction is allowed.



- ❖ CPT invariance prohibits Majorana neutrinos to have any electric and magnetic dipole moment. Loop effects also confirm this, following quantum statistics in $\gamma^* \rightarrow \nu \bar{\nu}$.

$$\text{edm} = 0 = \text{mdm}, \quad (\text{Majorana})$$

$$\text{edm} \neq 0 \neq \text{mdm}. \quad (\text{Dirac})$$

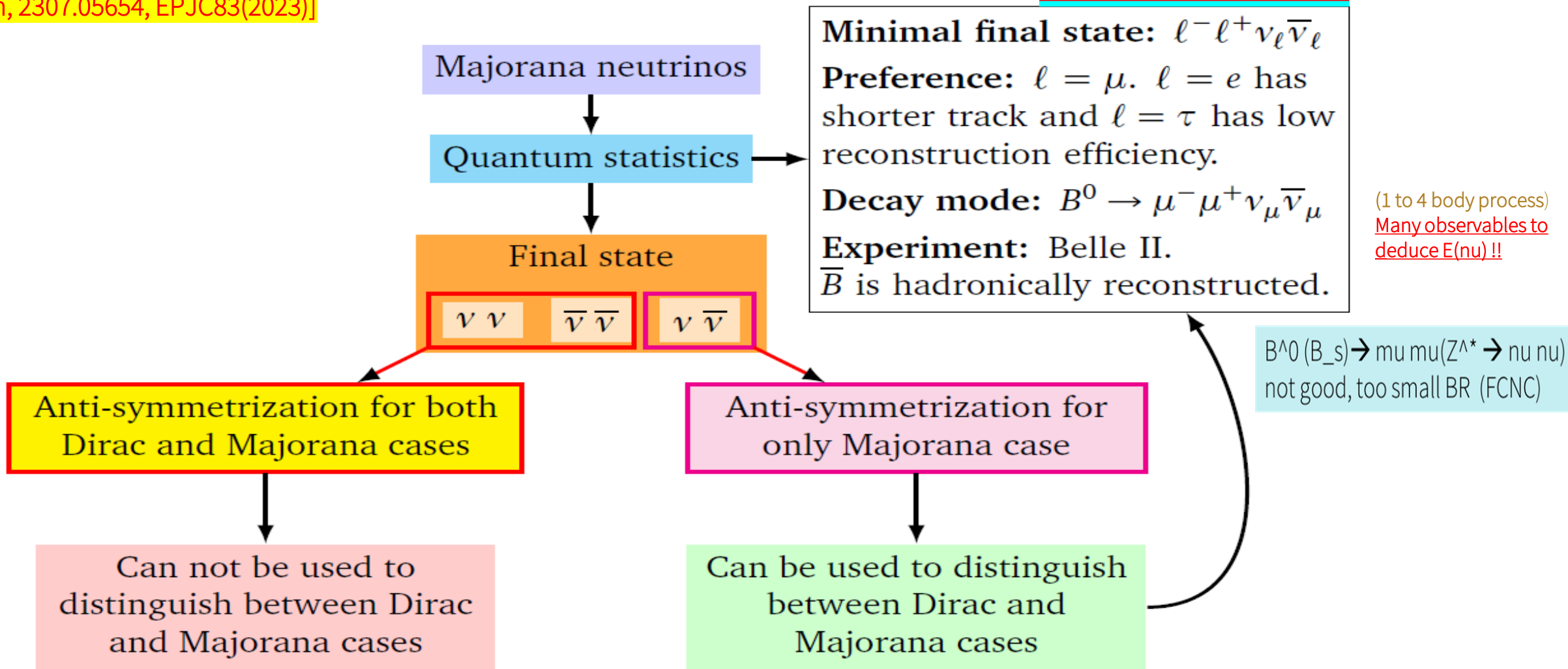
- ❖ Such striking differences in electromagnetic properties of Dirac and Majorana neutrinos vanishes as $m_\nu \rightarrow 0 \implies \text{DMCT}$.
[B. Kayser, Phys. Rev. D **26**, 1662 (1982).]

Alternative to 0nuBB (3) – Quantum Statistics (best option)

[CSK,MM,DS, 2106.11785, PRD105,113005(2022)]

[CS Kim, 2307.05654, EPJC83(2023)]

Allowed within SM



Alternative to 0nuBB (4) – Quantum Statistics + (2νββ)

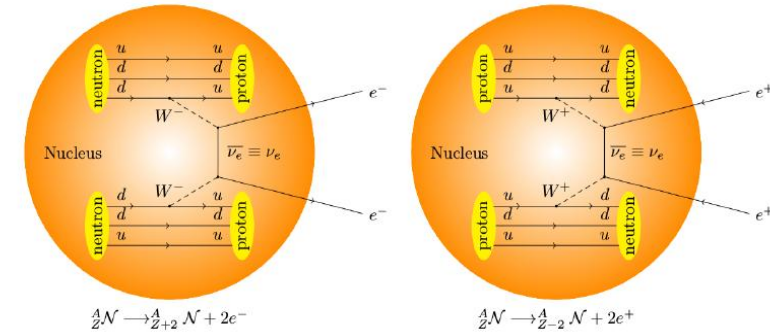
$0\nu\beta\beta$ is only for the evidence of “Lepton Number Violation”,

→ NOT unequivocal evidence of active (sub-eV) neutrino being Majorana neutrino.

eg. heavy sterile neutrinos can give the same effects

HOWEVER, $2\nu\beta^+\beta^-$ ($\nu\bar{\nu}$ not $\nu\nu, \bar{\nu}\bar{\nu}$) can be w/ Quantum Statistics !!

❖ Process:



$${}^A_Z N \rightarrow {}^A_{Z+2} N + 2e^-$$

$${}^A_Z N \rightarrow {}^A_{Z+2} N + 2e^- + (2\bar{\nu})$$

$${}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$$



$$e^- + {}^A_Z N \rightarrow {}^A_Z N + e^- + (\bar{\nu} + \nu)$$



$$e^+ + {}^A_Z N \rightarrow {}^A_Z N + e^+ + (\bar{\nu} + \nu)$$



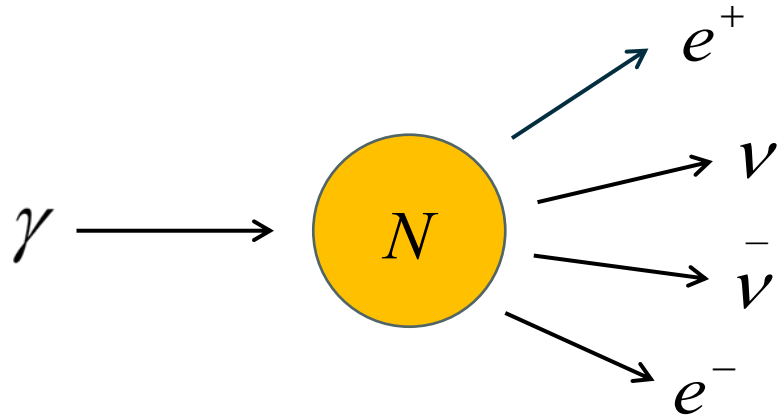
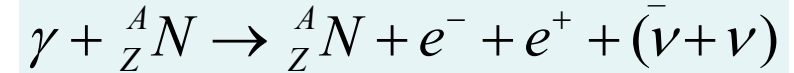
$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$$



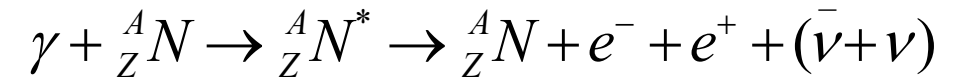
Alternative to 0nuBB (5) – Quantum Statistics + $(2\nu\beta\beta)$

HOWEVER, $2\nu\beta^+\beta^-$ ($\nu\bar{\nu}$ not $\nu\nu, \bar{\nu}\bar{\nu}$)
can be w/ Quantum Statistics !!

γ = (squeezed) Gamma-ray Laser Photon
[Compton back-scattered gamma ray]



$$E(\gamma) \geq M({}^A_ZN^*) - M({}^A_ZN)$$



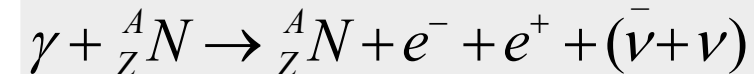
$$1.02\text{MeV} < E(\gamma) < M({}^A_ZN^*) - M({}^A_ZN)$$

$$E_\gamma = 3 \sim 5\text{MeV}$$

$$\lambda(\gamma) \sim 100\text{fm} \quad \lambda(N) = 10 \sim 100\text{fm} \quad \lambda(p,n) \sim 1\text{fm}$$

$$[\gamma + N \rightarrow N][N \rightarrow N + e^+e^-\nu\bar{\nu}]$$

$$[\gamma + N \rightarrow N][\gamma^* \rightarrow e^+e^-\nu\bar{\nu}]$$



Extra 4 []

- Issues in the process

$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$$

0. 2 Possible Processes (Double Beta vs. Electro-Magnetic X $Z^* \rightarrow \nu \nu$)

1. Scales and factorization

Physics Scale of gamma to N, and is the process factorizable? $\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$?

2. Probability of $\Gamma(\gamma + N \rightarrow N) \propto \text{Prob.}(\gamma N)$

3. Electro-magnetic X weak neutral

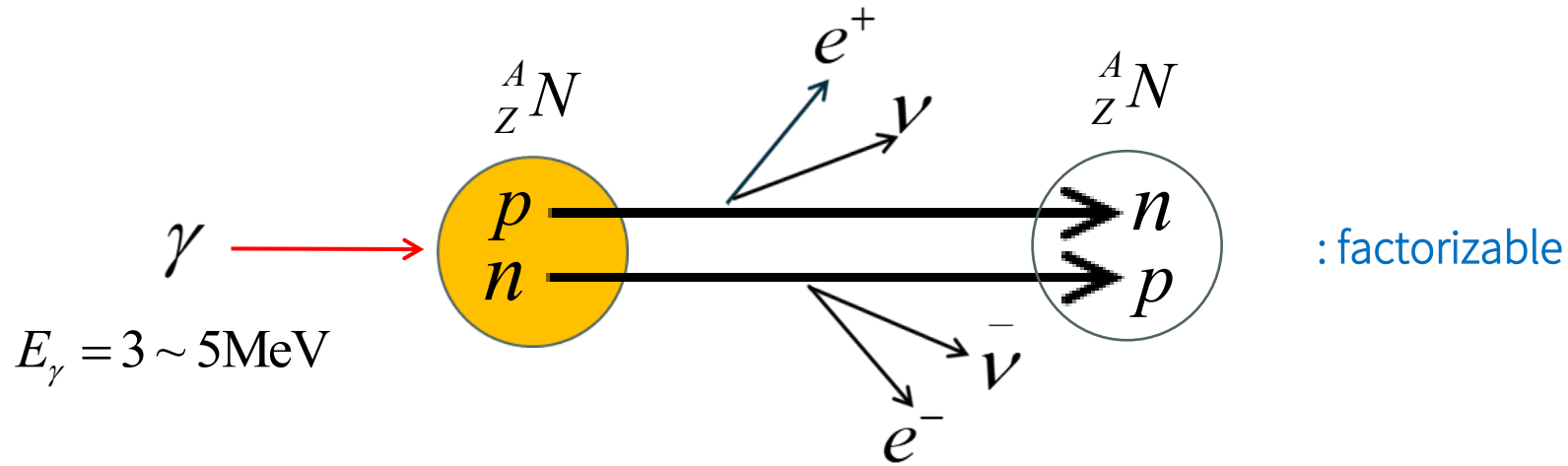
$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^+ + e^- + (\bar{\nu} + \nu)$$

3.1 Bethe-Heitler (BH) vs. Virtual Compton Scattering (VCS) of $e^+ e^-$ production

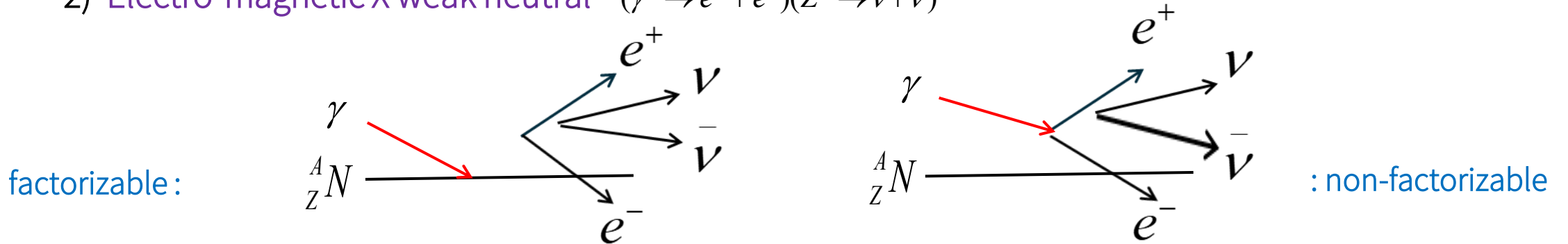
4. Is the final N moving as recoil motion ?

