

Status and prospects of CDEX the China Dark Matter Experiment

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(on behalf CDEX collaboration)

OUTLINE

- Dark Matter and CDEX Introduction
- CDEX-1 and CDEX-10
- CDEX next-stage plan
- Summary



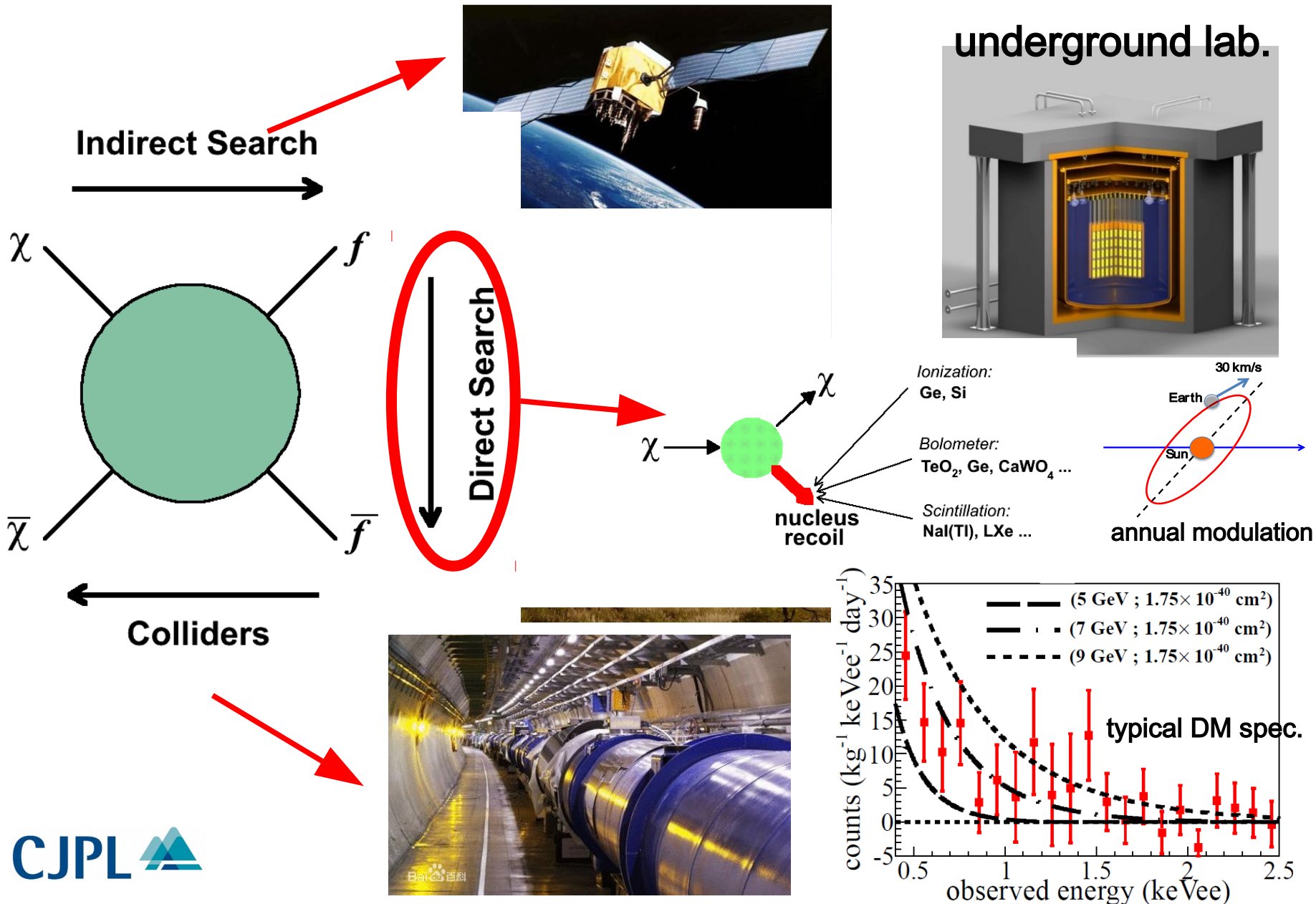
中国锦屏地下实验室
China Jinping Underground Laboratory

NDM 2018

6th Symposium on Neutrinos and Dark Matter in Nuclear Physics 2018

2018.6.30(Sat)

Dark Matter detection



CDEX: China Dark matter EXperiment



Established in 2009

- Tsinghua University (THU)
- Sichuan University (SCU)
- Nankai University (NKU)
- China Institute of Atomic Energy (CIAE)
- Beijing Normal University (BNU)
- Yalong River Company

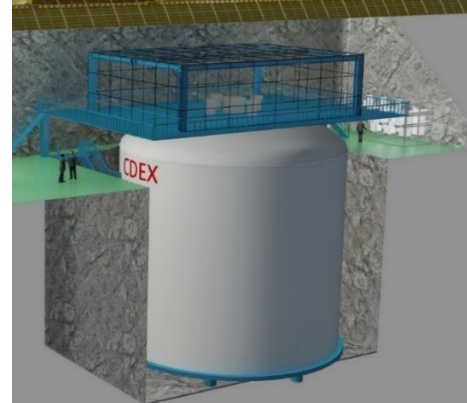
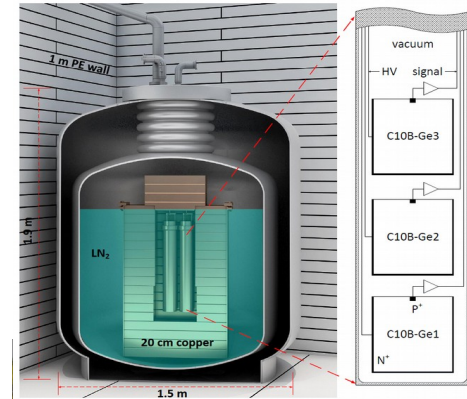
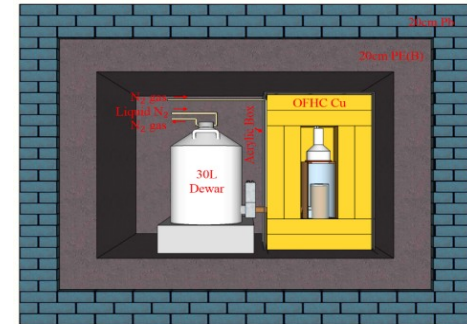
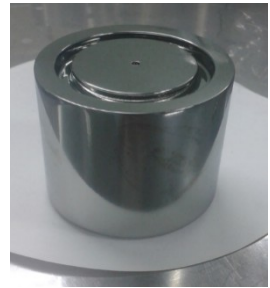
+

- Academia Sinica, Taiwan
- Banaras Hindu University, India
- Dokuz Eylül University, Turkey

