



Study of the neutrinoless double beta decay of ^{100}Mo with the CUPID-Mo demonstrator

Andrea Giuliani



Outline

- Challenges in neutrinoless double beta decay
- Scintillating bolometers and ^{100}Mo
- Results with $\text{Li}_2^{100}\text{MoO}_4$ crystals
- CUPID-Mo
- Prospects

Neutrinoless double beta decay ($0\nu2\beta$): standard and non-standard mechanisms

$0\nu2\beta$ is a test for « creation of leptons »:



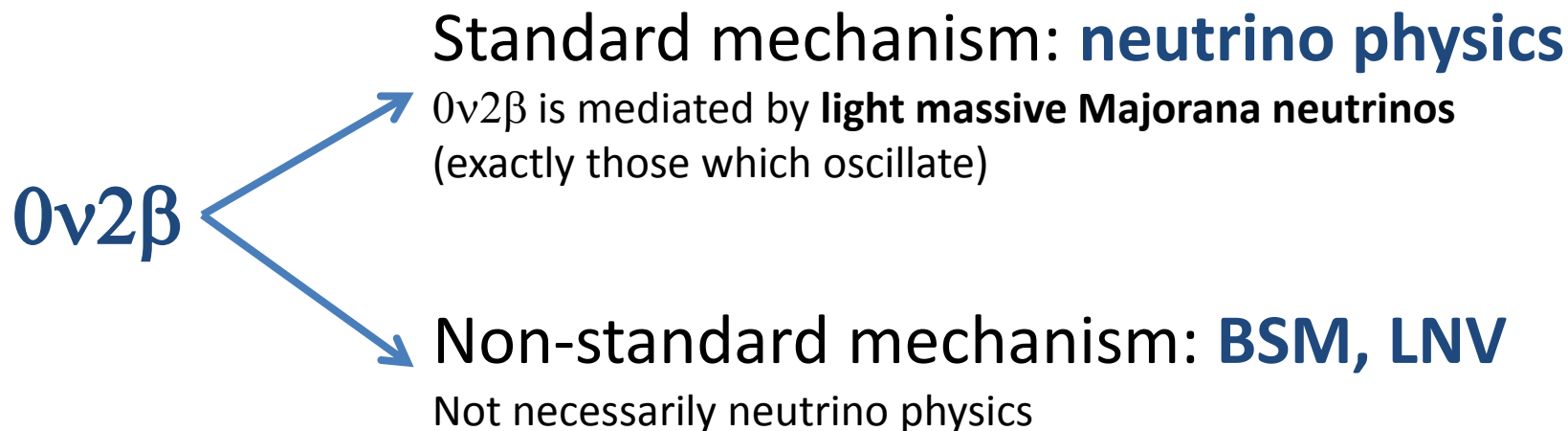
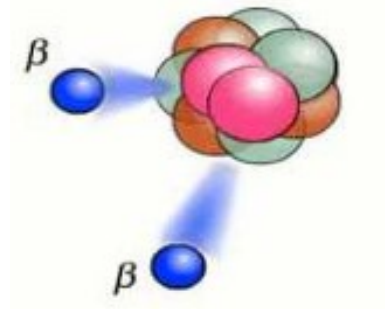
This test is implemented in the nuclear matter:



Energetically possible for **35 nuclei**

Only a few are experimentally relevant: ^{82}Se , ^{76}Ge , ^{100}Mo , ^{130}Te , ^{136}Xe

Enrichment is mandatory, **with the exception of ^{130}Te**



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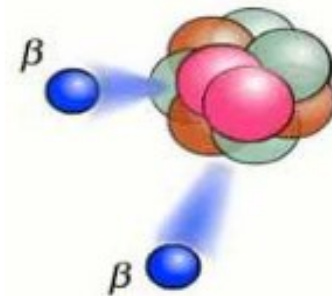
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Neutrinoless double beta decay is the only experimentally viable process that can ascertain the Majorana nature of neutrinos:

- New form of matter (fermions $\equiv$ antifermions)
- Naturally incorporated in Grand Unification Theories
- Explain smallness of neutrino masses (see-saw mechanism)
- Explain matter / antimatter asymmetry in the Universe (leptogenesis)

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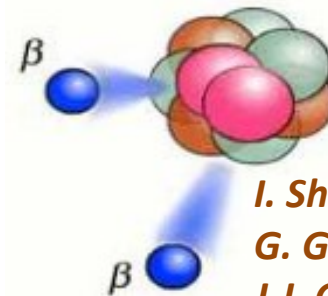
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I. Shimizu
G. Gratta
J.J. Gomez-Cadenas
Ke Han

S. Pirro
R. Saakyan
S. Schoenert
B. Zhu

HongJoo Kim

M. Chen
Ke Han

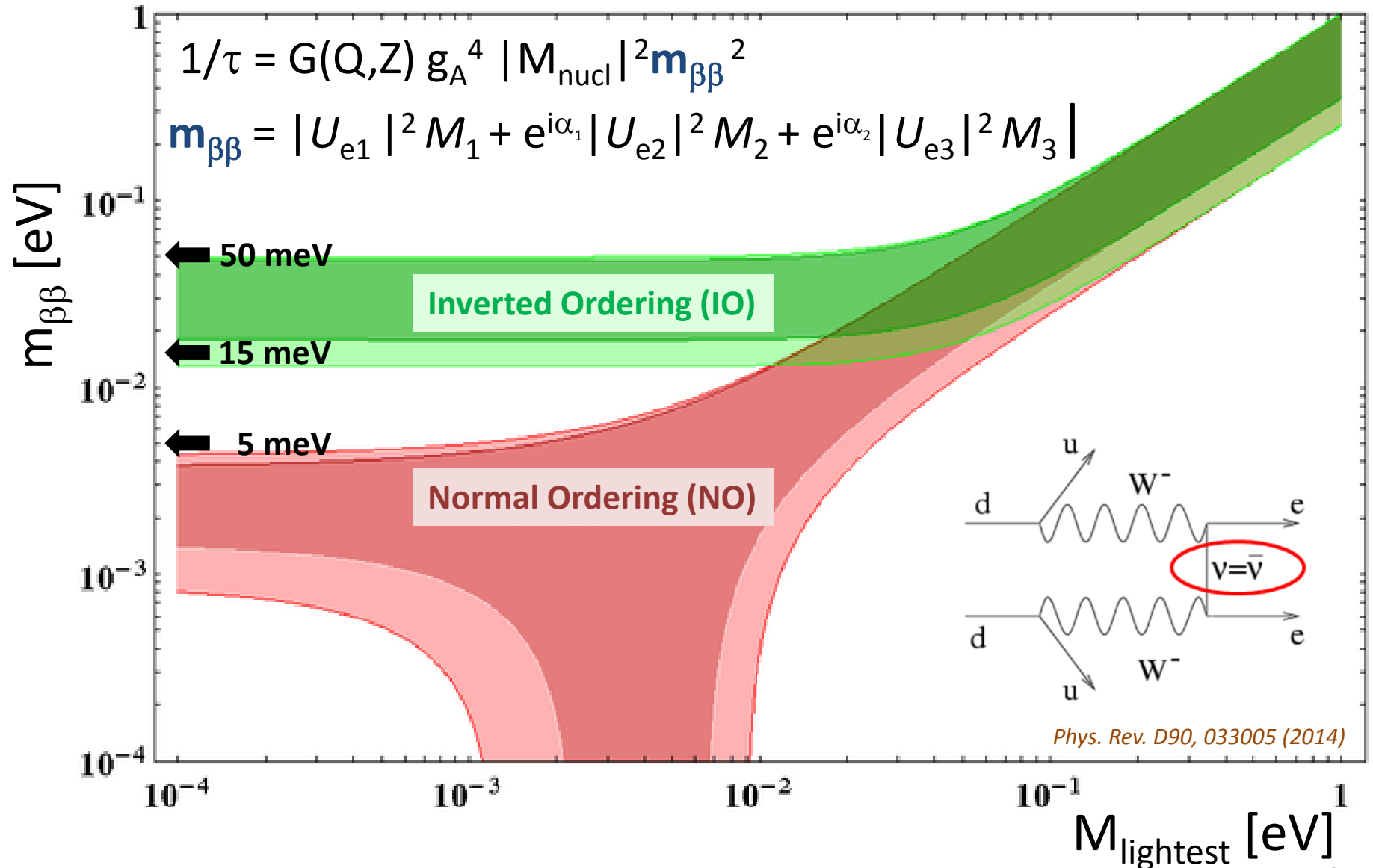
Neutrinoless double beta decay
that can ascertain the Majorana

- New form of matter (fermions)
- Naturally incorporated in GUTs
- Explain smallness of neutrino masses
- Explain matter / antimatter asymmetry

F. Simkovic
H. Ejiri
S. Stoica
R. Dvornicky
F. Cappuzello
J. Engel
J. Suhonen
S. Eliseev
J. Menendez

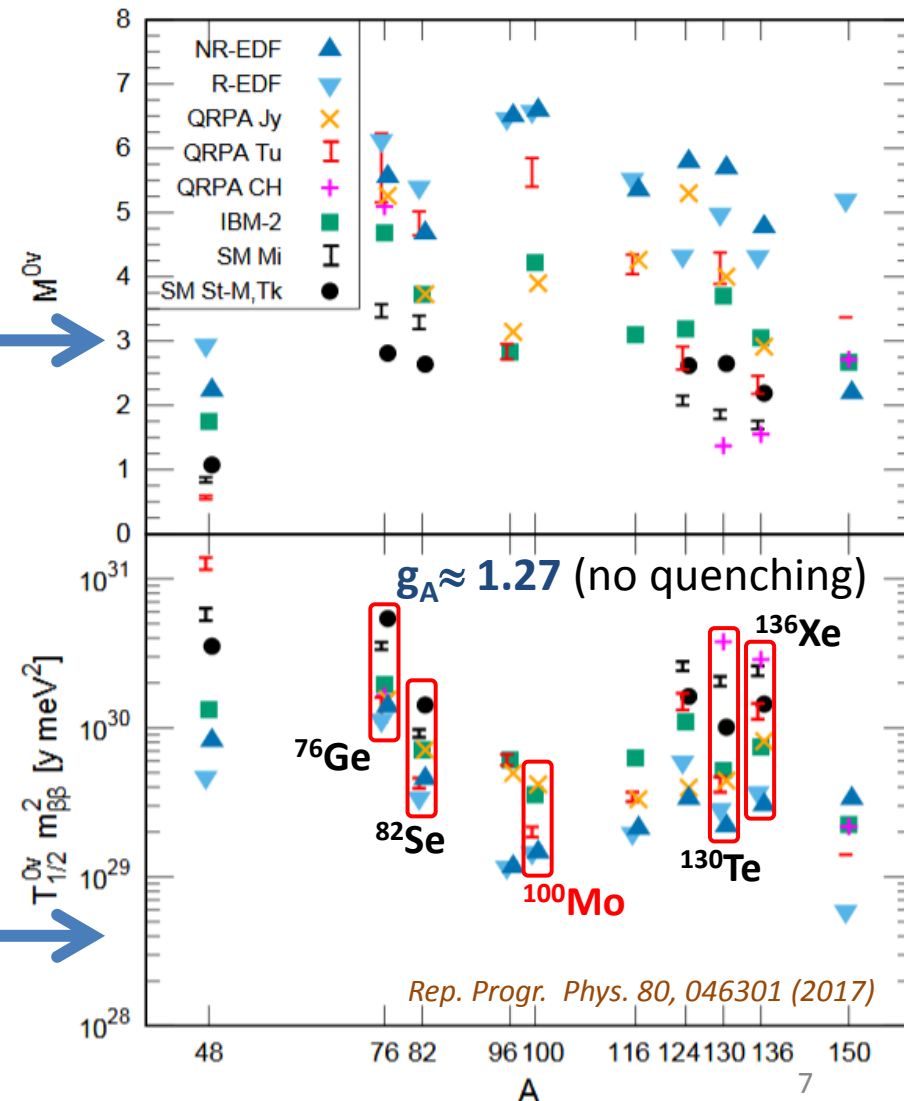
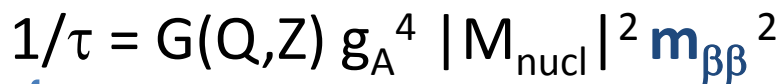
experimentally viable process
neutrinos:
- seesaw mechanism
- Majorana
- Theories
- saw mechanism)
- the Universe (leptogenesis)

Standard mechanism: $m_{\beta\beta}$ vs. lightest ν mass



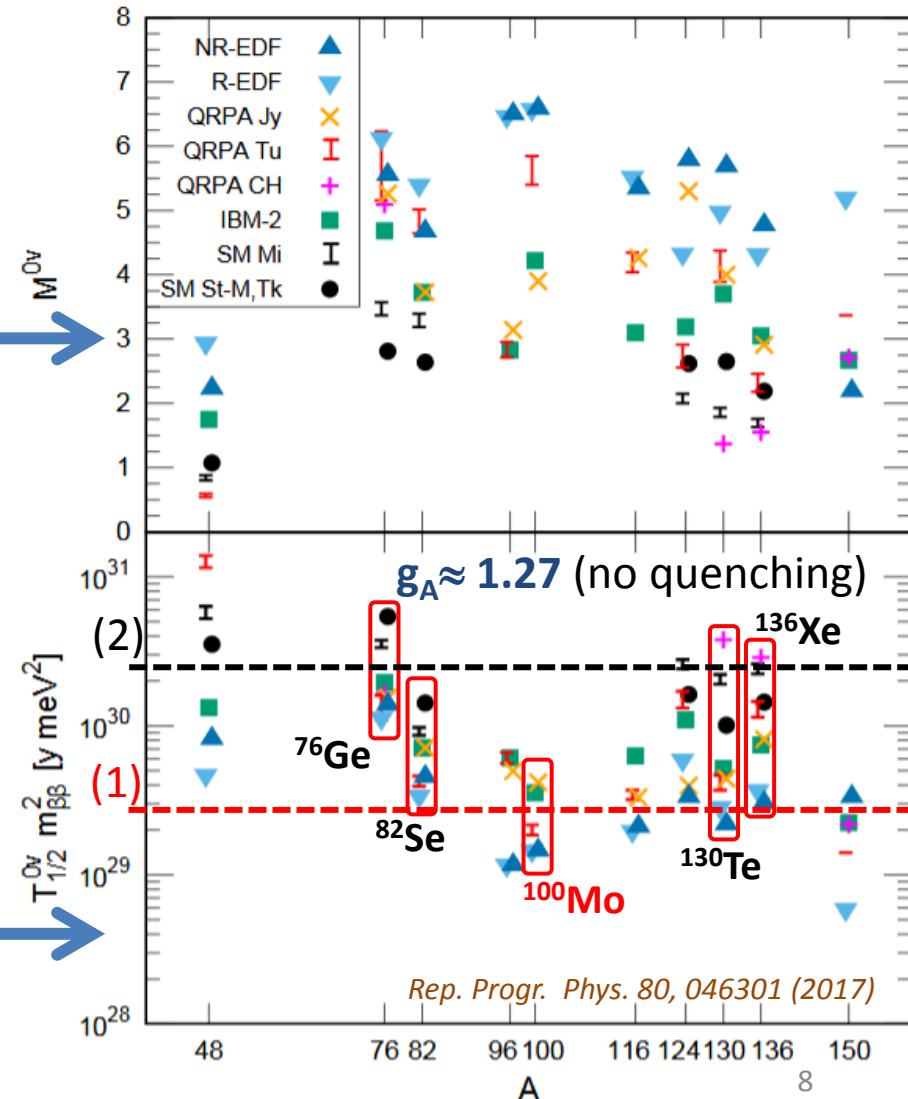
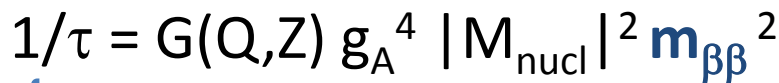
F. Simkovic R. Dvornicky
S. Stoica J. Suhonen
J. Engel J. Menendez

Nuclear matrix elements: several models

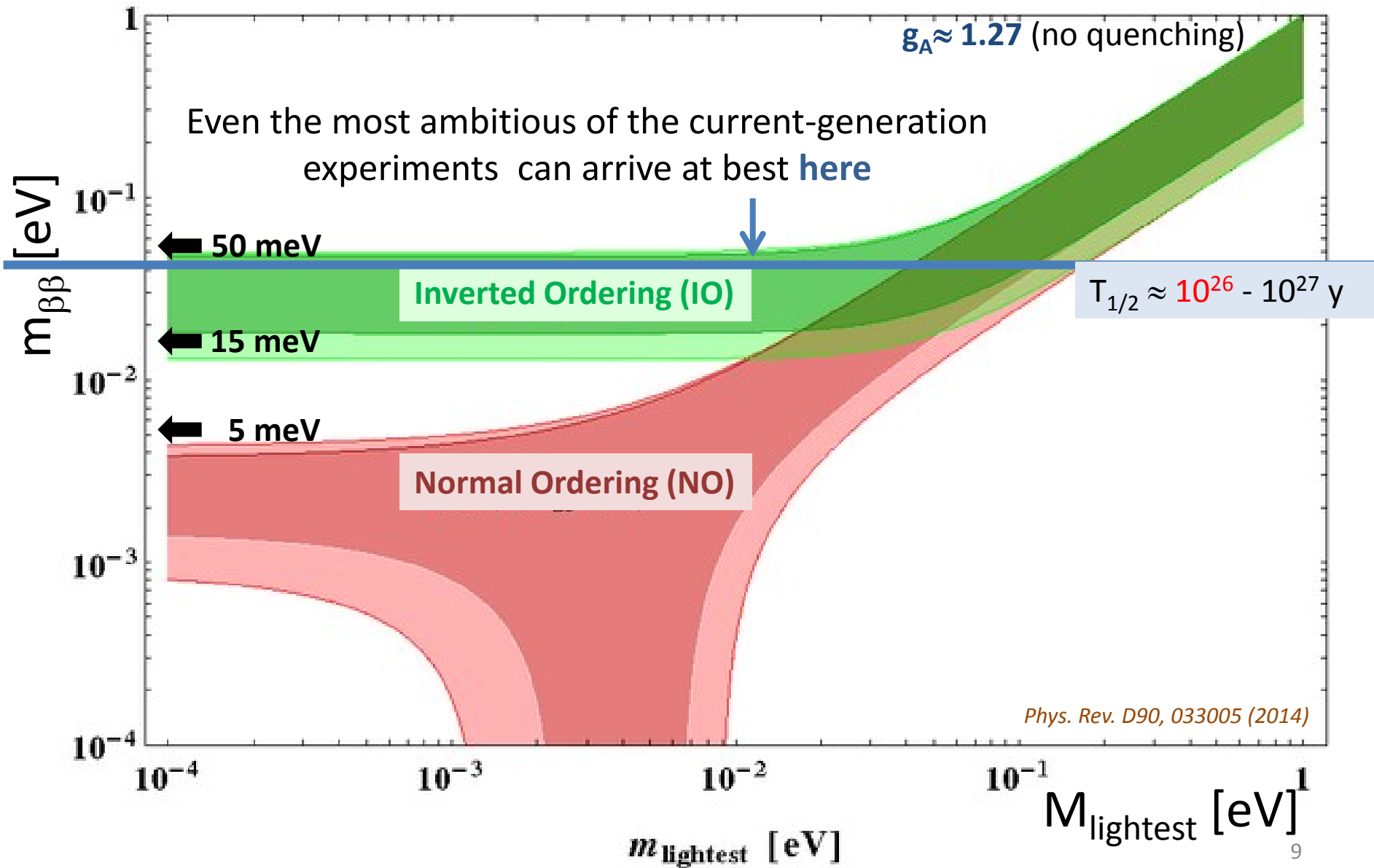


F. Simkovic R. Dvornicky
S. Stoica J. Suhonen
J. Engel J. Menendez

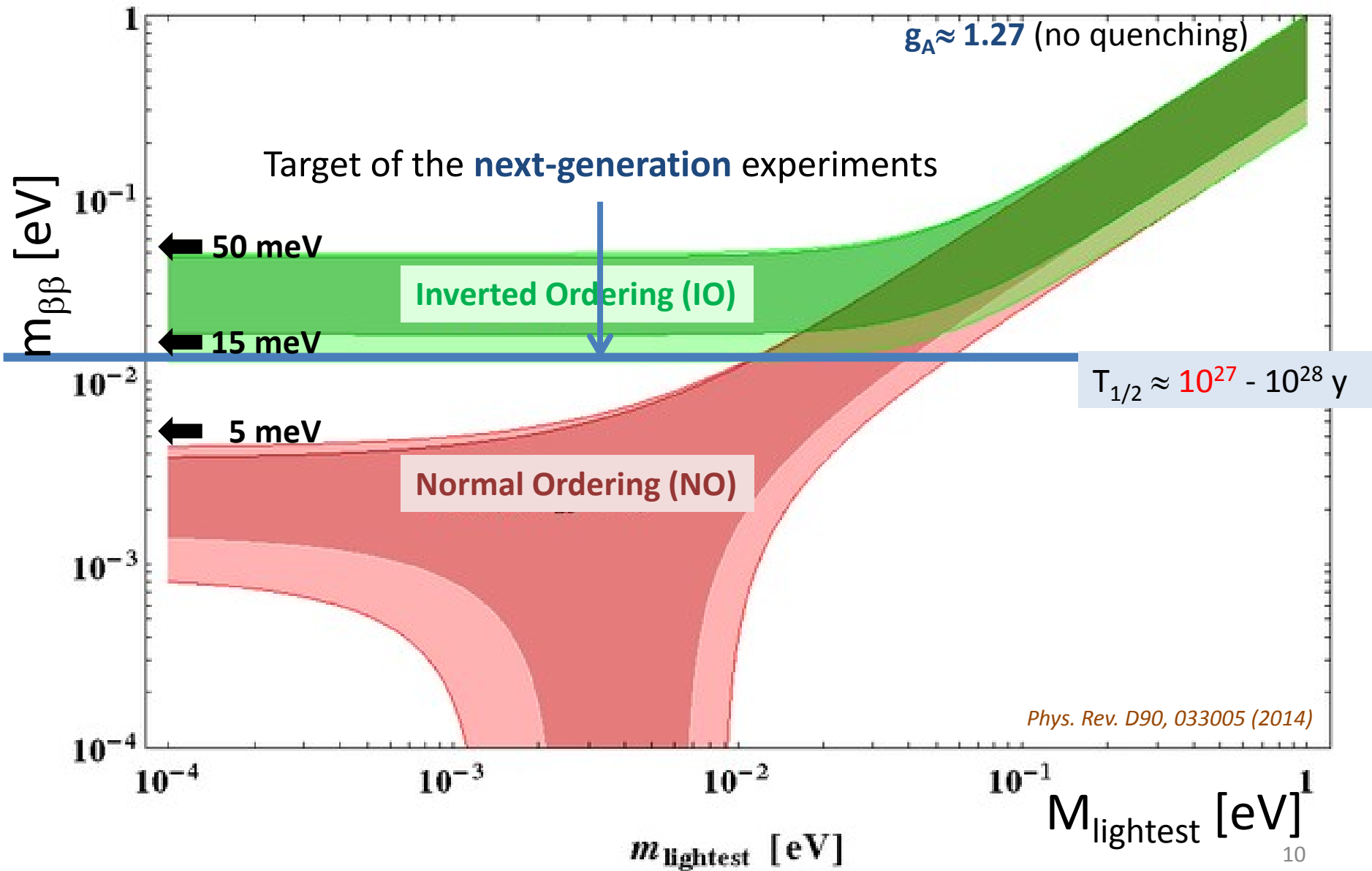
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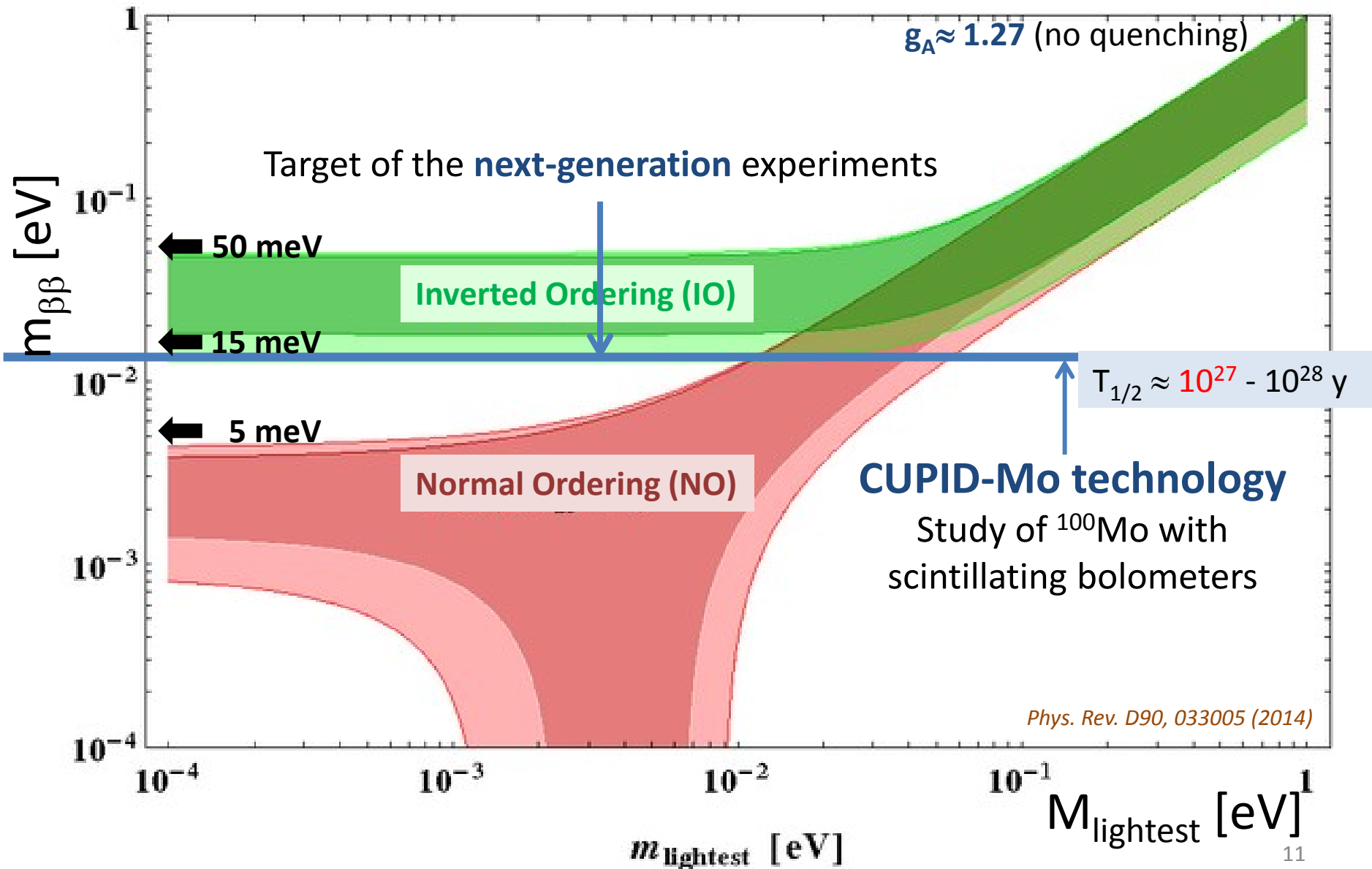
Current-generation experiments