0. 2 Possible Processes, $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$

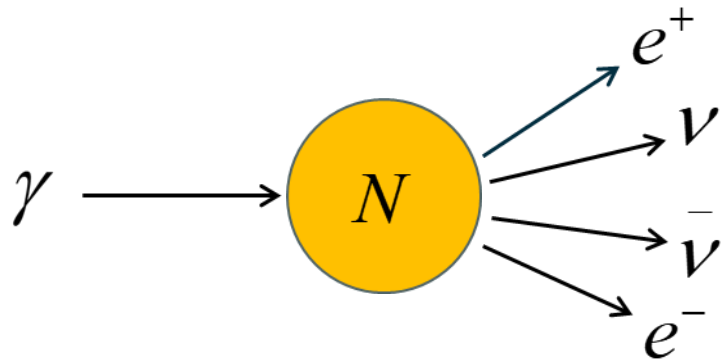
1) Double beta decays of energetic N



2) Electro-magnetic X weak neutral $(\gamma^* \rightarrow e^- + e^+)(Z^* \rightarrow \bar{\nu} + \nu)$



1. Scales and Factorization: $\Gamma(\gamma + {}^A_ZN \rightarrow {}^A_ZN + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_ZN \rightarrow {}^A_ZN + \dots)$



$$E_\gamma = 3 \sim 5 \text{ MeV}$$

$$\lambda(\gamma) \sim 100 \text{ fm} \quad \lambda(p, n) \sim 1 \text{ fm}$$



$$\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$$

(eg.) For double beta decay process:

Incoming photon **CANNOT** interact DIRECTLY to (n,p) of N,
nor to (e e nu nu) in beta decay process of (n,p) !!

$$1.02 \text{ MeV} < E(\gamma) < M({}^A_ZN^*) - M({}^A_ZN)$$

$$\gamma + {}^A_ZN \rightarrow (\text{energetic}) {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$$

$$\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$$



$$[\gamma + N \rightarrow N][N \rightarrow N + e e \nu \bar{\nu}] \quad \text{if factorized}$$

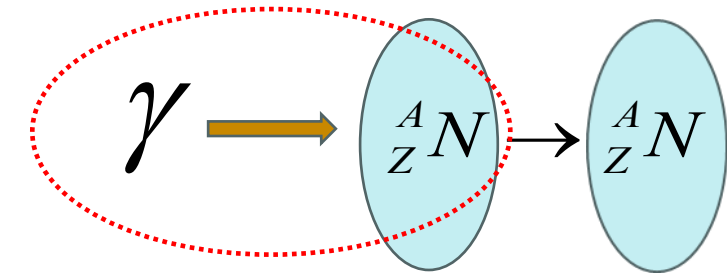
**** factorized due to MeV scale ($\gamma + N \rightarrow N$) and GeV scale ($n \rightarrow p + e^- + \bar{\nu}$)**

$$[\gamma + N \rightarrow N][N \rightarrow N + e e \nu \bar{\nu}]$$

$$[\gamma + N \rightarrow N][\gamma^* \rightarrow e e \nu \bar{\nu}]$$

$$\Gamma(\gamma + {}^A_ZN \rightarrow {}^A_ZN + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_ZN \rightarrow {}^A_ZN + \dots)$$

2. Probability of $\Gamma(\gamma + N \rightarrow N) \propto \text{Prob.}(\gamma N)$ in $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$



γ = (squeezed) Gamma-ray Laser Photon / Compton back-scattered gamma ray

$$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N})$$

$$E_\gamma = 3 \sim 5 \text{ MeV}$$

$$\text{Cross-section}(\sigma_{\gamma N}) \sim 2.1 \times 10^{-24} \text{ cm}^2 \quad (\text{XCOM Data})$$

Gamma rays cannot penetrate deep into target materials (at best $\sim O(5 \sim 10) \text{ cm}$)

[for Germanium and $\sim 5 \text{ MeV}$ gamma rays on density = 5.3 g/cm^3]

n_γ capable of creating $1 \text{ e}5/\text{sec}$ photons with their setup in NewSUBARU

[2304.08935]

$$\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$$

$$\text{Prob.}(\gamma N) = F_\gamma n_t \sigma_{\gamma N} \approx \left(\frac{10^5}{\text{sec}} \right) \left[\left(\frac{6 \times 10^{23}}{72.6 \text{ g}} \right) \left(\frac{5.326 \text{ g}}{\text{cm}^3} \right) (5 \text{ cm}) = 2.2 \times 10^{23} / \text{cm}^2 \right] [2.1 \times 10^{-24} \text{ cm}^2] = \frac{4.6 \times 10^4}{\text{sec}}$$

$$\Longrightarrow \Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) \approx \frac{4.6 \times 10^4}{\text{sec}} \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$$

3. Electron-Positron Production in

$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^+ + e^- + (\bar{\nu} + \nu)$$

For Electro-Magnetic X ($Z^* \rightarrow \nu \nu$):

In free space (vacuum):

$$\gamma \rightarrow e^+ + e^- + (\bar{\nu} + \nu)$$



With matter (nucleus):

$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^+ + e^- + (\bar{\nu} + \nu)$$



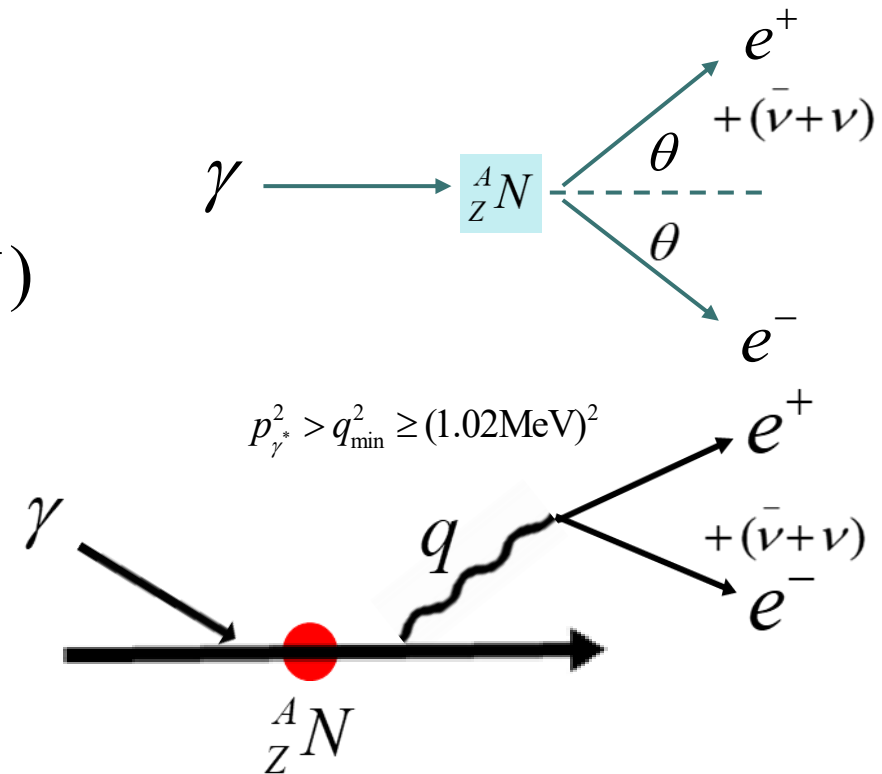
2 issues: 1) Is factorizable?

2) Is the final N moving?

$$E(\gamma) > 1.02 \text{ MeV}$$

$$\gamma + {}^A_Z N \rightarrow {}^A_Z N + \gamma^* (\rightarrow e^+ + e^-)$$

$$h\nu = 2m_e + E_{e^+} + E_{e^-}$$



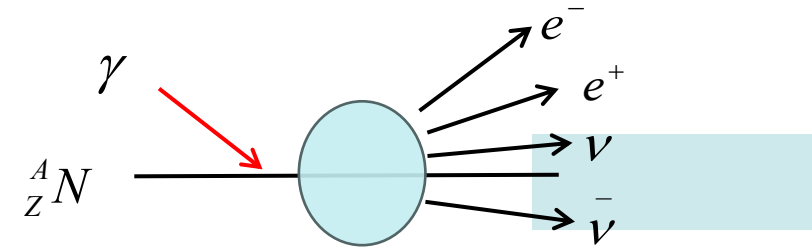
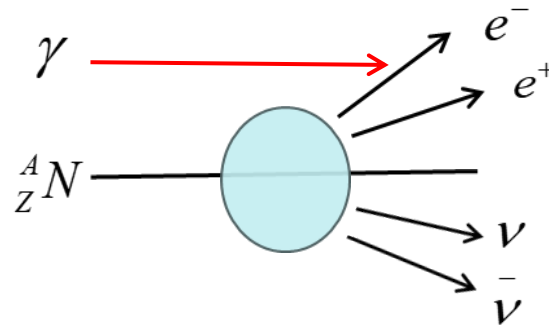
If factorized, as discussed, due to the physic scale of the process:

$$\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^+ + e^-) \approx \text{Prob.}(\gamma N) \times \Gamma(\gamma^*_{(p_{\gamma^*}^2 \sim q_{\min}^2)} \rightarrow e^- + e^+)$$

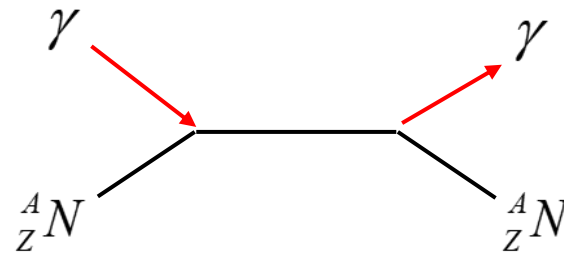
$$[\gamma + N \rightarrow N][\gamma^* \rightarrow e e \nu \bar{\nu}]$$