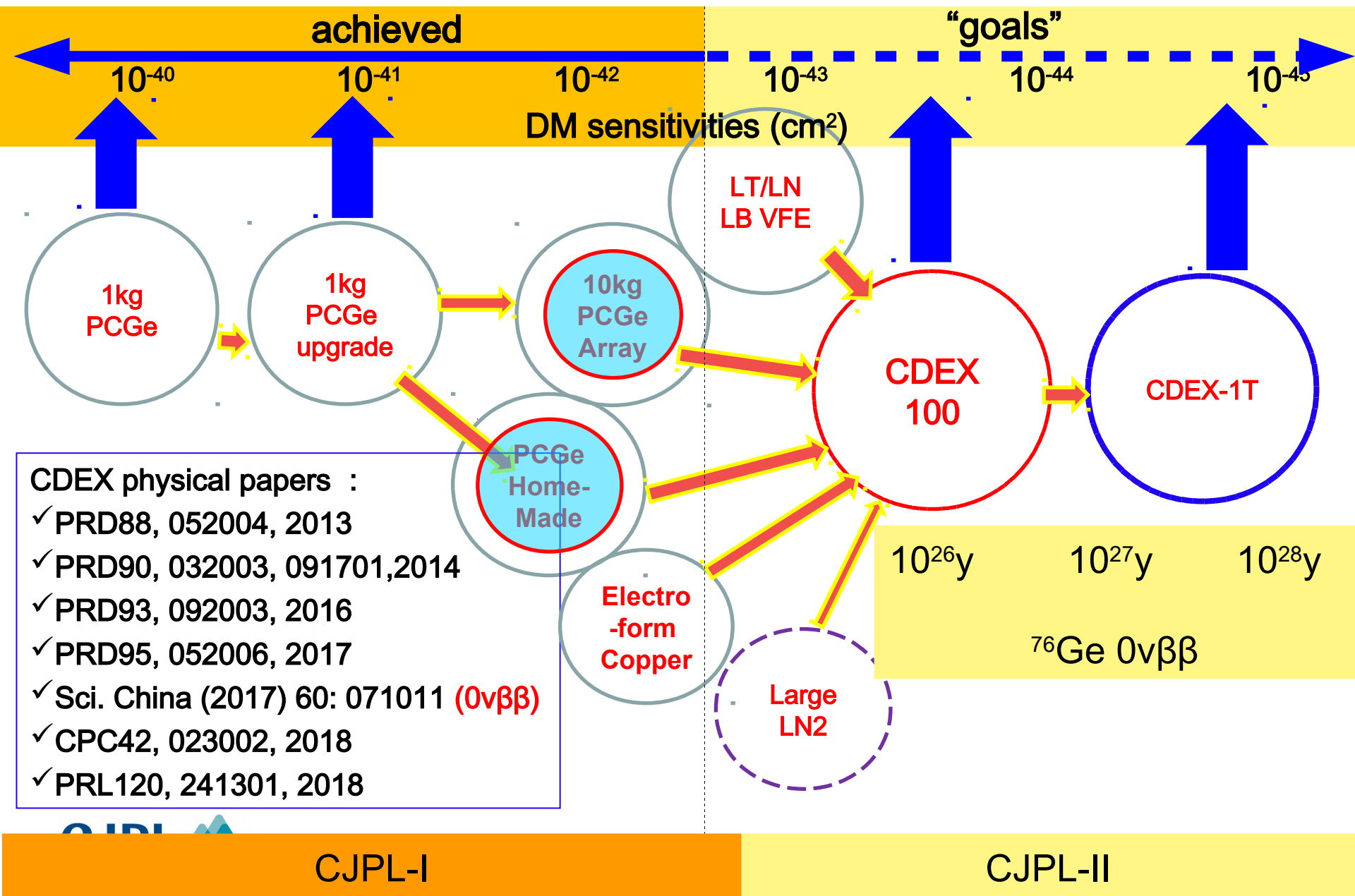
(as members of TEXONO collaboration)

CDEX stages

- Light WIMP mass searches on Ge
- CDEX-1: Development of pPC-HPGe detector, its background understanding, results published.
- CDEX-10: Performances of HPGe array detector system, results published.
- CDEX-10X: Fabrication of HPGe and Ge crystal growth by CDEX.
- CDEX-100: Ultra-low cosmogenic background and large LN_2 cooling and shielding system.
- CDEX-1T: Multi-purpose experiment for dark matter and double beta decay.



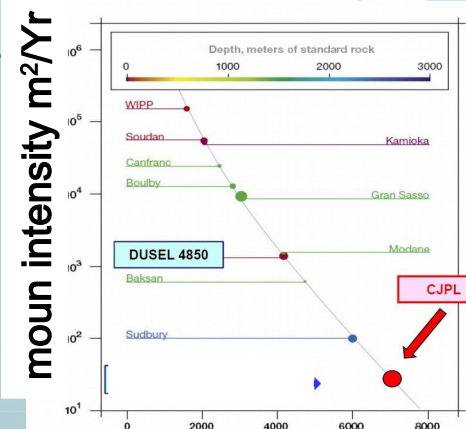
CDEX plan for DM & $0\nu\beta\beta$



Jinping Hydroelectric Power Plants

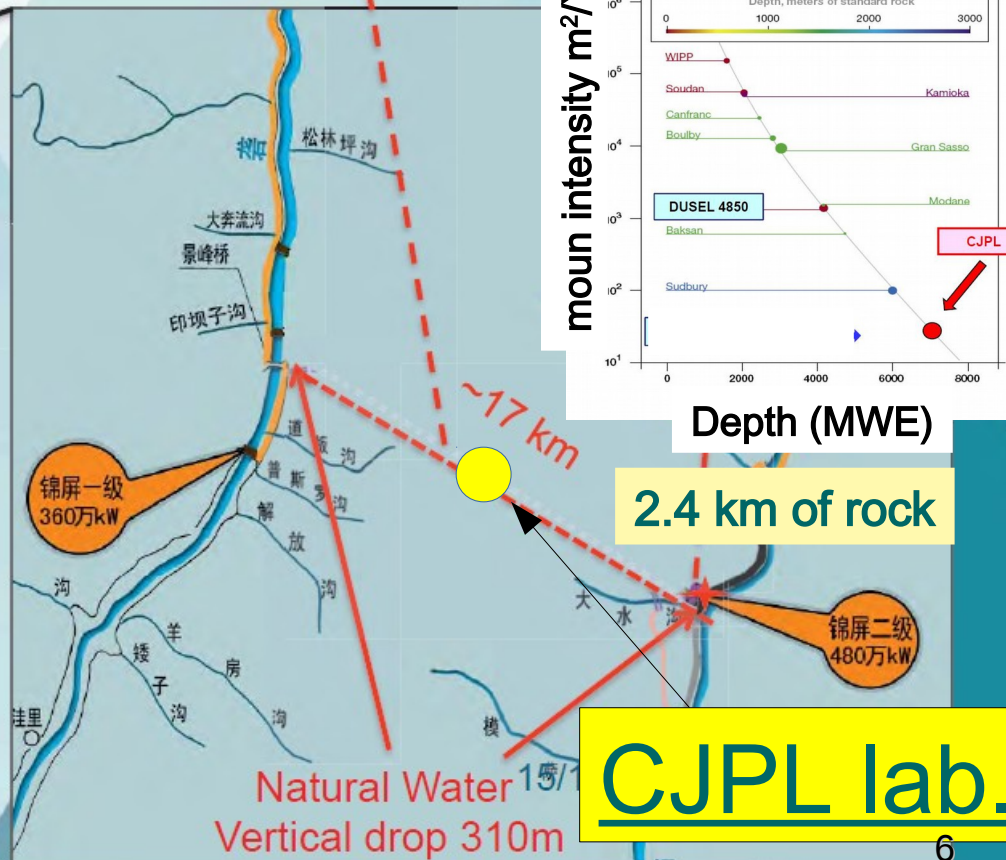
4 hydraulic
tunnels
 $\Phi 13\text{m} \times 16.6\text{km}$

Jinping-II
Power Plant
4800MW
(8×600MW)



Depth (MWE)

2.4 km of rock



Natural Water
Vertical drop 310m

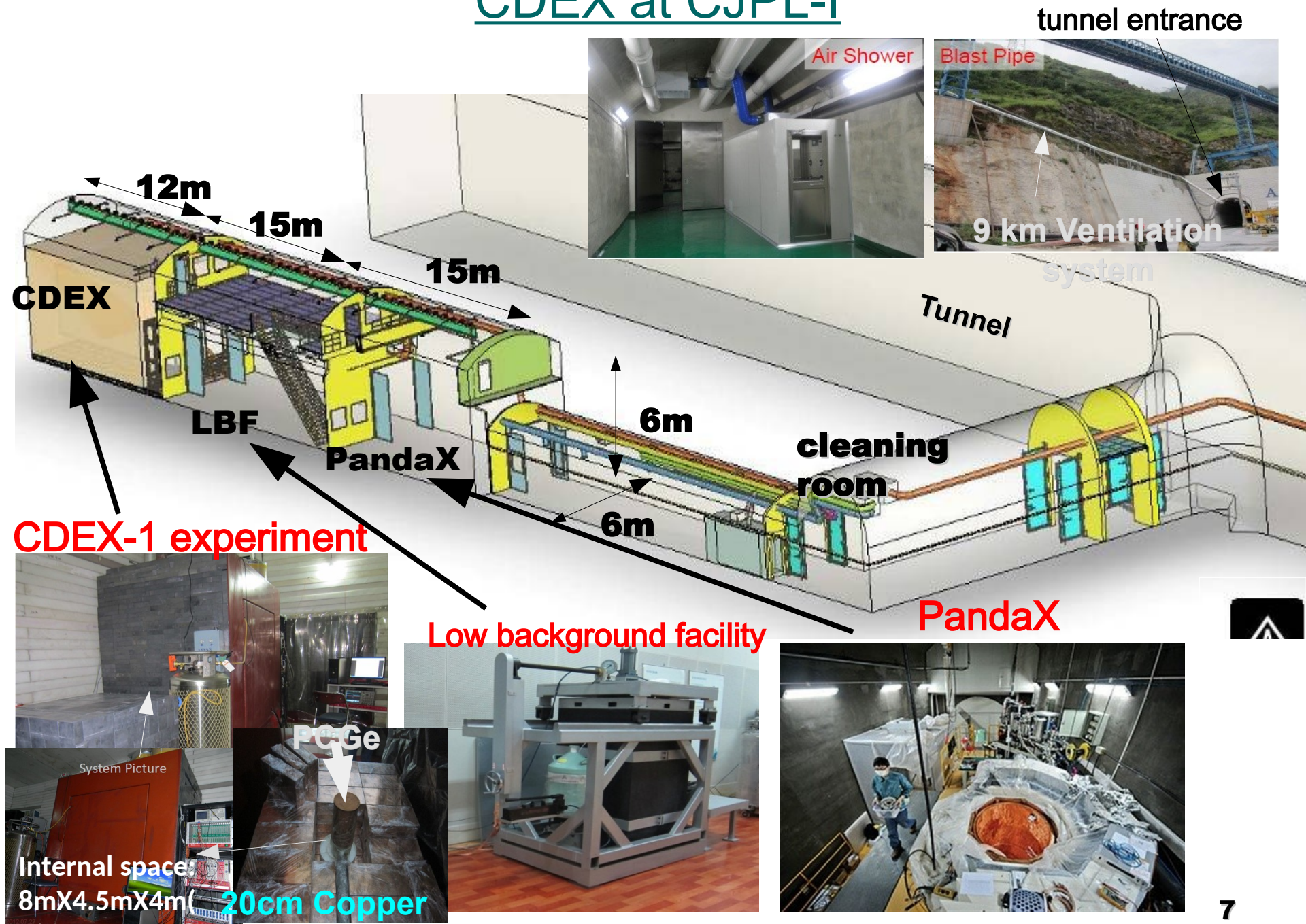
CJPL lab.