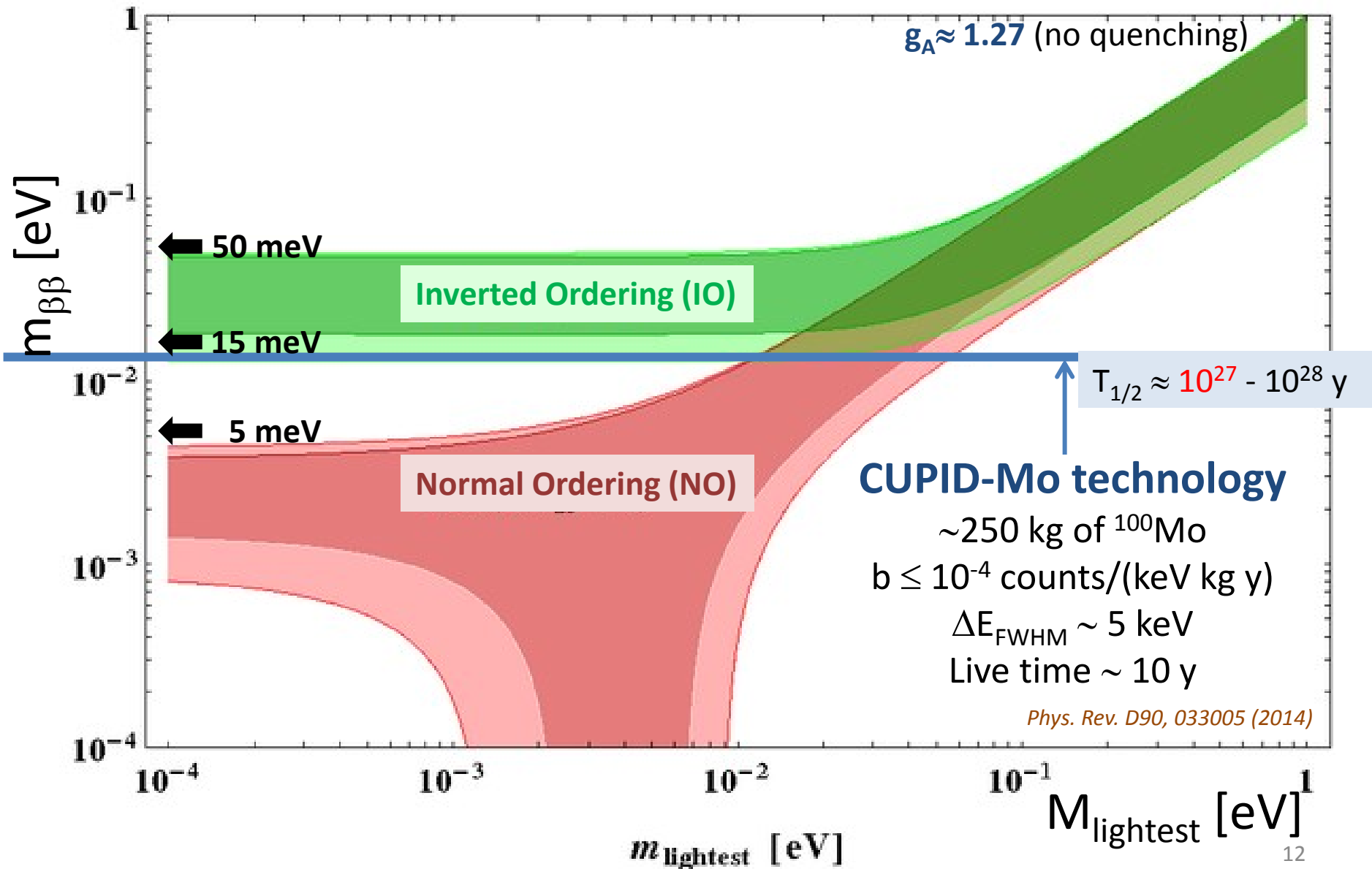
Next-generation experiments



Next-generation experiments



Next-generation experiments



Next-generation experiments

F. Simkovic

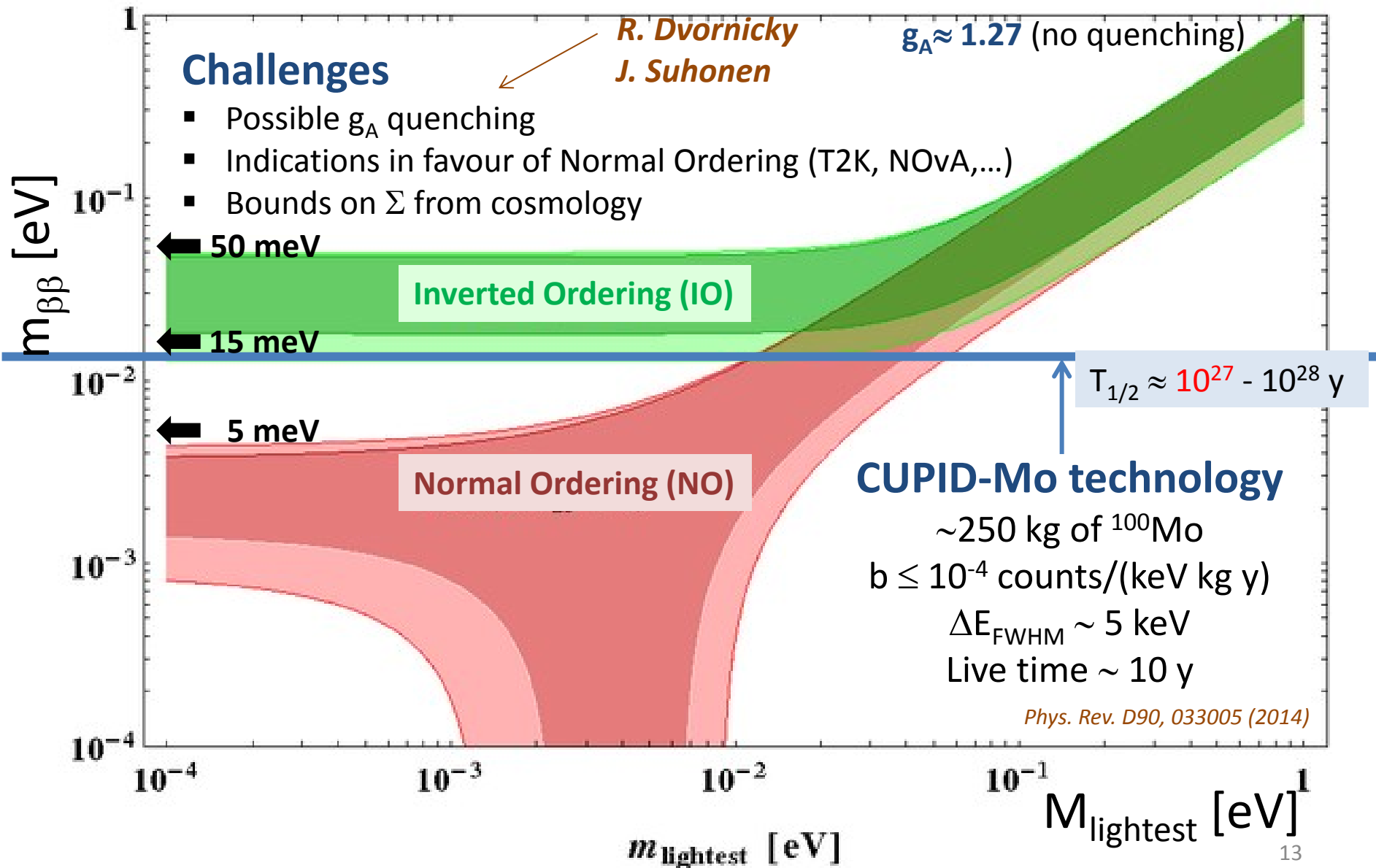
R. Dvornicky

J. Suhonen

$g_A \approx 1.27$ (no quenching)

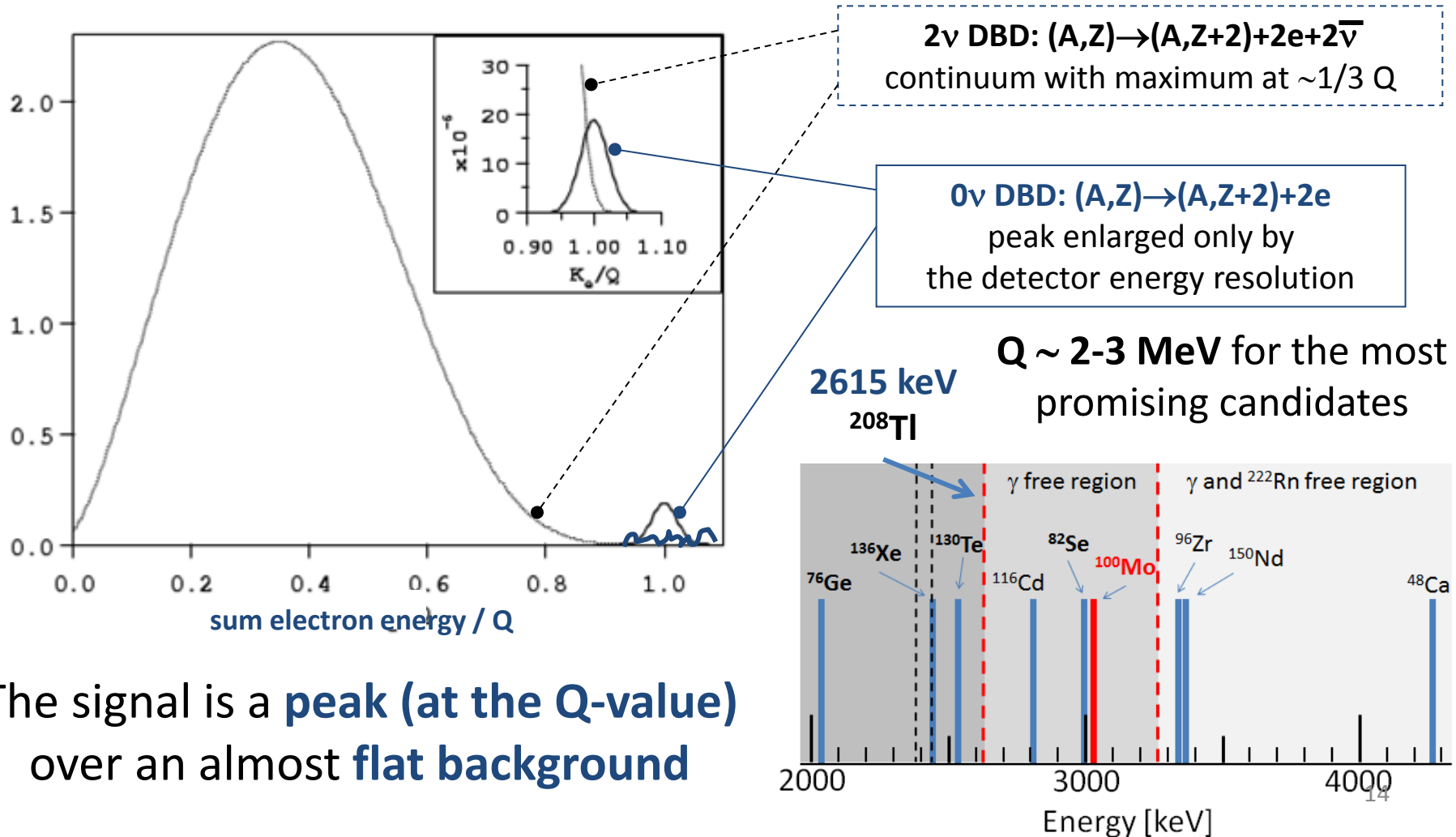
Challenges

- Possible g_A quenching
- Indications in favour of Normal Ordering (T2K, NOvA,...)
- Bounds on Σ from cosmology



What we are looking for

The shape of the two-electron sum-energy spectrum enables to distinguish between the 0ν (new physics) and the 2ν decay modes



How we do it: bolometers

Bolometric approach: the source is embedded in a crystal, which is cooled down to 10-20 mK and works as a **perfect calorimeter**

$$\Delta T = E/C$$

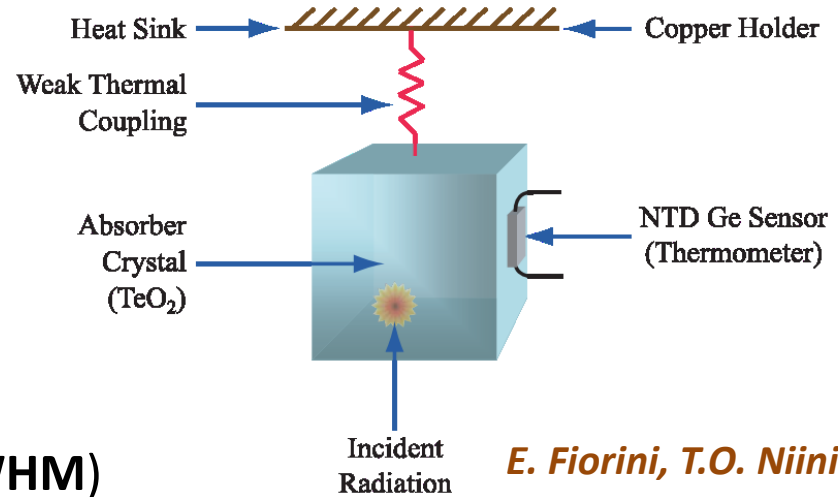
- High energy resolution (~ 5 keV FWHM)
- $\sim 0.1 - 0.5$ kg source in each crystal $\rightarrow$ arrays
- High efficiency ($\sim 70 - 90$ %)

Ke Han - CUORE

- Cuoricino – CUORE experiments $\rightarrow$ crystals of TeO_2 (isotope ^{130}Te)

- Large flexibility in the detector material choice: ^{130}Te , and three isotopes with $Q > 2615$ keV (^{82}Se , ^{100}Mo , ^{116}Cd) can be studied

Single multi-isotope experiment $\rightarrow$ *AG et al., Eur. Phys. J. C 78 (2018) 272*



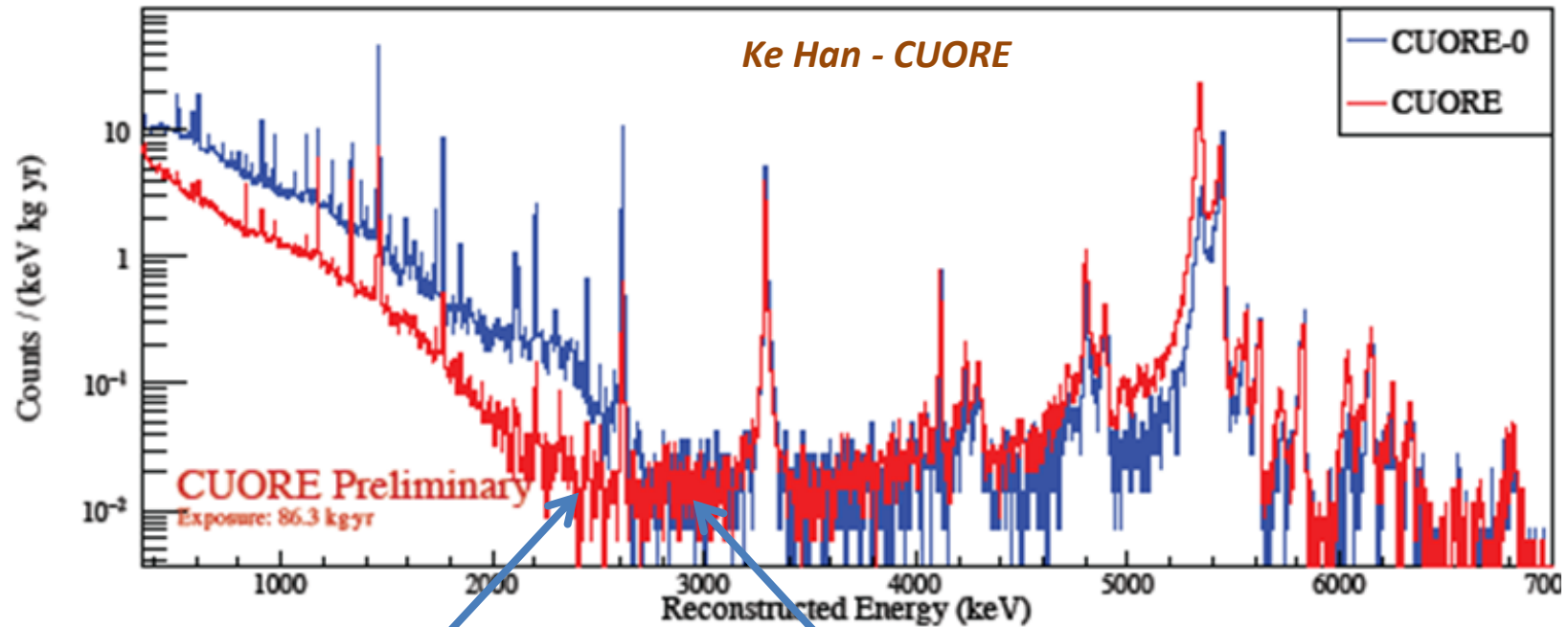
*E. Fiorini, T.O. Niinikoski
Low-temperature
calorimetry for rare decays
NIM 224 (1984) 83*

S. Pirro – CUPID-0

HongJoo Kim - AMoRE

CUORE-technology lessons about background

CUORE and its precursors are affected by alpha particle background



Irreducible background due to **alpha particles**, emitted at the **surfaces** and energy-degraded

$$b \sim 10^{-2} [\text{counts}/(\text{keV kg y})]$$

CUORE is not background free

Improvement by a factor 10^2 is required *M. Vignati – CALDER*

Current solution: scintillating / Cherenkov bolometers

Phys. Rev. C 97 (2018) 032501(R)

Alpha / beta separation

Alphas emit a different amount of light with respect to beta/gamma of the same energy (normally lower $\rightarrow \alpha \text{ QF} < 1$, but not in all cases – ZnSe is an exception).

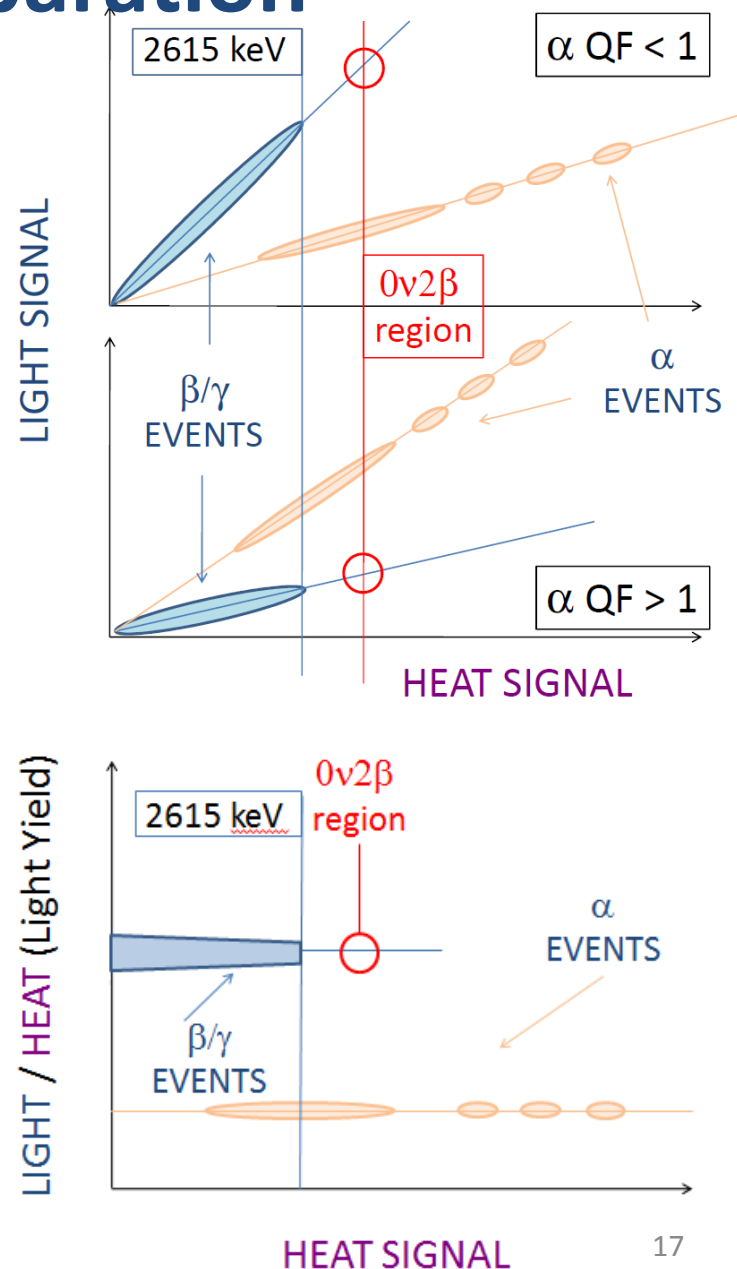
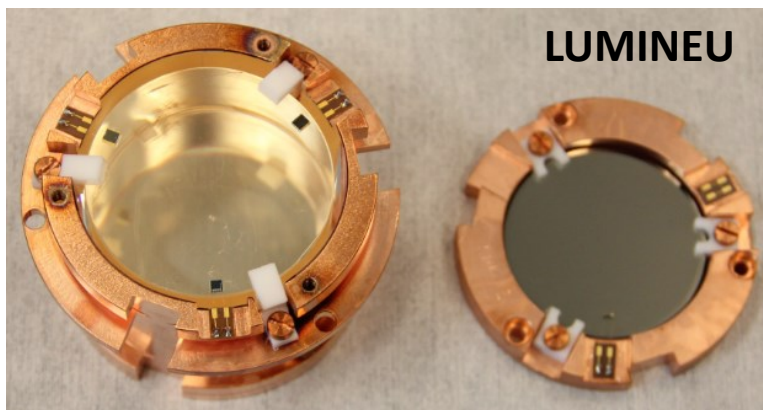
HongJoo Kim - AMoRE

A scatter plot light vs. heat or a plot light-yield vs. heat) separates alphas from betas / gammas.