3.1 BH vs. VCS

$$\gamma + N \rightarrow N + e^+ + e^- + (\bar{\nu} + \nu)$$

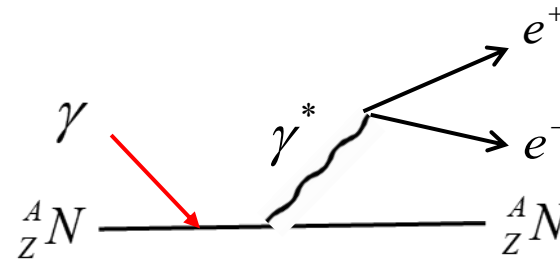


(A) Compton Scattering



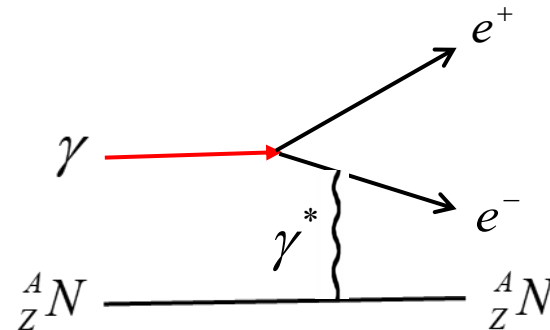
dominant for $E_\gamma < 2m_e$

(B) Virtual Compton Scattering
: factorizable, $\propto \alpha \times (Z_N)^2$



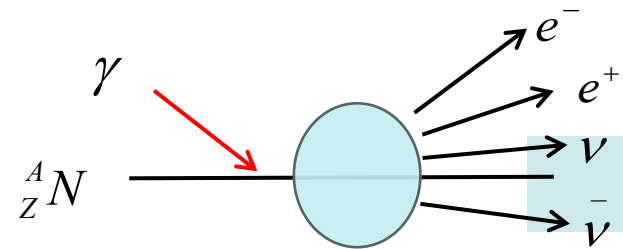
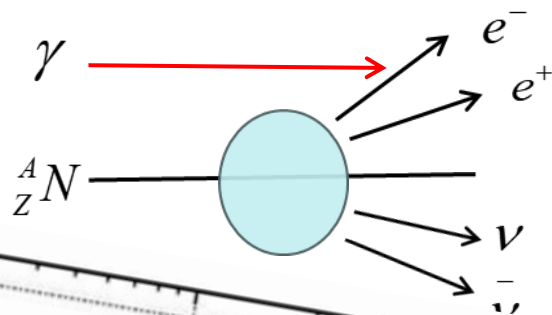
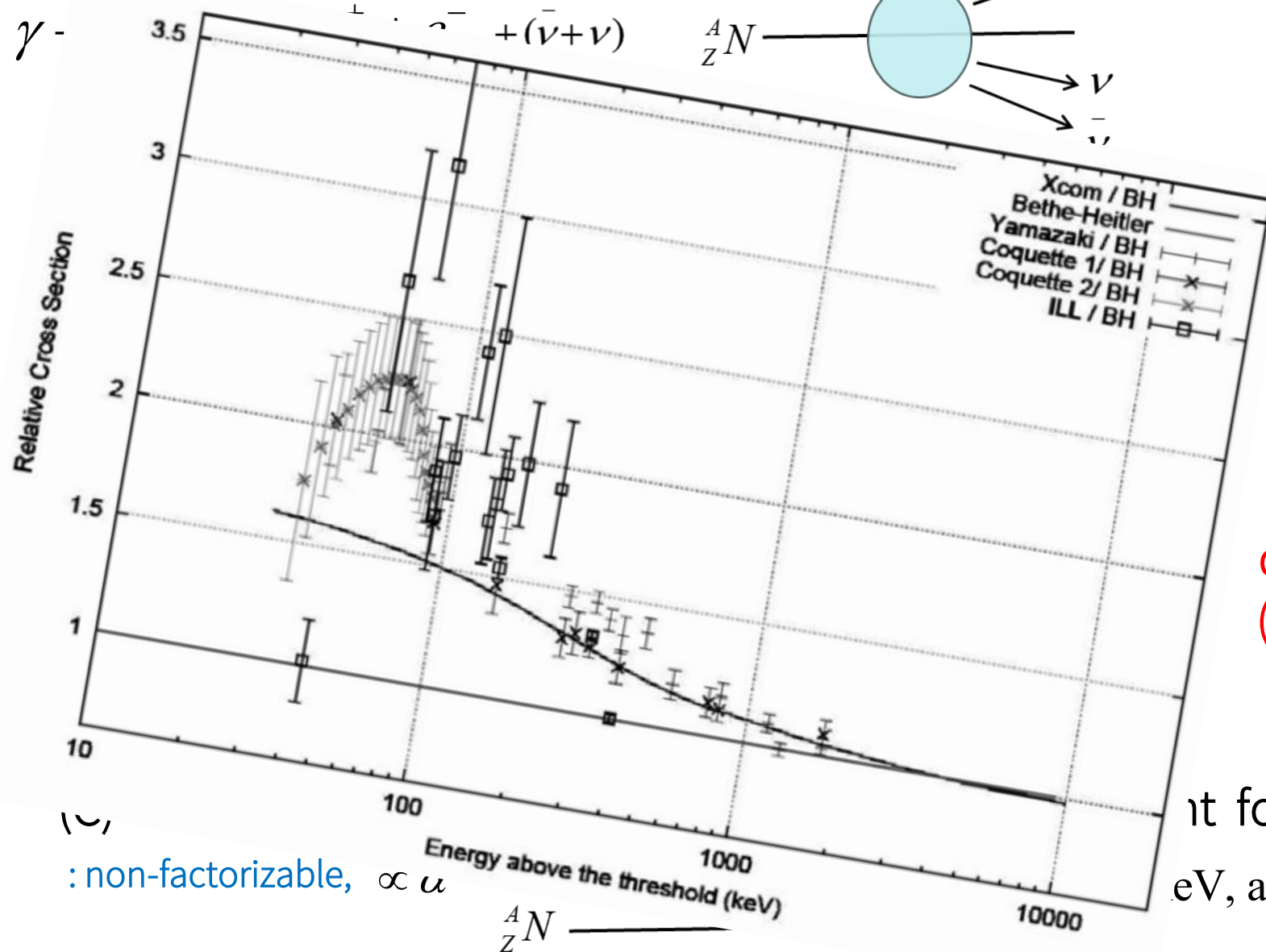
dominant for $E_\gamma \geq 2m_e$
(just above threshold $\propto 1/q^2$)

(C) Bethe-Heitler
: non-factorizable, $\propto \alpha$



dominant for $E_\gamma \gg 2m_e$
 $E_\gamma > 7\text{MeV}$, $\text{angle}(e^+, e^-) \sim \text{large}$

3.1 BH vs. VCS



inant for $E_\gamma < 2m_e$

dominant for $E_\gamma \geq 2m_e$
(just above threshold $\propto 1/q^2$)

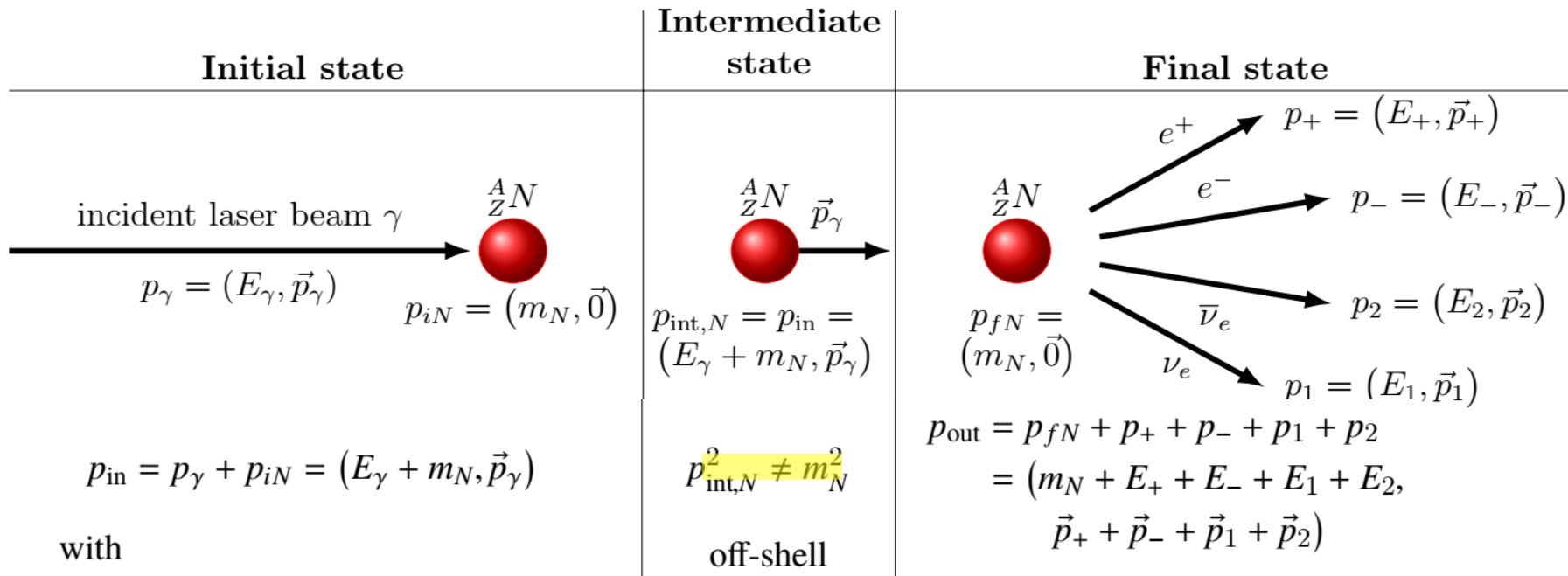
it for $E_\gamma \gg 2m_e$
eV, angle(e^+, e^-) $\sim$ large

: non-factorizable, $\propto \alpha$

4. Does the final N move as of recoil motion ?

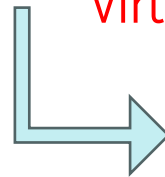
$$\gamma + \textcircled{{}_Z^A N} \rightarrow \textcircled{{}_Z^A N} \rightarrow \textcircled{{}_Z^A N} + e^- + e^+ + (\bar{\nu} + \nu)$$

Lab - frame



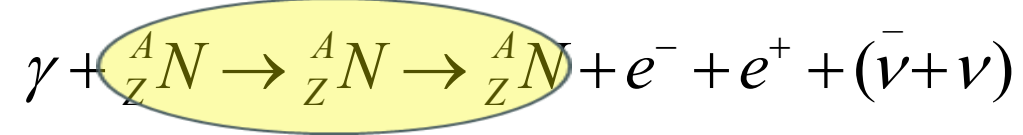
Source of Virtuality

Virtuality comes in to produce massive particles



$$p_\gamma (\text{incoming})^2 = 0 \quad p_{\gamma^*} (\rightarrow e e \nu \bar{\nu})^2 \succ 2m_e^2 \neq 0$$

Does the final N move as of recoil motion ?



Is the FINAL STATE N (in crystal) at rest or moving (after collision with photon) ??

With e.g. $E_\gamma = 3\text{MeV}$, $m(N) \sim 70\text{GeV}$, colliding photon CANNOT break NOR dislocate N, instead N oscillate !!

Time averaged VELOCITY $\langle \vec{V} \rangle = 0$, however ENERGY $\langle V^2 \rangle \neq 0 \rightarrow$

(a) the final N at rest and energetic to produce MASSIVE particles !!

(b) the final N gives the recoil energy to the produced final particles !!

$$E(e^+, e^-) = E_\gamma \rightarrow E(e^+, e^-) = E_\gamma + E_{\text{recoil}} = E_\gamma (1 + m_e / m_N) \approx E_\gamma$$

(c) In the experiment $\gamma + N \rightarrow N + e^+ + e^-$, never seen time-delayed energetic photon !!

(instead, target matter N would heat up via oscillation)

** N in crystal = (strong) spring interaction with neighboring N (unlike free N or N in cloud chamber detector)

C. S. Kim , D. Sahoo and G. Lopez (in progress)

BACK-TO-BACK ν and $\bar{\nu}$ in $2\nu\beta^+\beta^-$ $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$

- experimental observable exception to pDMCT ;
- back-to-back $\nu - \bar{\nu}$

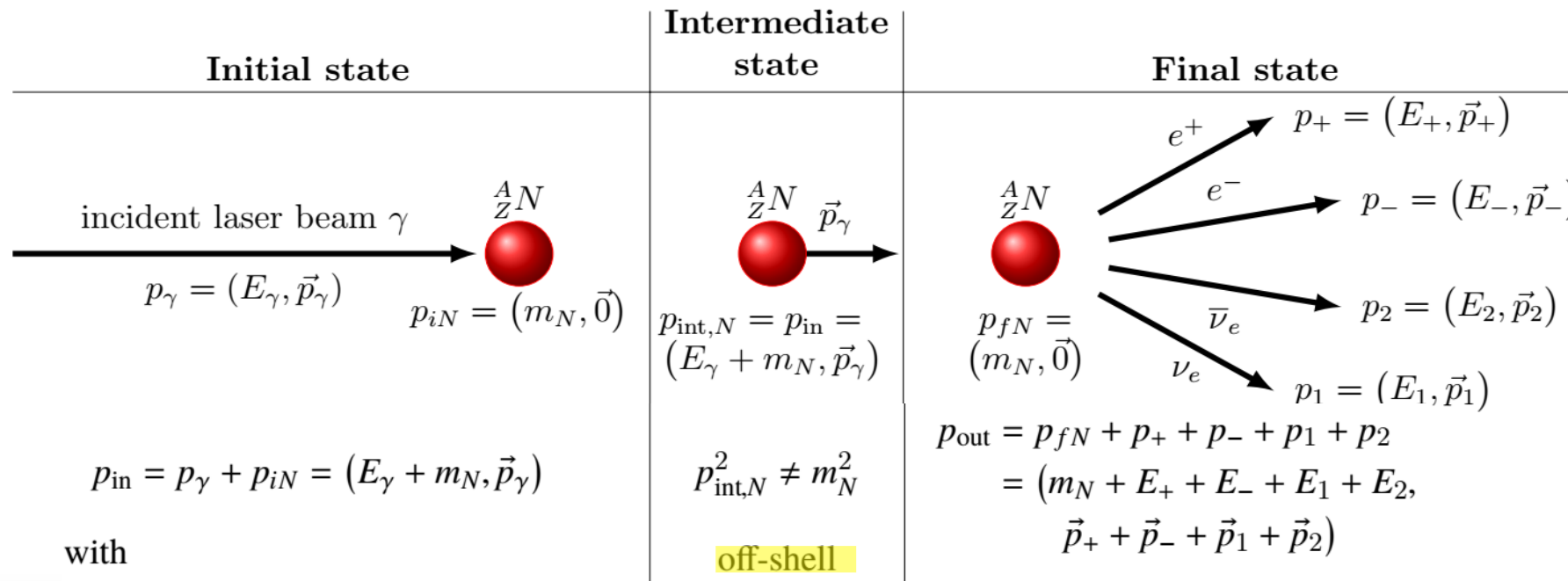
Kinematic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$
Back-to-back kinematics in $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$
Dynamic study of $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$
Background study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$



Kinematic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$ (1)

$$\gamma + {}^A_ZN \rightarrow {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$$