Jinping-I
Power Plant
3600MW
(6×600MW)

锦屏一级
360万kW

锦屏二级
480万kW

CDEX at CJPL-I



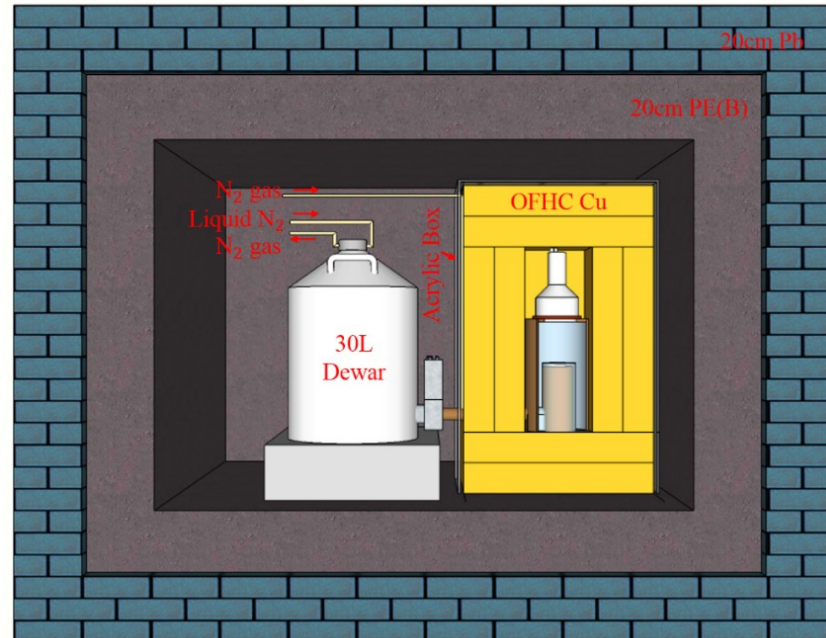
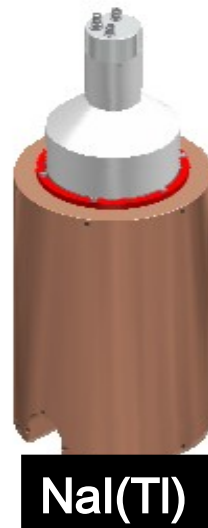
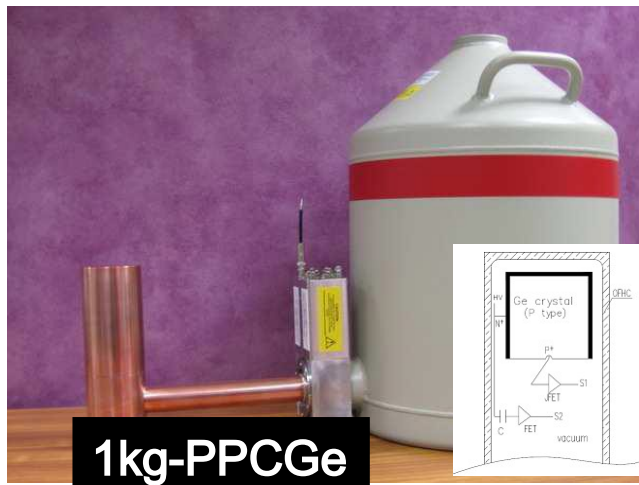
CDEX-1 experiment

1. HPGe technology

- ✓ Designed the first one single module 1kg-scale p-type point-contact Ge detector (1kg-PPCGe) “prototype” CDEX-1A
- ✓ Improved 1kg-PPCGe CDEX-1B

2. Active shielding technology: NaI(Tl) used as anti-Compton detector

- ✓ CDEX-1A 1kg-PPCGe run
- ✓ CDEX-1A 1kg-PPCGe + NaI(Tl) run
- ✓ CDEX-1B 1kg-PPCGe + NaI(Tl) run



CDEX-1

CDEX-1A 1kg PCGe



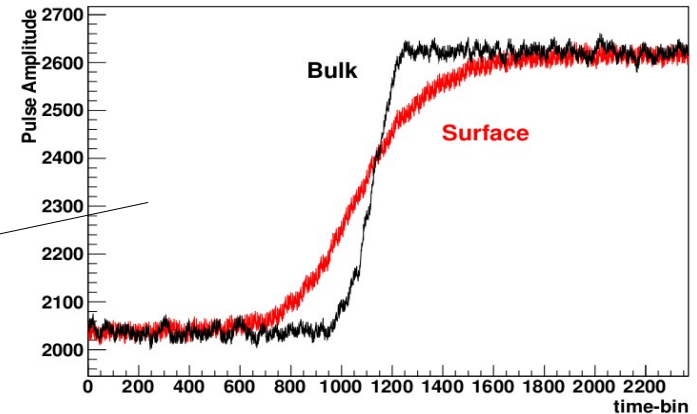
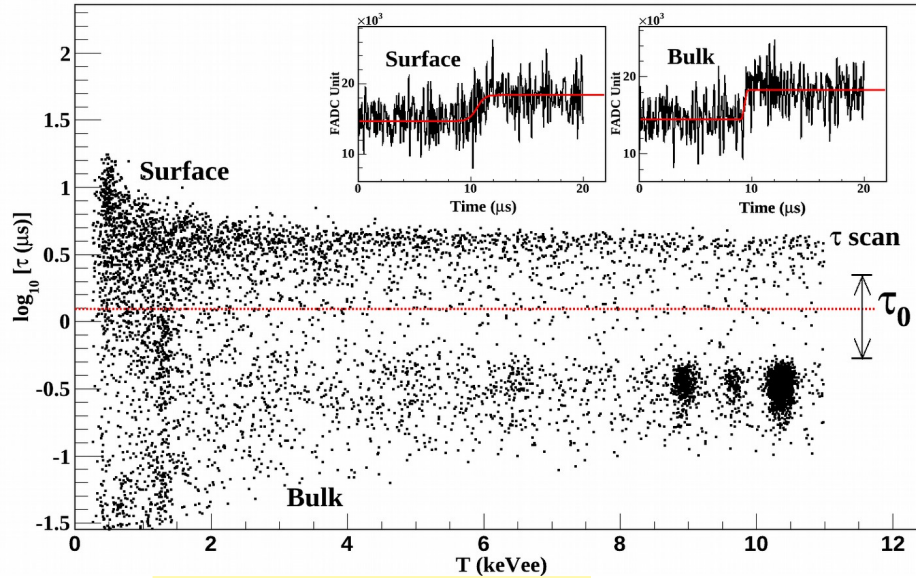
CDEX-1B 1kg PCGe



20cm OFHC Copper
+20cm Lead

pPCGe: bulk/surface

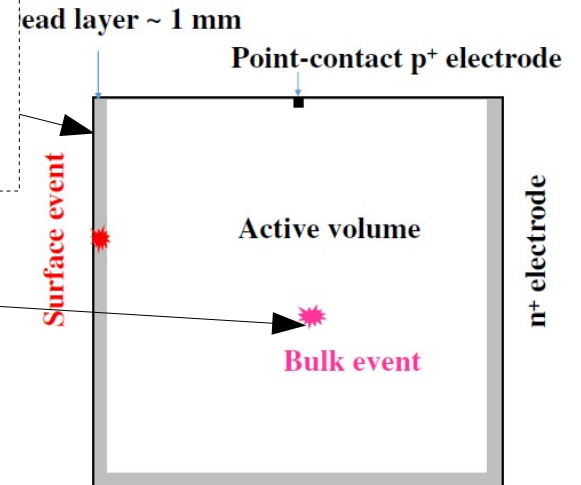
surface events : contamination, largest sources of uncertainties at low energy.



