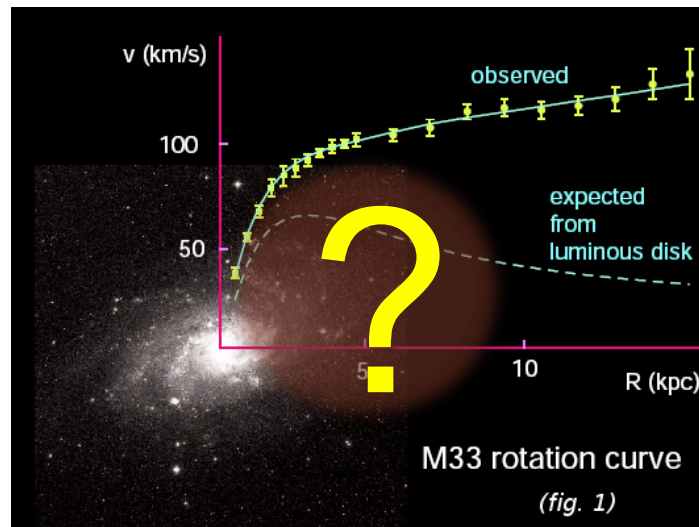
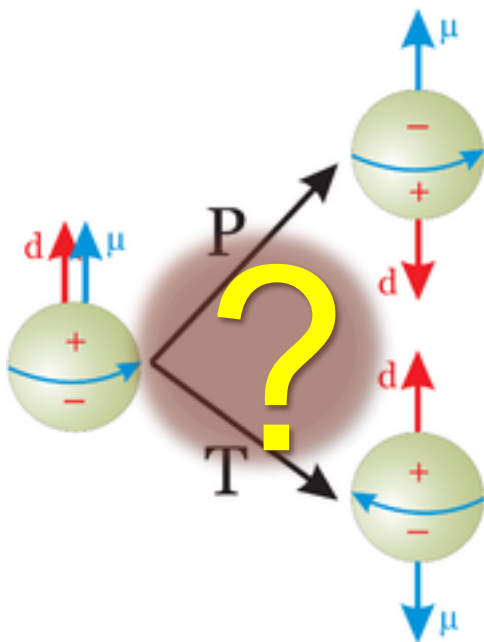


Axion Research at CAPP/IBS

SungWoo YOUN
CAPP/IBS

CosPA 2015, Daejeon, Korea
Oct. 12~16, 2015





CAPP/IBS

- *Center for Axion and Precision Physics Research*

- *Funded by the Institute for Basic Science*
- *Located at KAIST campus in Daejeon, Korea*
- *Led by Prof. Semertzidis Yannis*

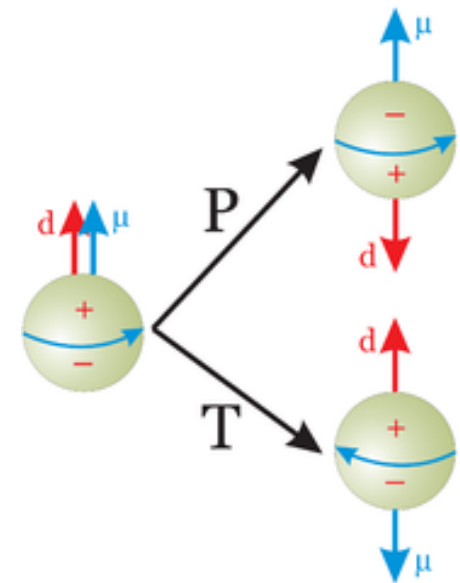


- *Experimental efforts to address the Strong-CP problem*
 - *State of the art dark matter axion experiment*
 - *Storage ring proton EDM*
- *Leading roles in international collaborations in intensity frontier*
 - *Muon $g-2$*
 - *COMET ($\mu \rightarrow e$ conversion)*
- *Plenary talk by Prof. Semertzidis tomorrow for more details*



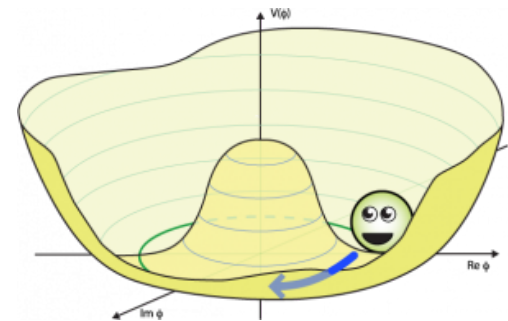
Strong-CP Problem

- *CP symmetry breaking (CPV)*
 - *Matter-antimatter imbalance in our universe*
- *CPV in electroweak interaction*
 - *Observations are consistent with predictions*
 - *CPV phase in CKM matrix*
- *CPV in strong interaction*
 - *Theory: natural CPV term due to non-zero QCD vacuum*
 - *Sizable neutron dipole moment (d_n)*
 - *At least one of the quarks are massless*
 - *Experiment: QCD respects CP symmetry*
 - $d_{n,exp} \leq 10^{-26} \text{ e*cm} \Rightarrow \Theta_{QCD} \leq 10^{-10}$
 - *None of the quarks are massless*
 - *This is known as **strong-CP problem***
 - *One of the most important but unsolved problems in physics*



Axion

- *Peccei-Quinn theory (1977)*
 - An elegant idea to solve the strong CP problem
 - A new global symmetry, $U(1)_{PQ}$, with a scalar field permeating all space
 - Spontaneous (explicit) PQ symmetry breaking involving a **new Goldstone boson: axion**
 - Similar to Higgs mechanism
 - Spontaneous EWSB involving the Higgs boson
 - Discovery in 2012 has enhanced the interest in the axion and the possibility of its existence