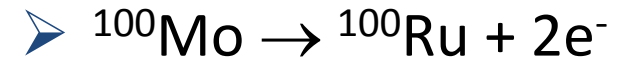
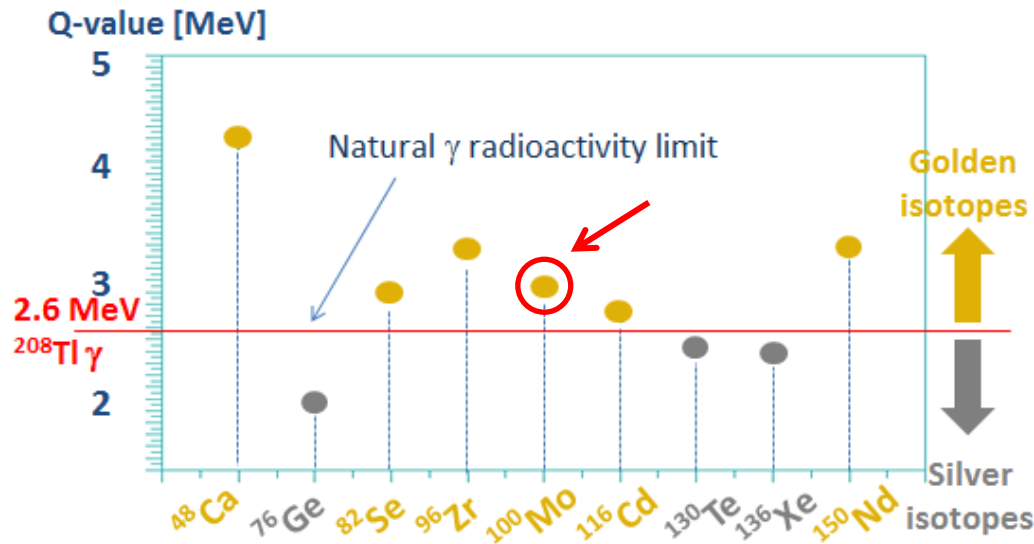
S. Pirro – CUPID-0

PSD can work as well

A bolometric light detector is needed, facing the main crystal



Some properties of ^{100}Mo



➤ $Q_{\beta\beta} = 3034 \text{ keV}$

➤ $I.A.(100) = 9.7 \%$

➤ enrichable by gas centrifugation

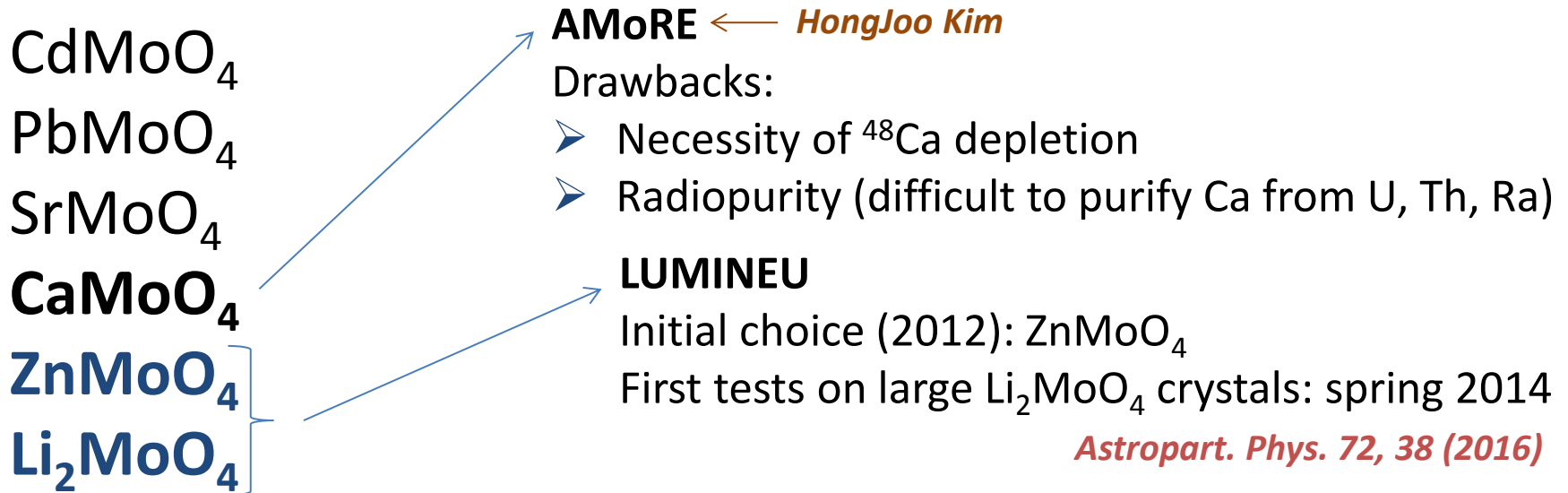
Caveats

➤ $T_{1/2}(2\nu) = 6.9 \times 10^{18} \text{ y}$ – the fastest one in all $0\nu 2\beta$ candidates

➤ ^{214}Bi line at 3054 keV – B.R. 0.021 % - Compton edge 2818 keV

Useful Mo-based crystals

Crystals successfully tested so far as scintillating bolometers:



LUMINEU: Selection of Li_2MoO_4 for a pilot experiment (March 2016)

- Better bolometric performance
- Easy crystallization / excellent quality
- Outstanding radiopurity

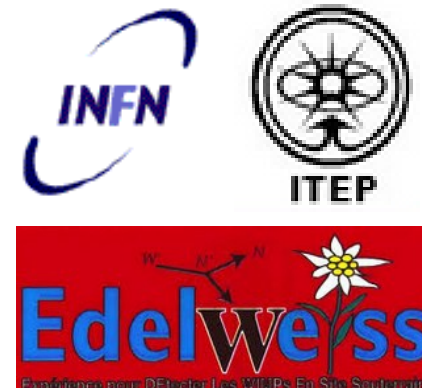
Caveats

- Hygroscopic material
- ^{40}K is natural contaminant
- Lower light yield (~ 0.8 keV/MeV)

Preparing a ^{100}Mo experiment: LUMINEU

Funding / resources from

- **ANR (France) – main fund provider (LUMINEU: 2012-2017)**
- CEA-Saclay – substantial funds / PhD
- CSNSM direction – funds for crystals (« AP interne »)
- EDELWEISS – underground facility, electronics & DAQ
- IN2P3 – dedicated personnel
- KINR Kiev – radiopure scintillator know-how, simulation, – enriched ^{100}Mo
- ITEP Moscow – enriched ^{100}Mo
- NIIC Novosibirsk - crystals
- INFN / LUCIFER – underground facility and manpower for R&D



Extension of the Mo collaboration: CUPID-Mo

New participants

LAL – Orsay France

MIT
UCB/LNBL } USA

Fudan Shanghai
USTC Hefei } China

<http://cupid-mo.mit.edu>



7 countries, 15 institutions, ~110 scientists

Strong interest in China

Large CUPID group.
Project for a “parallel CUPID”
at JinPing laboratory



Li_2MoO_4 : purification and crystallization

From 2013 to 2016, a series of important milestones were achieved:

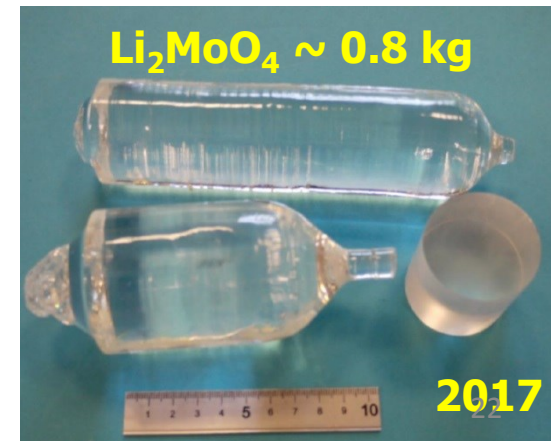
- **Mo purification / crystallization protocol** (NIIC, Novosibirsk, Russia) (**Mo irrecoverable losses < 4%**)
- Selection of the **appropriate Li_2CO_3 powder** for compound formation
- Successful program to **control internal content of ^{40}K** (from ~ 60 mBq/kg to < 5 mBq/kg)
 - Random coincidences: $2\nu 2\beta + ^{40}\text{K} \ll 2\nu 2\beta + 2\nu 2\beta$
- Efficient use of existing **~ 10 kg of ^{100}Mo** (~ 9 kg to ITEP-Moscow and ~ 1 kg to KINR-Kiev) (MoU among IN2P3 / INFN / ITEP – February 2015)
 - Natural isotopic abundance: **9.7%**

NIM A 729, 856 (2013)

JINST 9, P06004 (2014)

EPJC 74, 3133 (2014)

JINST 10, P05007 (2015)



EDELWEISS-III cryogenic facility at LSM (France)

Laboratoire Souterrain de Modane

1.7 km rock overburden (~ 4.8 km w.e.)

$5 \mu\text{/day/m}^2$; 10^{-6} n/day/cm 2 (> 1 MeV)

Deradonized air flow (~ 30 mBq/m 3)

EDELWEISS set-up

Clean room (ISO Class 4)

$^3\text{He}/^4\text{He}$ inverted wet cryostat

Passive shield

Modern lead (18 cm)

Roman lead (2 cm; 14 cm at 1 K plate)

Polyethylene (external $\sim 50+5$ cm and 10 cm at 1 K plate)

Background monitors

Muon veto (98.5% covering)

Neutron counter

Radon counter

Electronics, DAQ (Samba)

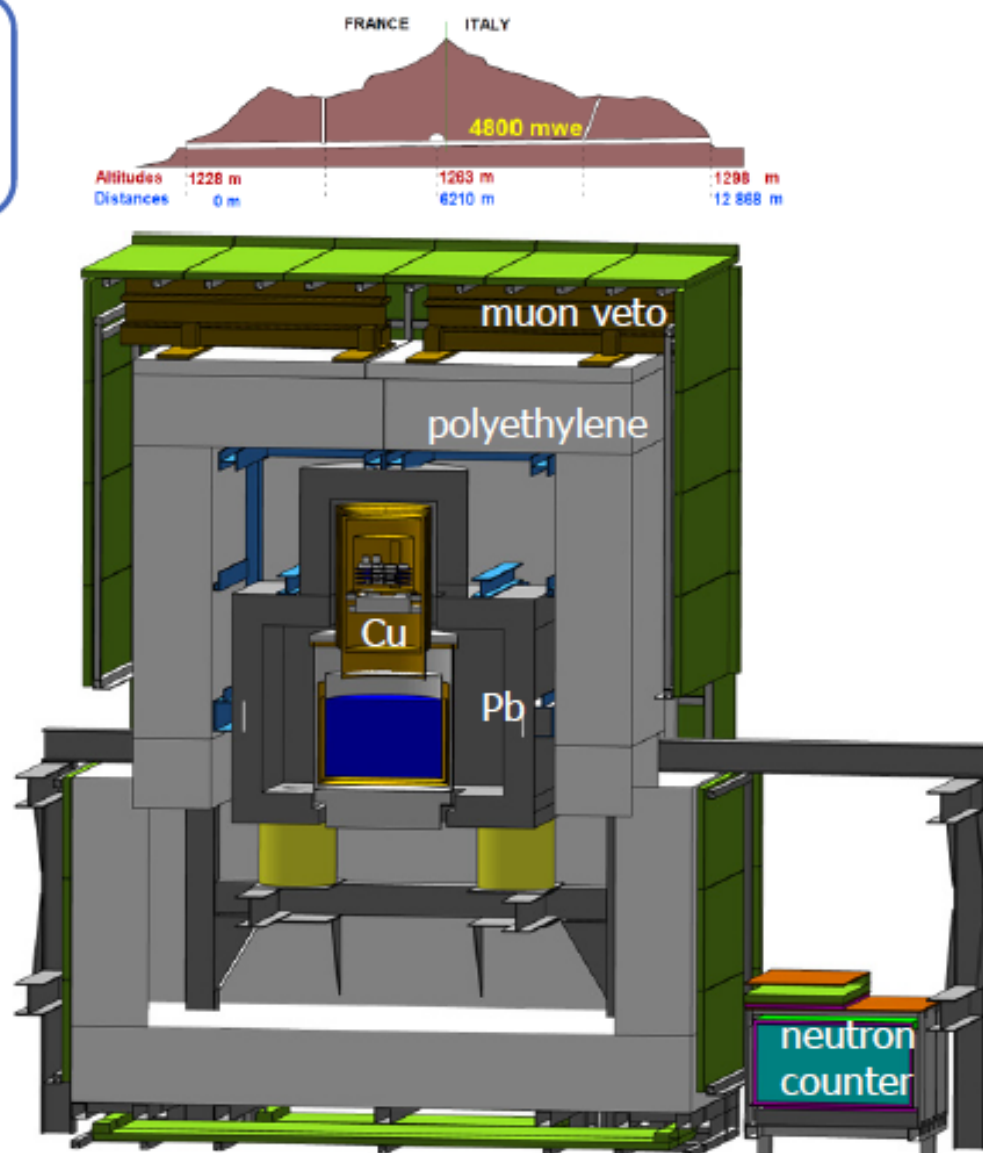
Low noise cold electronics

AC bias, modulation (100 kHz)

→ demodulation (up to 1 kHz)

16-bit or 14-bit ADC

Trigger and/or Stream data



PLB 702 (2011) 329; JINST 12 (2017) P08010; EPJC 77 (2017) 785

Tests of $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers

LUMINEU
03-07/2016
20-19.2 mK

ZMO nat-2

enrLMO-1

FID211

LUMINEU
12/2016-04/2017
17.0 mK

enrLMO-2

enrLMO-4

enrLMO-3

T1

LMO1b

204 g

Ge LD

LMO2b

207 g

Ge LD

LMO2t

213 g

Ge LD

186 g

CUPID-Mo
10-12/2017
21-19.3 mK

enrLMO-4

enrLMO-3

enrLMO-2

enrLMO-15

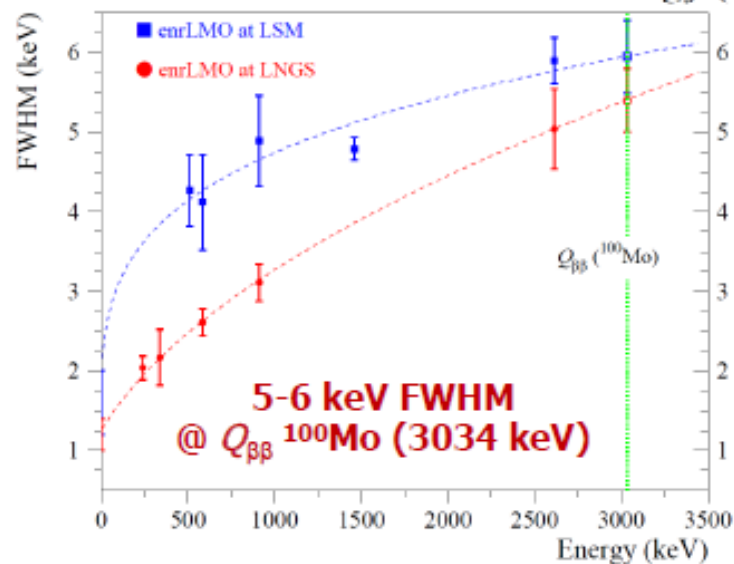
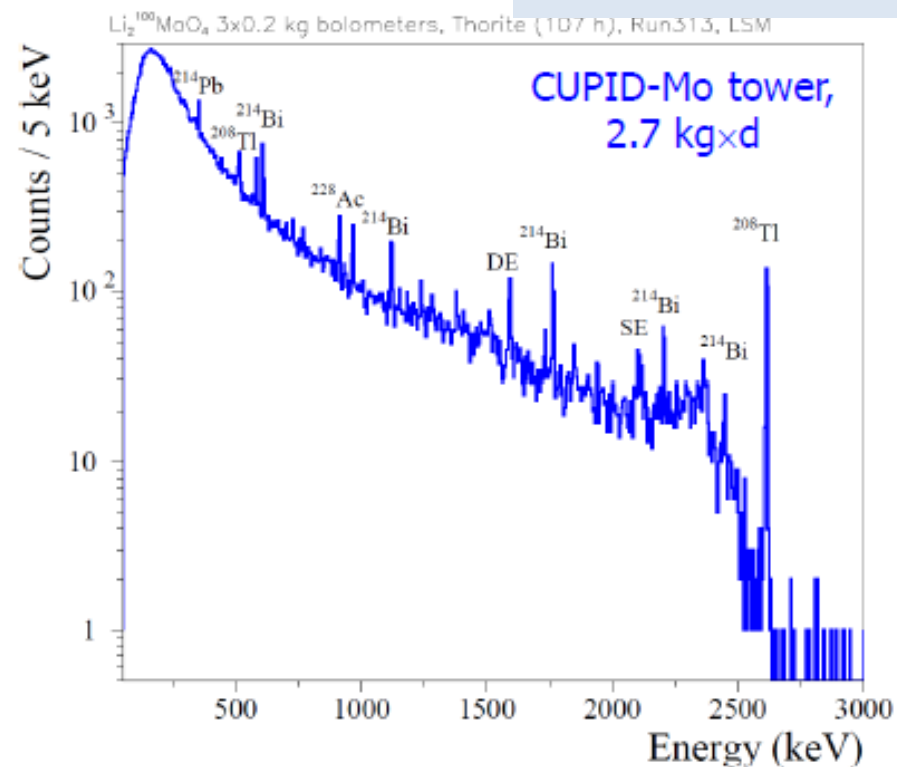
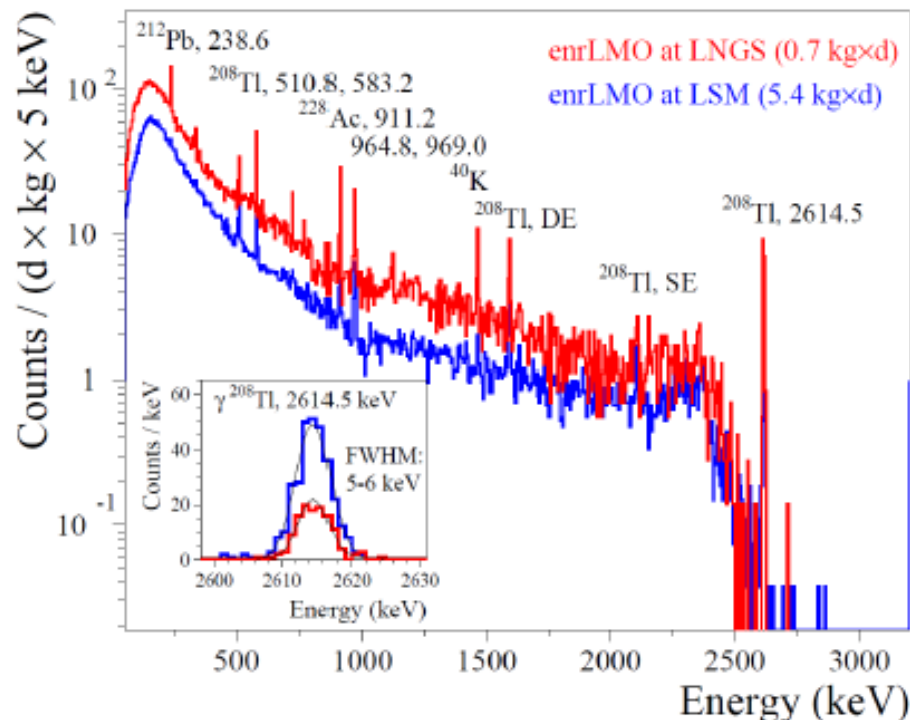
+ enrLMO-2
was tested at LNGS
05/2016, 12 mK

+ a tower T2
with enrLMO-1

D. Poda

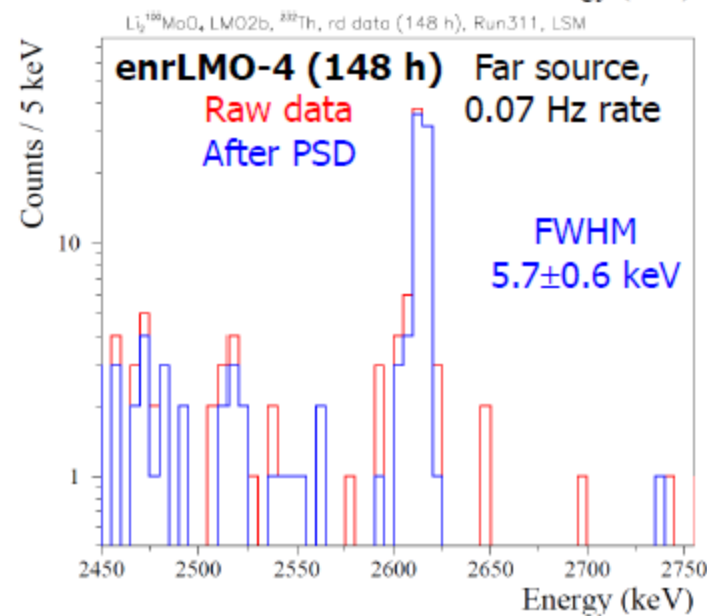
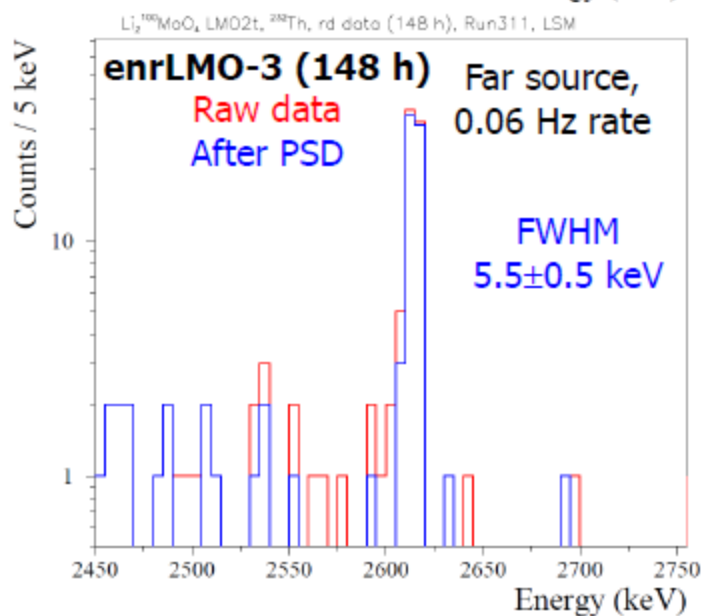
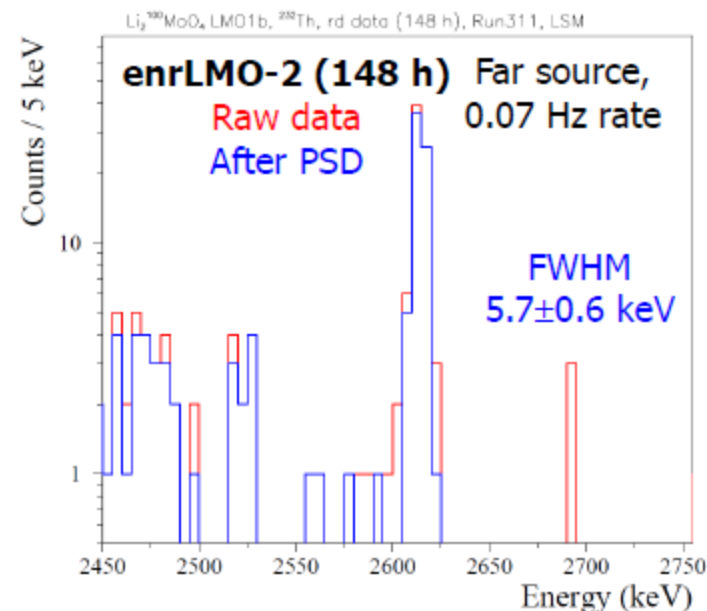
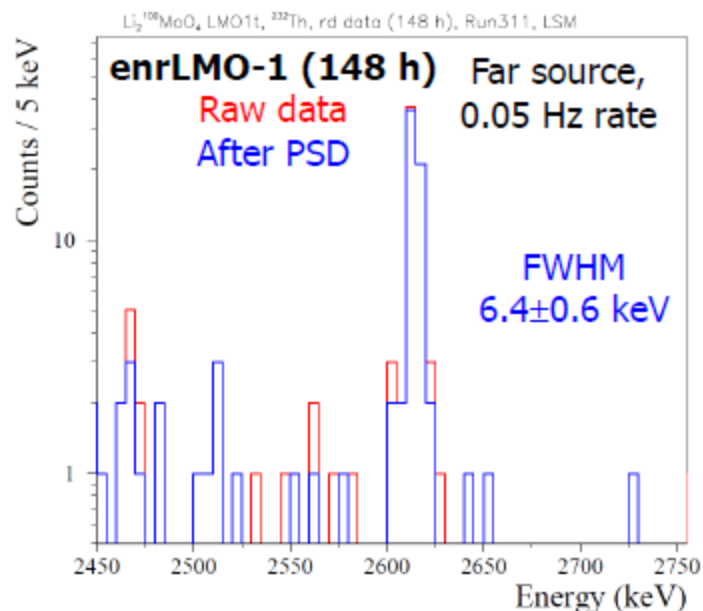
^{232}Th calibration

Heat channel



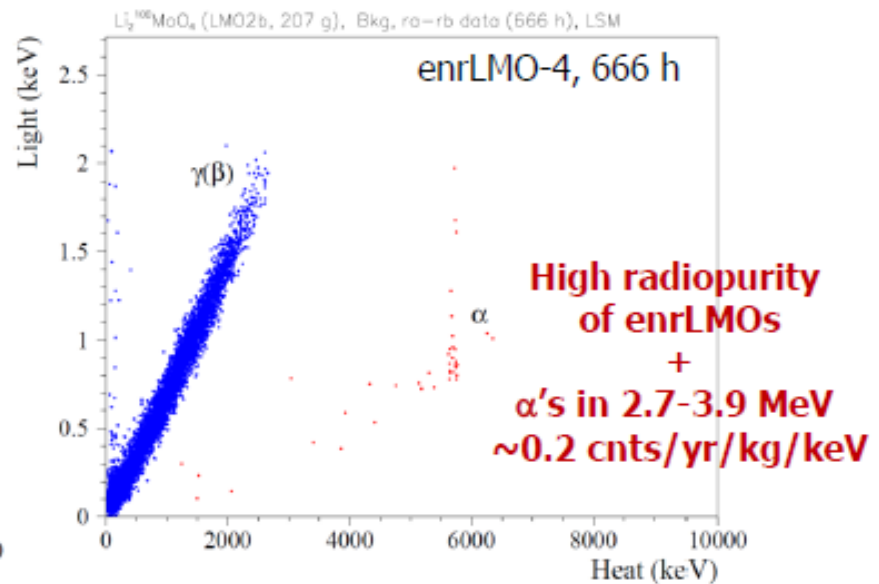
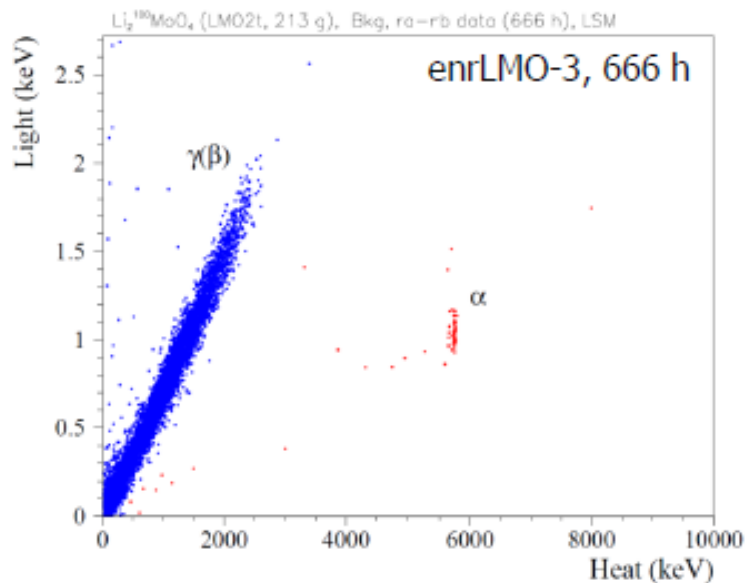
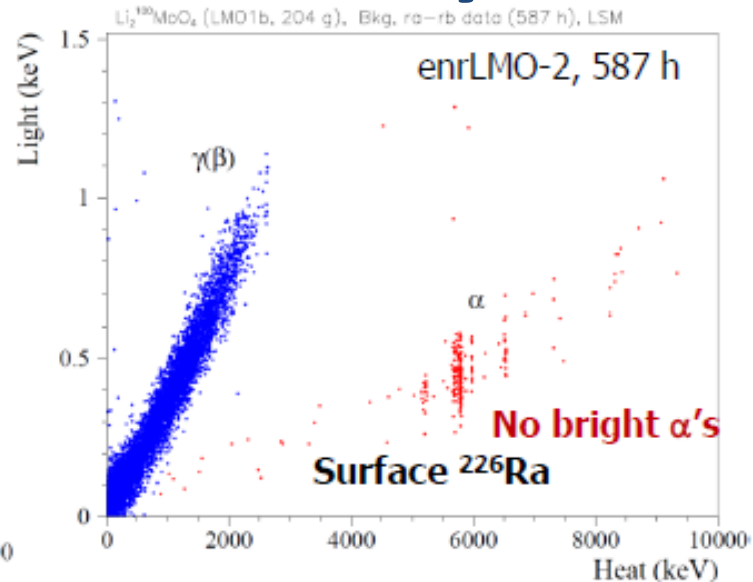
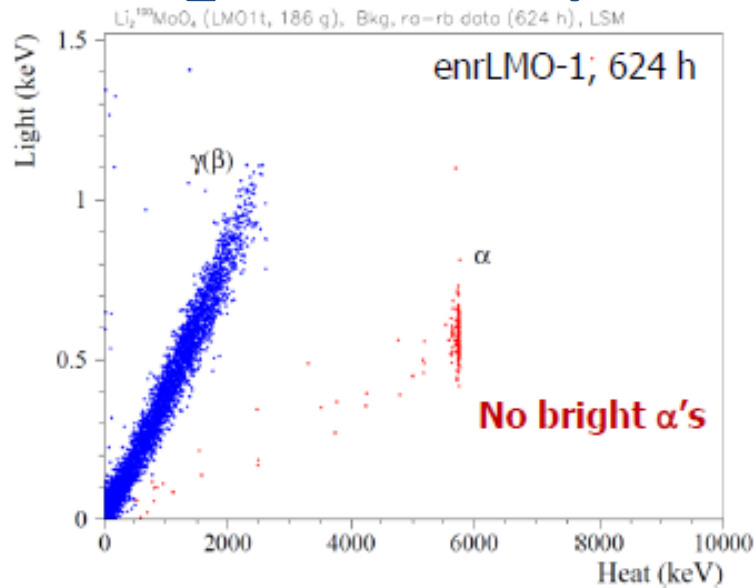
Mixed Th/U source is useful to estimate uncertainty of ^{100}Mo RoI determination