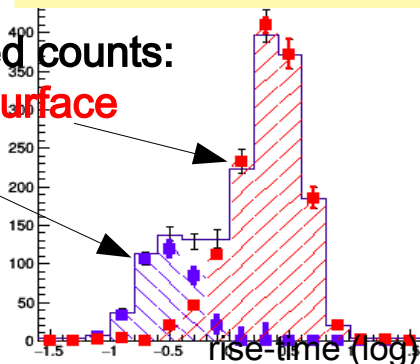
DM candidate events:
0.4-0.6 keV

surface
partial energy deposit
diffuse and drift \rightarrow slow

bulk
full energy deposit
drift \rightarrow fast



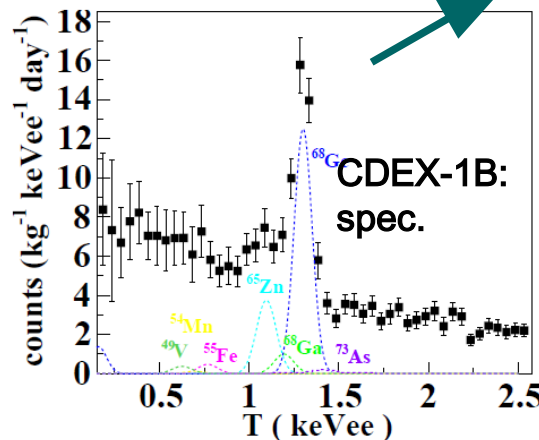
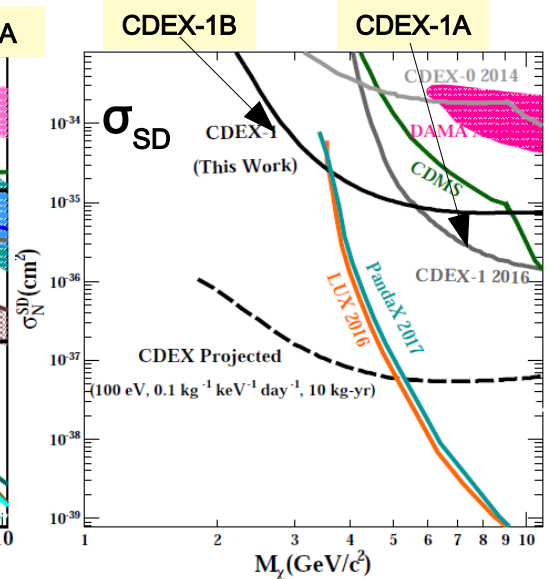
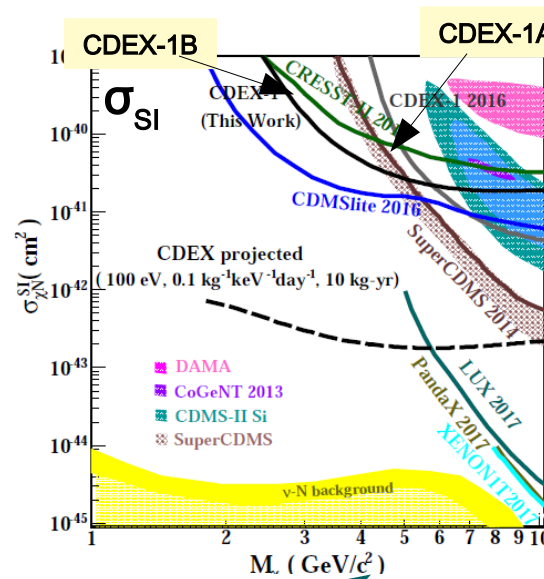
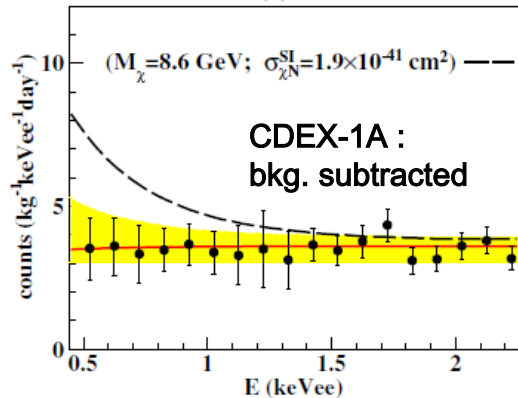
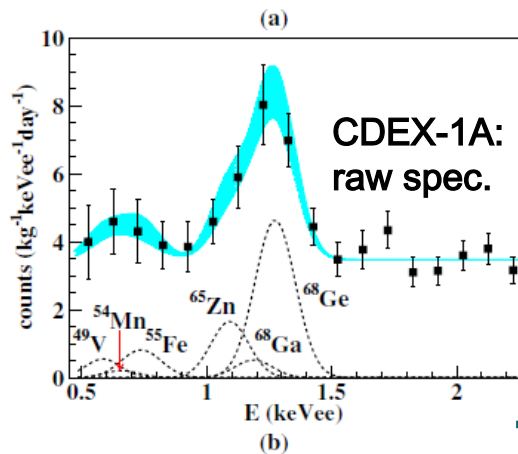
constructed counts:
bulk and surface



CDEX-1 results

Data set: CDEX-1A: ~ 500 kg-day; CDEX-1B: ~ 1000 kg-day

- Competitive SI/SD sensitivities pPCGe [PRD93, 092003, 2016](#), [CPC42, 023002, 2018](#)
- Competitive galactical Axion sensitivity below 1 keV [PRD95, 052006, 2017](#)
- threshold: CDEX-1A: 475 eV, CDEX-1B: 160 eV.

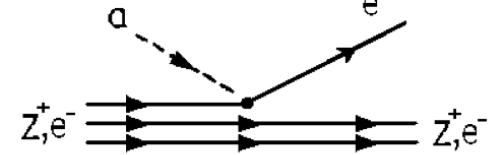


CDEX-1B:
threshold=160eV, ~ 2 yrs,
annual modulation results soon

CDEX-1 axion results

M1 transition from ^{57}Fe from Sun: $^{57}\text{Fe}^* \rightarrow ^{57}\text{Fe} + a$ [g_{AN}]

PRD95, 052006, 2017



Axioelectric
or Photoelectric-like

axion(a) from sun [g_{Ae}]

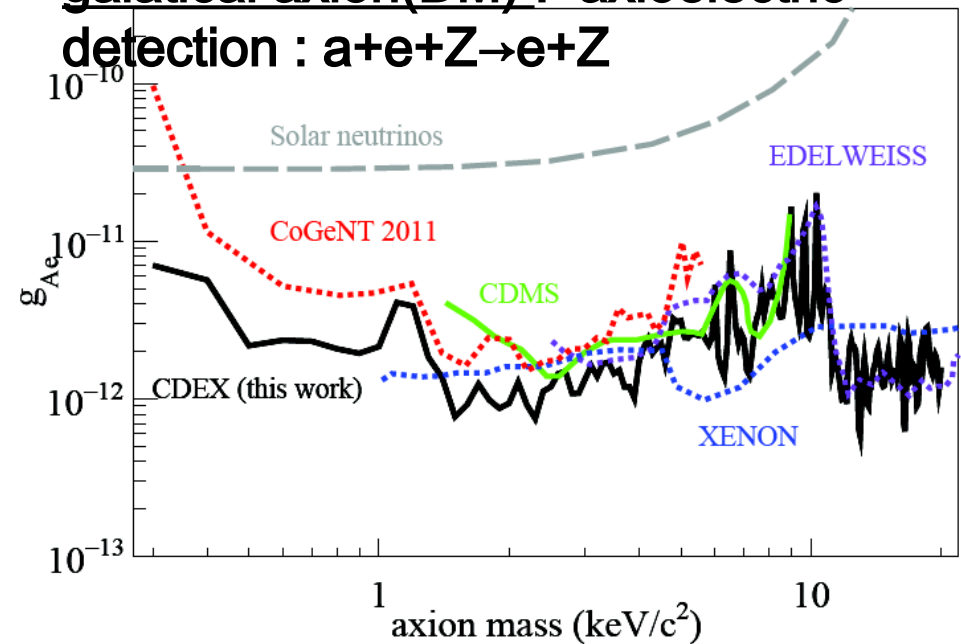
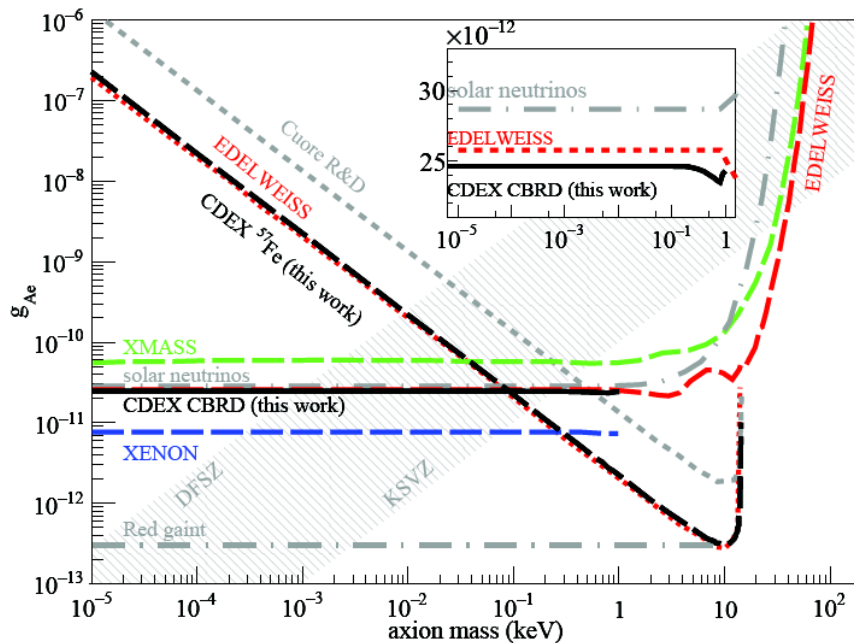
Compton(C): $\gamma + e \rightarrow e + a$