Lab - frame



$$\begin{aligned}
 p_{\text{in}} &= p_{\text{out}} \\
 \Rightarrow p_\gamma + p_{iN} &= p_{fN} + p_+ + p_- + p_1 + p_2 \\
 \Rightarrow \boxed{p_\gamma = p_+ + p_- + p_1 + p_2}
 \end{aligned}$$

$$p_\gamma(\text{incoming})^2 = 0$$

$$\boxed{\vec{p}_\gamma = \vec{p}_+ + \vec{p}_- + \vec{p}_1 + \vec{p}_2 \neq \vec{0}}$$

$$p_{\gamma^*}(\rightarrow e e \nu \bar{\nu})^2 \succ 2m_e^2 \neq 0$$

$$\text{Amp} \sim 1/q^2, \text{ so maximal contribution at } (p_+ + p_- + p_1 + p_2)^2 \sim q_{\text{min}}^2$$

Kinematic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$ (2)

Lab - frame

$$p_\gamma = p_+ + p_- + p_1 + p_2$$

$$\vec{p}_\gamma = \vec{p}_+ + \vec{p}_- + \vec{p}_1 + \vec{p}_2 \neq \vec{0}$$

Any Boosted - frame

$$p'_x = \Lambda p_x, \quad \text{for } x \in \{\gamma, iN, fN, +, -, 1, 2\}.$$

back-to-back neutrino frame

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when } \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$

Realizing back-to-back neutrino condition

$$p_{ee} \equiv (E_{ee}, \vec{p}_{ee}) = p_+ + p_- = (E_+ + E_-, \vec{p}_+ + \vec{p}_-),$$

$$p_{\nu\nu} \equiv (E_{\nu\nu}, \vec{p}_{\nu\nu}) = p_1 + p_2 = (E_1 + E_2, \vec{p}_1 + \vec{p}_2),$$

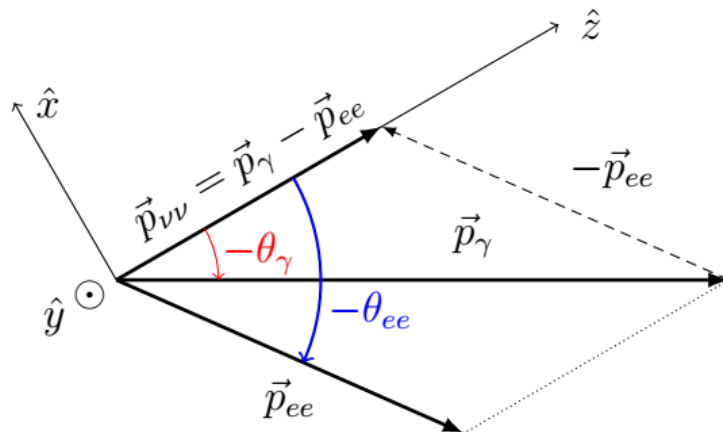
$$m_{ee}^2 \equiv p_{ee}^2 = E_{ee}^2 - |\vec{p}_{ee}|^2,$$

$$m_{\nu\nu}^2 \equiv p_{\nu\nu}^2 = E_{\nu\nu}^2 - |\vec{p}_{\nu\nu}|^2.$$

$$\vec{p}'_{\nu\nu} = \vec{0}$$

$$\vec{\beta}_{\nu\nu} = \frac{\vec{p}_{\nu\nu}}{E_{\nu\nu}}, \quad \gamma_{\nu\nu} = \frac{1}{\sqrt{1 - |\vec{\beta}_{\nu\nu}|^2}} = \frac{E_{\nu\nu}}{m_{\nu\nu}}.$$

$$p'_{\nu\nu} = \Lambda p_{\nu\nu} = \begin{pmatrix} \gamma_{\nu\nu} & 0 & 0 & -\gamma_{\nu\nu}\beta_{\nu\nu} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_{\nu\nu}\beta_{\nu\nu} & 0 & 0 & \gamma_{\nu\nu} \end{pmatrix} \begin{pmatrix} E_{\nu\nu} \\ 0 \\ 0 \\ |\vec{p}_{\nu\nu}| \end{pmatrix} = \begin{pmatrix} \frac{E_{\nu\nu}}{m_{\nu\nu}} & 0 & 0 & -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} & 0 & 0 & \frac{E_{\nu\nu}}{m_{\nu\nu}} \end{pmatrix} \begin{pmatrix} E_{\nu\nu} \\ 0 \\ 0 \\ |\vec{p}_{\nu\nu}| \end{pmatrix} = \begin{pmatrix} m_{\nu\nu} \\ 0 \\ 0 \\ 0 \end{pmatrix}$$



Lab - frame

Kinematic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$ (3)

Lab - frame

$$p_\gamma = p_+ + p_- + p_1 + p_2$$

Boosted (back-to-back neutrino) - frame

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when} \quad \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$

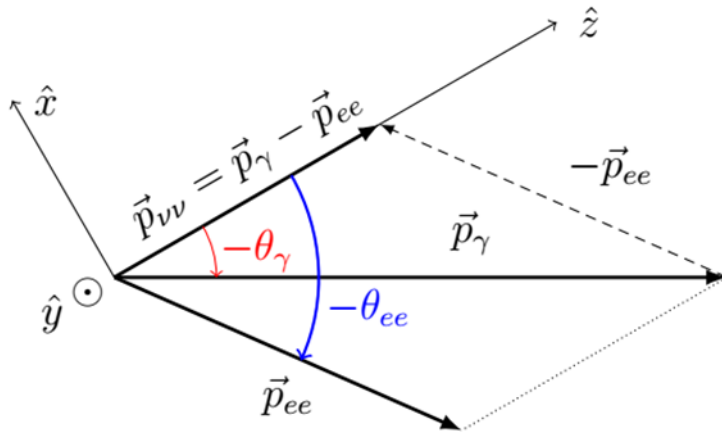
$$E'_\nu = m_{\nu\nu} / 2 = \frac{\sqrt{(E_\gamma - E_+ - E_-)^2 - (\vec{p}_\gamma - \vec{p}_+ - \vec{p}_-)^2}}{2}$$

$$m_{\nu\nu} \in \left\{ 2 m_\nu, \sqrt{(E_{ee} + |\vec{p}_{ee}| - 2 E_\gamma) (E_{ee} - |\vec{p}_{ee}|)} \right\}.$$

$$(E_\gamma, \vec{p}_\gamma), (E_+, \vec{p}_+), (E_-, \vec{p}_-), (E_{\nu\nu}, \vec{p}_{\nu\nu}), E'_\nu \quad \text{All observable, calculable !!}$$

$$\frac{d\Gamma_{D/M}}{dE'_\nu}$$

Measurable, calculable !!



Lab - frame

Kinematic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (4)

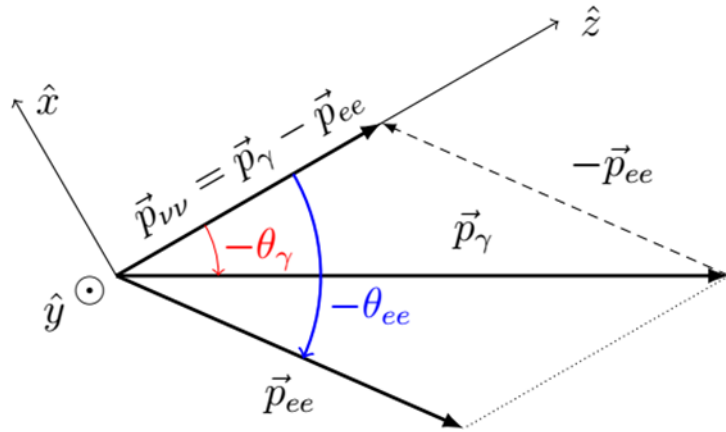
Lab - frame

$$p_\gamma = p_+ + p_- + p_1 + p_2$$

Boosted (back-to-back neutrino) - frame

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when} \quad \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$

$$p'_x = \Lambda p_x, \quad \text{for } x \in \{\gamma, iN, fN, +, -, 1, 2\}.$$



Lab - frame

$$\Lambda = \begin{pmatrix} \gamma_{\nu\nu} & 0 & 0 & -\gamma_{\nu\nu}|\vec{\beta}_{\nu\nu}| \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\gamma_{\nu\nu}|\vec{\beta}_{\nu\nu}| & 0 & 0 & \gamma_{\nu\nu} \end{pmatrix} = \begin{pmatrix} \frac{E_{\nu\nu}}{m_{\nu\nu}} & 0 & 0 & -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} & 0 & 0 & \frac{E_{\nu\nu}}{m_{\nu\nu}} \end{pmatrix}$$

e.g.

$$p'_{ee} = \Lambda p_{ee} = \begin{pmatrix} \frac{E_{\nu\nu}}{m_{\nu\nu}} & 0 & 0 & -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{|\vec{p}_{\nu\nu}|}{m_{\nu\nu}} & 0 & 0 & \frac{E_{\nu\nu}}{m_{\nu\nu}} \end{pmatrix} \begin{pmatrix} E_{ee} \\ -|\vec{p}_{ee}| \sin \theta_{ee} \\ 0 \\ |\vec{p}_{ee}| \cos \theta_{ee} \end{pmatrix} = \begin{pmatrix} \frac{E_{\nu\nu} E_{ee}}{m_{\nu\nu}} - \frac{|\vec{p}_{\nu\nu}| |\vec{p}_{ee}| \cos \theta_{ee}}{m_{\nu\nu}} \\ -|\vec{p}_{ee}| \sin \theta_{ee} \\ 0 \\ -\frac{|\vec{p}_{\nu\nu}| E_{ee}}{m_{\nu\nu}} + \frac{E_{\nu\nu} |\vec{p}_{ee}| \cos \theta_{ee}}{m_{\nu\nu}} \end{pmatrix}$$

$$\text{i.e. } Q' = E'_{e+} + E'_{e-} = \frac{E_{\nu\nu} E_{ee}}{m_{\nu\nu}} - \frac{|\vec{p}_{\nu\nu}| |\vec{p}_{ee}| \cos \theta_{ee}}{m_{\nu\nu}}$$

All observable, calculable !!

Back-to-back neutrinos, (exception to DMCT) in (1)

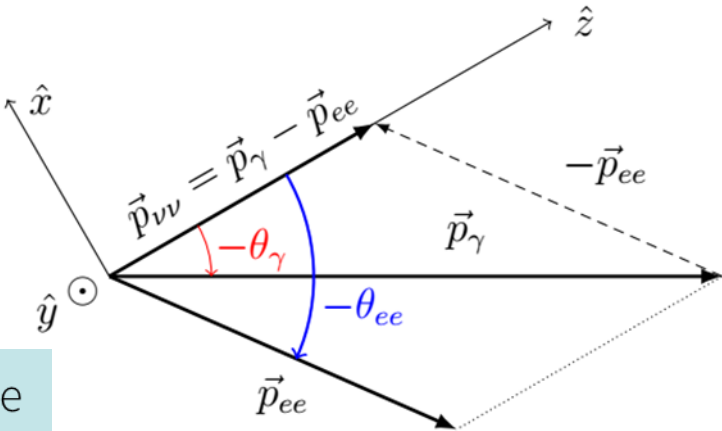
$$\gamma + \textcircled{{}^A_ZN} \rightarrow \textcircled{{}^A_ZN} + e^- + e^+ + (\bar{\nu} + \nu)$$

Boosted (back-to-back neutrino) - frame

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when} \quad \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$

Back-to-back (B2B) Constraints WITH Experimental Resolution, ΔE , $\Delta \vec{P}$.