Calibration points	2615 keV peak @
Up to 2615 (^{208}Tl)	2614.8
Up to 2204 (^{214}Bi)	2616.1



4 $\text{Li}_2^{100}\text{MoO}_4$ detector array

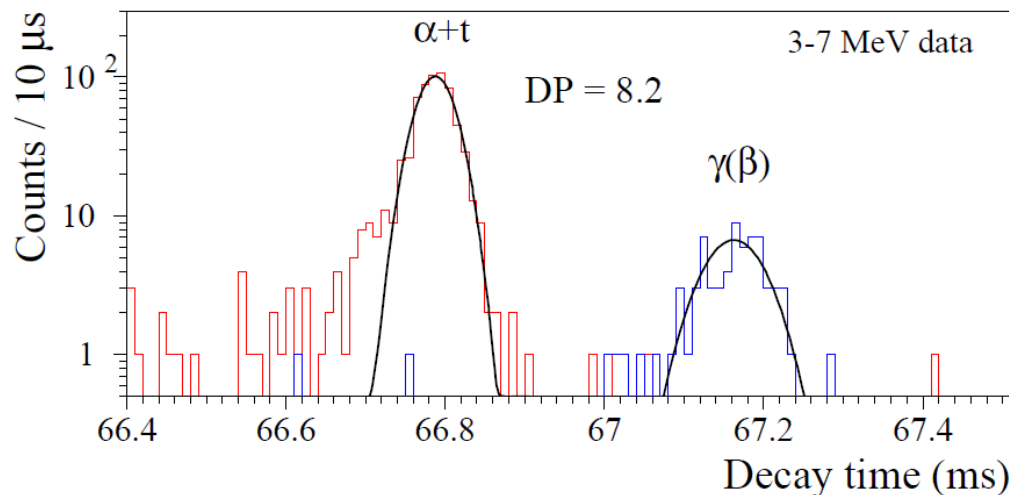
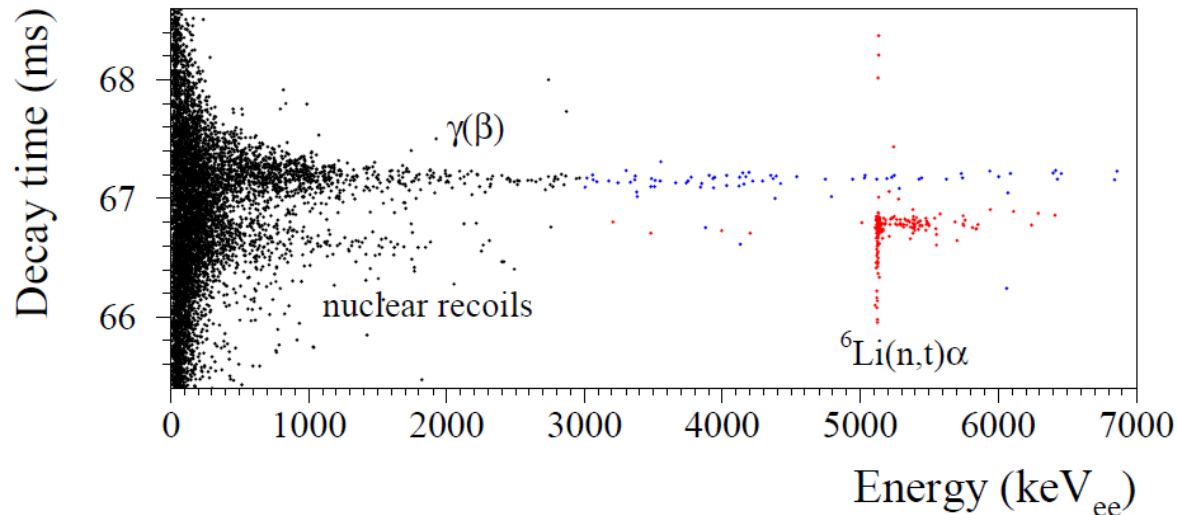
Heat + Light



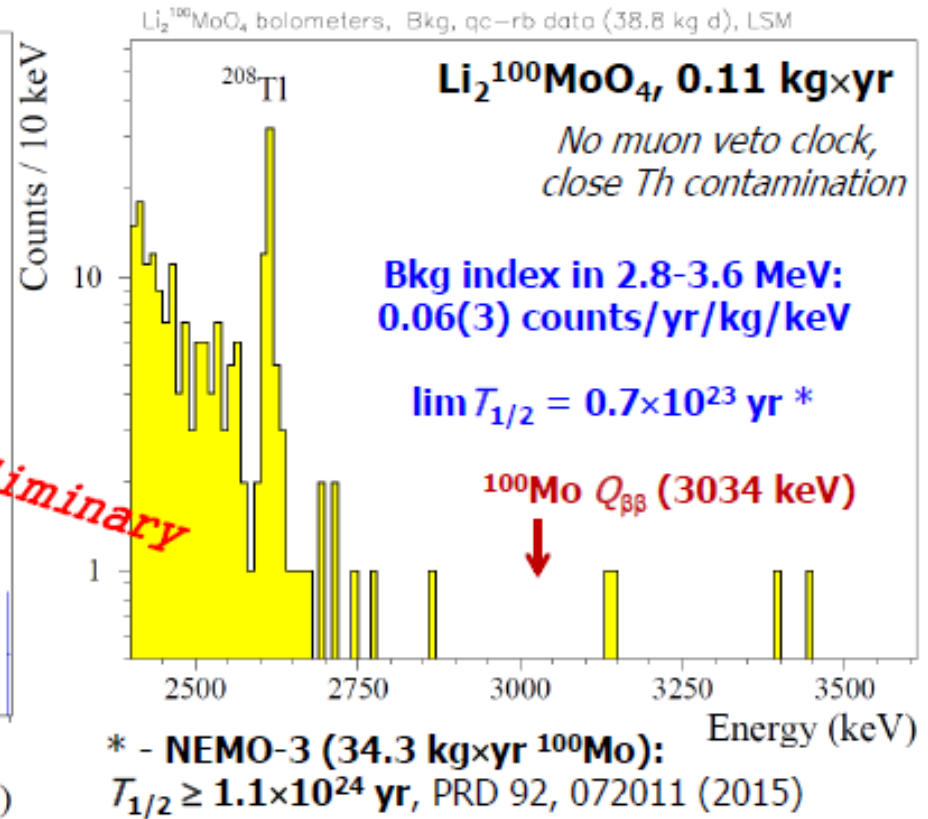
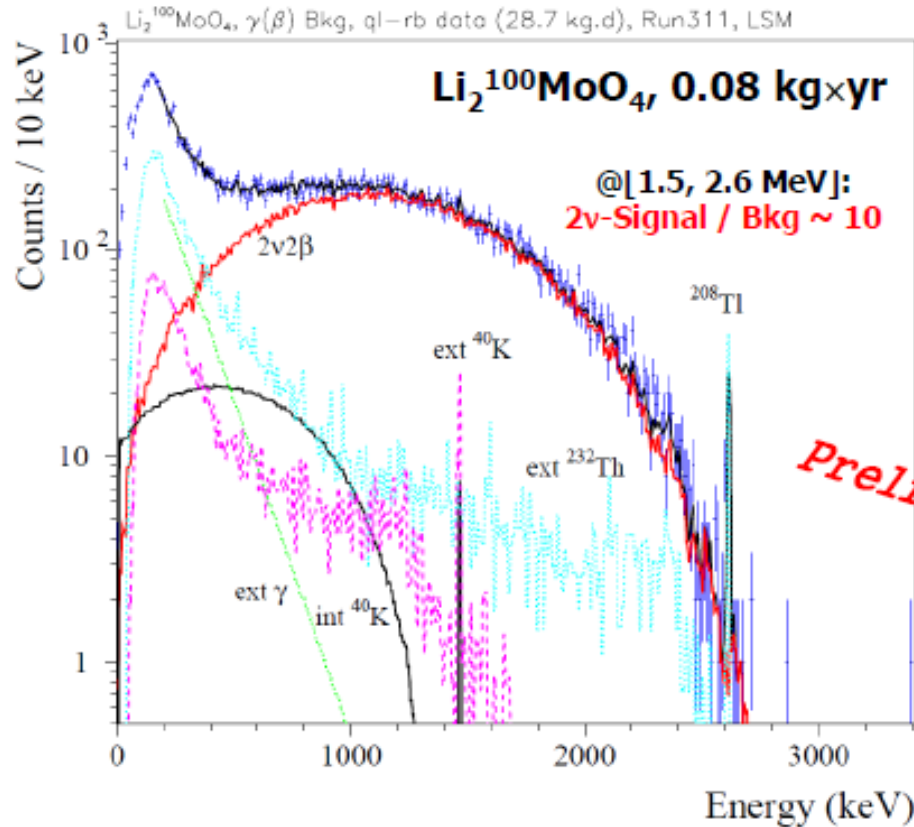
α/β separation without light detectors

We performed a specific study on a natural 150 g detector operated in LNGS.

A calibration with an AmBe source provided enough statistics in the β and α -like bands



LUMINEU: search for 2β decay of ^{100}Mo



$T_{1/2}$ [10^{18} yr]	^{100}Mo exposure	Experiment	Ref.
$7.11 \pm 0.02(\text{stat}) \pm 0.54(\text{syst})$	7.37 kg $\times$ yr	NEMO-3	PRL 95, 182302 (2005)
$7.15 \pm 0.37(\text{stat}) \pm 0.66(\text{syst})$	0.01 kg $\times$ yr	LUCIFER*	JPG 41, 075204 (2014)
$6.90 \pm 0.15(\text{stat}) \pm 0.37(\text{syst})$	0.02 kg$\times$yr	LUMINEU	EPJC 77, 785 (2017)

Data with higher statistics are available
CUPID-Mo paper in preparation

$\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers: a mature technology

LUMINEU has successfully developed the $\text{Li}_2^{100}\text{MoO}_4$ technology

Multiple tests with natural and enriched crystals (2014-2017) in LSM and LNGS
with outstanding results in terms of:

Reproducibility	→	excellent performance uniformity
Energy resolution	→	~ 4-6 keV FWHM in RoI
α/β separation power	→	> 99.9 %
Internal radiopurity	→	< 5 $\mu\text{Bq/kg}$ in ^{232}Th , ^{238}U ; < 5 mBq/kg in ^{40}K

→ Compatible with $b \leq 10^{-4}$ [counts/(keV kg y)]

LUMINEU
BB ON

THE EUROPEAN
PHYSICAL JOURNAL C



Eur. Phys. J. C (2017) 77:785
<https://doi.org/10.1140/epjc/s10052-017-5343-2>

Regular Article - Experimental Physics

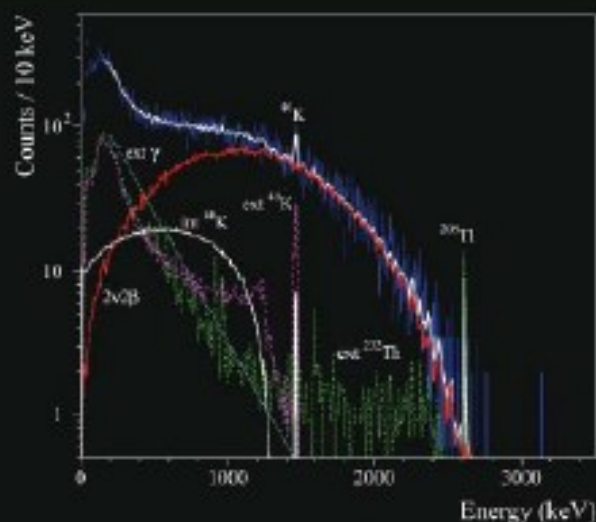
Development of ^{100}Mo -containing scintillating bolometers for a
high-sensitivity neutrinoless double-beta decay search

EPJ C



Recognized by European Physical Society

Particles and Fields



The $\gamma(\beta)$ background spectrum of a 186 g Li_2MoO_4 detector together with a simplified background model fit, built from the $2\nu 2\beta$ distribution of ^{100}Mo ($T_{1/2} = 6.9 \times 10^5$ years), internal ^{40}K , and external γ -quanta represented by exponential background of external ^{40}K and ^{232}Th . The $2\nu 2\beta$ signal-to-background ratio above 1.5 MeV is 8:1.

From E. Amengaud, C. Augier, A.S. Roshchek et al.: Development of ^{100}Mo -containing scintillating bolometers for a high-sensitivity neutrinoless double-beta decay search.



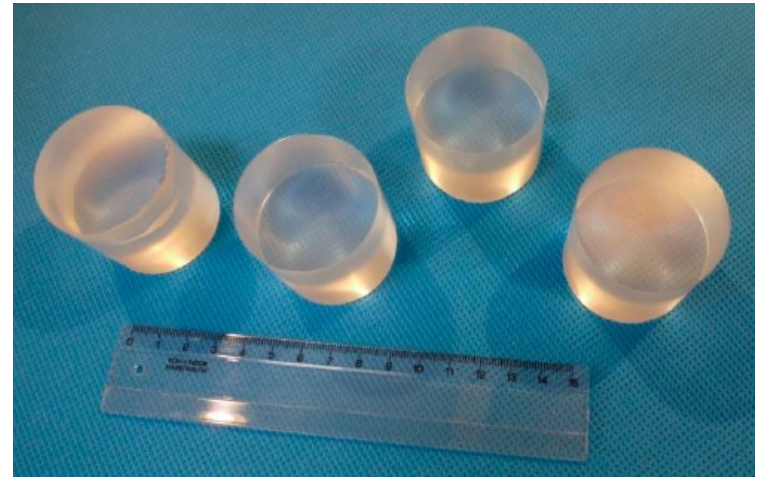
Springer

The CUPID-Mo demonstrators

CUPID-Mo Phase I (20 crystals):

- 20 ^{100}Mo -enriched (97%) Li_2MoO_4
($\varnothing 44 \times 45$ mm, 0.21 kg each; 4.18 kg total)
⇒ ~2.5 kg of ^{100}Mo
- 20 Ge light detectors ($\varnothing 44 \times 0.175$ mm)+SiO
- EDELWEISS set-up @ LSM (France)

START DATA TAKING: in the next weeks

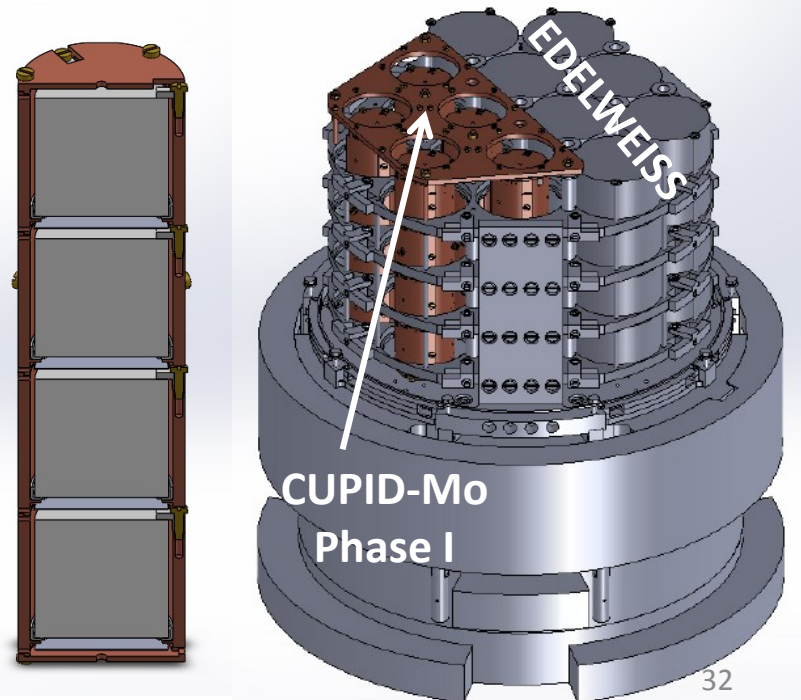
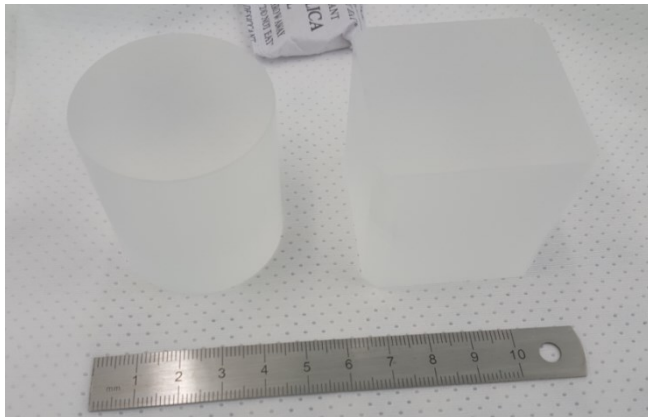


CUPID-Mo Phase II (20+26 crystals):

- Additional 26 cubic $\text{Li}_2^{100}\text{MoO}_4$
(45x45x45 mm, 0.28 g each)
⇒ ~5 kg of ^{100}Mo
- CUPID-0 set-up @ LNGS (Italy)

S. Pirro

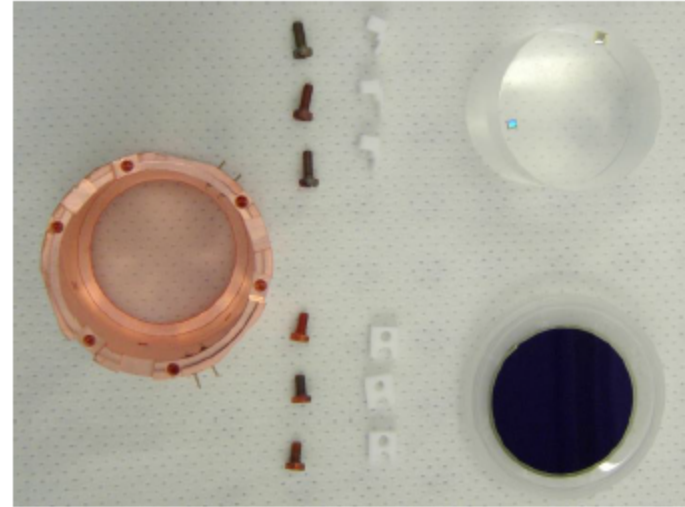
PLANNED START DATA TAKING: June 2019



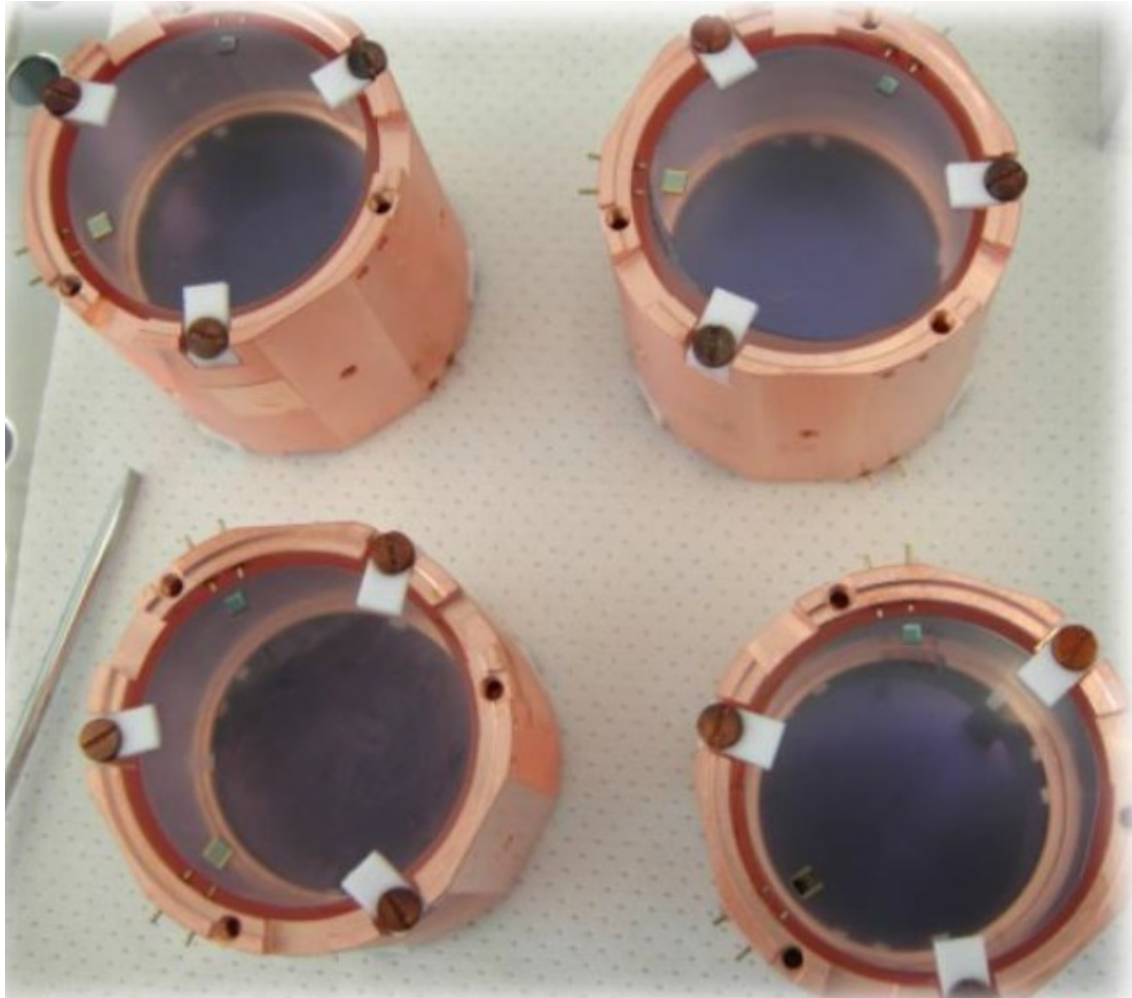


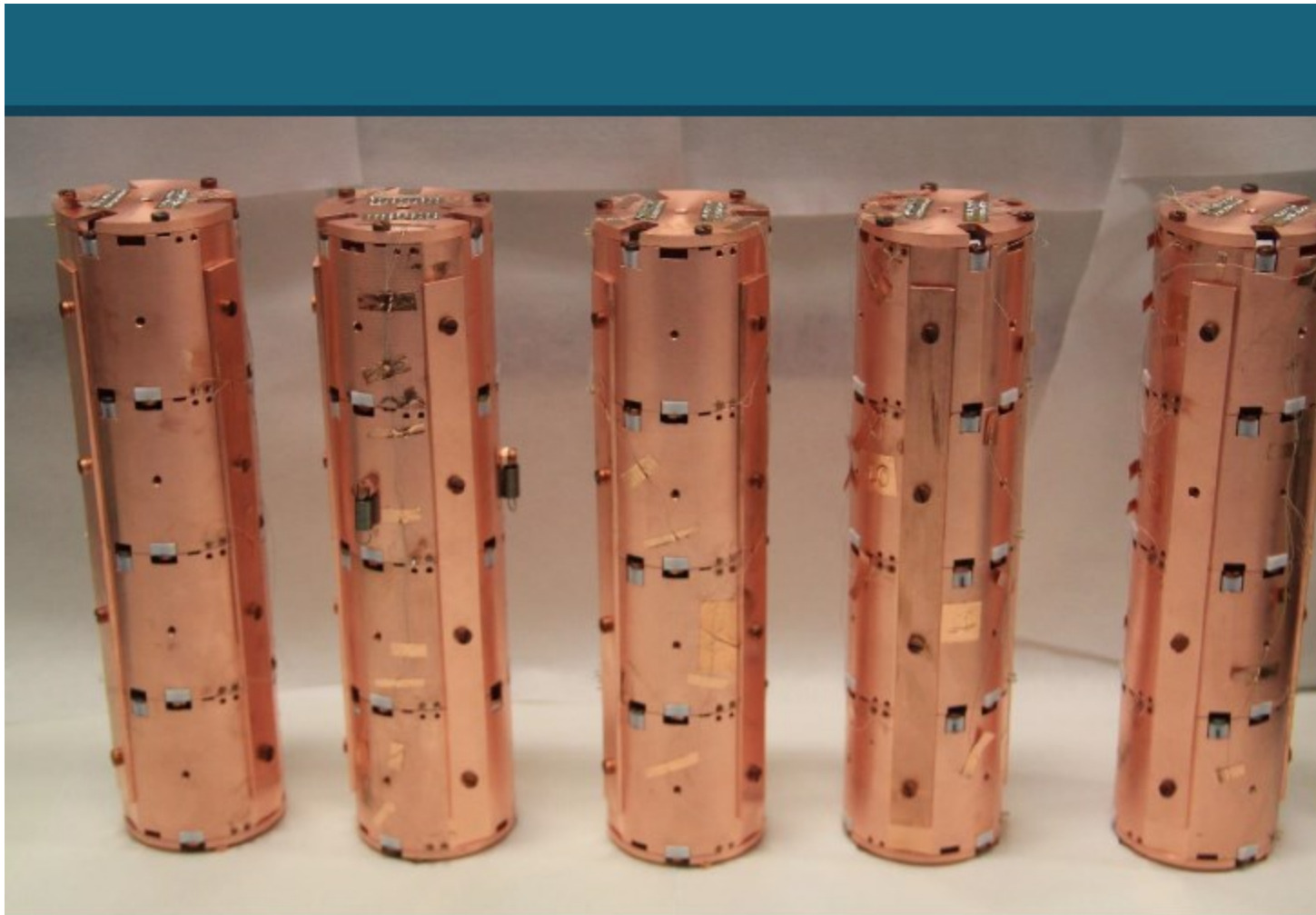
Structure

- Crystal: $\text{Li}_2^{100}\text{MoO}_4$
 $\varnothing 43.8 \times 45 \text{ mm}$
- Light detector: Ge wafer (UMICORE) $\varnothing 44.5 \text{ mm} \times 170 \text{ um}$ with SiO coating on both sides
- Copper holders: radiopure NOSV copper
- Spacers (thermal and mechanical link): PTFE
- Electrical connections: copper gold-plated tubes