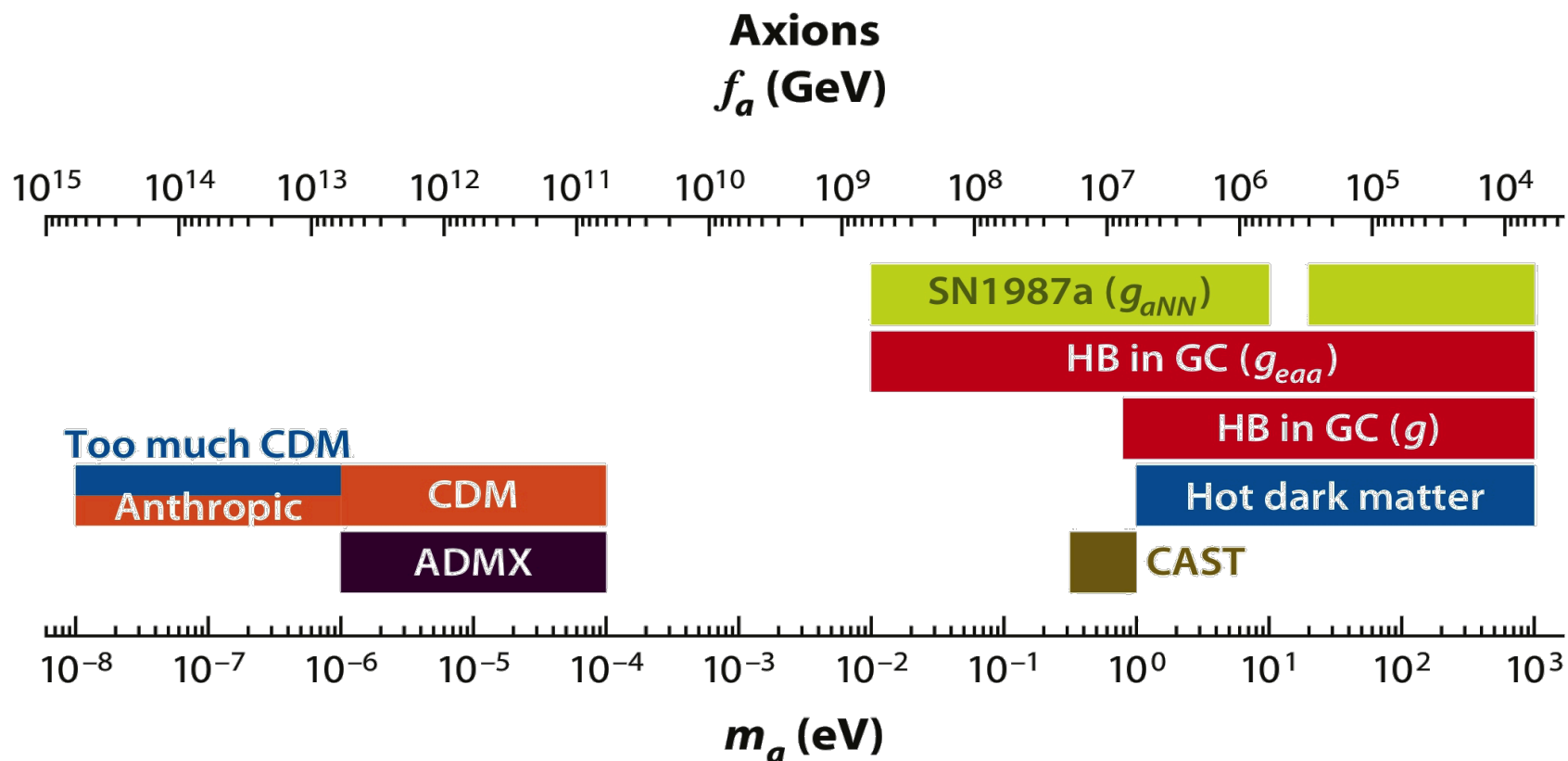
Axion properties			
Interaction	Gravity, EM	Mass	$\mu\text{eV to eV}$
C	0	ρ_{local}	0.45 GeV/cm^3
J^P	0^-	$\beta \sim 10^{-3} \rightarrow Q_a \sim 10^6$	



Axion and Dark Matter

- Attractive dark matter candidate*

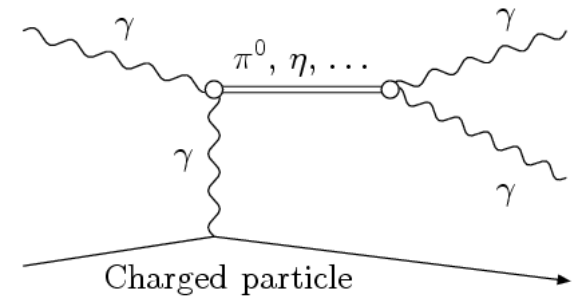
- If its mass lies between $1 \mu\text{eV}$ and $100 \mu\text{eV}$, the axion can fit the halo model*
- Referred to as **axion dark matter***



Detection Strategy

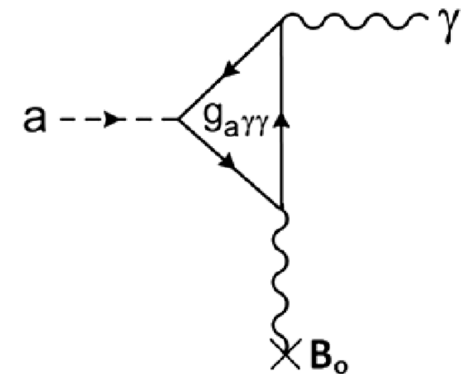
- *Primakoff Effect*

- Pseudoscalar production by a photon scattering with an EM field
- $\gamma\gamma \rightarrow \pi^0, \eta, \dots$



- *P. Sikivie's method (1983)*

- Reverse Primakoff effect
- Conversion of axions into microwave photons
 - $a \rightarrow \gamma\gamma$ (cf. $\pi^0 \rightarrow \gamma\gamma$)
- Detectable in a EM rasontor in the presence of a strong magnetic field
 - Principle of haloscope
 - Most promising technique for the faintest axion-photon coupling

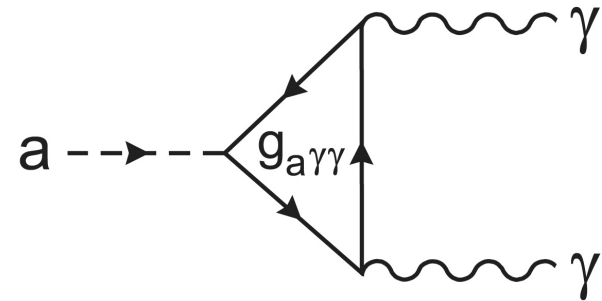


Axion Coupling to Photons

• Lagrangian

$$L_{a\gamma\gamma} = -\frac{g_{a\gamma\gamma}}{4} \varphi_a F_{\mu\nu} \tilde{F}^{\mu\nu} = -g_{a\gamma\gamma} \varphi_a \mathbf{E} \cdot \mathbf{B}, \quad g_{a\gamma\gamma} \equiv \frac{\alpha}{\pi} \frac{g_\gamma}{f_{PQ}}$$

- $g_{a\gamma\gamma}$: coupling constant
- φ_a : axion field
- F : EM field-strength tensor
- α : fine structure constant
- g_γ : model-dependent coefficient
 - 0.97 for KSVZ (hadronic)
 - -0.36 for DFSZ (minimum)