bremsstrahlung(B): $e + Q \rightarrow e + Q + a$

recombination(R): $e + I \rightarrow I + a$

de-excitation(D): $I^* \rightarrow I + a$

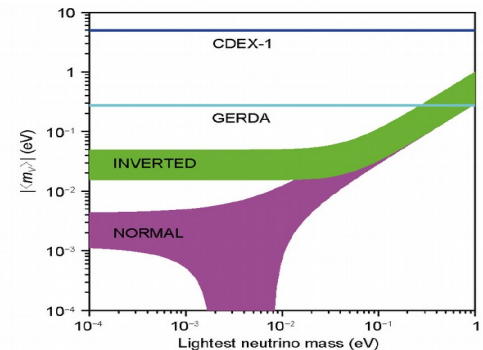
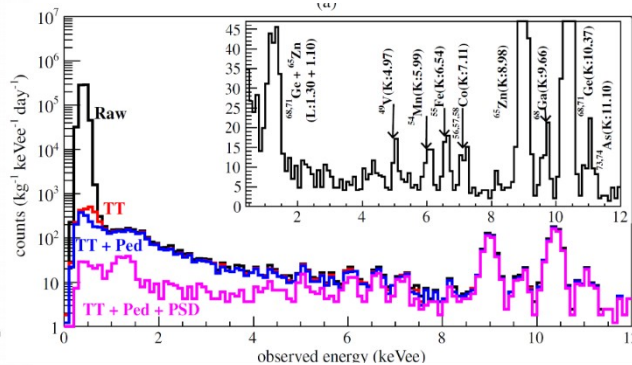
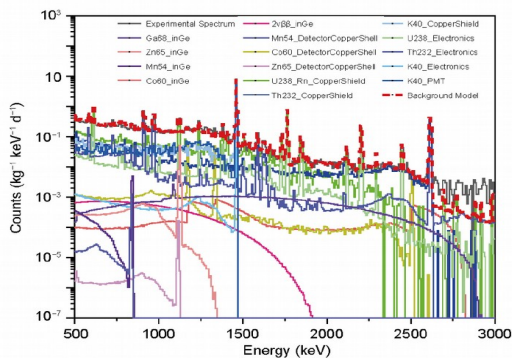
galactical axion(DM) : axioelectric
detection : $a + e + Z \rightarrow e + Z$



**Competitive results for DM axion
below the axion mass of 1 keV.**

CDEX-1 $0\nu\beta\beta$ result

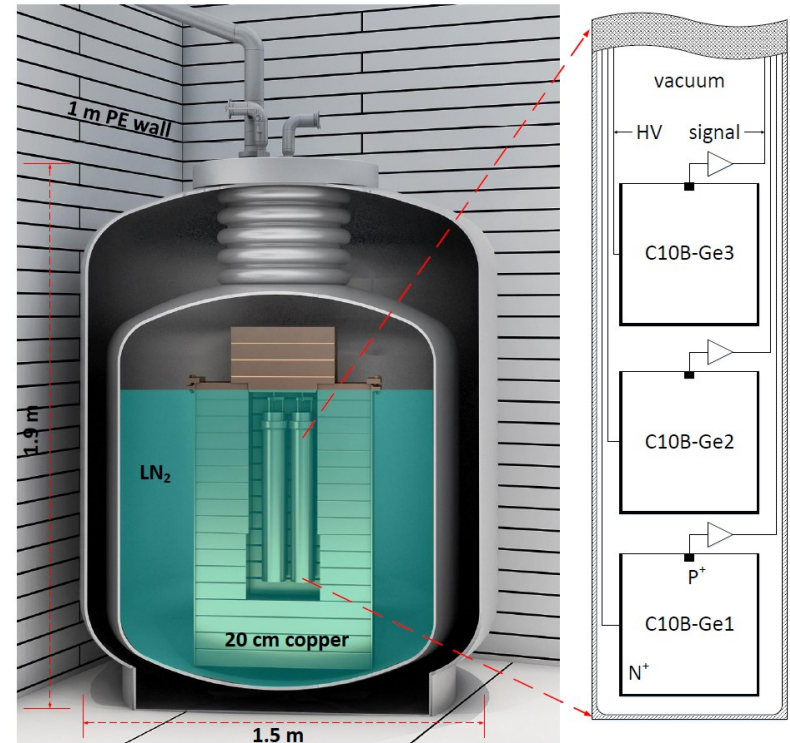
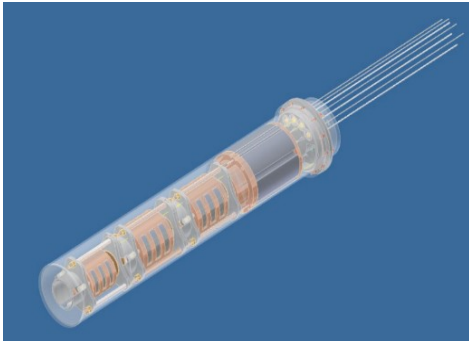
- Published ^{76}Ge $0\nu\beta\beta$ result based on CDEX-1A (natural Ge).
- Calculation of the level of cosmogenic events @ 2 MeV based on cosmogenic characteristic X-ray peaks $<10\text{keV}$.
- $\tau_{1/2} > 6.4 \times 10^{22} \text{ yrs}$