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when} \quad \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$



Lab - frame

$E_{e^+,e^-} \gg \Delta E$
 $E_{e^+} + E_{e^-} < E_\gamma - \Delta E$
 $\vec{p}'_+ + \vec{p}'_- \approx \vec{p}'_\gamma$

Measure

$$\frac{d\Gamma}{dE'_\nu}$$

$$\frac{d\Gamma}{dE_{e^-+e^+}} = \frac{d\Gamma}{dQ}$$

$$\frac{d\Gamma}{d\cos(\theta'_{ee})}$$

Back-to-back neutrinos, (exception to DMCT) in (2)

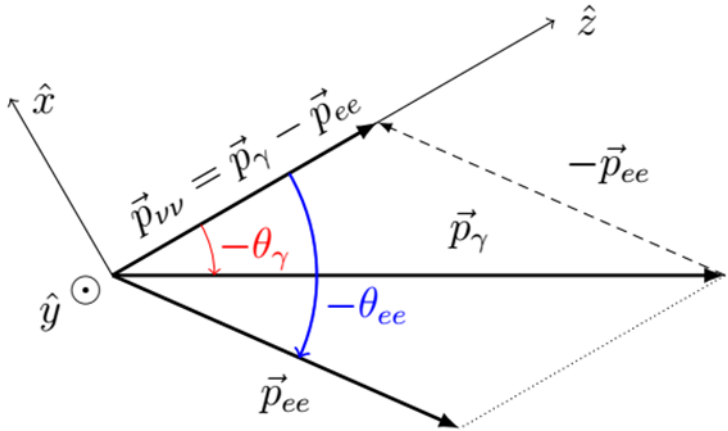
$$\gamma + \textcolor{yellow}{\textcircled{\gamma + \frac{A}{Z}N \rightarrow \frac{A}{Z}N}} + e^- + e^+ + (\bar{\nu} + \nu)$$

Lab - frame

$$p_\gamma = p_+ + p_- + p_1 + p_2$$

Boosted (back-to-back neutrino) - frame

$$\vec{p}'_1 + \vec{p}'_2 = \vec{0}, \quad \text{only when} \quad \vec{p}'_+ + \vec{p}'_- = \vec{p}'_\gamma \neq \vec{0}$$



$$d\Gamma_{D/M} = |M_{D/M}|^2 d_4(PS : \gamma \rightarrow e^- + e^+ + (\bar{\nu} + \nu))$$

[Barger, Phillips, Collider Physics, Appendix B]

$$d_4(PS : \gamma \rightarrow e^- + e^+ + (\bar{\nu} + \nu)) = d_2(PS : \gamma \rightarrow XY) dm_X^2 dm_Y^2 \\ \times d_2(PS : X \rightarrow e^- + e^+) d_2(PS : Y \rightarrow \bar{\nu} + \nu)$$

$$E'_\nu = m_{\nu\nu} / 2 = \frac{\sqrt{(E_\gamma - E_+ - E_-)^2 - (\vec{p}_\gamma - \vec{p}_+ - \vec{p}_-)^2}}{2}$$

Calculate

$$\frac{d\Gamma_{D/M}}{dE'_\nu}$$

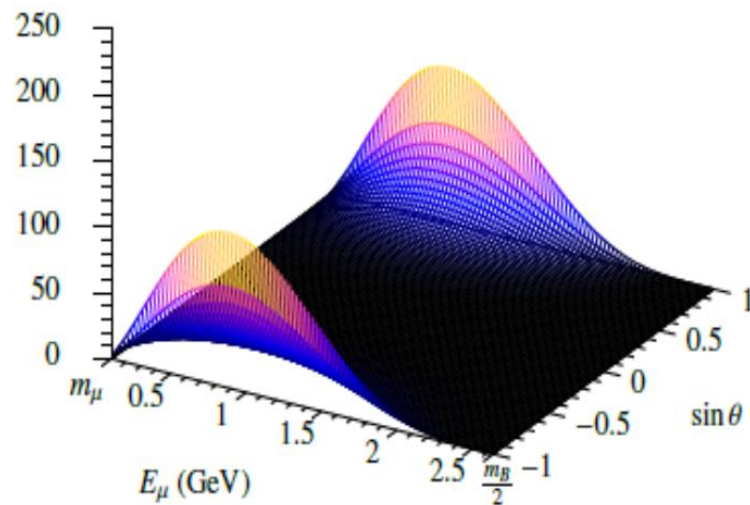
$$\frac{d\Gamma_{D/M}}{dE_{e^-+e^+}} = \frac{d\Gamma_{D/M}}{dQ}$$

$$\frac{d\Gamma_{D/M}}{d\cos(\theta'_{ee})} \text{ to compare D with M}$$

Back-to-back neutrinos in $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (3)

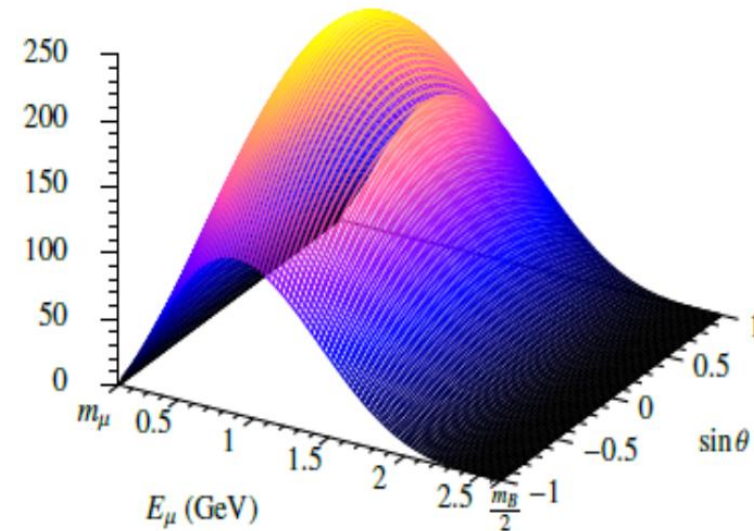
e.g.) B2B Muons in $B^0(p_B) \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2)$,

$$\frac{512\pi^6 m_B}{G_F^4 |\mathbb{F}_a|^2} \frac{d^3\Gamma_{\leftrightarrow}^D}{dE_\mu^2 d\sin\theta} (\text{GeV}^6)$$



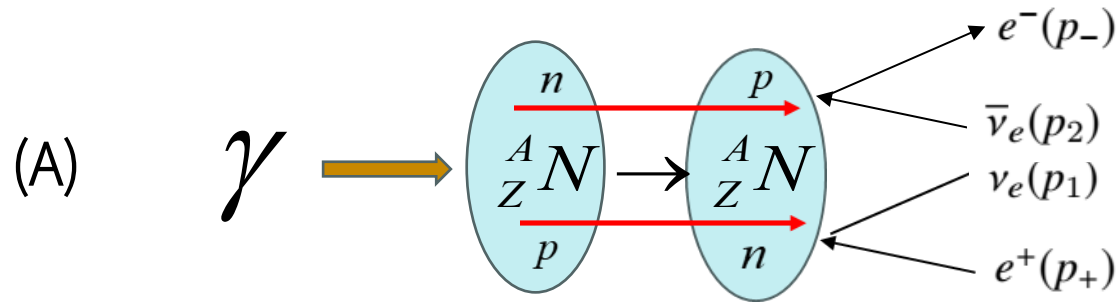
(a) Three dimensional view of the differential decay rate for Dirac case with an appropriate normalization as mentioned.

$$\frac{512\pi^6 m_B}{G_F^4 |\mathbb{F}_a|^2} \frac{d^3\Gamma_{\leftrightarrow}^M}{dE_\mu^2 d\sin\theta} (\text{GeV}^6)$$



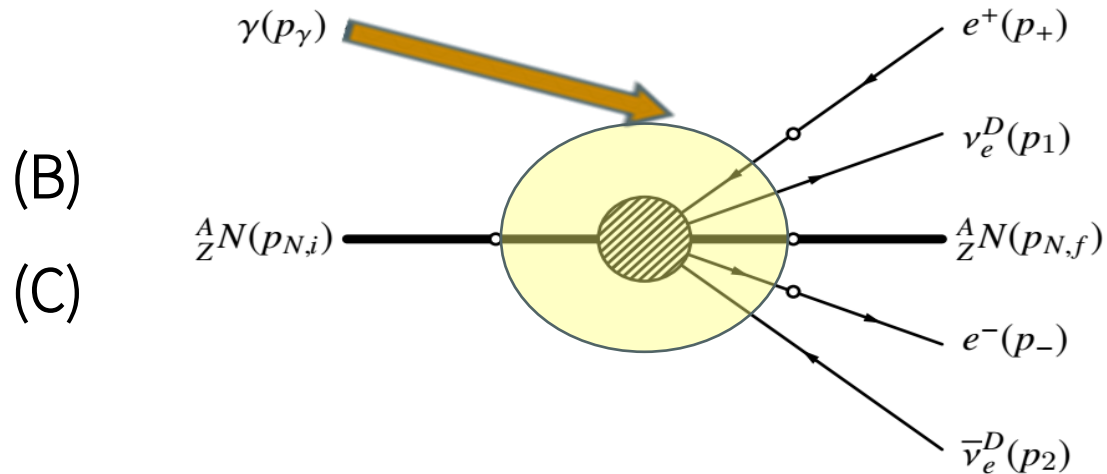
(b) Three dimensional view of the differential decay rate for Majorana case with an appropriate normalization as mentioned.

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (0) $|M^{D/M}|^2$



(A) $[\gamma + N \rightarrow N][N \rightarrow N + e e \nu \bar{\nu}]$

$E_\gamma = 3 \sim 5 \text{ MeV}$ $\lambda(\gamma) \sim 100 \text{ fm}$
 $\lambda(p, n) \sim 1 \text{ fm}$

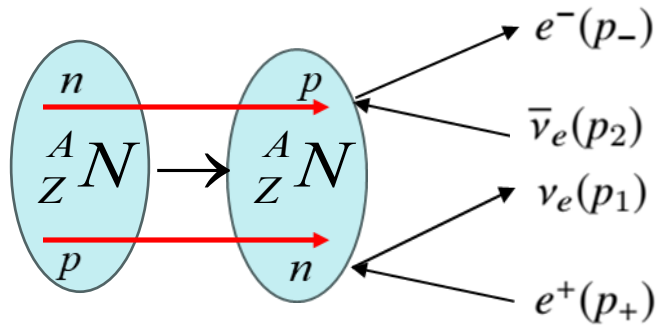


$[\gamma + N \rightarrow N][\gamma^* \rightarrow e e \nu \bar{\nu}]$

(B) $\gamma^* \rightarrow e^+ e^- (Z^* \rightarrow \nu \bar{\nu})$

(C) $\gamma^* \rightarrow W^+ (\rightarrow e^+ \nu) W^- (\rightarrow e^- \bar{\nu})$

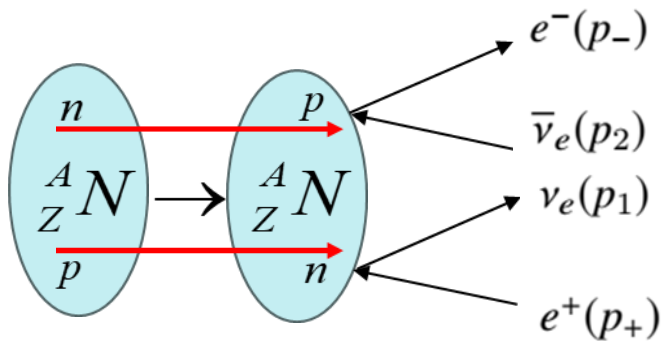
Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A-1)



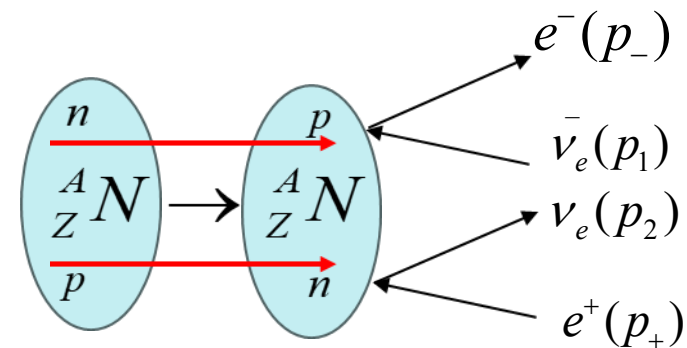
For Dirac neutrinos

$$\mathcal{M}^D = \frac{G_F^2}{2} H^{\alpha\beta} L_{\alpha\beta}$$

$$\mathcal{M}^M = \frac{G_F^2}{2\sqrt{2}} (H^{\alpha\beta} L_{\alpha\beta} - H'^{\alpha\beta} L'_{\alpha\beta})$$



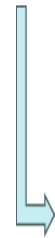
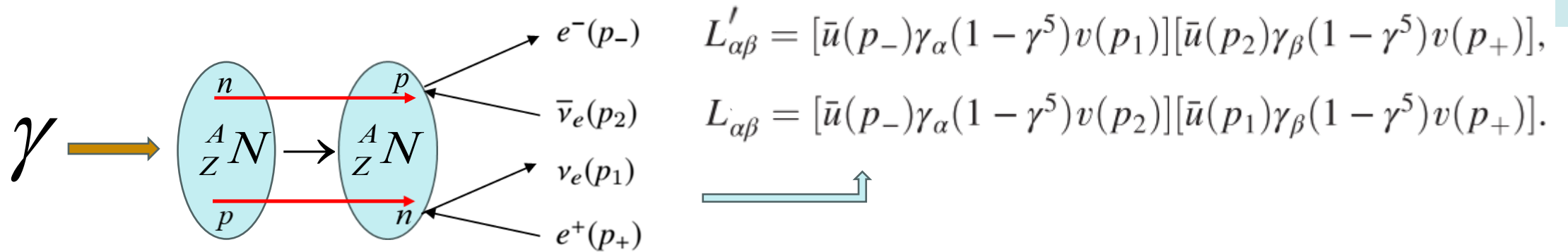
-