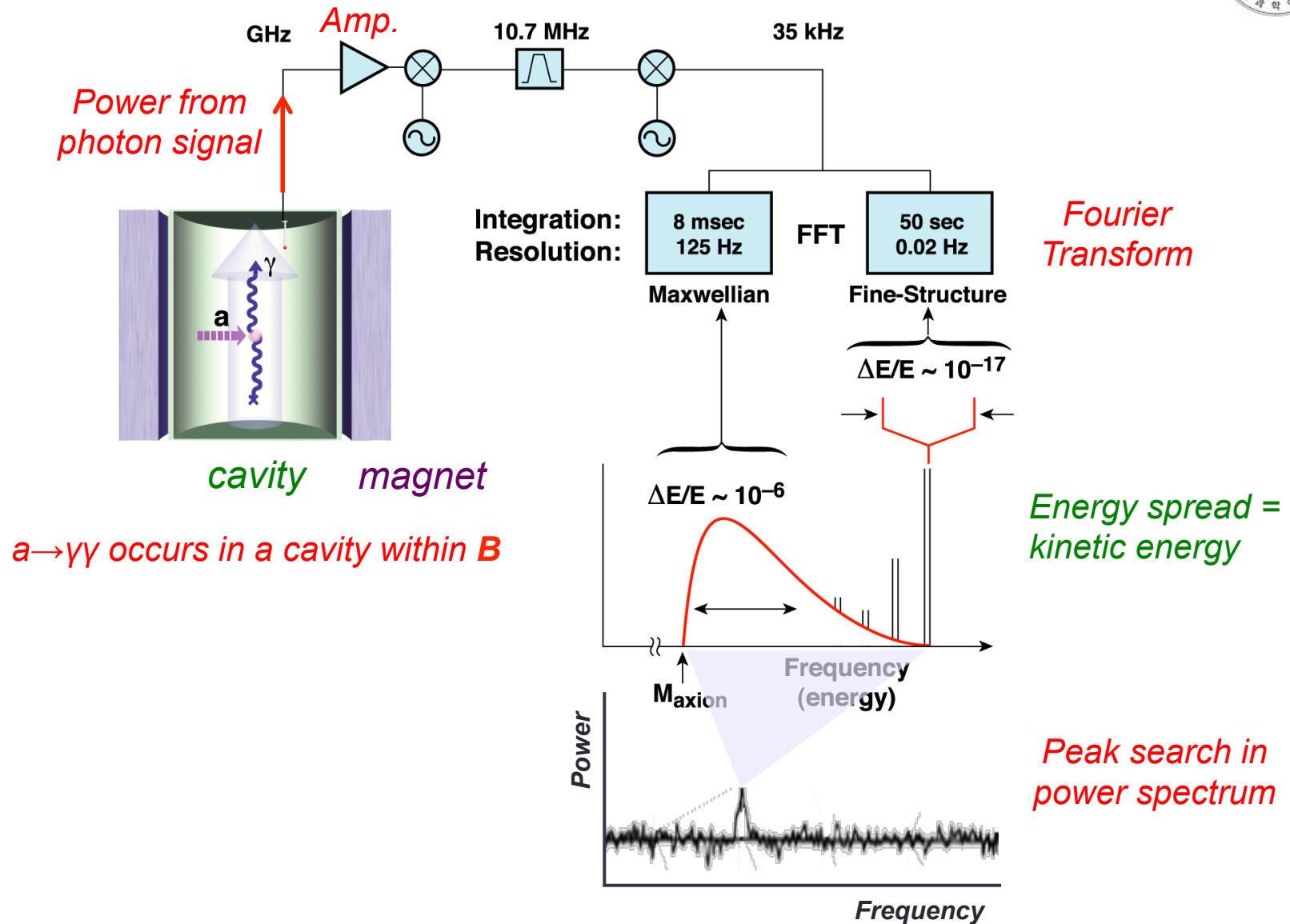


$$g_\gamma = \left(\frac{E}{N} - \frac{2}{3} \frac{4+z}{1+z} \right)$$

- f_{PQ} : PQ symmetry breaking scale ($10^{10} \sim 10^{12}$ GeV)

$$m_a = \frac{f_\pi m_\pi}{f_{PQ}} \frac{\sqrt{z}}{1+z} \simeq 6 \mu\text{eV} \frac{10^{12} \text{ GeV}}{f_{PQ}}, \quad z = \frac{m_u}{m_d}$$

Axion Detection





Conversion Power

- Conversion power*

- Theoretical parameters
- experimental parameters

$$P_{a \rightarrow \gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C_{mnp} \min(Q_L, Q_a)$$

Coupling constant ——— *Axion number density* ——— *Magnetic field* ——— *Effective volume* ——— *Cavity Q factor*
Axion Q factor

$$P_a = 5.2 \times 10^{-22} W \left(\frac{g_\gamma}{0.97} \right)^2 \left(\frac{\rho_a}{0.45 \text{ GeV/cc}} \right) \left(\frac{f_a}{6 \text{ GHz}} \right) \left(\frac{B}{8 \text{ T}} \right)^2 \left(\frac{V}{1 \text{ L}} \right) \left(\frac{C_{010}}{0.6} \right) \left(\frac{Q_L}{10^6} \right)$$



Conversion Power and Sensitivity

• Conversion power

- Theoretical parameters
- experimental parameters

$$P_{a \rightarrow \gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C_{mnp} \min(Q_L, Q_a)$$

Coupling constant $g_{a\gamma\gamma}$
 Axion number density ρ_a
 Magnetic field B
 Effective volume V
 Cavity Q factor Q_L
 Axion Q factor Q_a

• Signal-to-noise ratio (SNR)

$$SNR \equiv \frac{P_{\text{signal}}}{P_{\text{noise}}} = \frac{P_{a \rightarrow \gamma\gamma}}{k_B T_{\text{syst}}} \sqrt{\frac{t_{\text{int}}}{\Delta f_a}}$$

Integration time t_{int}
 Axion bandwidth Δf_a ($\sim 10^{-6} f$)
 System noise temperature T_{syst}

• Scan rate

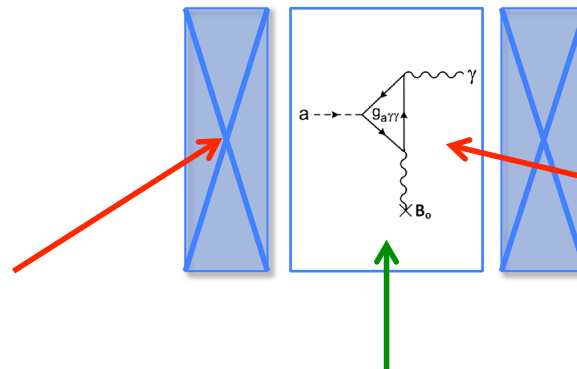
$$\frac{df}{dt} = \left(\frac{1}{SNR} \right)^2 \left(\frac{P(f)}{k_B T_{\text{syst}}} \right)^2 \cdot \frac{Q_a}{Q_L} \sim B^4 V^2 C^2 Q_L T_{\text{syst}}^{-2}$$

$$\frac{df_d}{dt} = \frac{140 \text{ MHz}}{\text{year}} \left(\frac{g_\gamma}{0.97} \right)^4 \left(\frac{\rho_a}{0.45 \text{ GeV/cc}} \right)^2 \left(\frac{B}{8 \text{ T}} \right)^4 \left(\frac{V}{1 \text{ L}} \right)^2 \left(\frac{C_{010}}{0.6} \right)^2 \frac{Q_L}{Q_a} \left(\frac{4}{SNR} \right)^2 \left(\frac{4.5 \text{ K}}{T_{\text{syst}}} \right)^2 \left(\frac{f}{6 \text{ GHz}} \right)^2$$

Experimental Setup

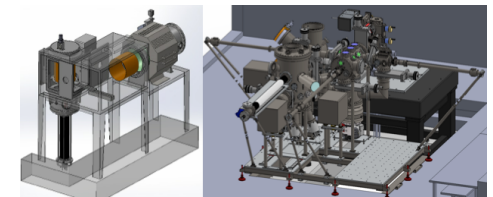
- Axion conversion power:*
$$P_{a \rightarrow \gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C_{mnp} \min(Q_L, Q_a)$$

High field SC magnet
25T and then 35T or 40T
From BNL (HTS Technology)



Primakoff Effect

High Q tunable cavity
Superconducting Coating
From Prof. Jhinhwan Lee of KAIST



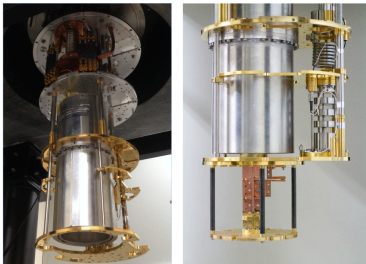
Experimental Setup

- Axion conversion power:*
$$P_{a \rightarrow \gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C_{mnp} \min(Q_L, Q_a)$$
- Signal-to-Noise Ratio:*
$$SNR \equiv \frac{P_{\text{signal}}}{P_{\text{noise}}} = \frac{P_{a \rightarrow \gamma\gamma}}{k_B T_{\text{syst}}} \sqrt{\frac{t_{\text{int}}}{\Delta f_a}}$$