Sci. China (2017) 60: 071011

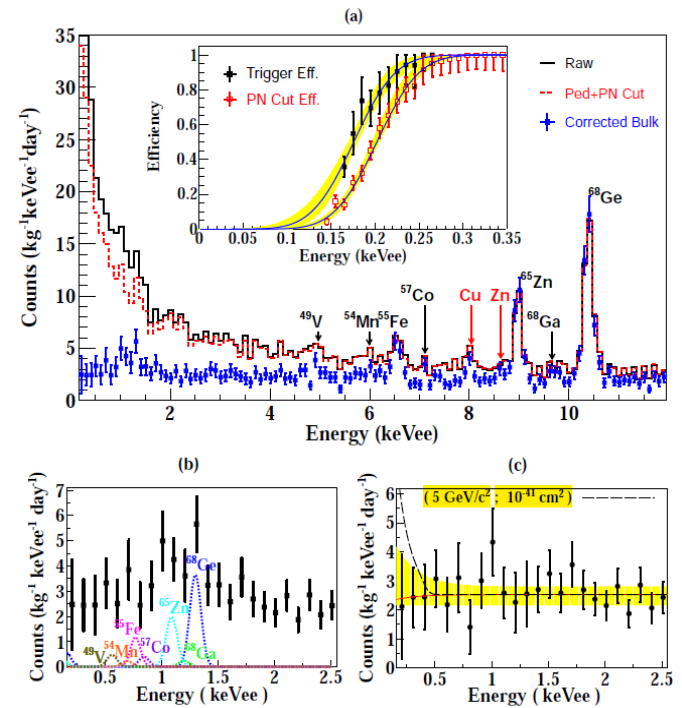
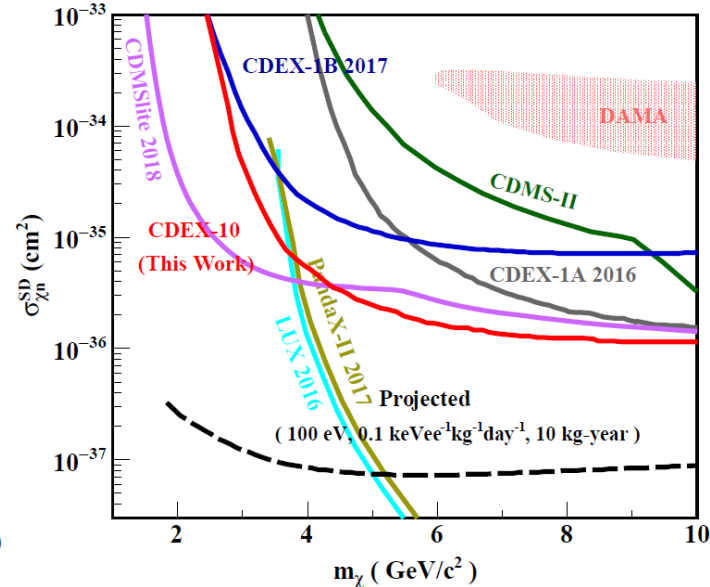
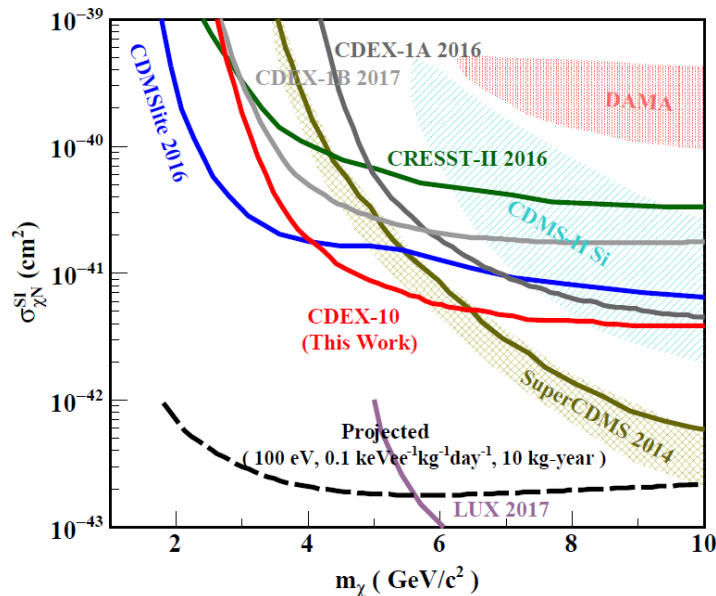
CDEX-10 experiment

- The important stage towards large-scale Ge experiment.
- Directly immersed into liquid nitrogen for cooling: operate at few K lower than “cold-finger” Ge.
- Dataset: 102.8 kg-day.



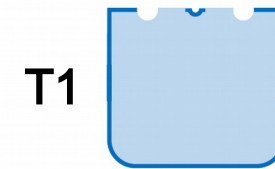
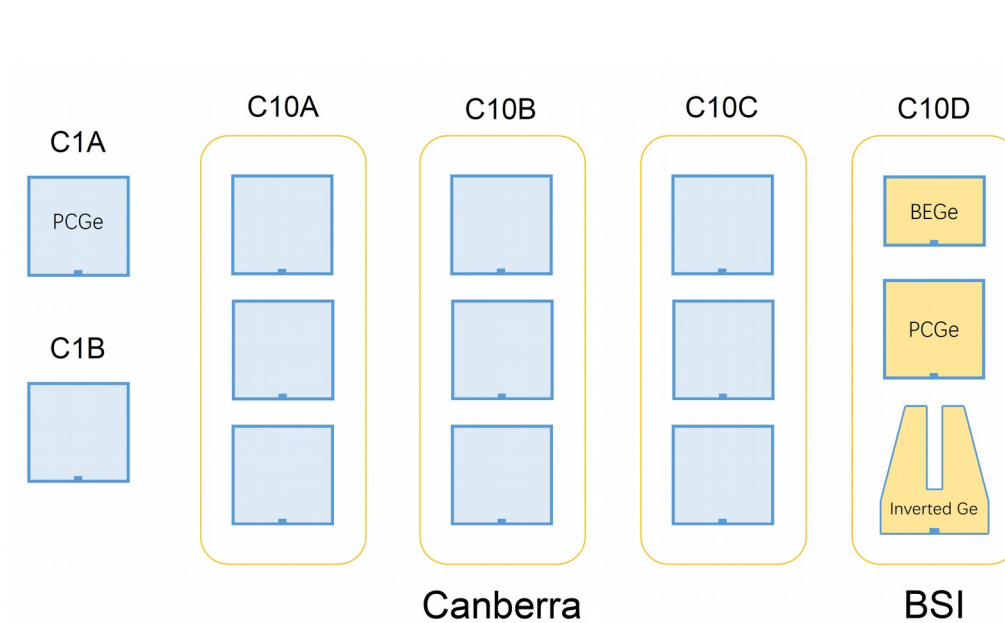
CDEX-10 results

- ✓ threshold: 160eV
- ✓ The competitive SI/SD results at light WIMP mass



CDEX-10 pPCGe detectors

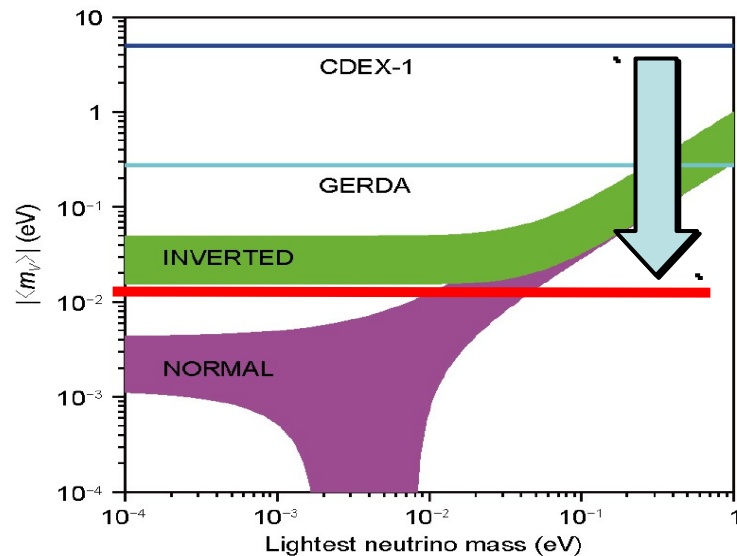
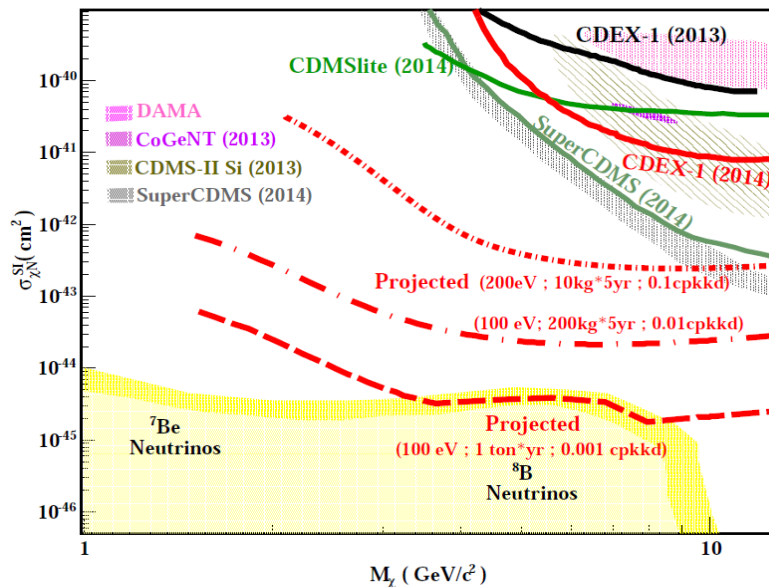
- Collaborate with commercial companies, such as Canberra, ORTEC, BSI and so on;
- Home-made pPCGe for the lower cosmogenic backgrounds prepared at CJPL now.