Momentum exchange
(not helicity flip)

For Majorana neutrinos

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A-2)



$$H^{\alpha\beta} = H'^{\alpha\beta} = F(q^2)g^{\alpha\beta},$$

$$|F(q^2)|^2 \equiv F_{NME}$$

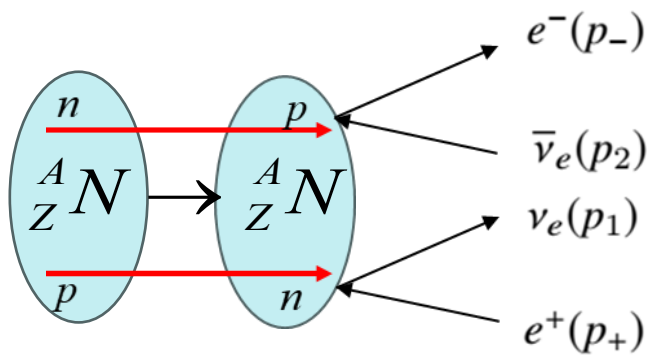
γ = (squeezed) Gamma-ray Laser Photon
[Compton back-scattered gamma ray]

$$1.02\text{MeV} < E(\gamma) < M({}^A_Z N^*) - M({}^A_Z N)$$

$$\mathcal{M}^D = \frac{G_F^2}{2} H^{\alpha\beta} L_{\alpha\beta}$$

$$\mathcal{M}^M = \frac{G_F^2}{2\sqrt{2}} (H^{\alpha\beta} L_{\alpha\beta} - H'^{\alpha\beta} L'_{\alpha\beta})$$

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A-3)



$$L'_{\alpha\beta} = [\bar{u}(p_-)\gamma_\alpha(1 - \gamma^5)v(p_1)][\bar{u}(p_2)\gamma_\beta(1 - \gamma^5)v(p_+)],$$

$$L_{\alpha\beta} = [\bar{u}(p_-)\gamma_\alpha(1 - \gamma^5)v(p_2)][\bar{u}(p_1)\gamma_\beta(1 - \gamma^5)v(p_+)].$$

$$\mathcal{M}^D = \frac{G_F^2}{2} H^{\alpha\beta} L_{\alpha\beta}$$

$$\mathcal{M}^M = \frac{G_F^2}{2\sqrt{2}} (H^{\alpha\beta} L_{\alpha\beta} - H'^{\alpha\beta} L'_{\alpha\beta})$$

$$H^{\alpha\beta} = H'^{\alpha\beta} = F(q^2)g^{\alpha\beta},$$

$$|F(q^2)|^2 \equiv F_{NME}$$

$$|M^D|^2 = G_F^4 F_{NME} [64(p_1 \cdot p_+)(p_2 \cdot p_-)]$$

$$|M^M|^2 = G_F^4 F_{NME} [64(p_1 \cdot p_+)(p_2 \cdot p_-) + 64(p_1 \cdot p_-)(p_2 \cdot p_+)] / 2 + O(m_\nu^2)$$



$$|M^D|^2 - |M^M|^2 \neq 0$$

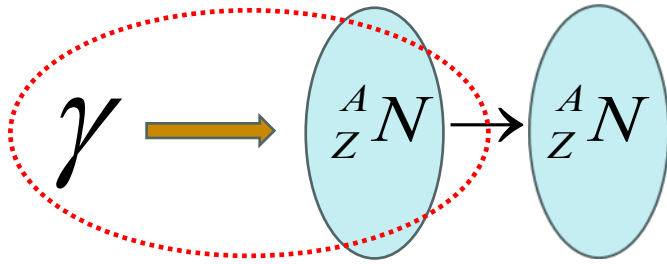
$$\int (|M^D|^2 - |M^M|^2) dp_1 dp_2 = 0$$

(whole neutrino phase space)

$$\int (|M^D|^2 - |M^M|^2) dp_1 dp_2 \neq 0$$

(constrained phase space,
w/ back-to-back neutrinos)

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A-4)



γ = (squeezed) Gamma-ray Laser Photon
[Compton back-scattered gamma ray]

$$1.02\text{MeV} < E(\gamma) < M({}^A_Z N^*) - M({}^A_Z N)$$

1) For $\gamma - {}^A_Z N$ Interaction Probability :

$$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N})$$

2) For Total Decay Width $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$:

$$\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$$

$$\approx n_\gamma \times \sigma_{\gamma N} \times G_F^4 F_{NME}$$

• For Comparison, $0\nu\beta\beta$

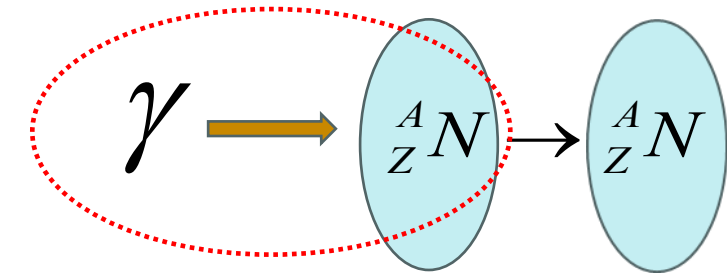
$$\Gamma({}^A_Z N \rightarrow {}^A_{Z+2} N + e^- + e^-)$$

$$\approx G_F^4 F_{NME}^2 m_{\nu_1}^2$$

$$\rightarrow [\dots] \times \text{Av}(=6.022 \times 10^{23}) \sim G_F^4 F_{NME}^2$$

** factorized due to MeV scale ($\gamma + N \rightarrow N$) and GeV scale ($n \rightarrow p + e^- + \bar{\nu}$)

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A-5)



γ = (squeezed) Gamma-ray Laser Photon : Compton back-scattered gamma ray

$$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N})$$

$$E_\gamma = 3 \sim 5 \text{ MeV}$$

$$\text{Cross-section}(\sigma_{\gamma N}) \sim 2.1 \times 10^{-24} \text{ cm}^2 \quad (\text{XCOM Data})$$

Gamma rays cannot penetrate deep into target materials (at best $\sim O(5 \sim 10) \text{ cm}$)

[for Germanium and $\sim 5 \text{ MeV}$ gamma rays on density = 5.3 g/cm^3]

n_γ capable of creating $1 \text{e}5/\text{sec}$ photons with their setup in NewSUBARU

[2304.08935]

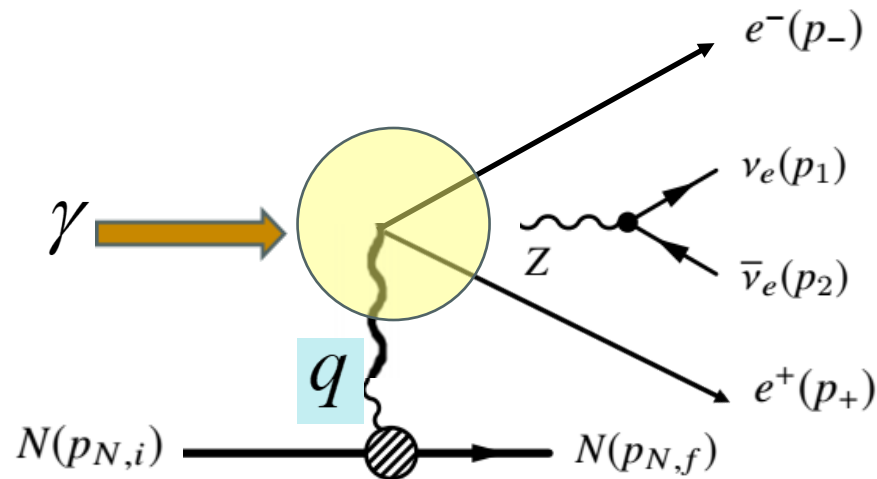
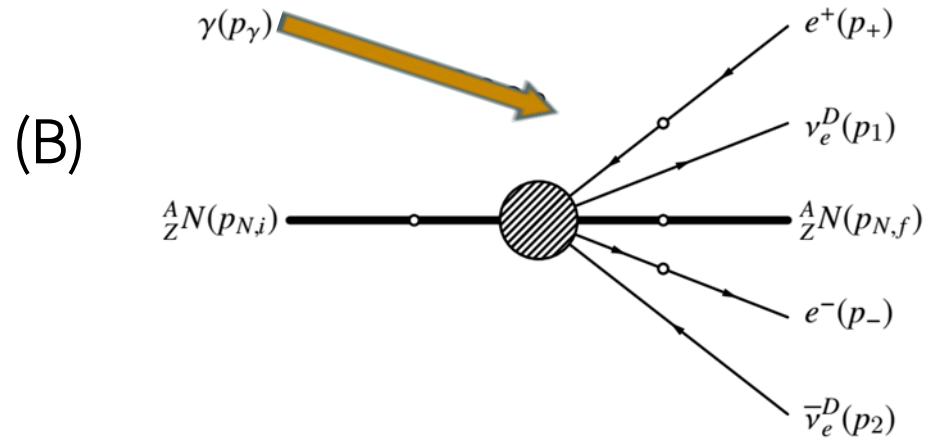
$$\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) = \text{Prob.}(\gamma N) \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$$

$$\text{Prob.}(\gamma N) = F_\gamma n_t \sigma_{\gamma N} \approx \left(\frac{10^5}{\text{sec}} \right) \left[\left(\frac{6 \times 10^{23}}{72.6 \text{ g}} \right) \left(\frac{5.326 \text{ g}}{\text{cm}^3} \right) (5 \text{ cm}) = 2.2 \times 10^{23} / \text{cm}^2 \right] [2.1 \times 10^{-24} \text{ cm}^2] = \frac{4.6 \times 10^4}{\text{sec}}$$

$$\longrightarrow \Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots) \approx \frac{4.6 \times 10^4}{\text{sec}} \times \Gamma({}^A_Z N \rightarrow {}^A_Z N + \dots)$$

$$\Gamma({}^A_Z N \rightarrow {}^A_{Z+2} N + e^- + e^-) \quad 0\nu\beta\beta \\ \approx G_F^4 F_{NME}^2 m_{\nu_1}^2$$

Dynamic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$ (B-1)



(B) $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$

$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N})$
 $= \frac{4.6 \times 10^4}{\text{sec}}$

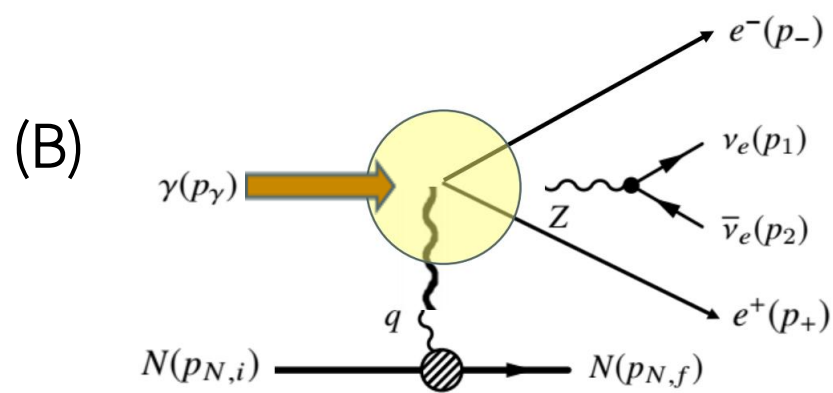
$\approx [\gamma + {}^A_ZN \rightarrow {}^A_ZN] \cdot [\gamma_{(p_\gamma^2 \sim q_{\min}^2)}^* \rightarrow e^- + e^+ + \bar{\nu} + \nu]$