Cryogenics

<100mK

Prof. Hyungssoon Choi from KAIST



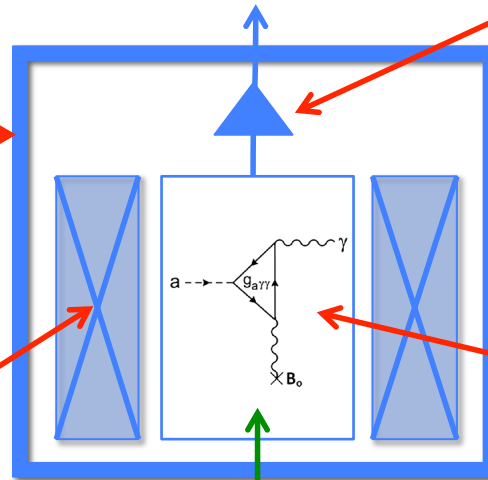
High field SC magnet

25T and then 35T or 40T

From BNL (HTS Technology)



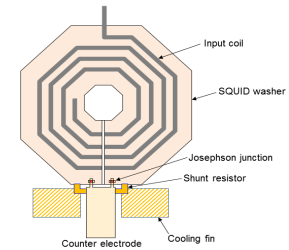
To RF Receiver



Primakoff Effect

SQUID amplifier

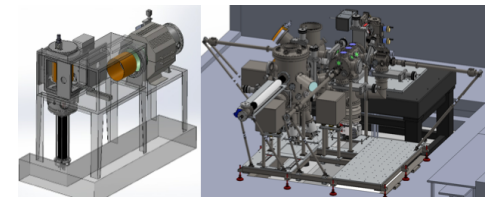
Outsourced Research from KRISS



High Q tunable cavity

Superconducting Coating

From Prof. Jhinhwan Lee of KAIST





CAPP/IBS Axion Research

- *Focuses on improvements in experimental parameters*

$$\frac{df}{dt} \sim B^4 V^2 C^2 Q_L T_{\text{syst}}^{-2}$$

- *Magnetic Field*
 - *Critical (4th power), but technically challenging and expensive*
- *Cavity volume*
 - *Limited by the magnet bore size*
 - *Higher frequencies require smaller cavities*
- *Quality factor*
 - *Several ideas in community – pure metal, superconductor, ...*
- *System temperature*
 - *Dilution refrigerator technique*
 - *SQUID amplifier*

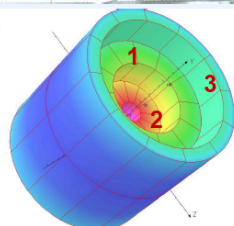
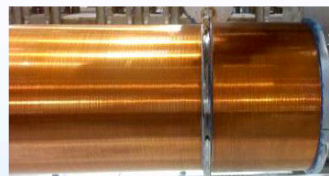
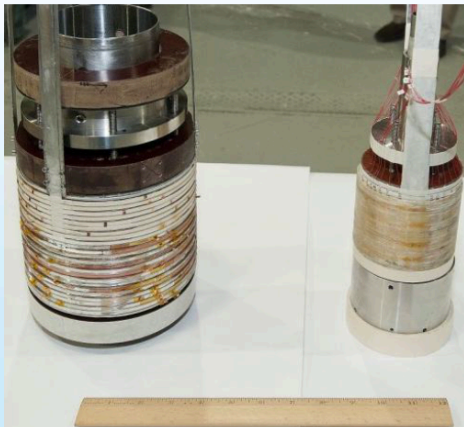
High Magnetic Field: B

- *Launched a R&D program*
 - *BNL's Superconducting Magnet R&D Group (Dr. R. Gupta)*
 - *Determining the cable for the final design within a year*
- *Two stage program*
 - *25 T with 10 cm bore (HTS)*
 - *35(40) T with 10 cm bore – hybrid design HTS(inner)/LTS(outer)*
 - *cf. current axion experiments use < 10 T*

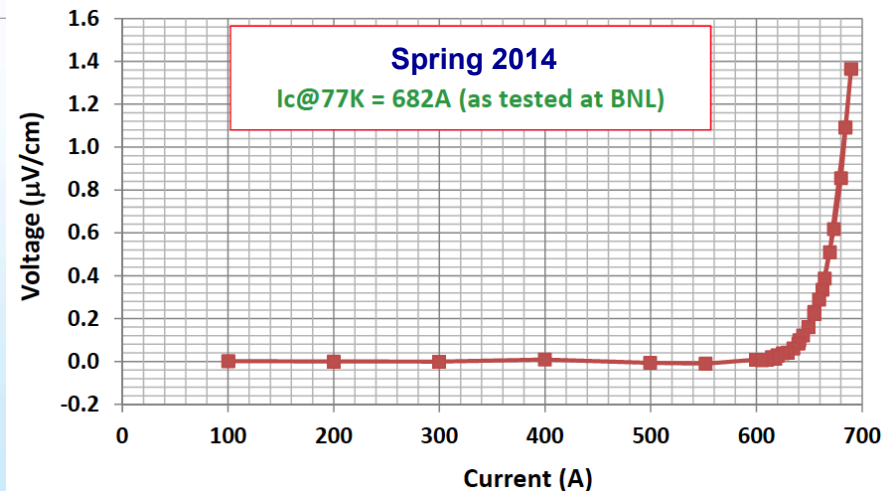
BROOKHAVEN
NATIONAL LABORATORY
Superconducting
Magnet Division

Status of High Field MAP Solenoids

Two HTS coils together made with SuperPower
HTS is expected to create 20-25 T, if successful



~30 T with NbTi outer
(40 T with Nb₃Sn or more HTS)

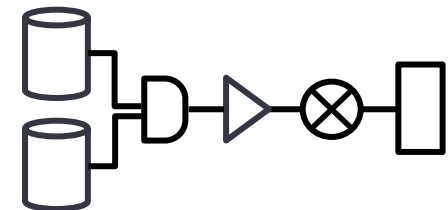
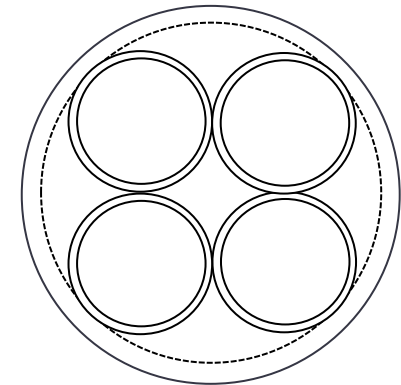
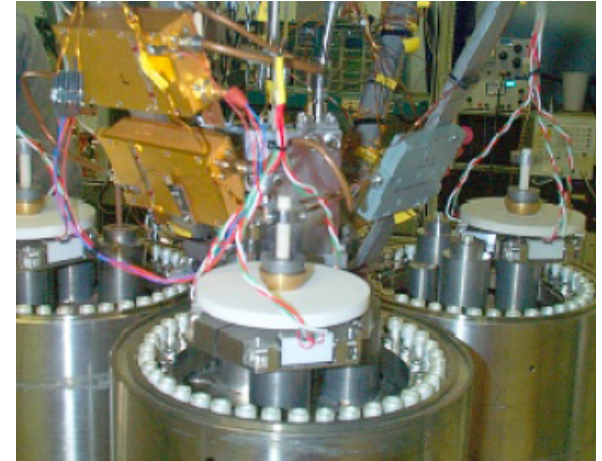


Large Cavity Volume: V

- *Multiple cavities in-phase*
 - *First experimental trial by ADMX (2000) using a 4 cavity system*
- *Improves sensitivity by up to N w.r.t. a single cavity system*

$$SNR: \frac{N \cdot P_S}{\sqrt{N \cdot P_C + P_A}} \quad \text{v.s.} \quad \frac{P_S}{P_C + P_A}$$