- Commercial Ge crystal;
- **Structure machining;**
- **Li-drift and B-implanted;**
- **Home-made ULB PreAmp;**
- **Underground EF-Cu;**
- **Underground assembly;**
- **Underground testing...**

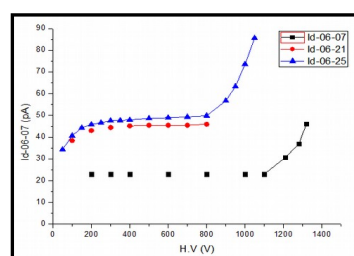
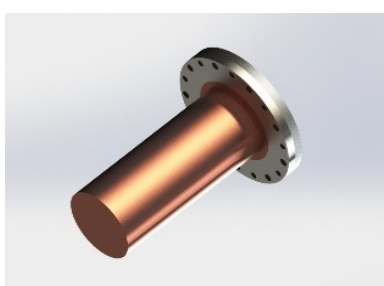
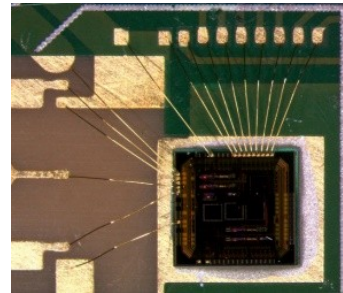
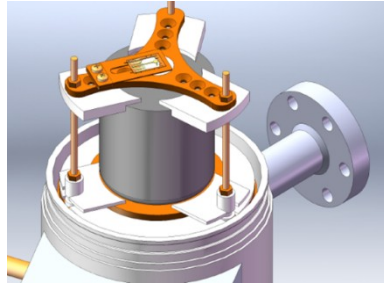
Toward CDEX-1T@CJPL-II

- A future Ge detector composed of the PCGe detector array and LN shielding and cooling system in the CJPL-II.
- Both Dark matter and Double Beta Decay (CDEX groups are part of LEGEND programs)



Key technologies towards CDEX-1T

- Ge purification and crystal growth;
- HPGe detector fabrication;
- Ultra-low background VFE;
- Ultra-pure Cu for structure and cables;
- Large-volume cooling tank.



CDEX-1T Ge crystal growth



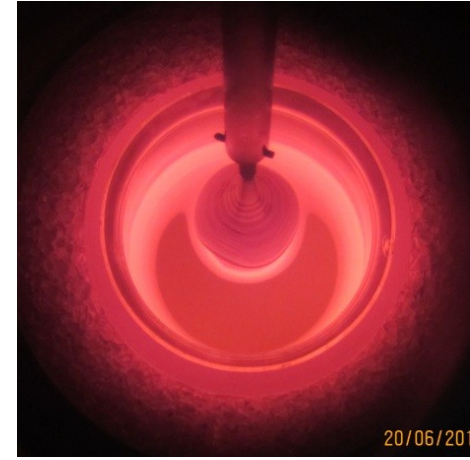
**Zone refining
machine**



**Czochralski
machine**



**Cutting &
Polishing**



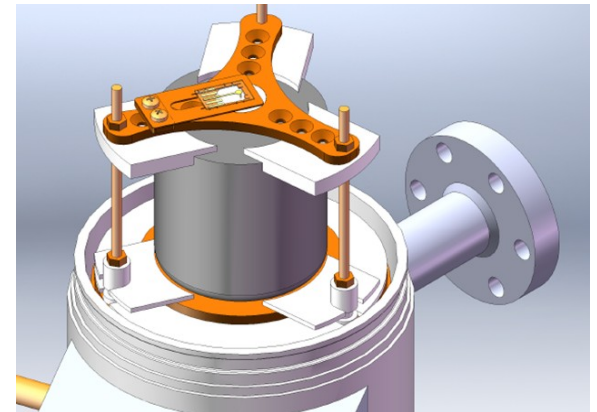
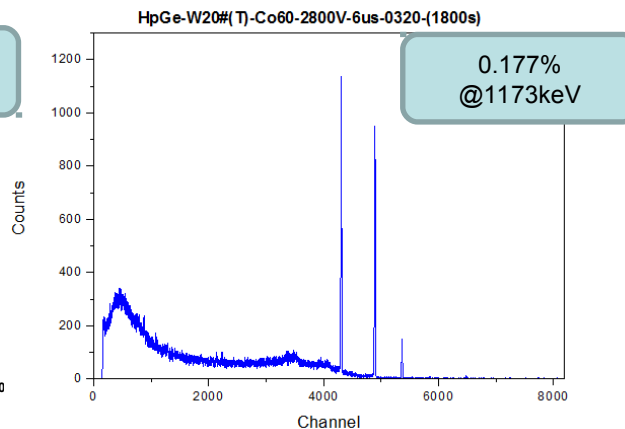
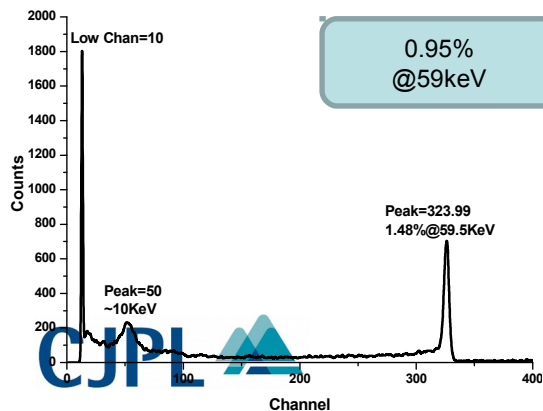
Crystal growth



- ✓ The requirement for making P-type Ge detector
 - ✓ Impurity density: $\sim 10^{10} \text{ cm}^{-3}$
 - ✓ Dislocation: $< 5000 \text{ cm}^{-2}$
- ✓ CDEX are working on this two points.

HPGe detector fabrication

- First 500g home-made pPCGe+ASIC finished testing, energy resolution and energy threshold compared with commercial one.



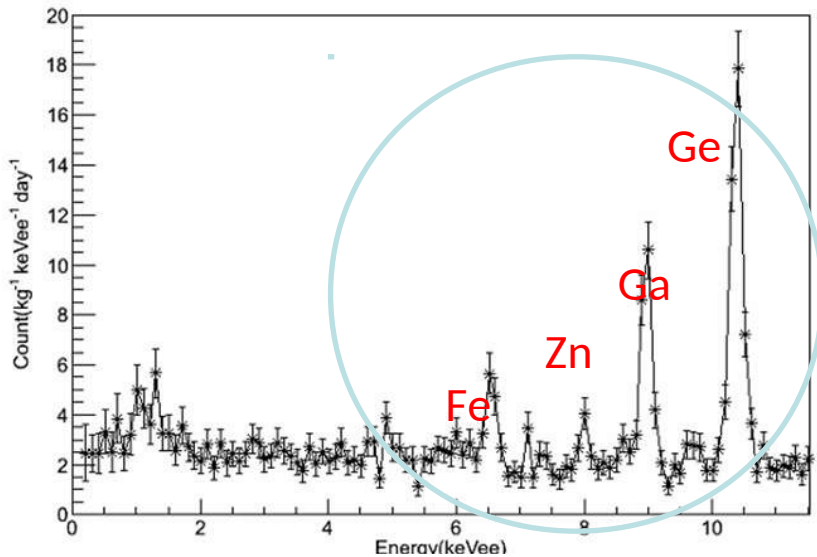
CDEX ULB-Cu @ CJPL

- Setting up the facilities for ULB-Cu production;
- CDEX copper goal will be the Majorana EFCu purification:
 $\text{Th} < 0.06 \mu\text{Bq/kg}$ 、 $\text{U} < 0.17 \mu\text{Bq/kg}$.
- Shielded by LN2, Structure materials used as little as possible in order to lower the background contribution.

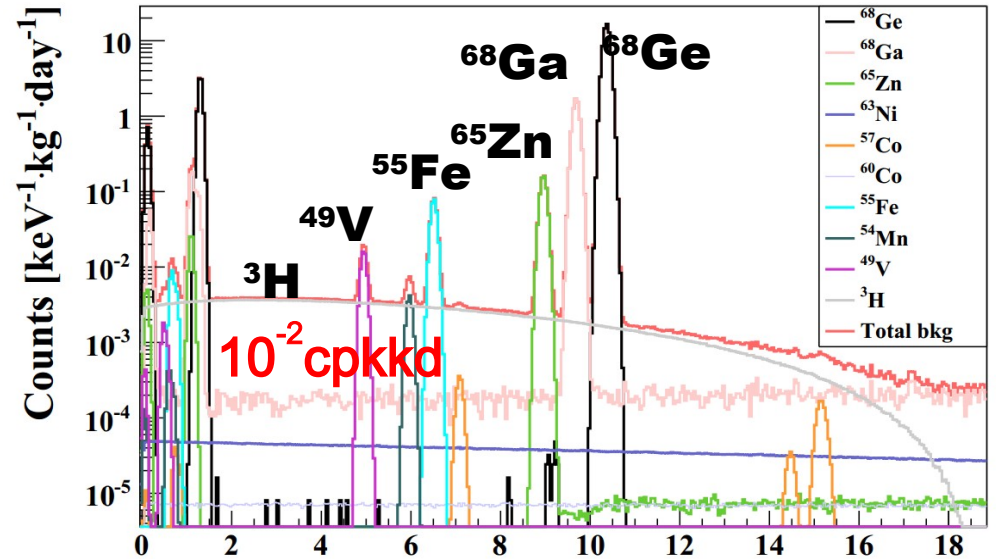


Cosmogenic Background of Ge detector

- Long-time ground preparation of detector induces high cosmogenic background level ($\sim 2 \text{ cpkkd @ } 2\text{-}4 \text{ keV}$);
- Based on simulation, **2 months ground fabrication and transportation** could decrease the ^3H continuous background level to **$\sim 10^{-2} \text{ cpkkd @ } 2\text{-}4 \text{ keV}$** .