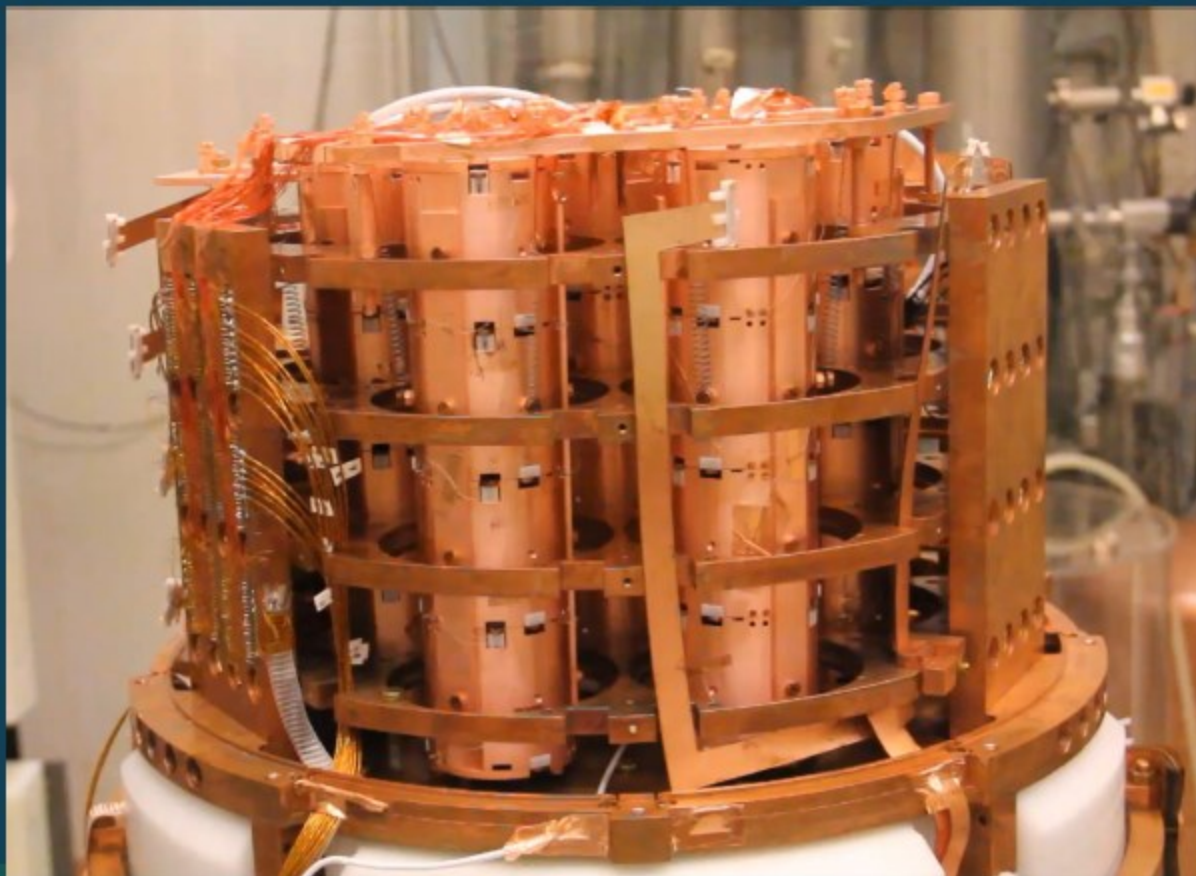


Four modules





Installation

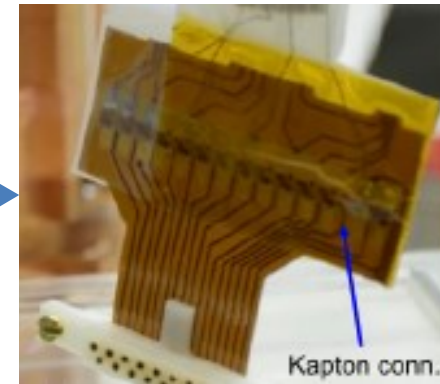


Suspension

Background consideration

Background index in the CUPID-Mo precursor $\rightarrow \mathbf{b = 0.06(3) \text{ c/(keV kg)}}$

This value is compatible with the presence of **^{232}Th -contaminated connectors** close to the detectors



Simultaneous absorption of **2615 keV γ** and coincident **583 keV γ** (^{208}Tl)

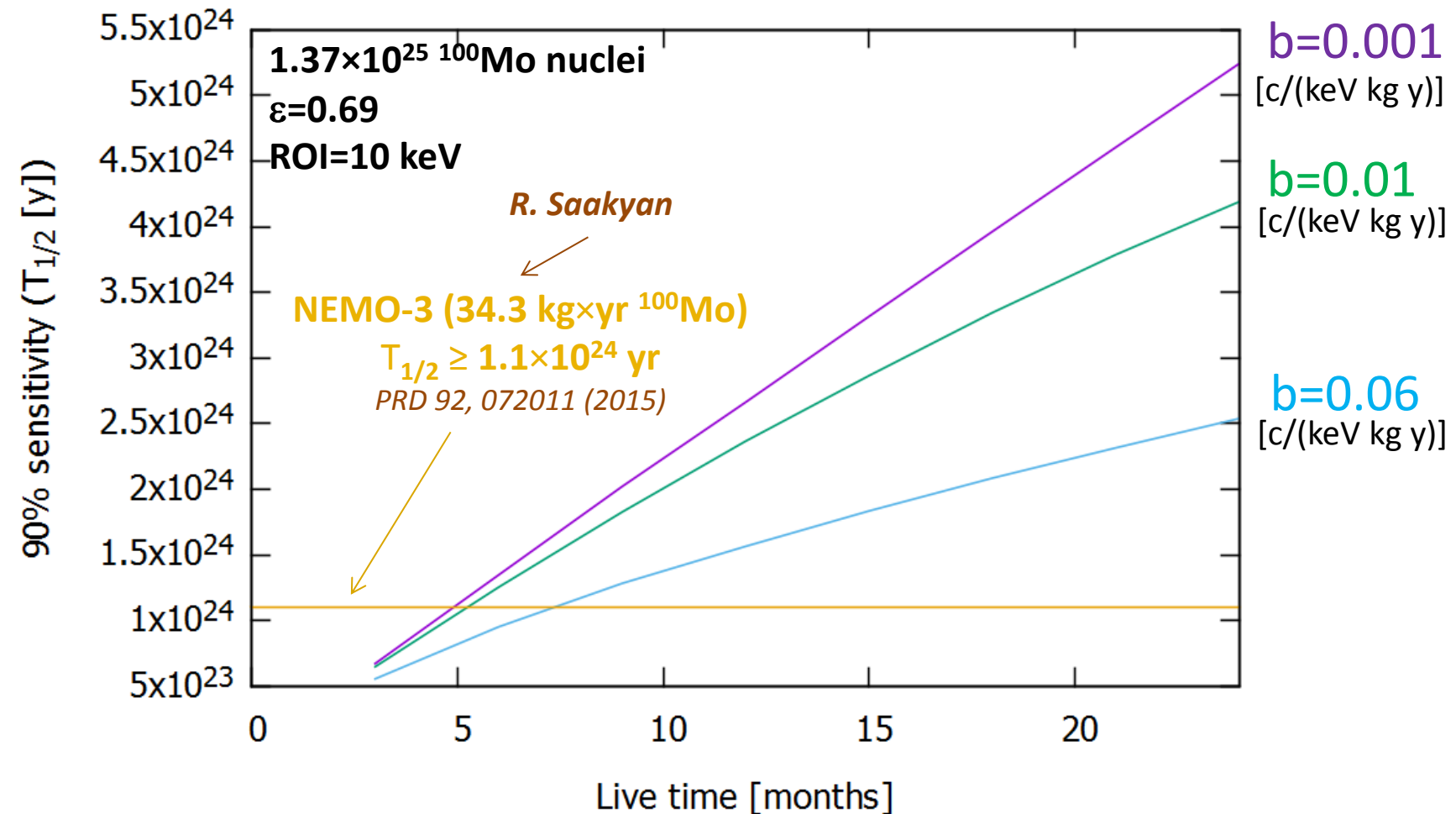
The contaminated connectors were substantially reduced for the CUPID-Mo run
(improvement by $\div 10$ is expected)

Full estimation of the background is in progress

Reasonable expectation: $\mathbf{b \sim 10^{-2} - 10^{-3} \text{ c/(keV kg)}}$

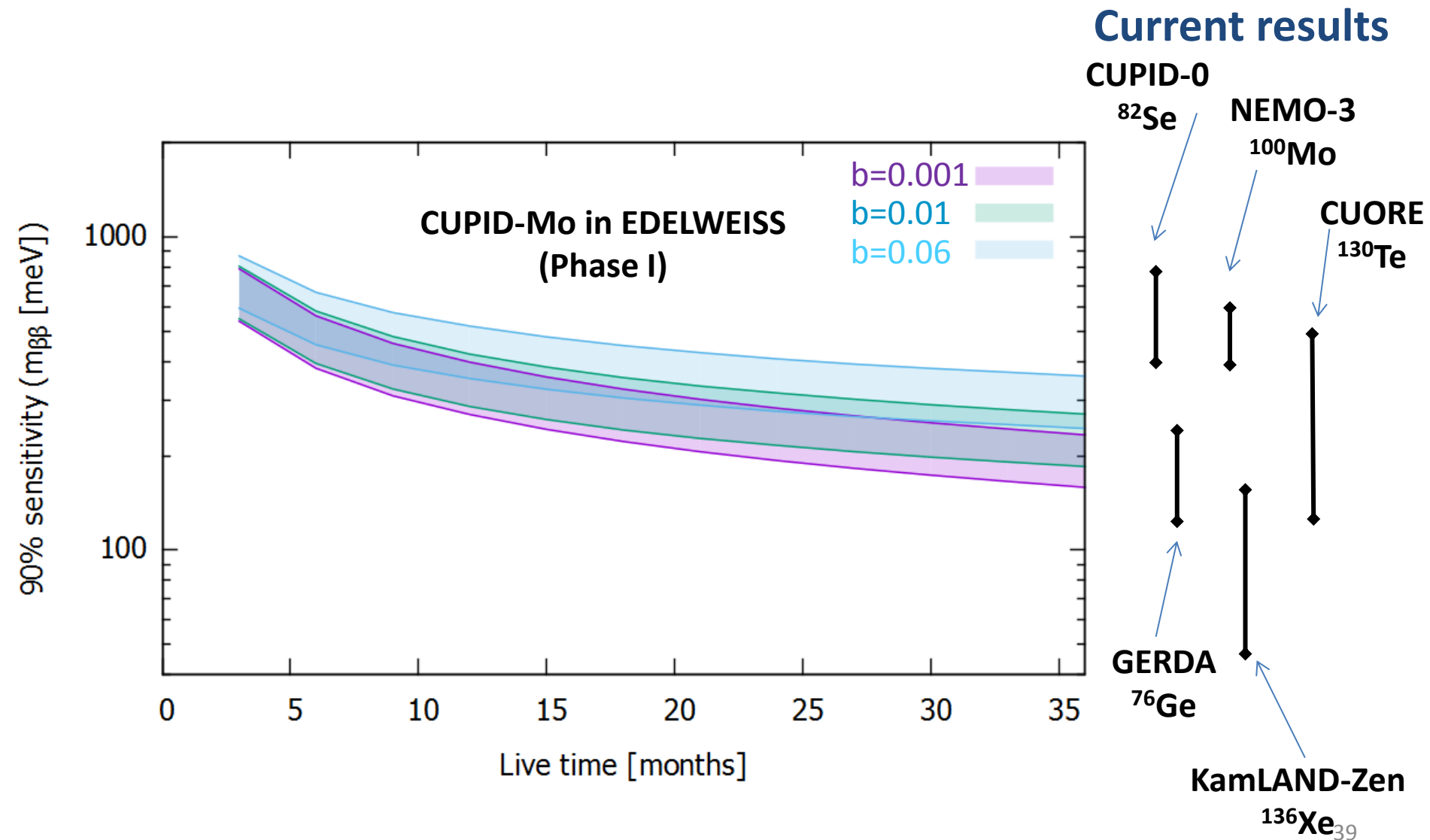
CUPID-Mo Modane

Evolution of the half-life sensitivity



CUPID-Mo Modane (Phase I)

Evolution of the Majorana mass sensitivity



CUPID-Mo Modane (Phase I)

Preliminary results at 22 mK

Preliminary data acquired at **22 mK**

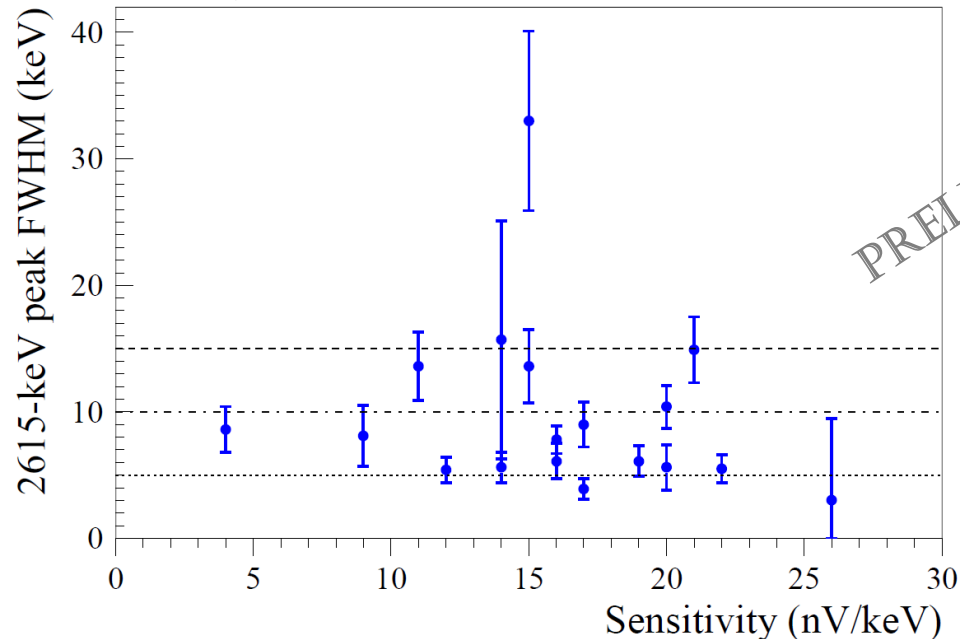
PLOTS RECEIVED YESTERDAY !

- This is not the final configuration – temperature is still decreasing (now ~20 mK)
- Large microphonic noise, work to improve is in progress
- 2 heat channels are lost

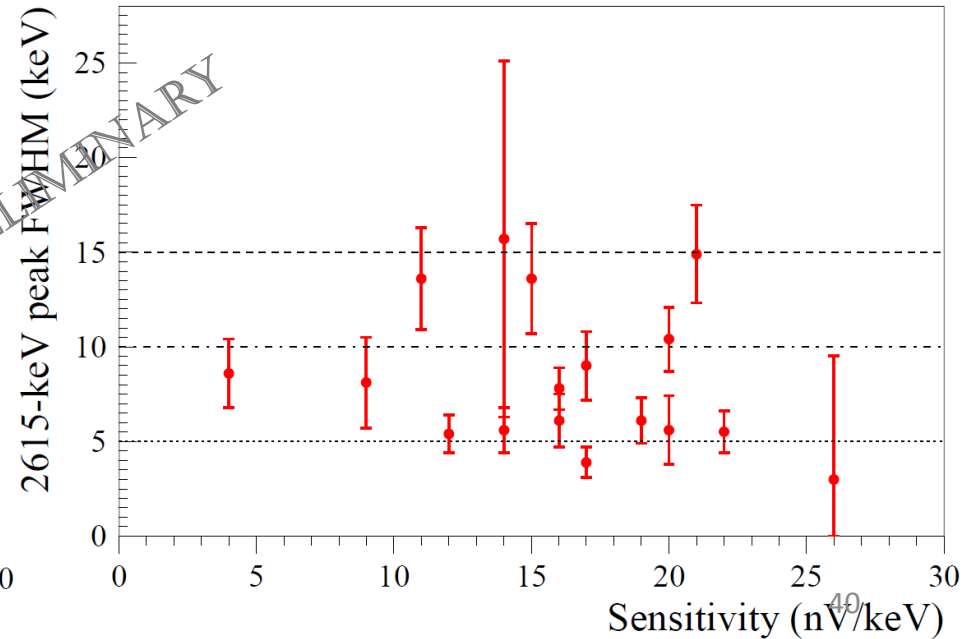
Calibration with thorite source – Th + U



$\text{Li}_2^{100}\text{MoO}_4$ performance at 22 mK, CUPID-Mo, Run316, LSM

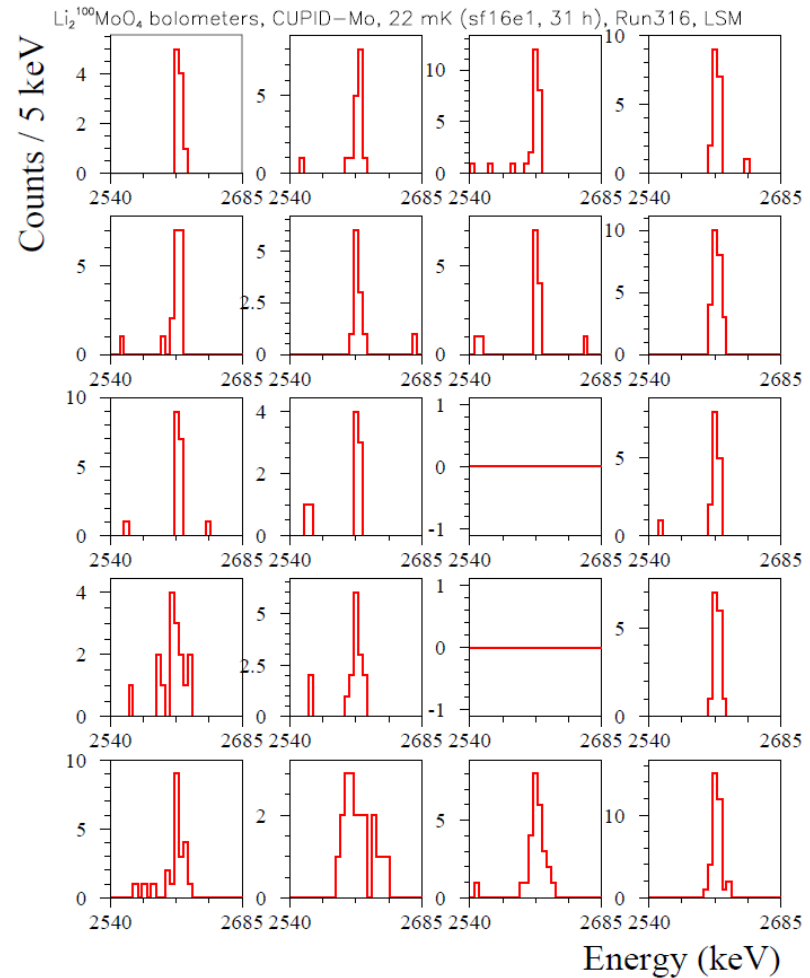
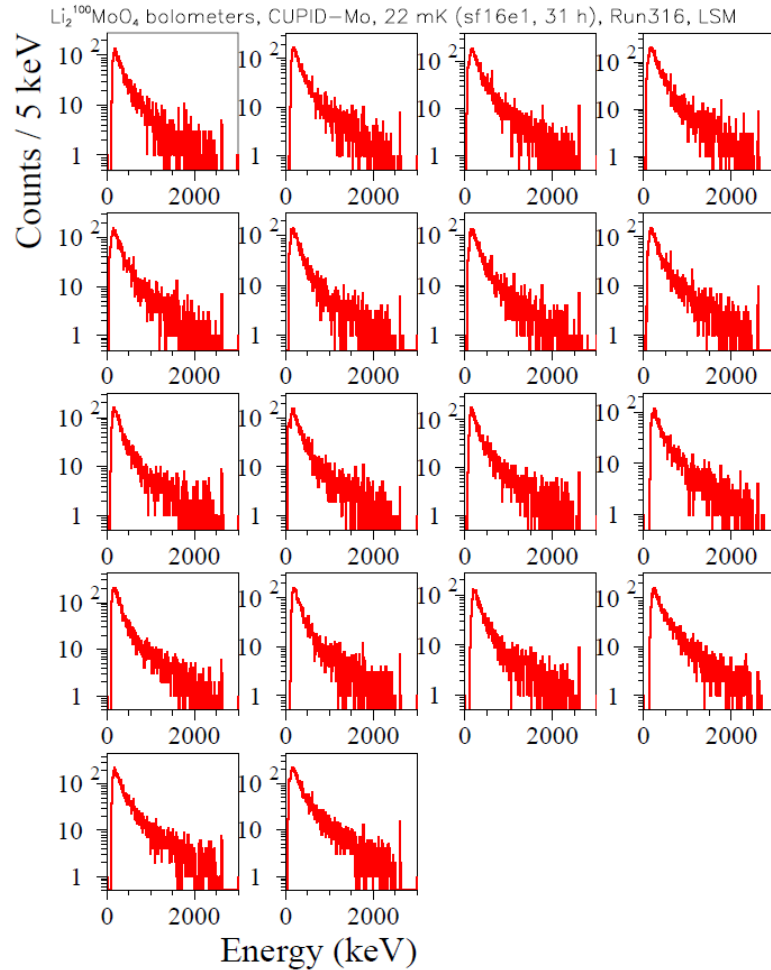