- *N: cavity multiplicity*
- *Key issue is phase-locking*
 - *Challenging*
 - *Phase matching in both frequency and time domains*
- *I have received the 5-year IBS Young-Scientist award with CAPP to develop this system*



High Quality Factor: Q

- Typical Cu cavity*

- $Q_{max} \sim 10^5$ with annealing
- cf. $Q_a \sim 10^6$

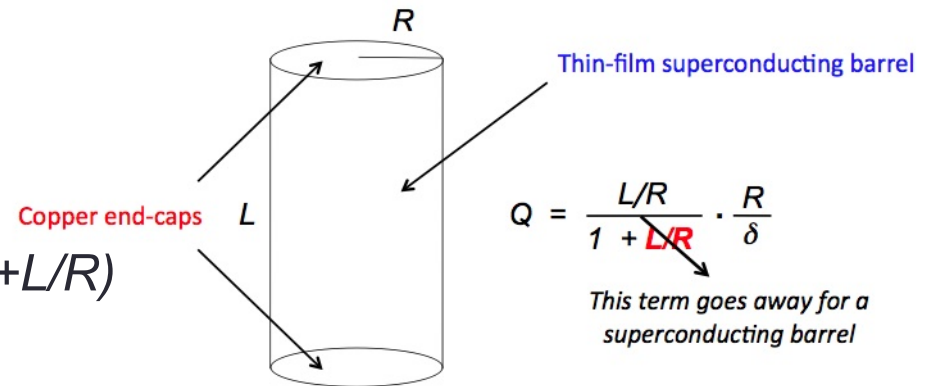
- Superconducting cavity*

- Hybrid design improves Q by $(1+L/R)$

$$Q = \omega \frac{\text{Stored energy}}{\text{Power loss}}$$

$$= \left(\frac{2V}{S\delta} \right) = \frac{2\pi R^2 L}{2\pi R^2 + 2\pi RL} \frac{1}{\delta} = \frac{L}{R+L} \frac{R}{\delta}$$

The concept of a hybrid superconducting cavity:




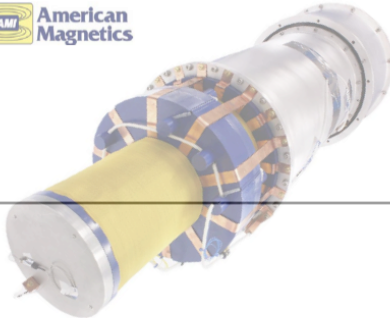


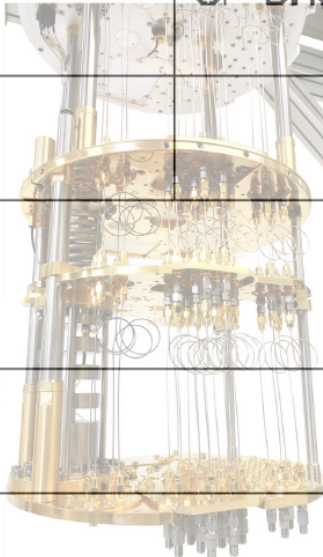
$$Q_{hybrid} = (1 + L/R) \cdot Q_{cu}$$

For typical ADMX cavity, $L/R = 5$, enhancement factor = 6

- R&D program with Prof. J. Lee (KAIST/IBS) for SC cavity*

- SC walls including top/bottom plates
- CAPP goal: up to $Q \sim 10^7$ in high B-field

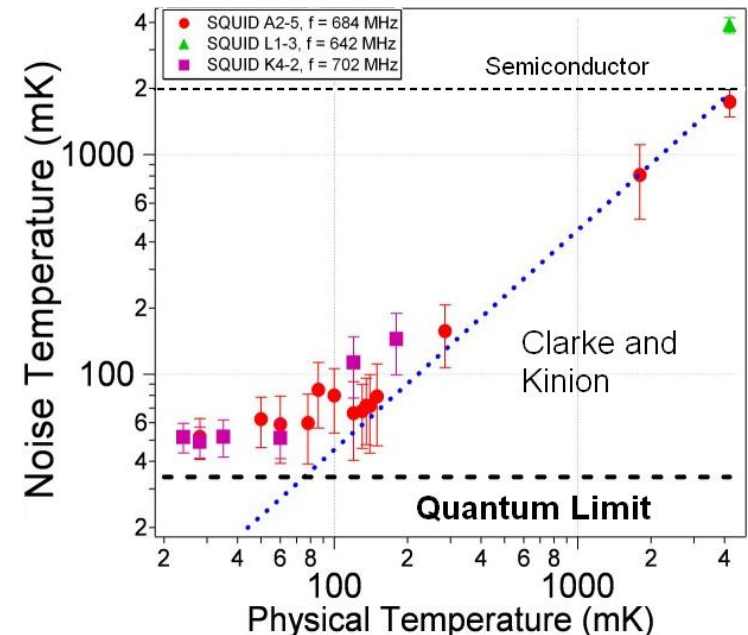
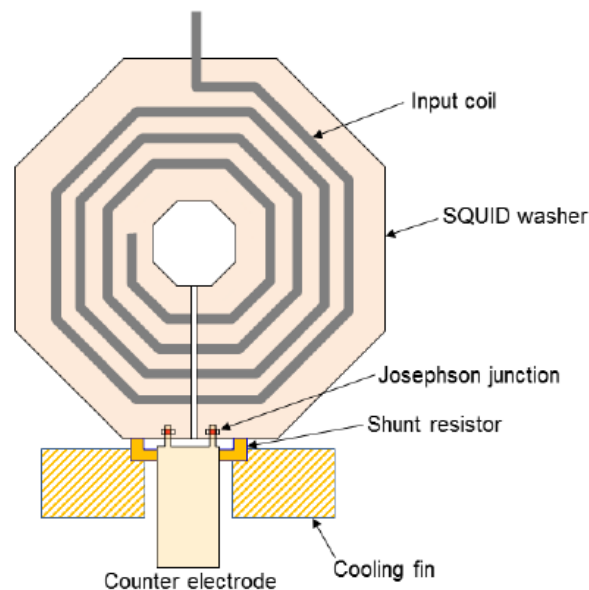
Low System Temperature: T – Cryostats

	2014	2015	2016	2017	2018
Essential Equipments	CF-DR(RF1) CF-DR(magnet) Wet-He3(large bore)			<div> American Magnetics</div> <div></div>		
Quantum Amplifier Research	<div></div>	<div> QUOTATION</div> <div>CF-DR(RF2)</div>				
Small-scale Integration	<div></div>			CF-DR(testbed) <div><p>Example of a 7T-100mm cryogen free AMI magnet integrated with a BF-LD system (Protective aluminum magnet cover shield not shown in picture).</p></div>		
Low-noise Experiments				Wet-DR1(precision) Wet-DR2(precision)		
AxionDetector main				Main DR (Axion Detector)		
Helium Liquefier	<div><p>On the left 4x superconducting (none-attenuated) RF lines in a KF40 LOS-port, in the middle 12x CuNi lines in a KF63 LOS-port and right 4x CuNi (attenuated) lines in KF40 LOS-port.</p></div>			Helium Liquefier		

On the left 4x superconducting (none-attenuated) RF lines in a KF40 LOS-port, in the middle 12x CuNi lines in a KF63 LOS-port and right 4x CuNi (attenuated) lines in KF40 LOS-port.