$\hookrightarrow M^D(p_-, p_+, p_1, p_2)$

$M^M = [M^D(p_1, p_2) - M^D(p_2, p_1)] / \sqrt{2}$

$|M(p_1, p_2)|^2 \neq |M(p_2, p_1)|^2 \quad ?$

Dynamic study on $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$ (B-2)



$$\Gamma(\gamma^*_{(p_\gamma^2 \sim q_{\min}^2)} \rightarrow e^- + e^+ + \bar{\nu} + \nu)$$

$$(B) \quad \gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$$

$$\Gamma(\gamma + {}^A_ZN \rightarrow {}^A_ZN + \dots)_B \approx \text{Prob.}(\gamma N) \times \Gamma(\gamma^*_{(p_\gamma^2 \sim q_{\min}^2)} \rightarrow e^- + e^+ + \bar{\nu} + \nu)$$

$$\approx n_\gamma \times \sigma_{\gamma N} \times \alpha^2 Z_N^2 G_F^2$$

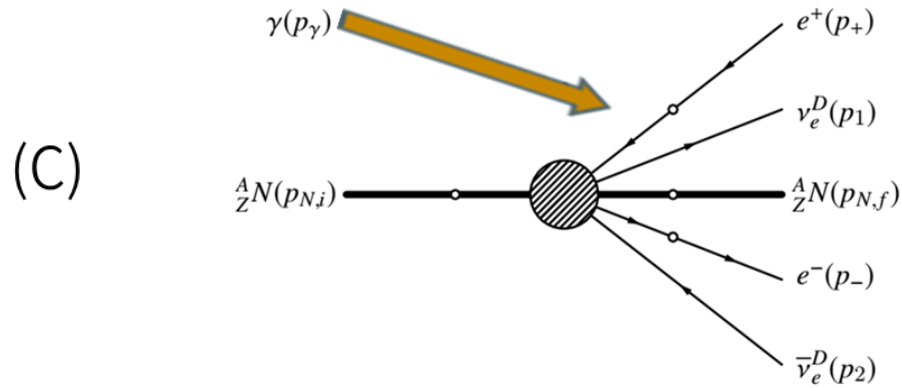
$$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N}) = \frac{4.6 \times 10^4}{\text{sec}}$$

$$M^D(p_-, p_+, p_1, p_2)$$

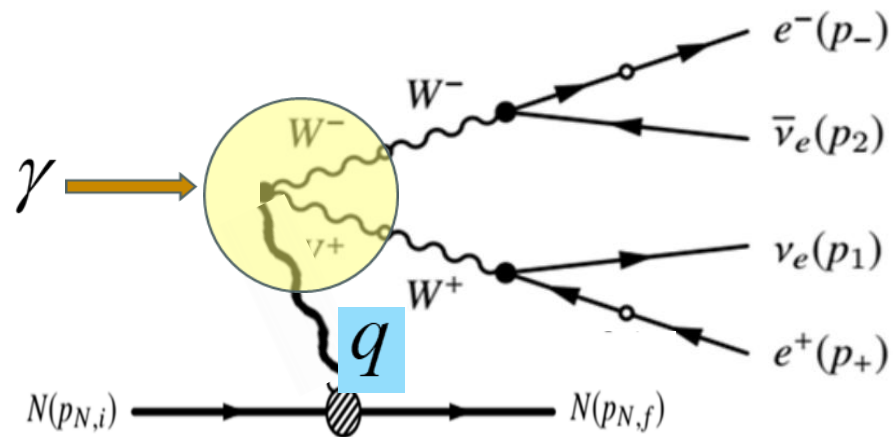
$$M^M = [M^D(p_1, p_2) - M^D(p_2, p_1)] / \sqrt{2}$$

$$|M(p_1, p_2)|^2 \neq |M(p_2, p_1)|^2 \quad ?$$

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (C)



$$(C) \quad \gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$$



$$\approx [\gamma + {}^A_Z N \rightarrow {}^A_Z N] \cdot [\gamma^*_{(p_\gamma^2 \sim q_{\min}^2)} \rightarrow e^- + e^+ + \bar{\nu} + \nu)]$$

Decay Width of (C): $\Gamma(\gamma + {}^A_Z N \rightarrow {}^A_Z N + \dots)_C \approx \text{Prob.}(\gamma N) \times \Gamma(\gamma^*_{(p_\gamma^2 \sim q_{\min}^2)} \rightarrow e^- + e^+ + \bar{\nu} + \nu) \ll (B)$

$$\approx n_\gamma \times \sigma_{\gamma N} \times \alpha^2 Z_N^2 G_F^4$$

Dynamic study on $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$ (A,B,C)

(A,B,C) $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$

(A) $\approx [\gamma + {}^A_Z N \rightarrow {}^A_Z N] \cdot [(p \rightarrow n e^+ \nu)(n \rightarrow p e^- \bar{\nu})]$

$$\approx n_\gamma \times \sigma_{\gamma N} \times G_F^4 F_{NME}$$

(B) $\approx [\gamma + {}^A_Z N \rightarrow {}^A_Z N] \cdot [\gamma_{(p_\gamma^2 \sim q_{\min}^2)}^* \rightarrow e^- + e^+ + \bar{\nu} + \nu]$

$$\approx n_\gamma \times \sigma_{\gamma N} \times \alpha^2 Z_N^2 G_F^2 \quad \leftarrow \text{Most promising !!}$$

(C) $\approx [\gamma + {}^A_Z N \rightarrow {}^A_Z N] \cdot [\gamma_{(p_\gamma^2 \sim q_{\min}^2)}^* \rightarrow e^- + e^+ + \bar{\nu} + \nu]$

$$\approx n_\gamma \times \sigma_{\gamma N} \times \alpha^2 Z_N^2 G_F^4$$

$$\text{Prob.}(\gamma N) = \text{Luminosity}(\propto n_\gamma) \times \text{Cross-section}(\sigma_{\gamma N}) = n_\gamma \times \sigma_{\gamma N} = \frac{4.6 \times 10^4}{\text{sec}}$$

$$0\nu\beta\beta \quad \Gamma({}^A_Z N \rightarrow {}^A_{Z+2} N + e^- + e^-) \\ \approx G_F^4 F_{NME} m_{\nu_1}^2$$

$$\rightarrow [\dots] \times \text{Av}(=6\text{e}+23) \sim G_F^4 F_{NME}$$

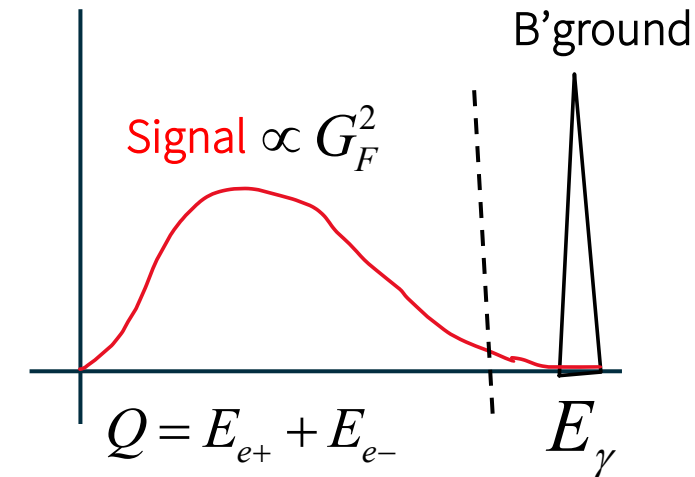
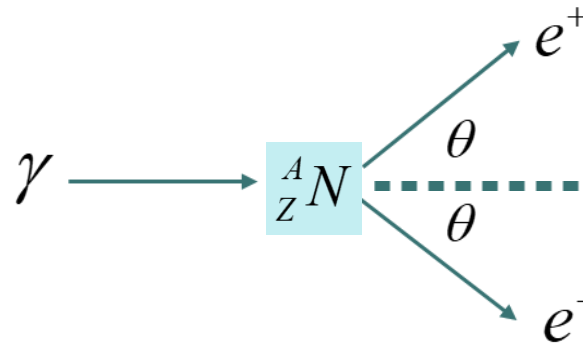
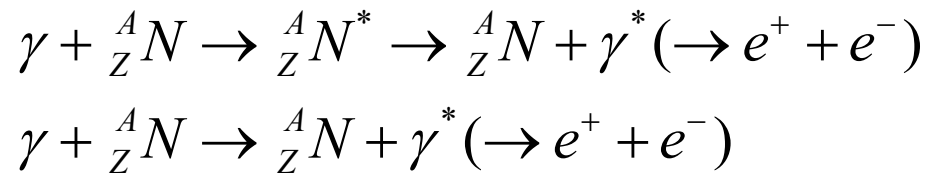
$$2\nu\beta\beta \quad \approx G_F^4 F_{NME}$$

$$\rightarrow G_F^4 F_{NME} \times \text{Av}(=6\text{e}+23)$$

Background Study of $\gamma + {}^A_Z N \rightarrow {}^A_Z N + e^- + e^+ + (\bar{\nu} + \nu)$

WITH Experimental Resolution, ΔE , $\Delta \vec{P}$.

1) Electron-positron creation



2) $\gamma \rightarrow$ Atomic ionization (e^-) + Single beta decay (e^+)

Requires e^+ , e^- from the same production point

Back-to-back Constraint and Q-value will Almost Clear All possible Backgrounds

i.e. $\vec{p}_+ + \vec{p}_- \approx \vec{p}_\gamma$ (back-to-back ν constraint)

e^+ , e^- from the same production point

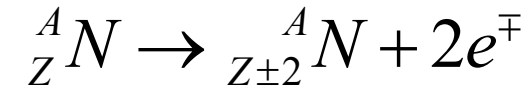
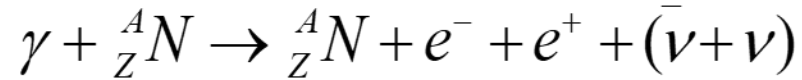
$E_{e^+} + E_{e^-} \ll E_\gamma - n \times \Delta E$ (Q-value constraint)

Distinguish $e^+ e^- \nu \bar{\nu}$ from $e^+ e^-$

$$2\nu\beta^+\beta^-$$

vs.

$$0\nu\beta^\pm\beta^\pm$$



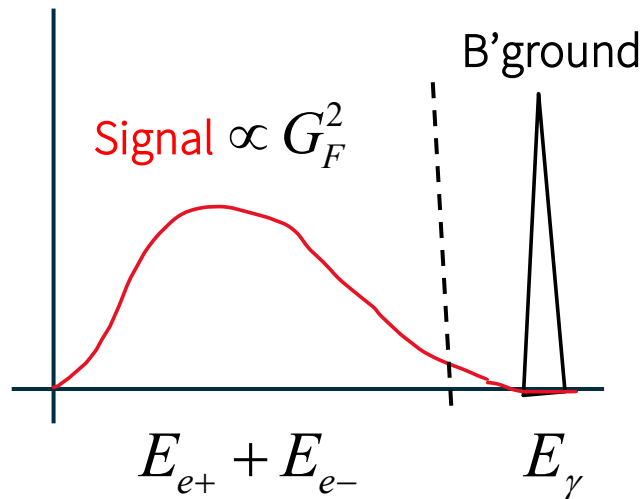
$$\approx n_\gamma \times \sigma_{\gamma N} \times \alpha^2 Z_N^2 G_F^2$$

$$\Gamma({}^A_ZN \rightarrow {}^A_{Z+2}N + e^- + e^-)$$

$$\approx G_F^4 F_{NME}^2 m_{\nu_1}^2$$

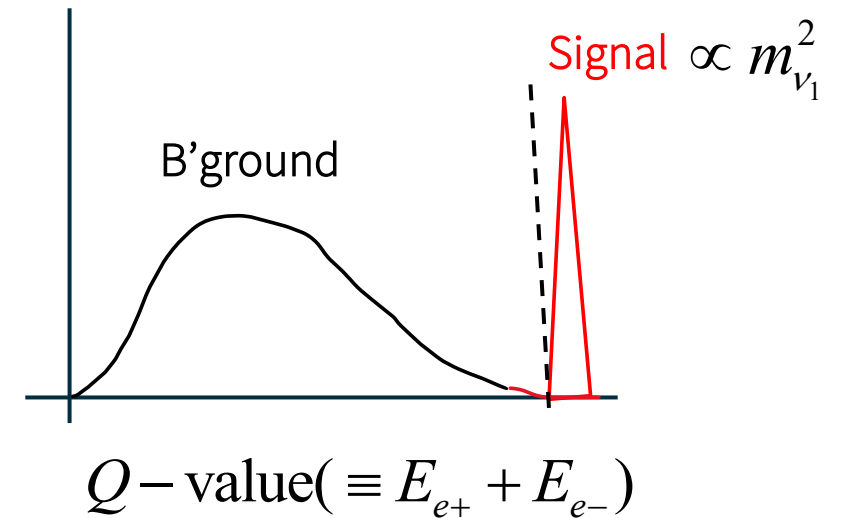
$$\rightarrow [\dots] \times \text{Av}(=6\text{e}+23) \sim G_F^4 F_{NME}$$

$$n_\gamma \times \sigma_{\gamma N} \approx \frac{4.6 \times 10^4}{\text{sec}}$$



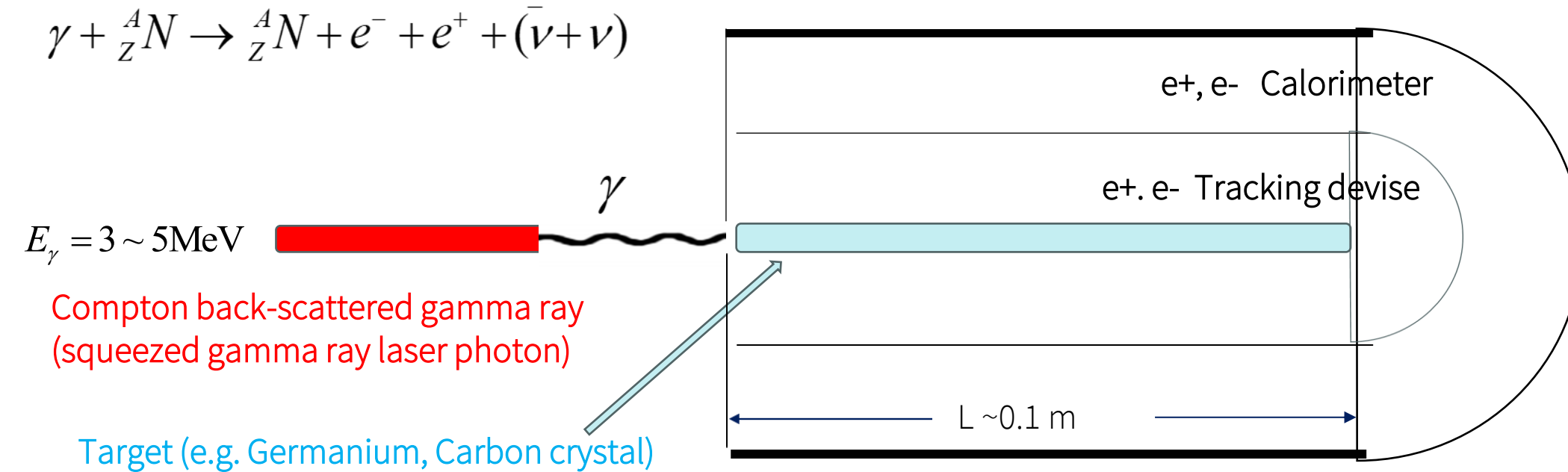
$$E_{e+} + E_{e-} \ll E_\gamma - n \times \Delta E$$