$\text{Li}_2^{100}\text{MoO}_4$ performance at 22 mK, CUPID-Mo, Run316, LSM



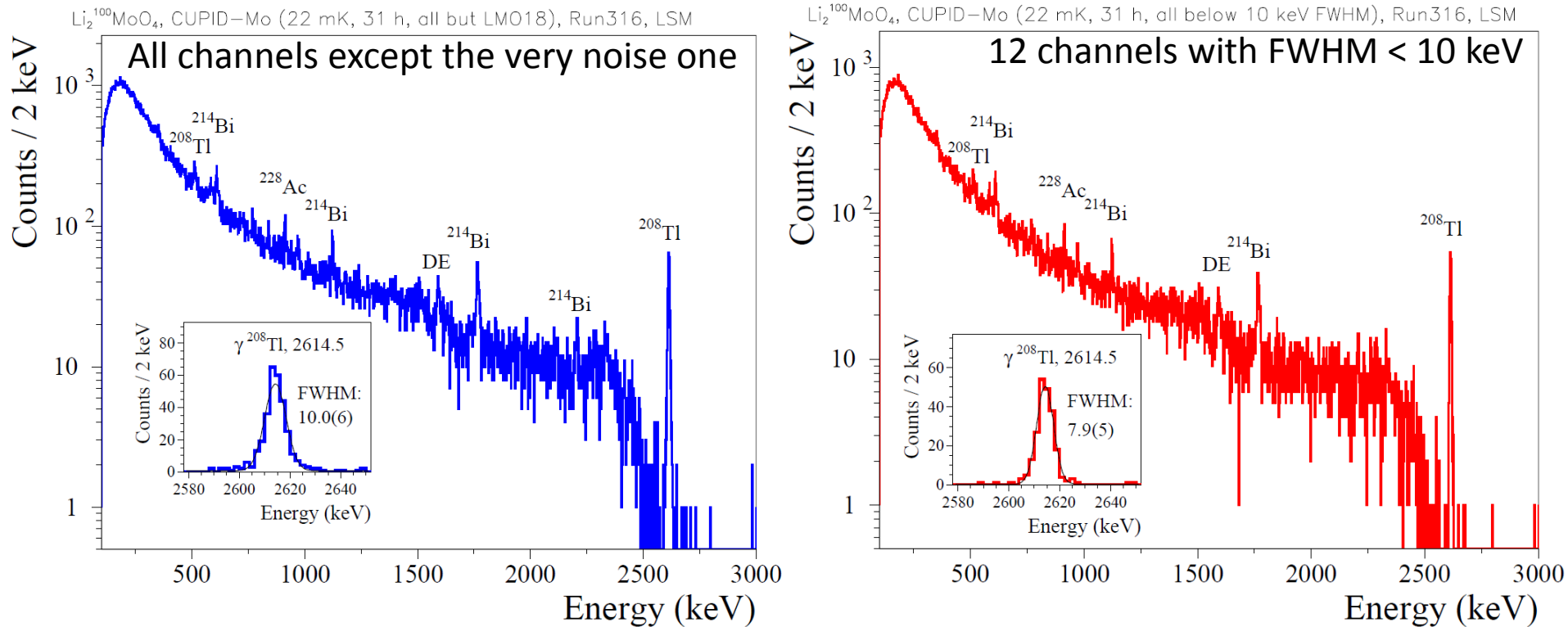
CUPID-Mo Modane(Phase I)

Preliminary results at 22 mK



CUPID-Mo Modane (Phase I)

Preliminary results at 22 mK



CUORE → CUPID

Ke Han - CUORE

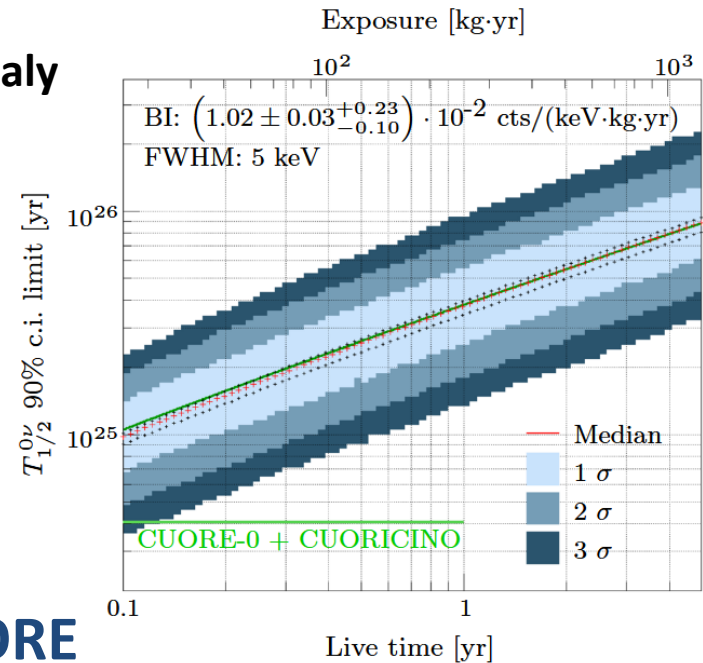
CUORE is collecting data successfully

LNGS – Italy

➤ 5 y projected half-life sensitivity: $\sim 10^{26}$ y

➤ $m_{\beta\beta} < 50 - 190$ meV

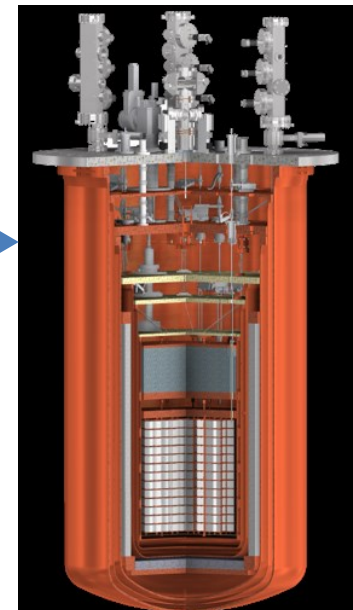
- Background according to expectations:
 $1.4 \pm 0.2 \times 10^{-2}$ c/(keV·kg·yr)
- Energy resolution close to expectations:
 ~ 7.7 keV FWHM → margins for improvement



Three important messages from CUORE

1. A tonne-scale bolometric detector is technically feasible
2. Analysis of ~ 1000 individual bolometers is handable
3. An infrastructure to host a bolometric next-generation $0\nu\beta\beta$ experiment is already available

CUPID is the natural evolution of CUORE



Prospects for CUPID

Results of the ongoing R&D and demonstrators + CUORE background model



1. $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers → promising **baseline option** for CUPID

2. $^{130}\text{TeO}_2$ Cherenkov bolometers → mature viable alternative
Phys. Rev. C 97 (2018) 032501(R)

Fast and high-sensitivity light detectors are a common R&D

- Detection of Cherenkov light in TeO_2 *M. Vignati – CALDER*
- Rejection of $2\nu 2\beta$ random coincidences in $\text{Li}_2^{100}\text{MoO}_4$ *EPJ C 77, 3 (2017)*

The purpose of CUPID is to fully explore the IO region

Mission: **half-life sensitivity higher than 10^{27} y** *arXiv:1504.03599*

With background < 0.1 counts/(ton y) in the ROI, ^{100}Mo sensitivity is **2.1×10^{27} y**
 $m_{\beta\beta} < 6 - 17$ meV

- **CUPID collaboration will be formed in the near future**
- **CUPID kick-off meeting is being planned in fall 2018**

Beyond light detectors: CROSS

ERC advanced grant CROSS

Cryogenic Rare-event Observatory with Surface Sensitivity

CROSS develops an innovative bolometric technology to search for 0ν -DBD



➤ **Core of the project** (high risk / high gain)

Background rejection through **pulse shape discrimination**

- **Surface sensitivity** through **superconductive Al film coating**
 - **Fast NbSi high-impedance TES** to replace / complement NTDs if necessary
- **get rid of light detectors**

➤ Complete crystallization of available **^{100}Mo (10 kg)** in Li_2MoO_4 elements → **CUPID-Mo demonstrator in LNGS (Phase II)**

➤ Purchase / crystallize **^{130}Te (up to 10 kg)** in TeO_2 elements

➤ Run demonstrator in a dedicated cryostat (**LSC – Spain**)

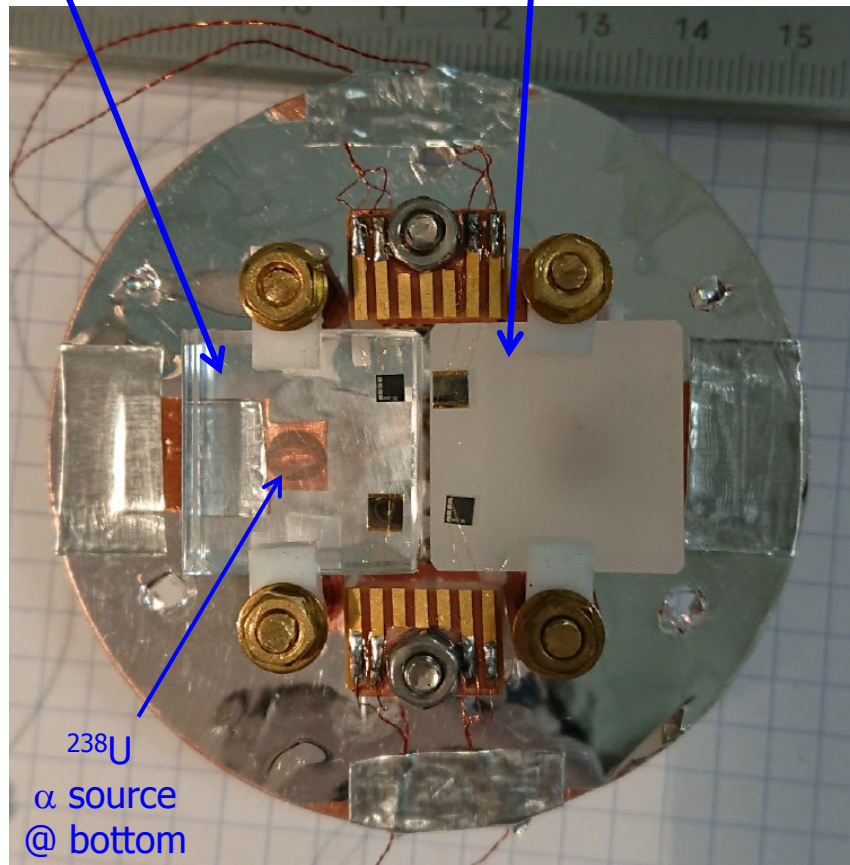
CROSS R&D at CSNSM

Above-ground runs at CSNSM

CROSS #1

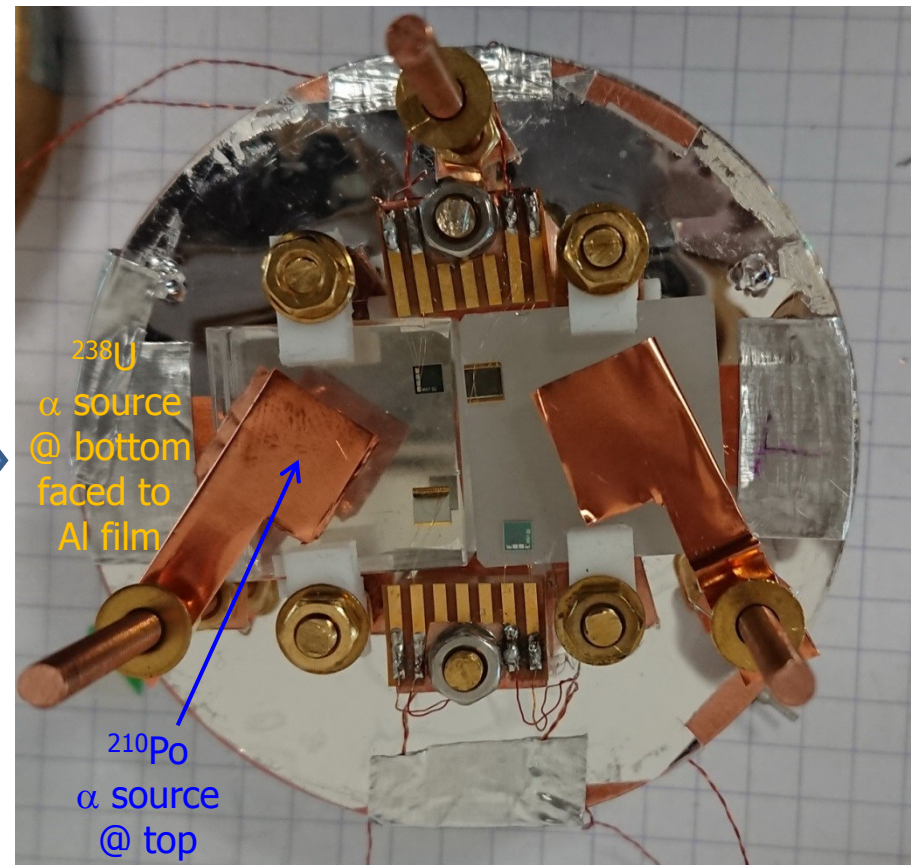
Li_2MoO_4 (2×2×1 cm, 12 g)

TeO_2 (2×2×1 cm, 25 g)

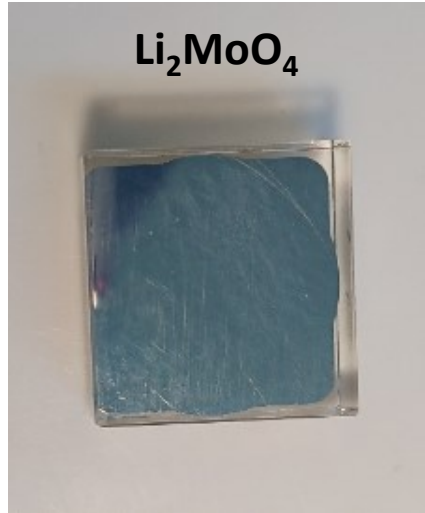


CROSS #2

Li_2MoO_4 & TeO_2 (both with 10 μm Al @ bottom)



Preparation of the Al film test

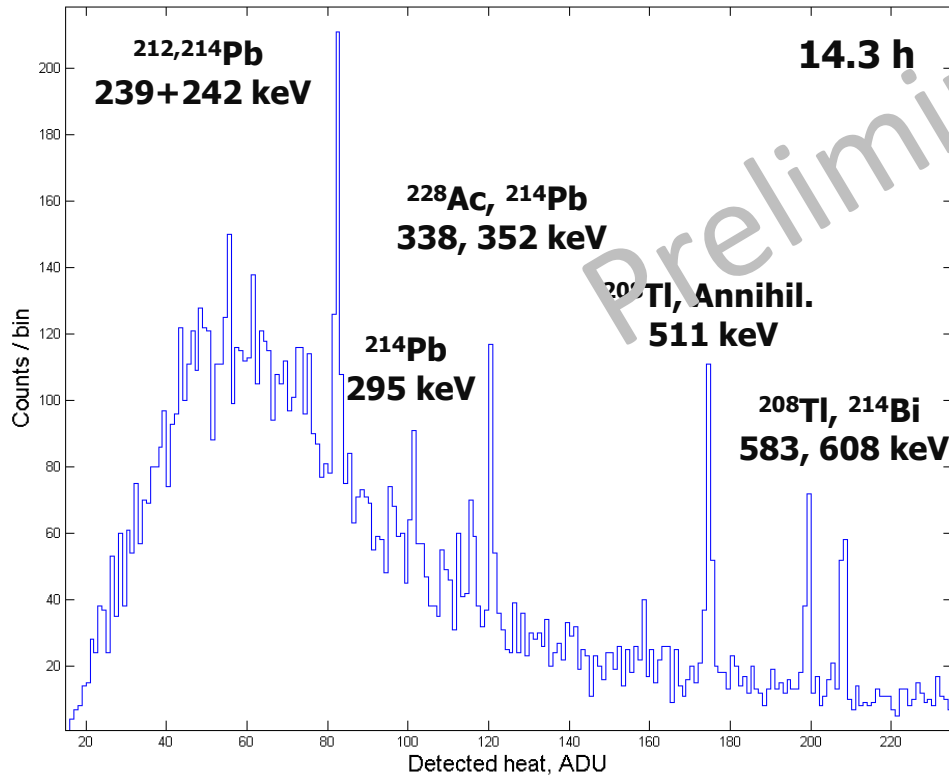


- **2x2x1 cm crystals**
- **Al coating** by evaporation on a **2x2 cm face** for both crystals (thickness: 10 μ m) – (**1/4 of crystal surface is covered**)

TeO₂ behaviour with Al coating

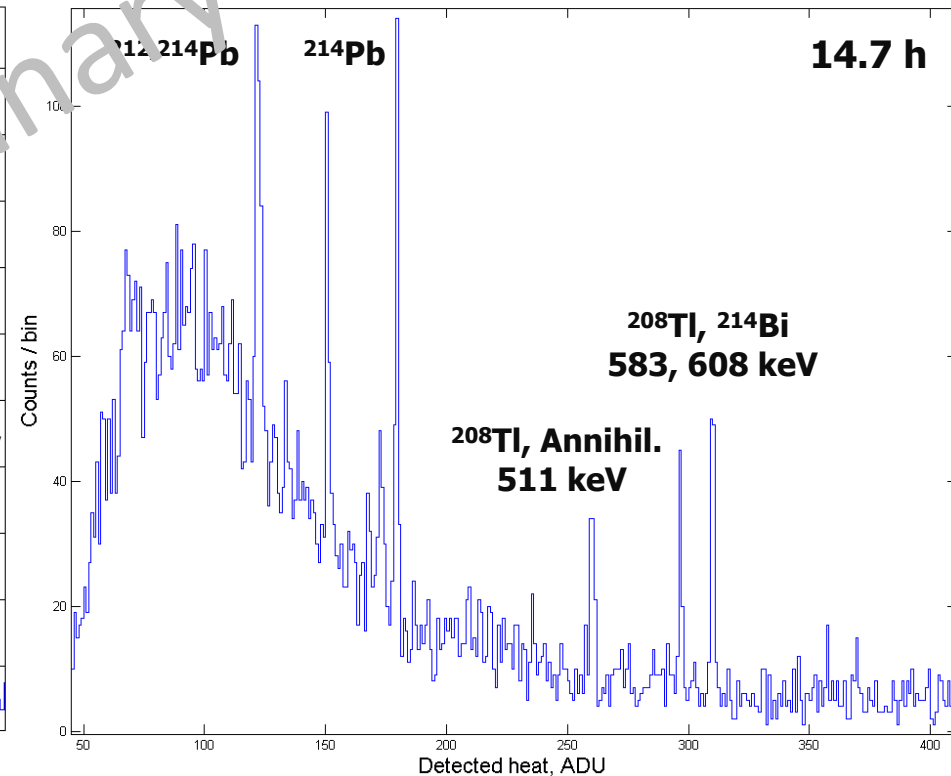
No Al film

TeO₂ without Al film, Run40 in Ulysse, CSNSM



With Al film

TeO₂ with Al film, Run44 in Ulysse, CSNSM



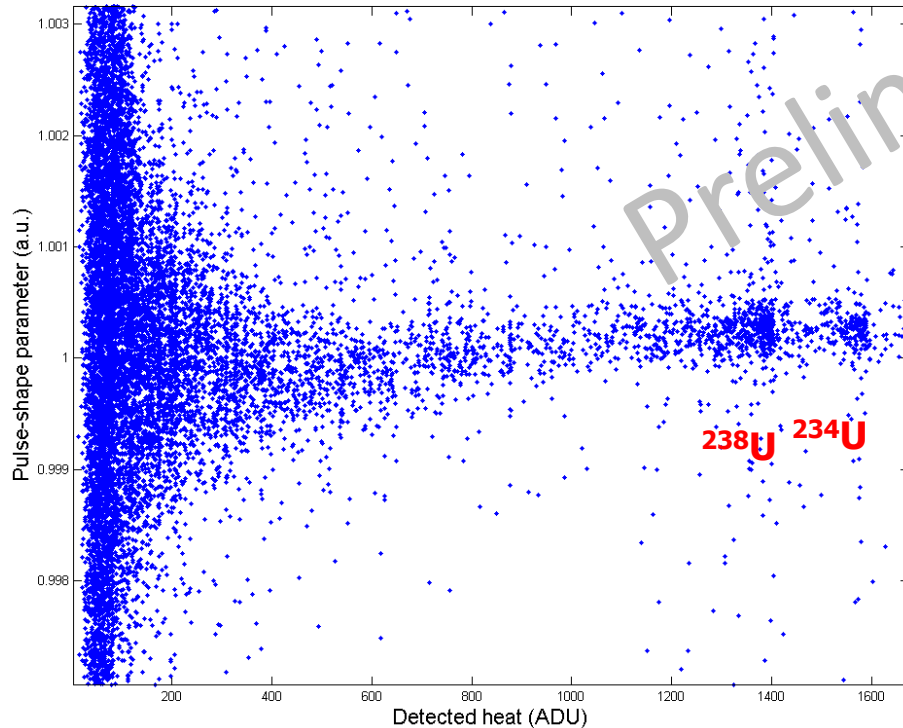
The 10 μm -thick aluminum film affects **neither the sensitivity nor the energy resolution** of the TeO₂ bolometer.

TeO: Pulse shape discrimination

Pulse shape parameter vs. Pulse amplitude

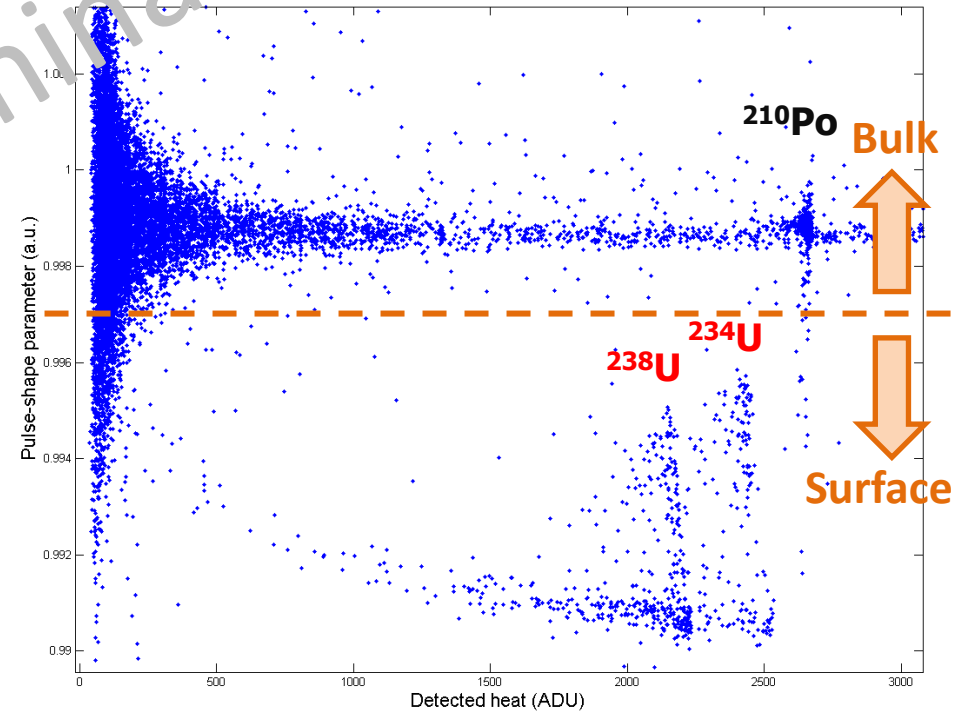
No Al film

TeO₂ without Al film, Run40 in Ulysse, CSNSM



With Al film

TeO₂ with Al film, Run44 in Ulysse, CSNSM



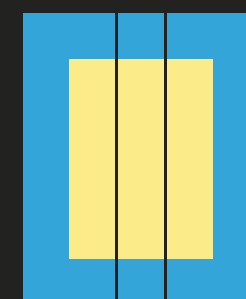
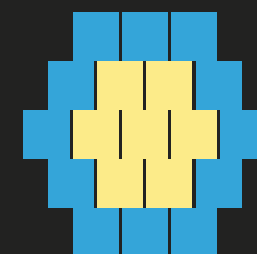
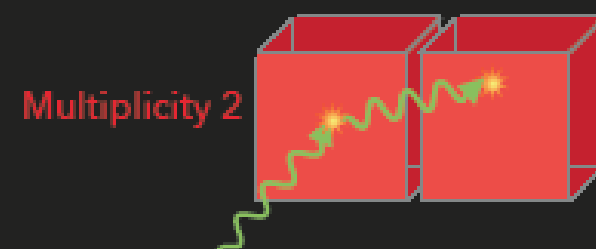
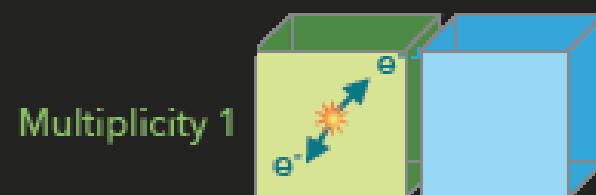
Alphas impacting on the film side are clearly discriminated

Conclusions

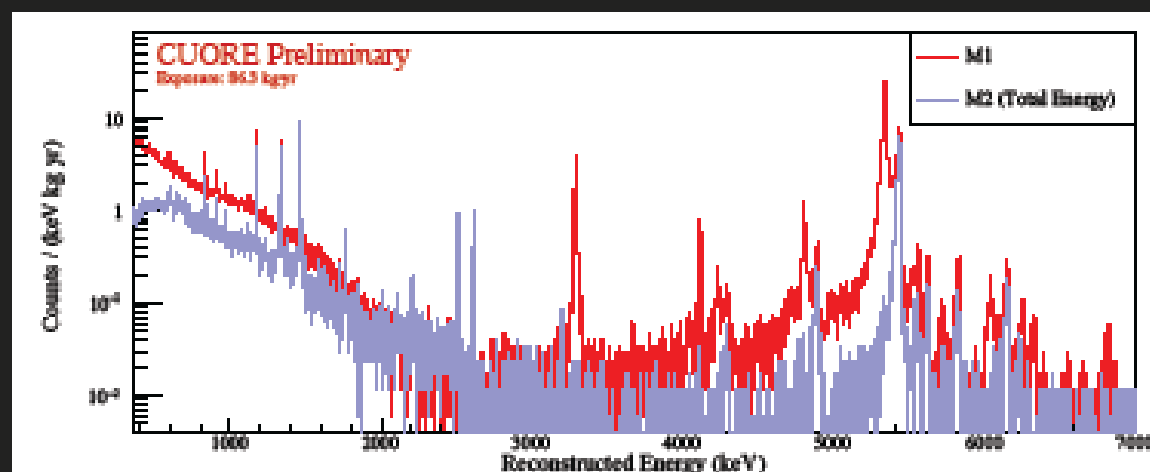
- Study of neutrinoless double beta decay is one of the most urgent topics in particle physics and cosmology
- The bolometric approach is a viable technique, confirmed at large scale by the CUORE results
- A promising technology based on enriched $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers was developed and is now applied to the CUPID-Mo demonstrators
- Innovative approaches for surface background are under study
- CUPID (CUORE follow up) implemented with Li_2MoO_4 scintillating bolometers is one of the most promising next-generation searches

Building the CUORE Background Model

- ▶ 86.3 kg·yr of TeO_2 from summer 2017
- ▶ Split data into inner and outer layers
- ▶ Split data into Multiplicity 1 (M1), Multiplicity 2 (M2), and Multiplicity 2 Sum ($\Sigma 2$) spectra
- ▶ Higher multiplicity spectra sensitive to backgrounds