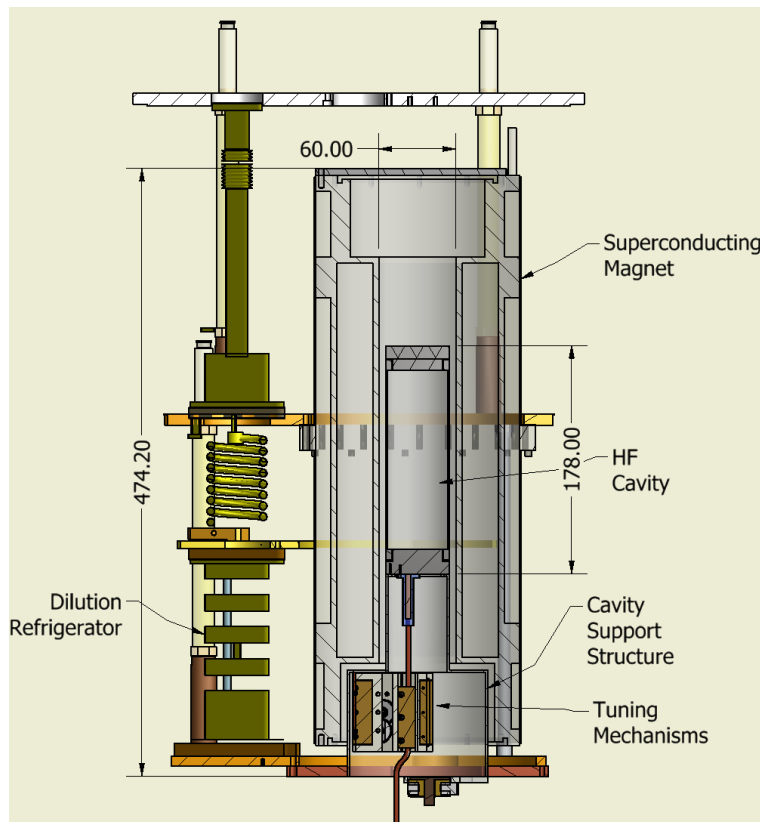
Low System Temperature: T – SQUID

- *Started a development program with KRISS (5 year)*
 - *Korea Research Institute of Standards and Science*
 - *Providing us with (near) quantum noise limited SQUID amplifiers in the 1-20 GHz range*
 - $T_{\text{quantum}} \approx 50 \text{ mK} * f \text{ (GHz)}$
 - *Evaluating methods for higher frequencies*



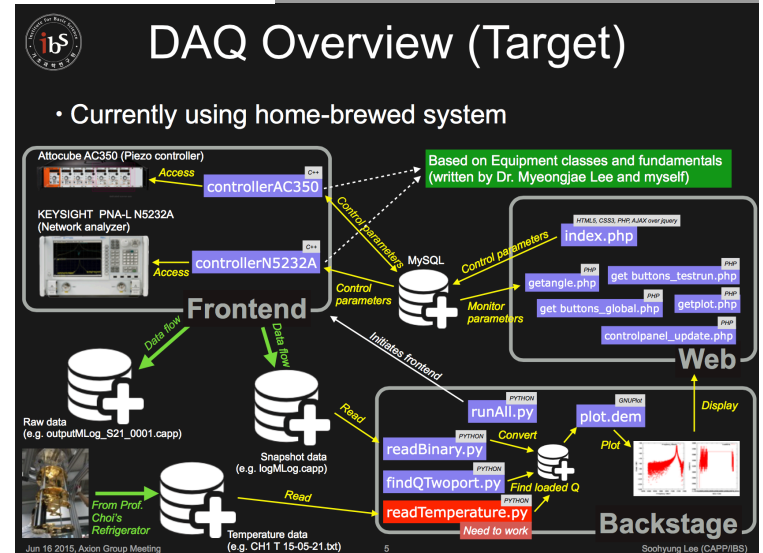
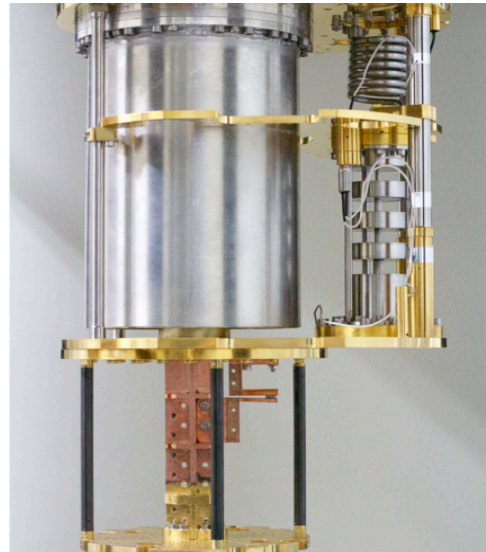
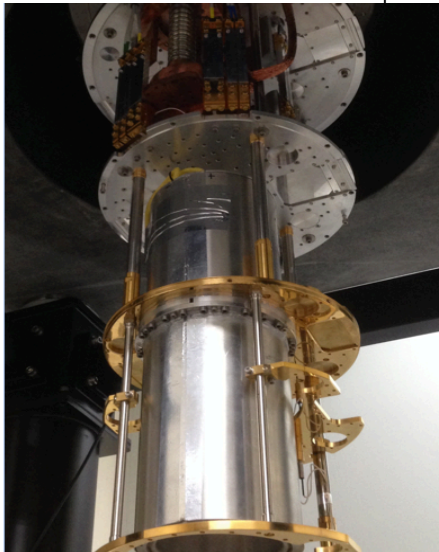
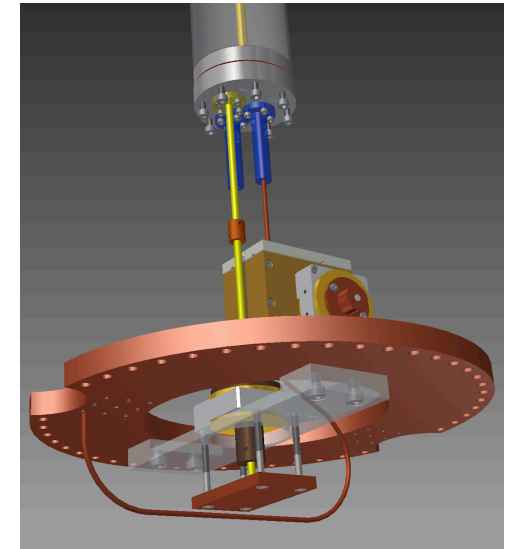
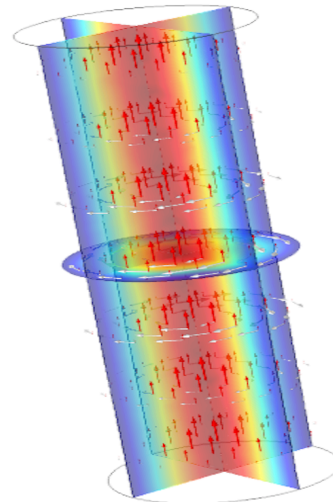
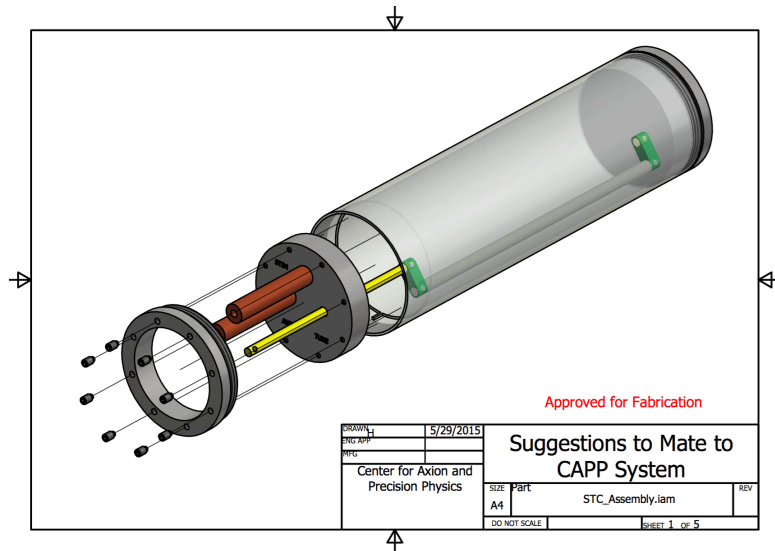
CULTASK