- 1) soft photon radiation from B'ground
- 2) $\gamma^* \rightarrow e^+ e^-$ will give same hemisphere



Experimental set-up of $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$

To measure $\vec{p}_{e^+,e^-}, Q\text{-value} (\equiv E_{e^+} + E_{e^-})$



Back-to-back Constraint and Q-value will Almost Clear All possible Backgrounds

$$\vec{p}_+ + \vec{p}_- \approx \vec{p}_\gamma \quad (\text{back-to-back } \nu \text{ constraint})$$

Requires e^+, e^- from the same production point

$$E_{e^+} + E_{e^-} \ll E_\gamma - n \times \Delta E \quad (Q\text{-value constraint})$$

Distinguish $e^+ e^- \nu \bar{\nu}$ from $e^+ e^-$



SUMMARY

- 1) $0\nu\beta\beta$, suppressed by $|m_{\beta\beta}|^2$, can be a evidence of **LNV**, but not an unequivocal evidence of sub-eV active neutrino as Majorana.
- 2) Back-to-back $\nu - \bar{\nu}$ in $\gamma + {}^A_ZN \rightarrow {}^A_ZN + e^- + e^+ + (\bar{\nu} + \nu)$, can be an alternative to $0\nu\beta\beta$, and possibly the best way to distinguish Dirac from Majorana for sub-eV active neutrinos.

Conclusion – Final Comment

** The neutrino-less double beta decay (NDBD) has a limitation that it is dependent on the unknown tiny mass of the neutrino as well as it is not an unequivocal evidence of sub-eV active neutrino being Majorana. Our proposals are very interesting viable alternatives to NDBD as far as probing Majorana nature of sub-eV active neutrinos is concerned.

“ urgent to explore whether there are any SM allowed processes and kinematic observables that can directly probe the Majorana nature of neutrinos avoiding this pDMCT constraint, instead of using $0\nu\beta\beta$ or LNV. “

Extra Comments:

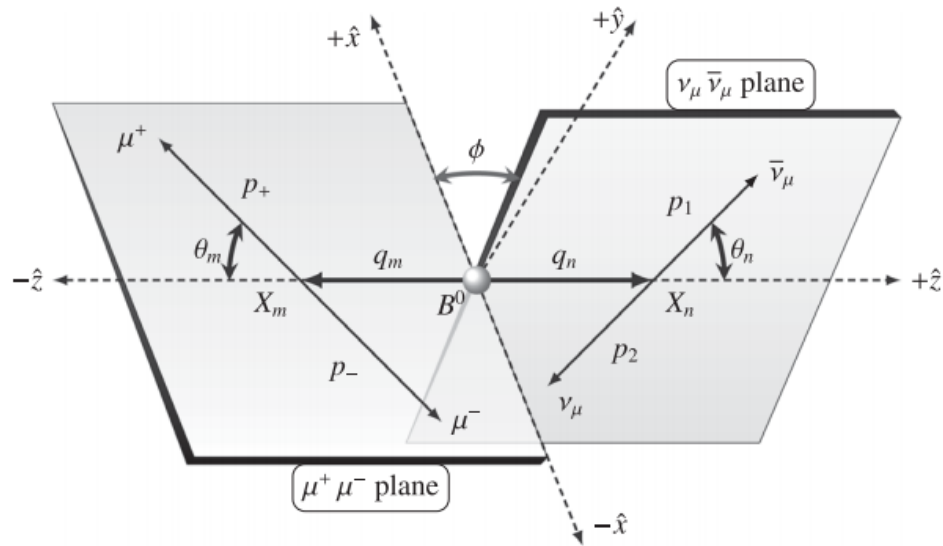
1. Comments on “On the Dirac-Majorana neutrinos distinction in four-body decays” (arXiv:2305.14140) [CSK et al, arXiv:2308.08464]
2. Comments on “Can quantum statistics help distinguish Dirac from Majorana neutrinos?” (arXiv:2402.05172) [CSK et al, arXiv:2402.11386]
3. Comments on “Exploring Quantum Statistics for Dirac and Majorana Neutrinos using Spinor-Helicity techniques” (arXiv:2507.07180) [CSK et al, arXiv:2407.07180]



1. Comments on “On the Dirac-Majorana neutrinos distinction in four-body decays” (arXiv:2305.14140)

[CS Kim et.al, arXiv:2308.08464]

On angle ϕ :



$$B^0 \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2)$$

1) With Back-to-back muons (& neutrinos)

$$\vec{p}_+ + \vec{p}_- = \vec{0} = \vec{p}_1 + \vec{p}_2$$

two straight lines, $\vec{p}_+ + \vec{p}_-$ and $\vec{p}_1 + \vec{p}_2$ meet at one point, making one plane,

$$\phi = 0$$

2) Back-to-back constraints give 3 constraints in phase space integral, only 2 independent variables,

$$E_\mu \quad \text{and} \quad \cos(\theta_m)$$

$$\frac{d^5 \Gamma^{D/M}}{dm_{\mu\mu}^2 dm_{\nu\nu}^2 d\cos\theta_m d\cos\theta_n d\phi} = \frac{Y Y_m Y_n \langle |\mathcal{M}^{D/M}|^2 \rangle}{(4\pi)^6 m_B^2 m_{\mu\mu} m_{\nu\nu}},$$

2. Comments on “Can quantum statistics help distinguish Dirac from Majorana neutrinos?” (arXiv:2402.05172) [CS Kim et.al, arXiv:2402.11386]

1. Decay rate of $B^0 \rightarrow \mu^-(p_-) \mu^+(p_+) \bar{\nu}_\mu(p_1) \nu_\mu(p_2)$ is given by

$$\Gamma^{D/M} = \frac{1}{2m_B} \int |\mathcal{M}^{D/M}(p_+, p_-, p_1, p_2)|^2 d\text{LIPS} (2\pi)^4 \delta^{(4)}(p_B - p_+ - p_- - p_1 - p_2),$$

$$d\text{LIPS} = \left(\frac{d^3 \vec{p}_+}{(2\pi)^3 2E_+} \right) \left(\frac{d^3 \vec{p}_-}{(2\pi)^3 2E_-} \right) \left(\frac{d^3 \vec{p}_1}{(2\pi)^3 2E_1} \right) \left(\frac{d^3 \vec{p}_2}{(2\pi)^3 2E_2} \right),$$

$$\text{mass dim}[\Gamma^{D/M}] = 1.$$

Imposing b2b kinematic condition, $\vec{p}_+ + \vec{p}_- = \vec{0} = \vec{p}_1 + \vec{p}_2$, instead of following suitable variable changes and constraining them, multiplying Dirac delta function $\delta^{(3)}(\vec{p}_1 + \vec{p}_2)$

$$X^{D/M} = \frac{1}{2m_B} \int |\mathcal{M}^{D/M}(p_+, p_-, p_1, p_2)|^2 d\text{LIPS} (2\pi)^4 \delta^{(4)}(p_B - p_+ - p_- - p_1 - p_2) \delta^{(3)}(\vec{p}_1 + \vec{p}_2)$$

neglecting m_μ, m_ν

$$X^{D/M} \approx \frac{1}{32 (2\pi)^6 m_B} \int |\mathcal{M}^{D/M}(\underbrace{p_+, p_-}_{\vec{p}_- = -\vec{p}_+}, \underbrace{p_1, p_2}_{\vec{p}_2 = -\vec{p}_1})|^2 dE_\mu d\cos\theta,$$

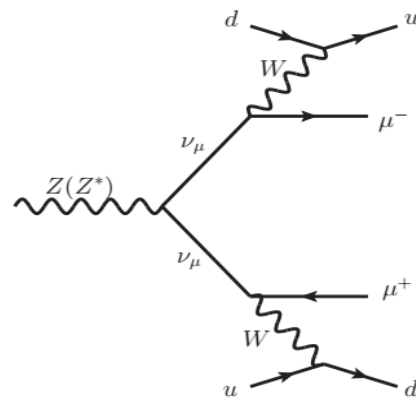
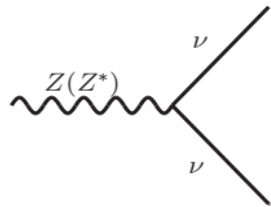
$$\text{mass dim}[X^{D/M}] = -2,$$

due to the fact that Dirac delta function is dimensionful quantity.

Anti-symmetrization of Amplitude : F-D Statistics

****** Should the amplitude be anti-symmetrized for pair of Majorana neutrinos of the same flavor ??

- 1) Quantum statistics requires **absolutely identical indistinguishable particles**.
- 2) **Quantum measurement effect** is analogous to the well-known fact that the interference pattern in a double slit experiment is destroyed if one could somehow know through which slit the photon has passed.



→ requires a method deducing E (or p) of neutrino without destroying quantum effects.

Need **Anti-symmetrization**

No **Anti-symmetrization** (→ similar to the double slit experiment)
once we know $\nu_1 \neq \nu_2$

3. Comments on “Exploring Quantum Statistics for Dirac and Majorana Neutrinos using Spinor-Helicity techniques” (arXiv:2507.07180)

[CS Kim et.al, arXiv:2507.14463]

Consider $B^0(p_B) \rightarrow \mu^-(k)\mu^+(\bar{k})\bar{\nu}(p_1)\nu(p_2)$ $\mathcal{M}^D = \mathcal{M}(p_1, p_2)$ $\mathcal{M}^M = \frac{1}{\sqrt{2}}(\mathcal{M}(p_1, p_2) - \mathcal{M}(p_2, p_1))$

$$\langle |\mathcal{M}^D|^2 \rangle \sim (k \cdot p_1)(\bar{k} \cdot p_2)$$

However, arXiv:2507.07180 added a term by hand to symmetrize

$$\langle |\mathcal{M}^D|^2 \rangle \sim (k \cdot p_1)(\bar{k} \cdot p_2) + (k \cdot p_2)(k \cdot p_1)$$

$$W^+ \rightarrow \ell^+ \bar{\nu}_\ell (\Delta L = -2) \text{ and } W^- \rightarrow \ell^- \nu_\ell (\Delta L = +2)$$

$$\langle |\mathcal{M}^M|^2 \rangle \sim (\bar{k} \cdot p_1)(k \cdot p_2) + (\bar{k} \cdot p_2)(k \cdot p_1) + (\text{interference terms})$$

- 1. There is no fundamental principle or law of physics which requires that the amplitude square for a process having two distinguishable particles in the final state ought to be symmetrized with respect to the exchange of the 4-momenta of the concerned particles, when the two particles are not detected in the detector
- 2. In the standard model (SM) the Dirac neutrino and antineutrino carry lepton numbers;
 $W^+ \rightarrow \ell^+ \nu_\ell$ and $W^- \rightarrow \ell^- \bar{\nu}_\ell$ for $\ell = e, \mu, \tau$, to ensure lepton number conservation
- 3. For Majorana neutrinos, for which $\nu = \bar{\nu}$ and the lepton number is not a good quantum number.