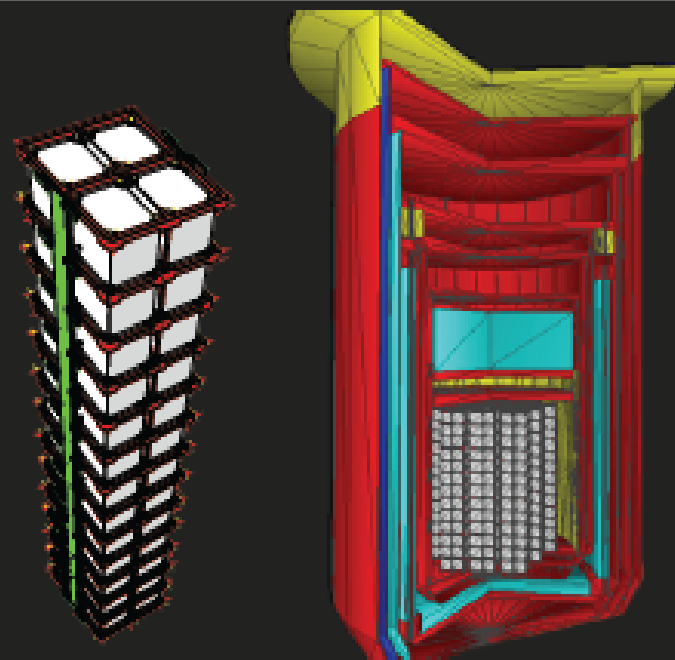


Inner Layer
Outer Layer



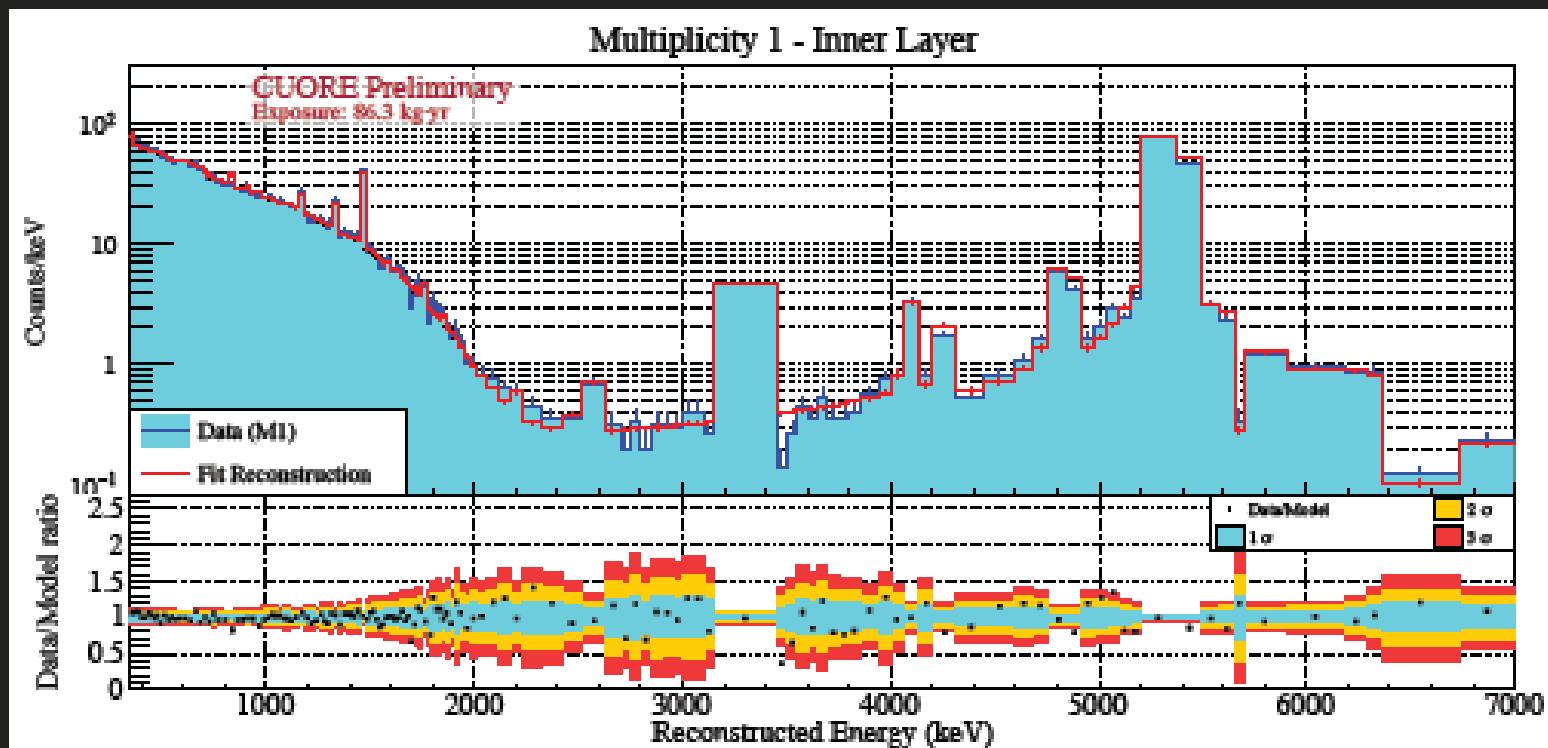
Fitting the CUORE Background Model

- ▶ Simulate the contaminations coming from different cryostat components using a detailed Geant4 MC simulation
- ▶ ~60 independent parameters representing various contaminations that could contribute to the CUORE background model
- ▶ Perform a large Bayesian fit to the data using a MCMC Gibbs sampler
- ▶ Flat priors on all parameters except muons which come from a cosmogenic analysis



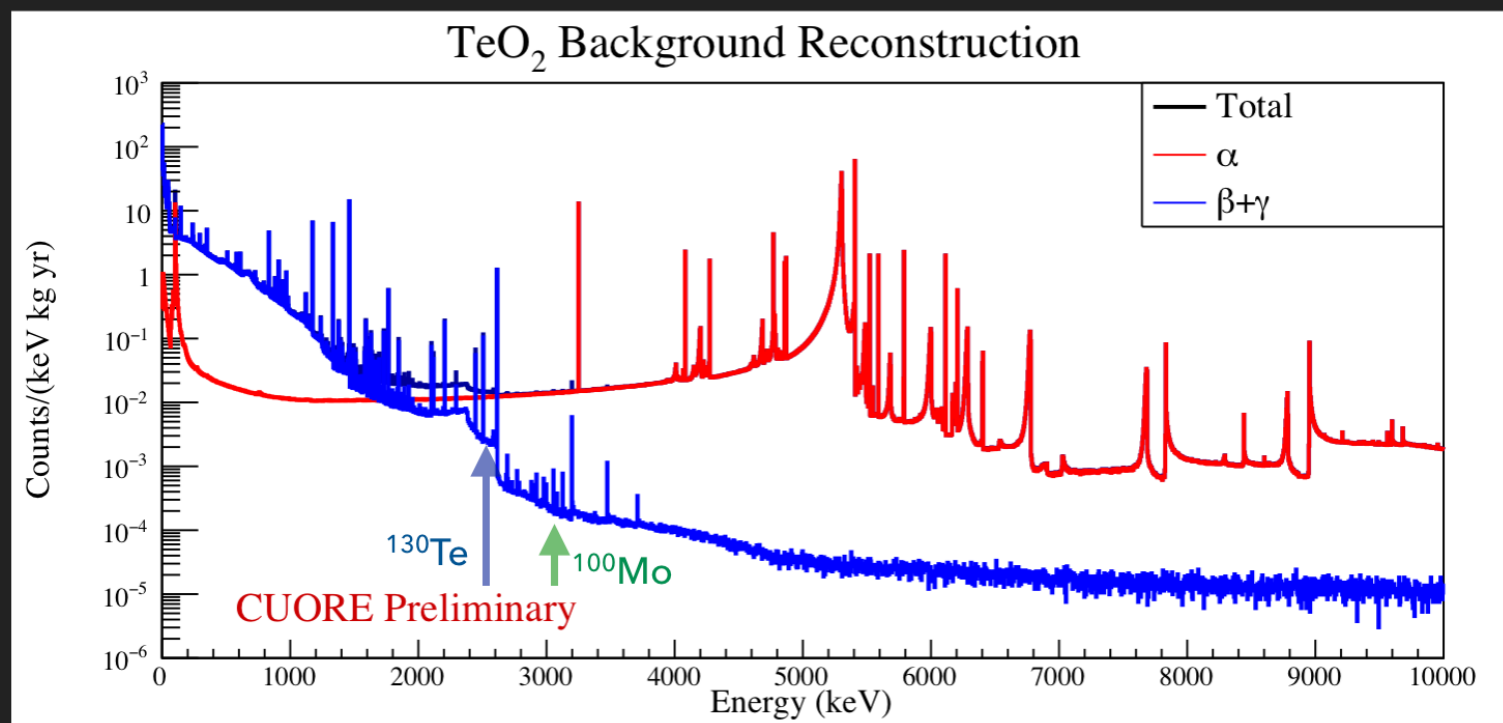
Volume	Type	Components
TiO ₂	Bulk	²³² Th, ²³² Pb, ²³² Th, ²³² Ra, ²³² Pb, ²³² U, ²³² Th, ²³² Ra, ²³² Pb, ⁴⁰ K, ⁶⁰ Co, ¹³⁷ Ba, ²³² Pb
TiO ₂	Surface (0.01 μm)	²³² Th, ²³² Ra, ²³² Pb, ²³² U, ²³² Th, ²³² Ra, ²³² Pb, ²³² Pb
TiO ₂	Surface (1 μm)	²³² Pb
TiO ₂	Surface (10 μm)	²³² Pb, ²³² Th, ²³² U
CuNOSV	Bulk	²³² Th, ²³² U, ⁴⁰ K, ⁶⁰ Co, ⁶⁴ Mn
CuNOSV	Surface (0.01 μm)	²³² Pb, ²³² Th, ²³² U
CuNOSV	Surface (1 μm)	²³² Pb, ²³² Th, ²³² U
CuNOSV	Surface (10 μm)	²³² Pb, ²³² Th, ²³² U
Roman lead	Bulk	²³² Th, ²³² U, ^{109m} Ag
Top lead	Bulk	²³² Th, ²³² U, ²¹⁰ Pb
Ext. lead	Bulk	²¹⁰ Pb
CUORE	Bulk	²³² Th, ²³² U, ⁶⁰ Co
External	-	Cosmic muons

Fitting the CUORE Background



➡ Able to reconstruct the major features of the observed spectrum in CUORE

Background Model Contamination Levels (Very Preliminary!)



TeO₂ γ background at 2527 keV $\sim 2.6 \times 10^{-3}$ cts/(keV·kg·yr)

Li₂MoO₄ γ background at 3034 keV $\sim 1 \times 10^{-4}$ cts/(keV·kg·yr) (Based on assumptions)

$2\nu\beta\beta$ decay random coincidences

Contribution to the background index in the ROI: $\varnothing 50 \text{ mm} \times 50 \text{ mm}$ (300 g)

$$\text{BKG(rc)} [\text{counts}/(\text{keV kg y})] = 3 \times 10^{-4} [T_R / 1 \text{ ms}] [M / 300 \text{ g}]$$

Our approach (partial simulation + PSD)

- Take a large value for T_R (typically $T_R \sim 3 \times \text{rise time}$)
- Use real-shape pulses
- Use real noise baselines
- Generate pulses with correct 2ν pulse amplitude distribution
- Calculated rejection efficiency by PSD of pulse-pair separated by less than T_R
- Multiply the above formula by rejection efficiency

EPJ C 72, 1989 (2012)

EPJ C 74, 2913 (2014)

EPJ C 77, 3 (2017)

In a real case (heat channel):

$$T_R = 45 \text{ ms}$$

Rejection efficiency by PSD = 99.3 % (using the so-called **mean-time method**)
(95% acceptance)

$$\text{BKG(rc)} [\text{counts}/(\text{keV kg y})] \sim 1 \times 10^{-4}$$

It can be improved:

- Full simulation
- Advanced pulse shape parameters (after optimum filtering)
- Using the light channel with high signal-to-noise ratio light detectors → Neganov-Luke

$2\nu\beta\beta$ decay random coincidences

Contribution to the background index in the ROI:

$\varnothing 50 \text{ mm} \times 50 \text{ mm}$ (300 g)

$$\text{BKG(rc)} [\text{counts}/(\text{keV kg y})] = 3 \times 10^{-4} [\text{counts}/(\text{keV kg y})]$$

Our approach (partial simulation + PSD)

- Take a large value for T_R (typically 100 ns)
- Use real-shape pulses
- Use real noise background
- Generate random coincidences
- Calculate the background index

Eur. Phys. J. C (2017) 77:3
DOI 10.1140/epjc/s10052-016-4565-z

Regular Article - Experimental Physics

Rejection of randomly coinciding events in $\text{Li}_2^{100}\text{MoO}_4$ scintillating bolometers using light detectors based on the Neganov-Luke effect

D.M. Chernyak^{1,6}, F.A. Danevich¹, L. Dumoulin², A. Giuliani^{2,3,4}, M. Mancuso^{2,3,7}, P. de Marcillac², S. Marmieros², C. Nones⁴, E. Olivieri², D. V. Poda^{1,2}, V. I. Tretyak^{1,5}

the so-called mean-time method)

$$b \sim 6 \times 10^{-5} \text{ c}/(\text{keV kg y})$$



It can be

- Full simulation
- Advanced pulse shape parameters (after optimum filtering)
- Using the light channel with high signal-to-noise ratio light detectors → Neganov-Luke

Neutrons



$$\sigma = 940 \text{ barn}$$

(thermal neutrons)

$$Q = 4.78 \text{ MeV}$$

Harmless

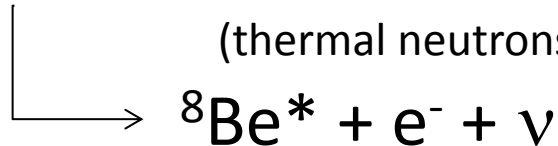
No associated β radiation

Huge internal energy deposition

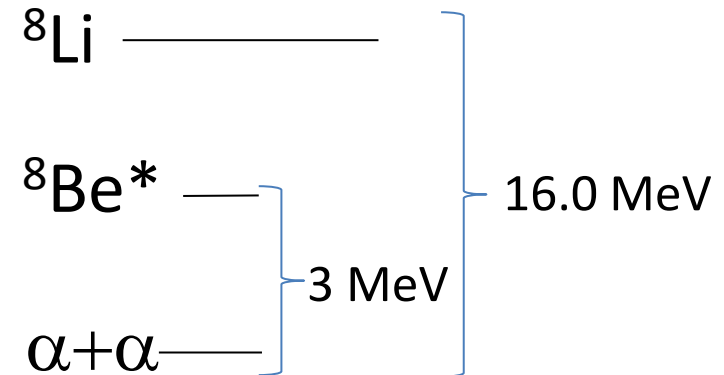
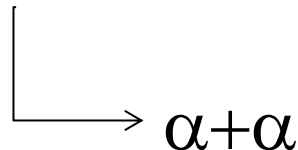


$$\sigma = 45.4 \text{ mbarn}$$

(thermal neutrons)



Prompt ($\Gamma \sim 1.5 \text{ MeV}$)

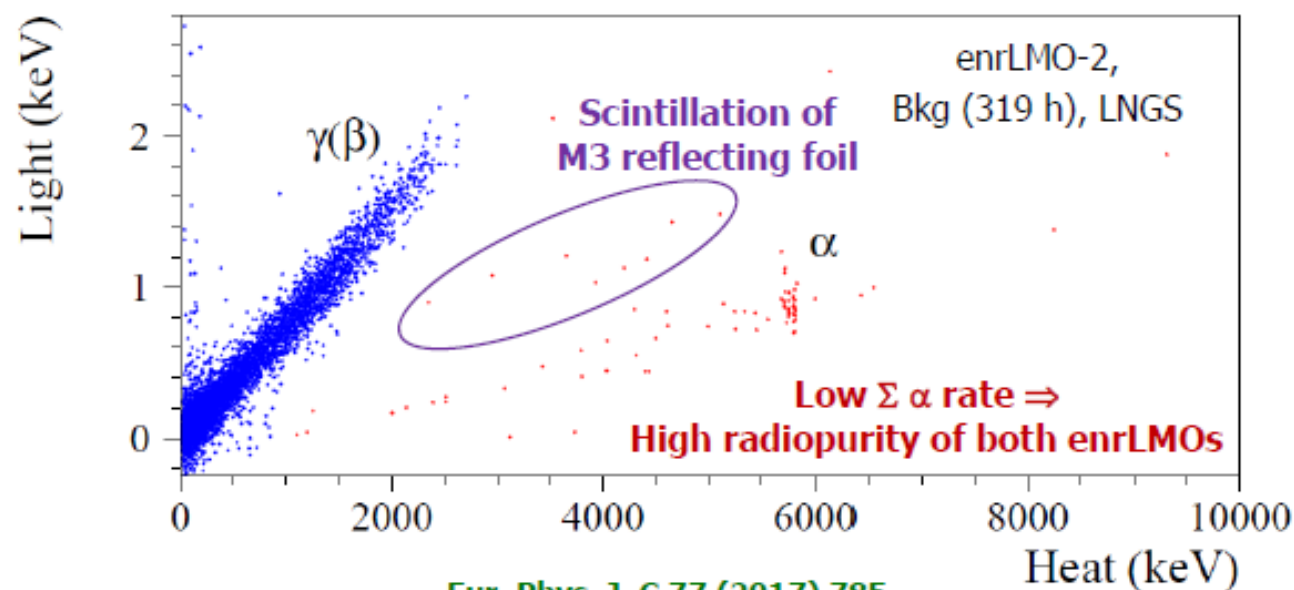
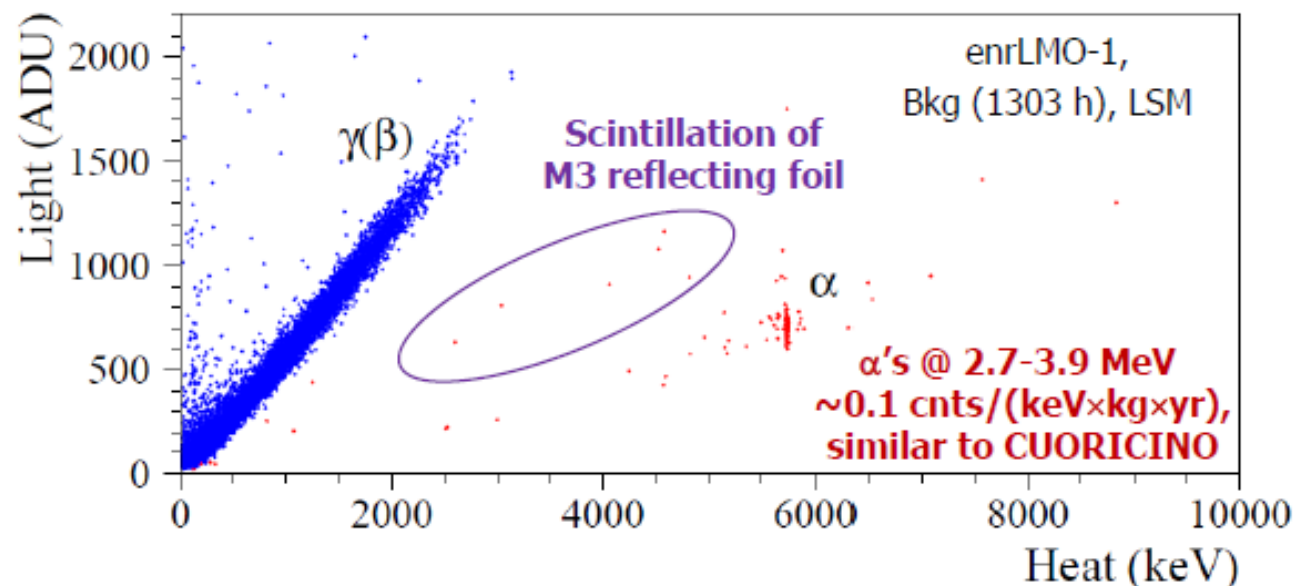


Harmless

Very low cross section

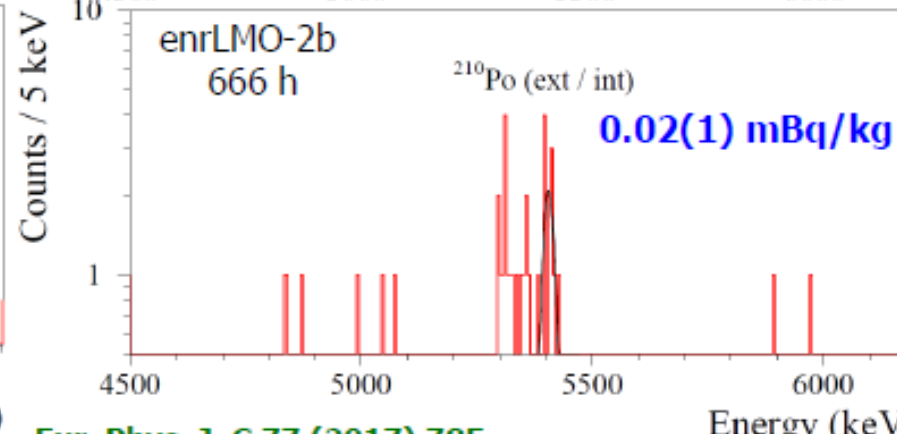
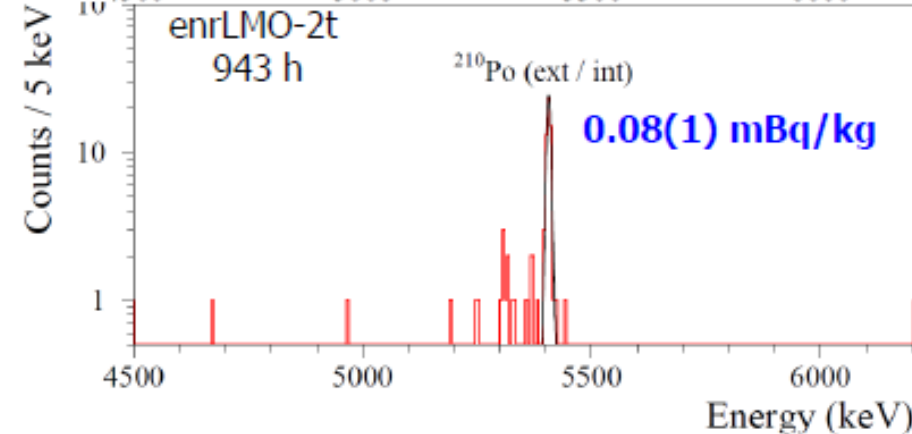
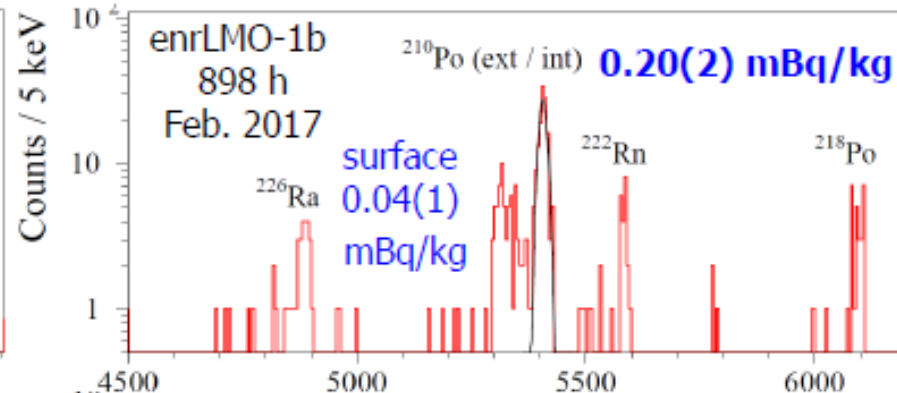
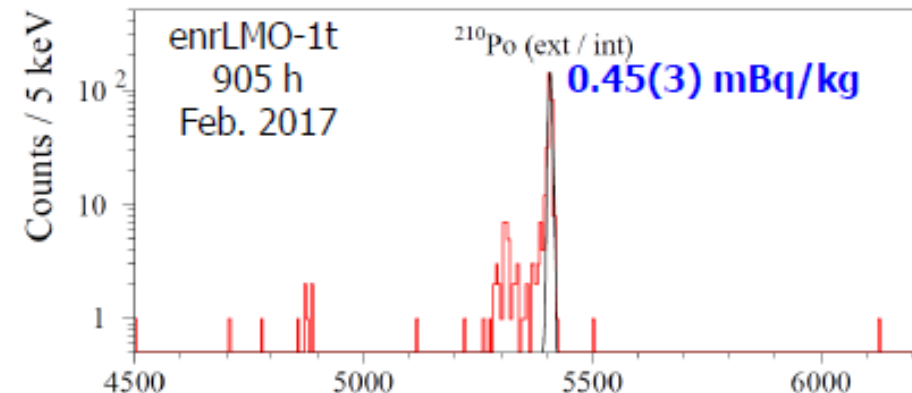
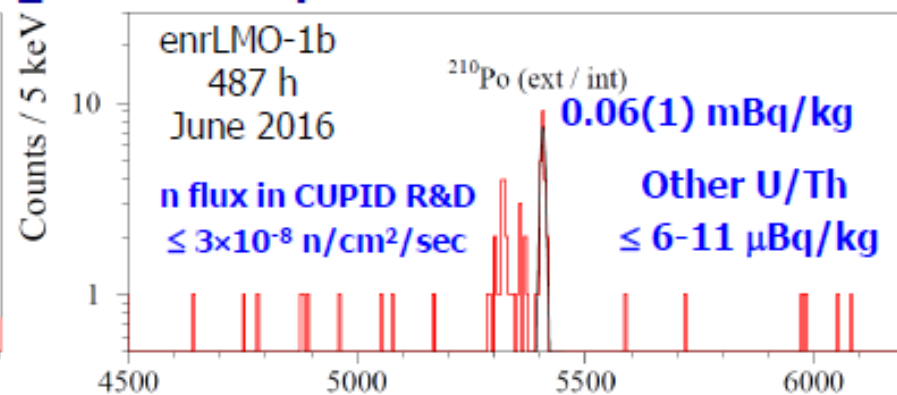
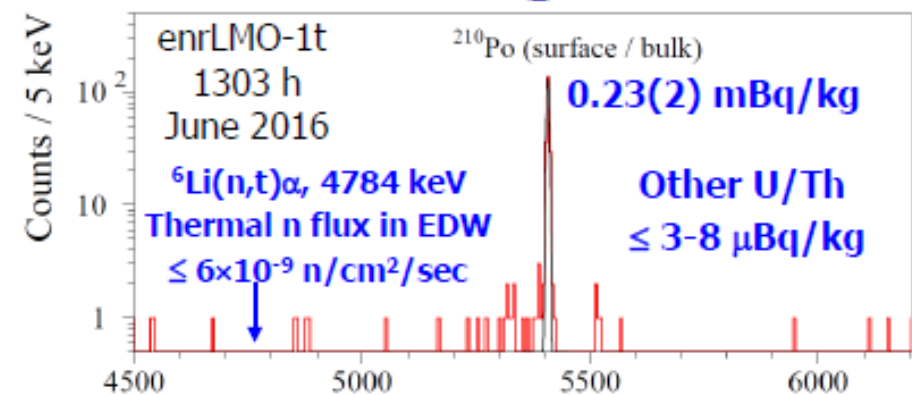
Mixed events with α component