- *CAPP Ultra Low Temp Axion Search in Korea*
 - *Corporate body of axion research efforts at CAPP*
 - *Coldest axion experiment (<100 mK)*



- *Currently in engineering run*
 - *Building infrastructure*
- *Status*
 - *Designing cavities with tuning mechanism*
 - *Completed RT electronics test*
 - *Cryogenic RF circuits test ongoing*
 - *DAQ system ready*
 - *Start with ~ 5 GHz (ID: 4.5 cm)*

CULTASK

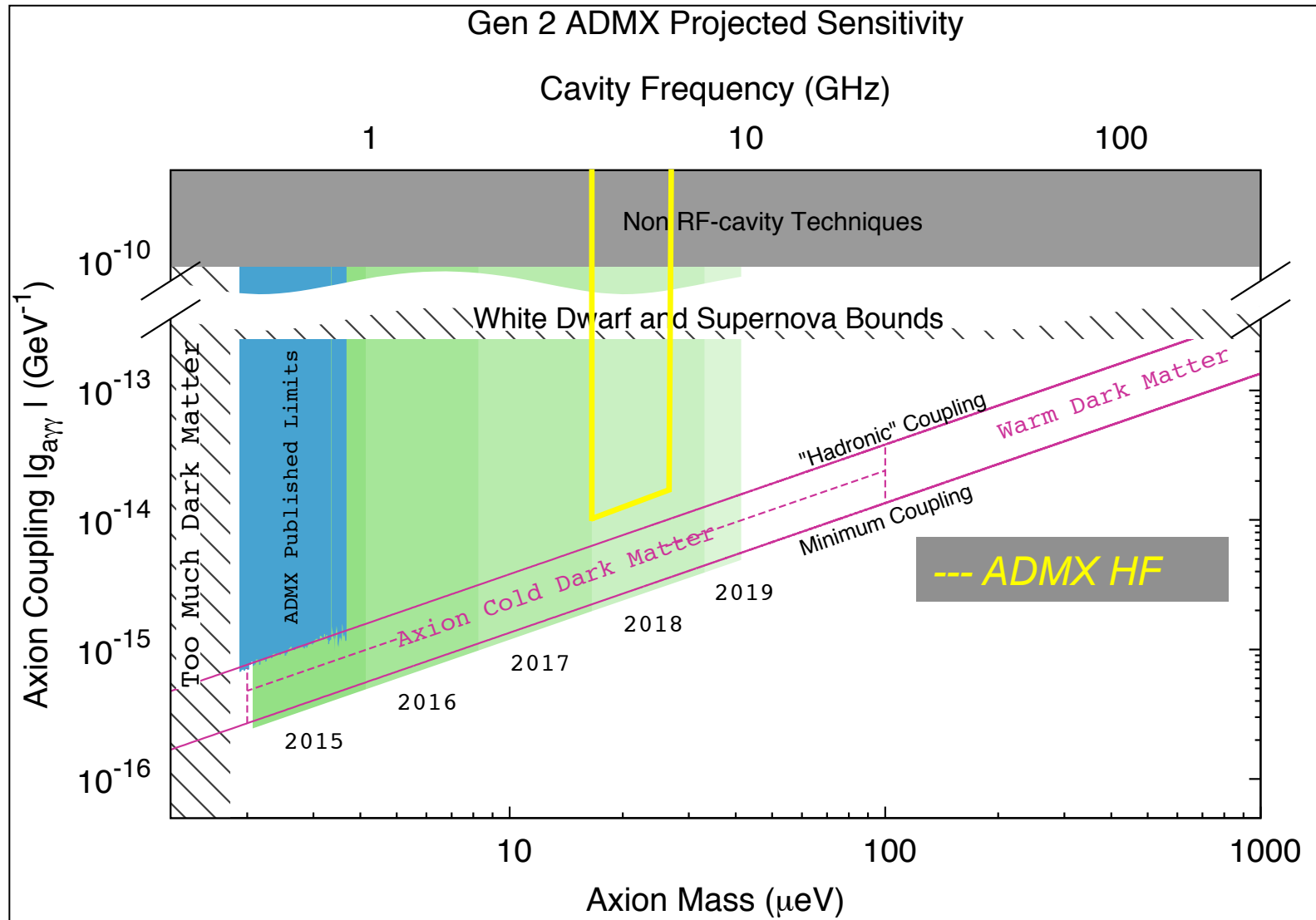




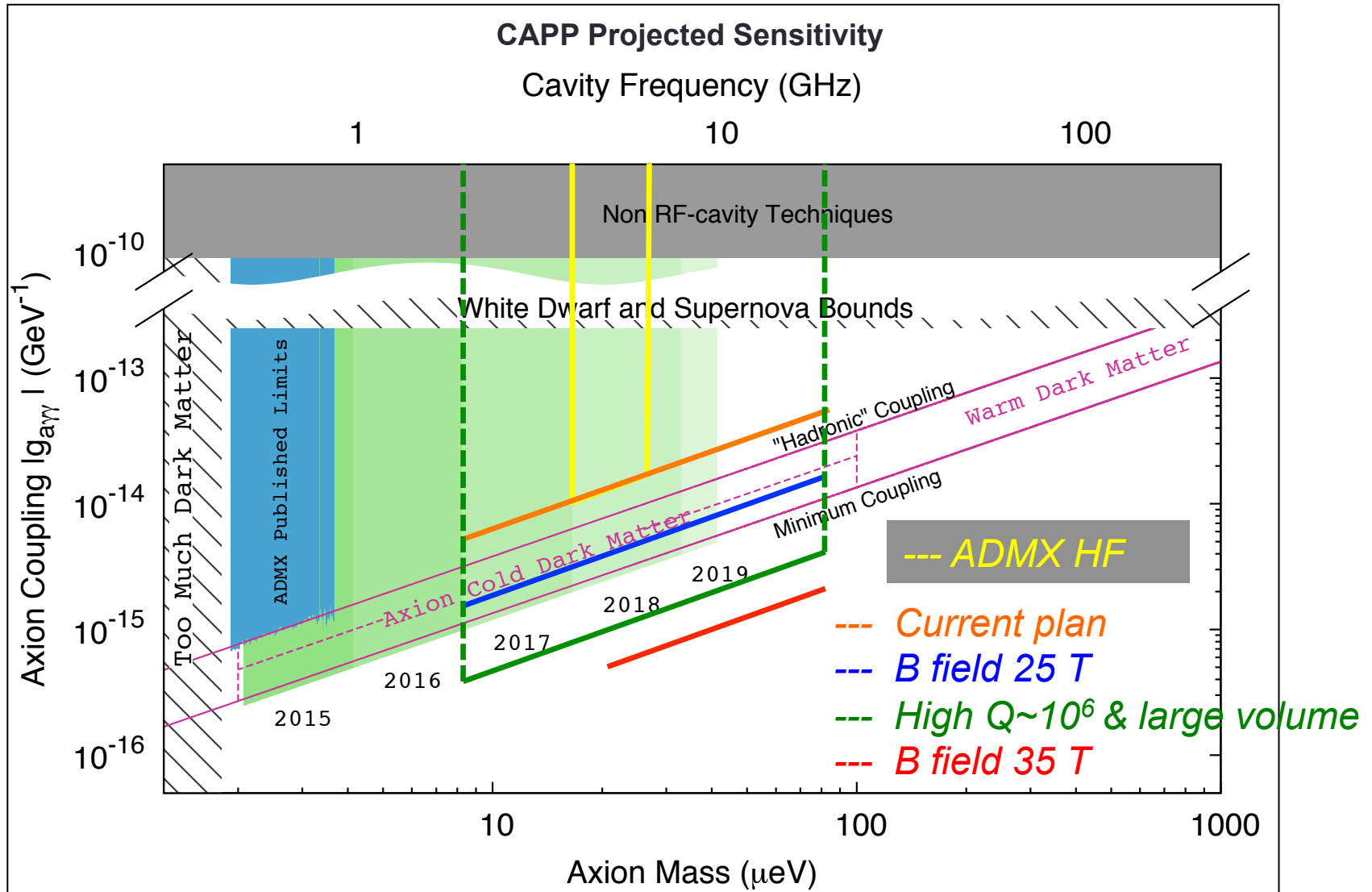
Timeline

	2015	2016	2017	2018	2019
Lab Space	Munji Campus Design & Renovation	Occupation			
High Field Magnet	Prototype, testing of SuperC cables	25T, 10cm bore SuperC Magnet design	Work on 35T, 10 cm bore SuperC magnet	Magnet Delivery	
SC Cavity Development	Procure Equip. Study res. and geom.	Development of high Q SC resonator	Production of high Q resonator		
SQUID Amplifier	Design and production of prototype SQUID for 1-10 GHz Acquire JPA and test		SQUID delivery from KRISS Develop higher freq. amplifier		
Axion Cavity Experiment	Building infrastructure. Engineering Run at KAIST	Experimental Setup at Munji Test Runs		High Field Magnet + SQUID + SC Cavity	