First background measurements with $\text{Li}_2^{100}\text{MoO}_4$



Eur. Phys. J. C 77 (2017) 785

α Background of $\text{Li}_2^{100}\text{MoO}_4$ detectors



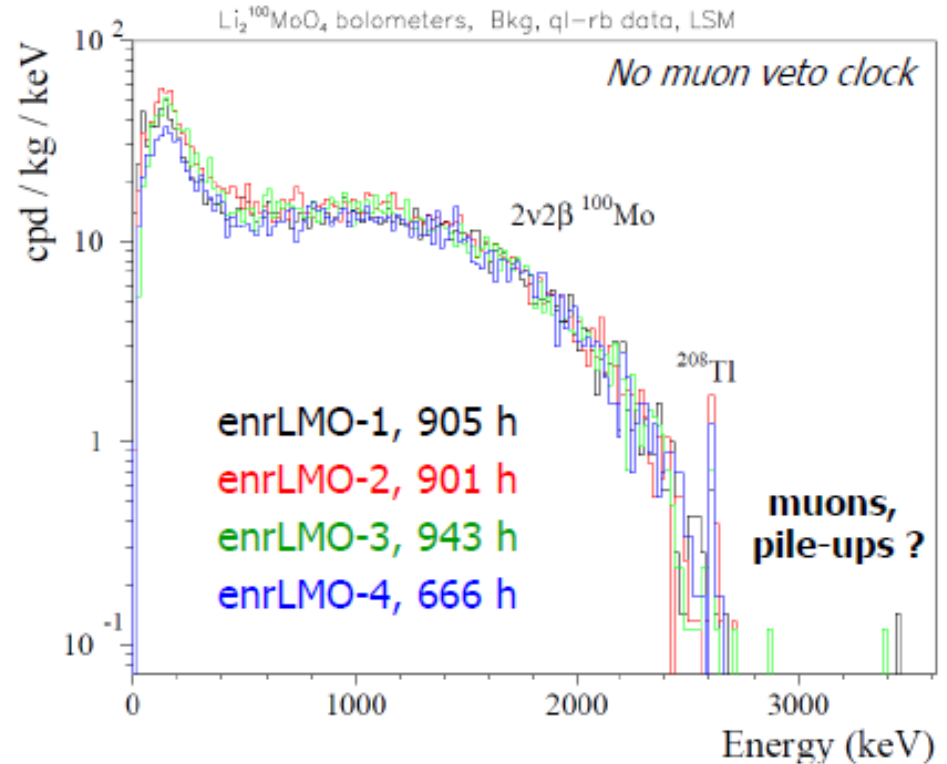
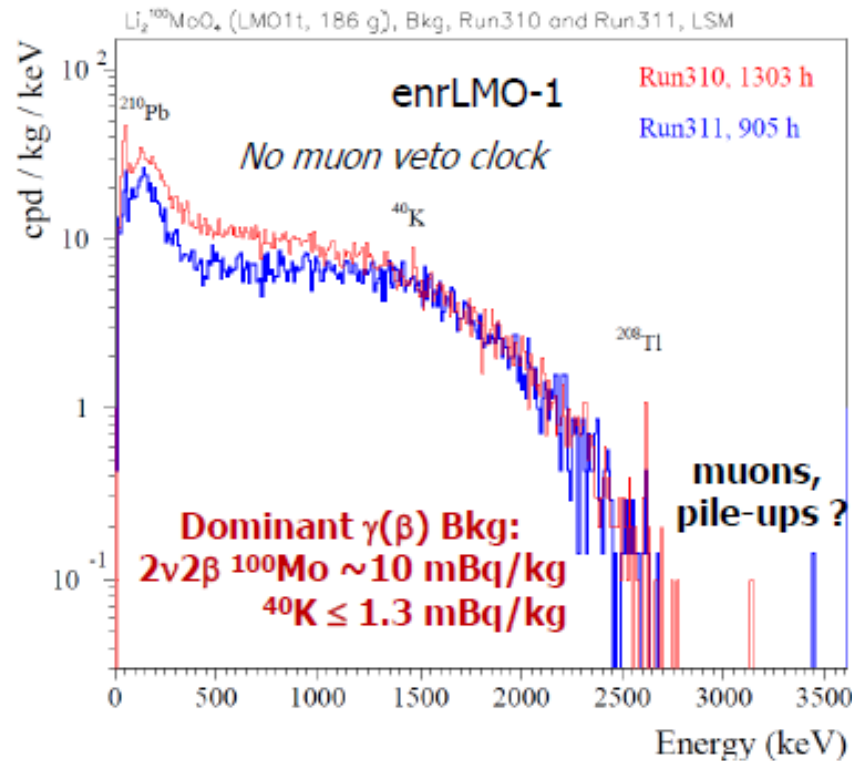
Eur. Phys. J. C 77 (2017) 785

AIP Conf. Proc. 1894 (2017) 020017

D. Poda

59 ²⁶

$\gamma(\beta)$ Background of $\text{Li}_2^{100}\text{MoO}_4$ detectors

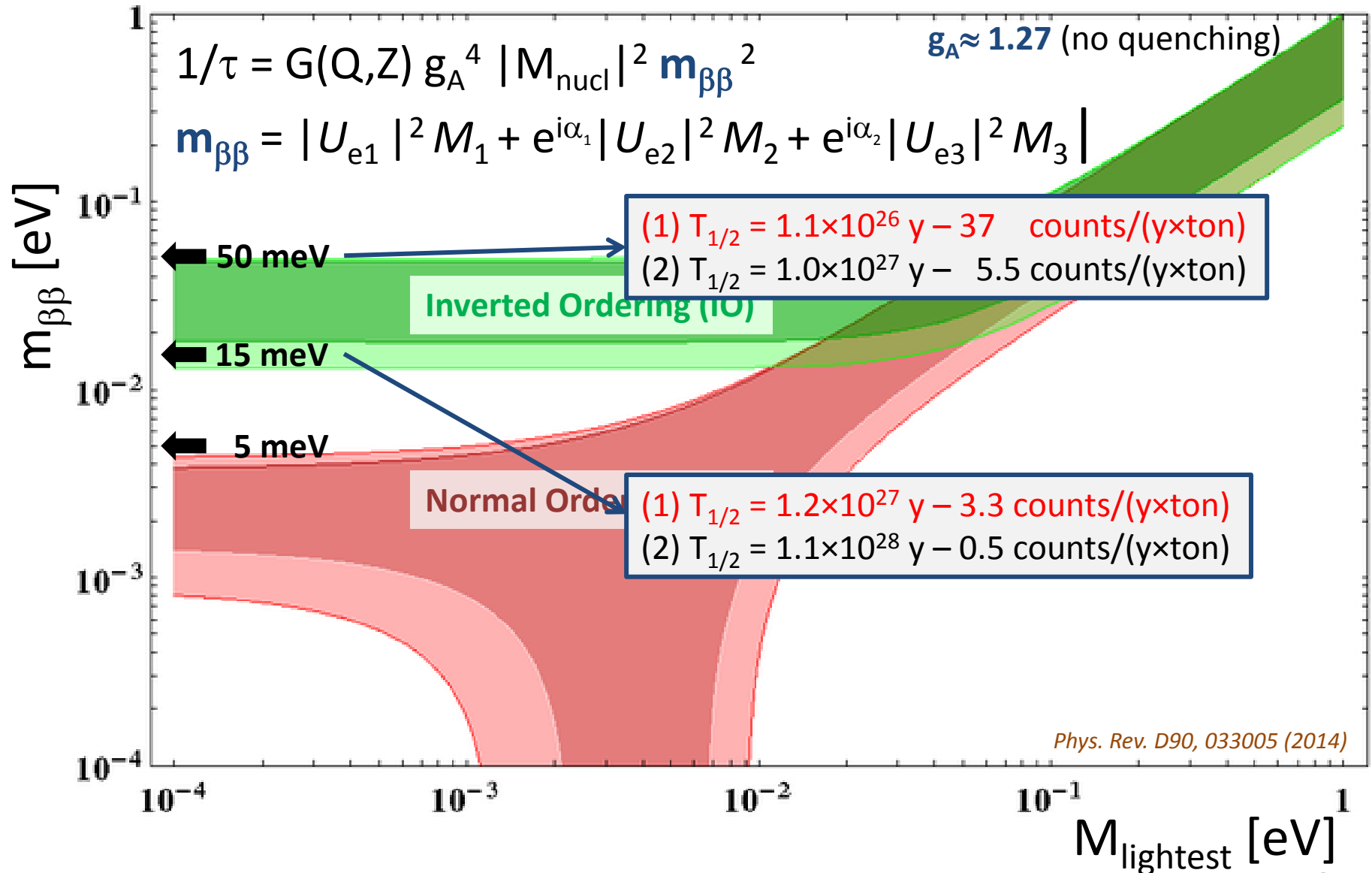


Rate [cnts/day/kg] of 2615 keV γ 's	LUMINEU	CUPID-0
Spread	0.7(3) - 2.1(5)	0.03 - 0.33
Average	1.1(2)	0.14

L. Cardani, private communication

External ^{232}Th background is expected to be reduced thanks to a new wiring

How difficult is it?

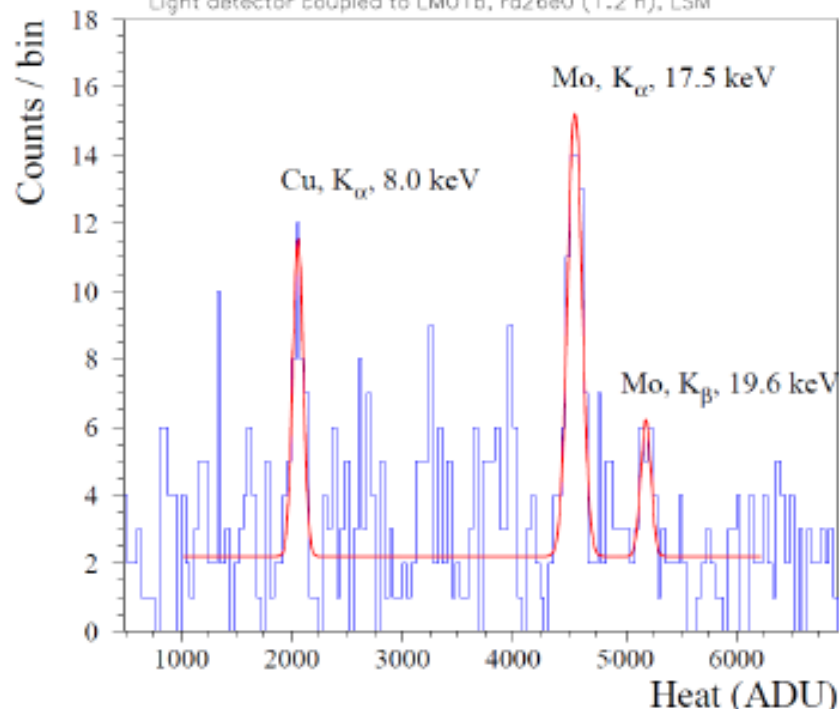


Light detectors calibration by external γ source

A high activity ^{60}Co source (~ 100 kBq) is used by EDELWEISS to regenerate Ge bolometers...

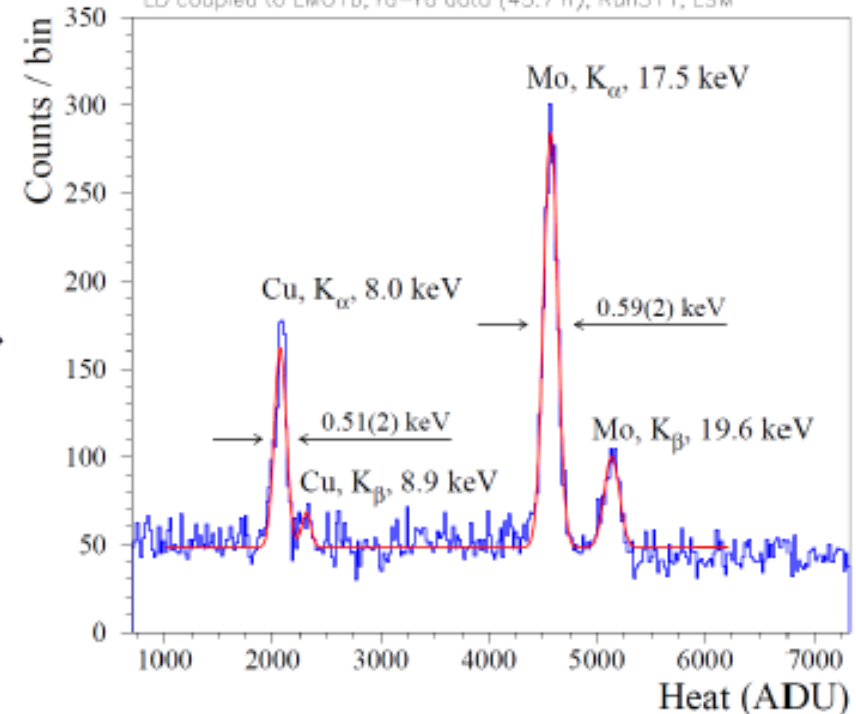
^{60}Co "regen" run, 1.2 h

Light detector coupled to LM01b, ra26e0 (1.2 h), LSM



^{60}Co "regen" runs, 45.7 h

LD coupled to LM01b, ra-rd data (45.7 h), Run311, LSM



A simple & fast way to calibrate LDs by γ source excited X-ray fluorescence

LUMINEU light detectors performance

More colder temperature conditions require stronger NTD polarization to overcome operation instability of high resistivity NTDs (5 mg) with AC electronics

Light detector	T_{bolo} (mK)	I_{bolo} (nA)	R_{bolo} (M Ω)	Signal ($\mu\text{V/keV}$)	$\text{FWHM}_{\text{baseline}}$ (keV)
M3	17.0	6.8	0.55	0.86	0.11
Lum10	19.3	2.3	1.52	2.13	0.07
	17.0	6.8	0.39	1.06	0.12
Lum11	19.3	2.3	1.32	1.78	0.07
	17.0	7.3	0.51	1.15	0.07
Lum12	19.3	2.3	1.09	1.50	0.14
	17.0	6.8	0.31	0.73	0.11
LUMINEU LD	19.3	2.3	1.3	1.8	0.08
	17.0	6.9	0.4	0.9	0.10
CUPID-0 LD EPJC 76(2016)364	20.0	n.a.	0.8	1.3	0.10

Good reproducibility of high performance of NTD-instrumented LDs

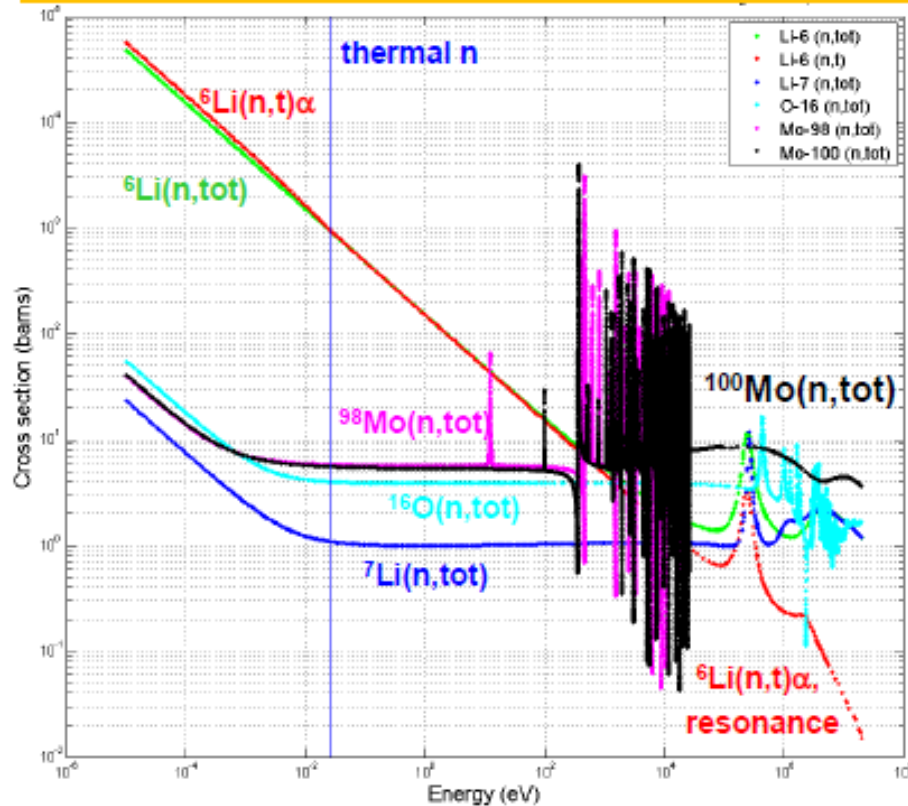
Neutron spectroscopy with $\text{Li}_2^{100}\text{MoO}_4$ bolometers

$\text{Li}_2^{100}\text{MoO}_4 \Rightarrow 7.6\%$ of ^6Li

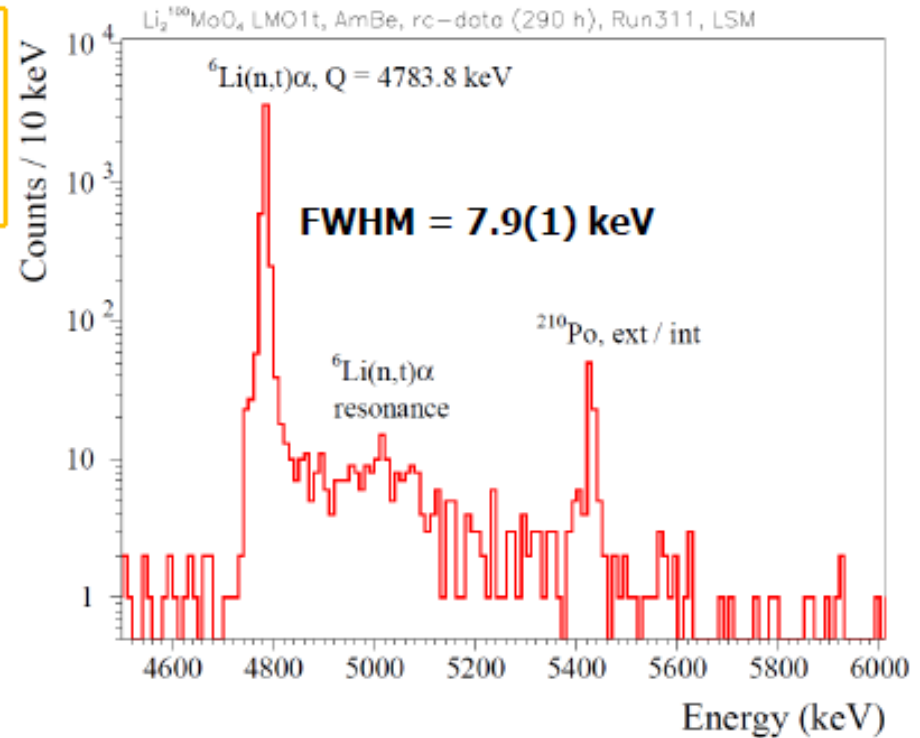
$^6\text{Li} + n \rightarrow t + \alpha$ ($Q = 4783$ keV)

$E(t+\alpha) = 4783$ keV (thermal n, ~ 25 meV)

$E(t+\alpha) = 5022$ keV (resonance @ ~ 240 keV)



enrLMO-1, AmBe (290 h)

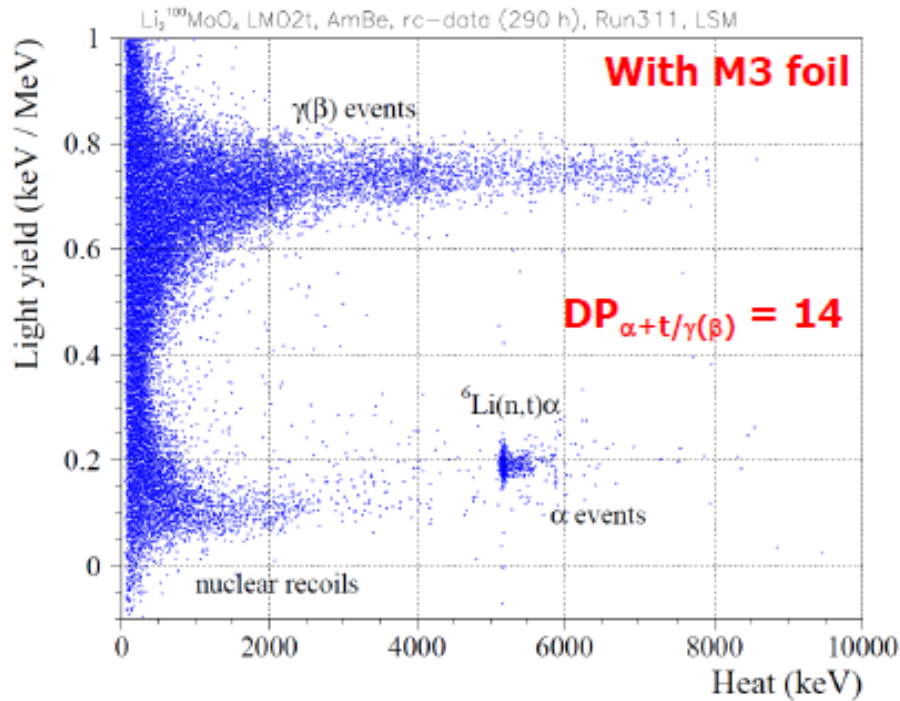


Advantages

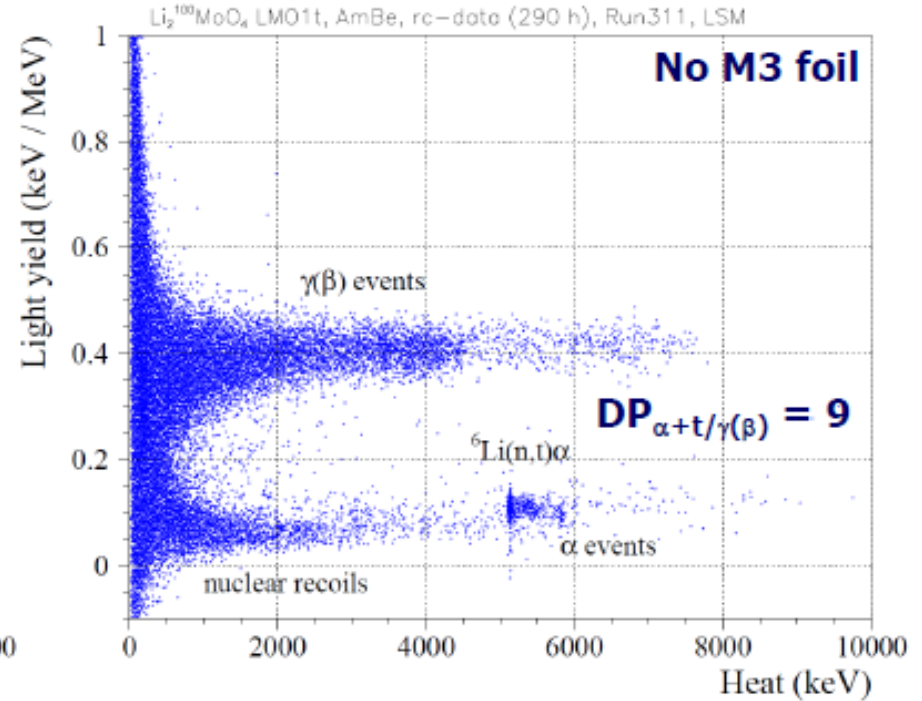
- ✓ $\sim 100\%$ detection of thermal n
- ✓ clear $\alpha+t$ signature @ $Q+E_n$
- ✓ $\gamma(\beta)$ background-free ROI
- ✓ world record resolution of thermal n capture on ^6Li (6-11 keV FWHM @ 4783 keV)

Prospects for in-situ neutron detection

enrLMO-3, AmBe (290 h)



enrLMO-1, AmBe (290 h)



Efficient $\alpha/\gamma(\beta)$ separation even for detectors operated without a reflecting foil

$$DP = \frac{|\mu_{\beta/\gamma} - \mu_{\alpha}|}{\sqrt{\sigma_{\beta/\gamma}^2 + \sigma_{\alpha}^2}}$$

CROSS #3 – Al film in LMO crystal – 10 μm thick