Cryogenics	Setup plan, acquisition	DR's(2) + Wet He3 & DR	DR's(2)	Main DR and He liquefier	

Projected Sensitivity – ADMX Gen 2



Projected Sensitivity – CAPP





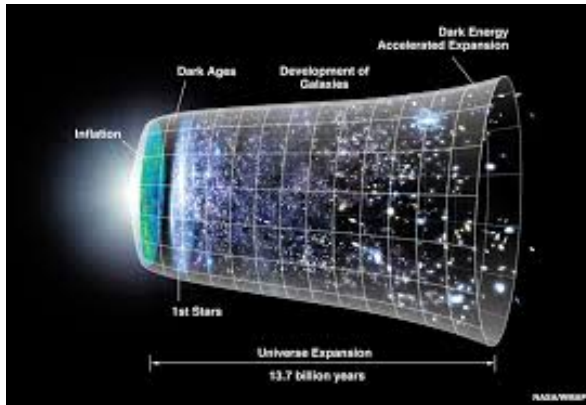
Summary

- *Axion is the key particle to solve the strong-CP problem and dark matter issue*
- *Cavity-based approach is promising to detect the challenging axion-photon coupling*
- *CAPP/IBS is an experimental particle physics group dedicated to dark matter axion searches in Korea*
- *Exciting results are expected in a few years*

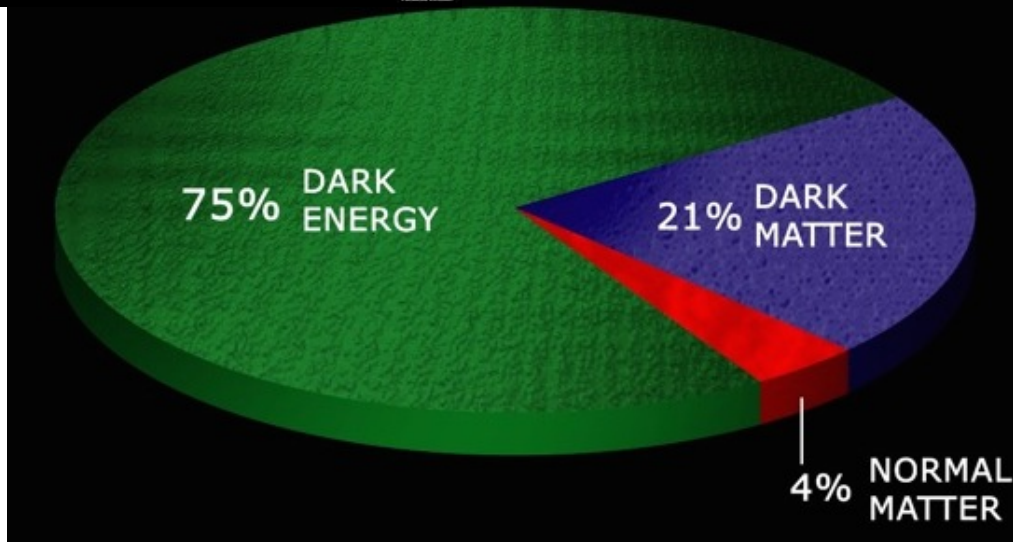




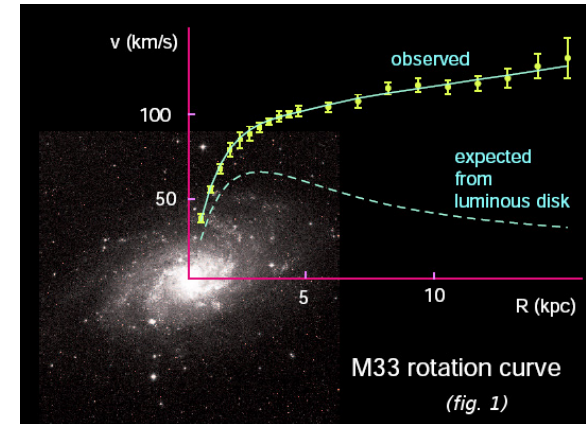
Our Universe



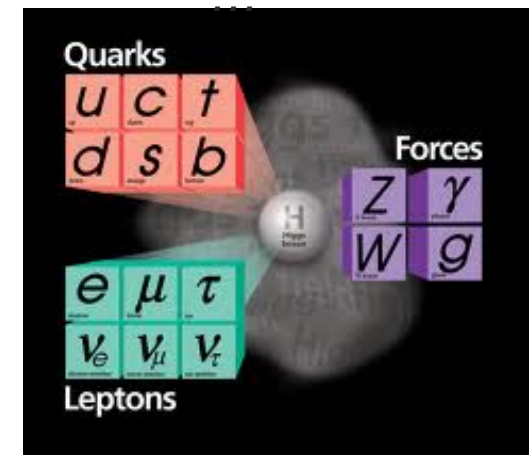
Dark Energy
Expanding universe



Normal Matter
*Standard Model
(EM, weak, strong)*



Dark Matter
*Rotation velocity of galaxies
Gravitational Lensing
Cosmic Microwave Background*



High Quality Factor: Q (II)

- RRR (Residual Resistance Ratio)*

$$Q \sim \frac{V}{S\delta}, \quad \delta = \sqrt{\frac{2}{\omega\mu\sigma}}$$

- Comprehensive studies for Cu and Al
- Magnetic field effect

- DC vs. AC*

- Anomalous behavior of skin depth at high frequencies at low temperatures
- High purity Al is preferable (?)
- Dedicated study is planned