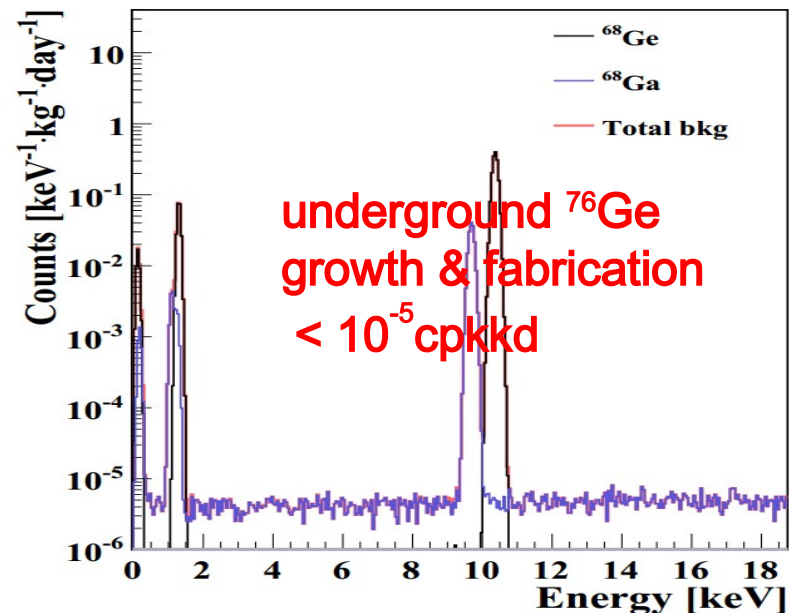
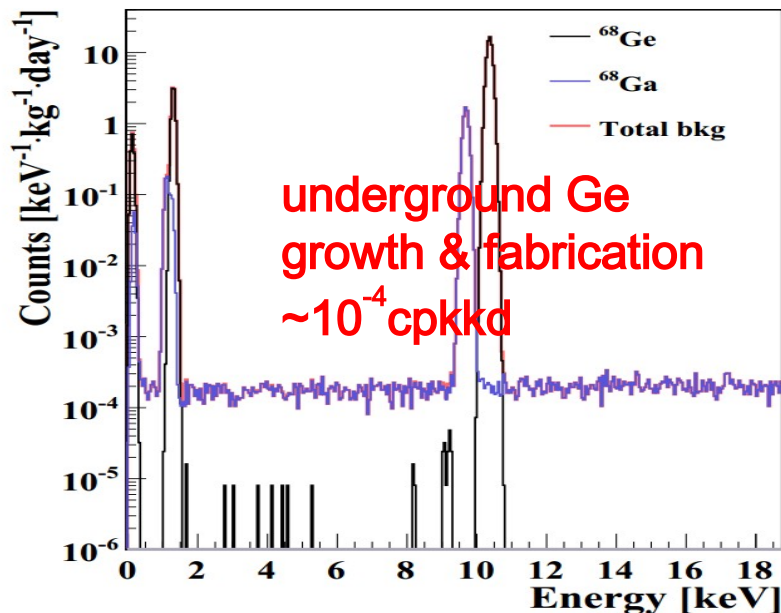
CDEX-10 background spectrum



Simulation spectrum

Cosmogenic Background of U-Ge detector

- Underground germanium crystal growth and detector fabrication could dramatically decrease the cosmogenic backgrounds from non-Ge isotopes, such as ^3He , ^{65}Zn ;
- ^{76}Ge Enriched germanium material could help to decrease $^{68}\text{Ge}(^{68}\text{Ga})$ cosmogenic backgrounds too.



The layout of CJPL-II

- 4 main halls : 14m(H)×14m(W)×130m(L);
- Total Volume: 300K m³;
- Two expanded spaces:

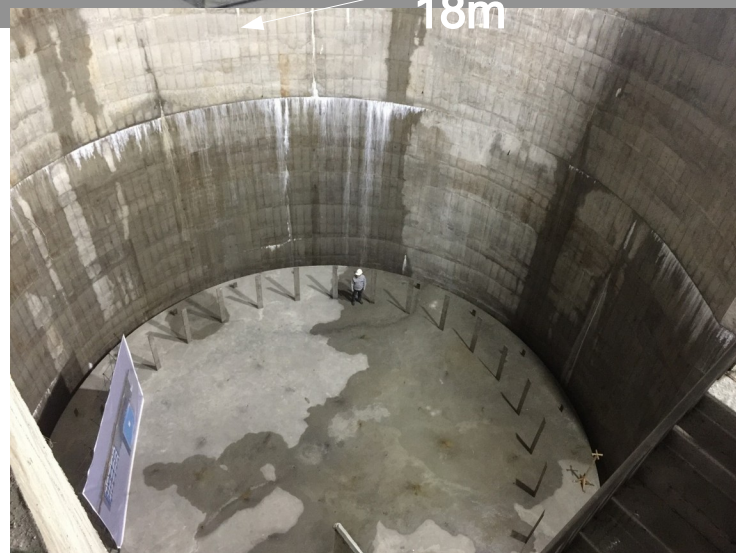
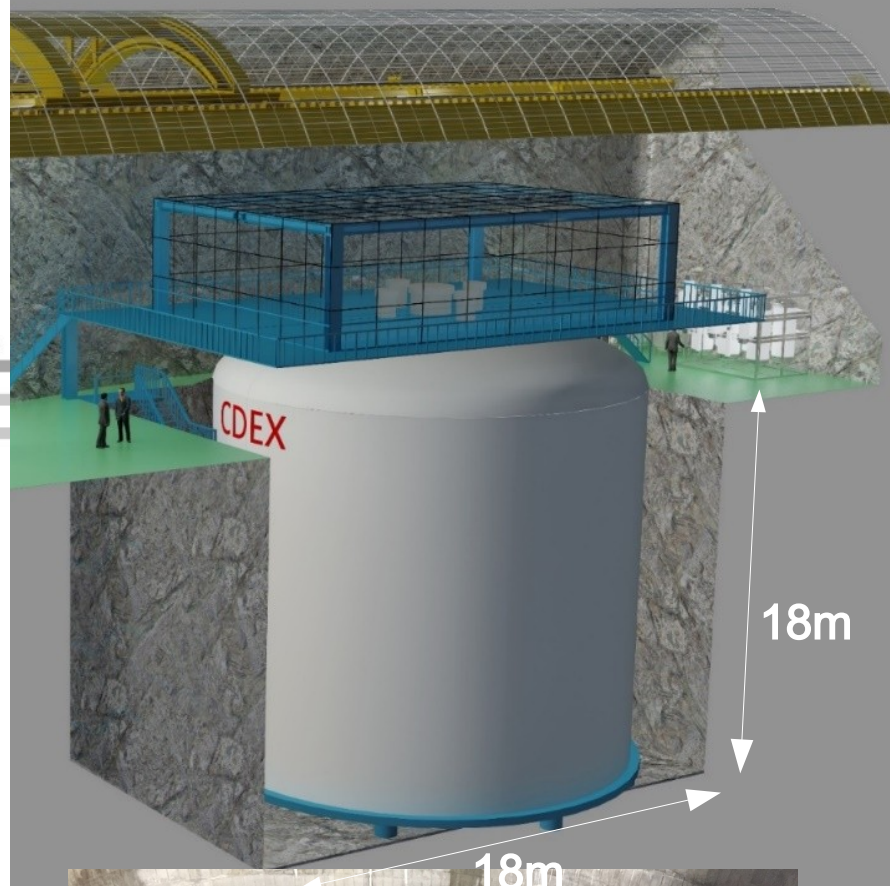
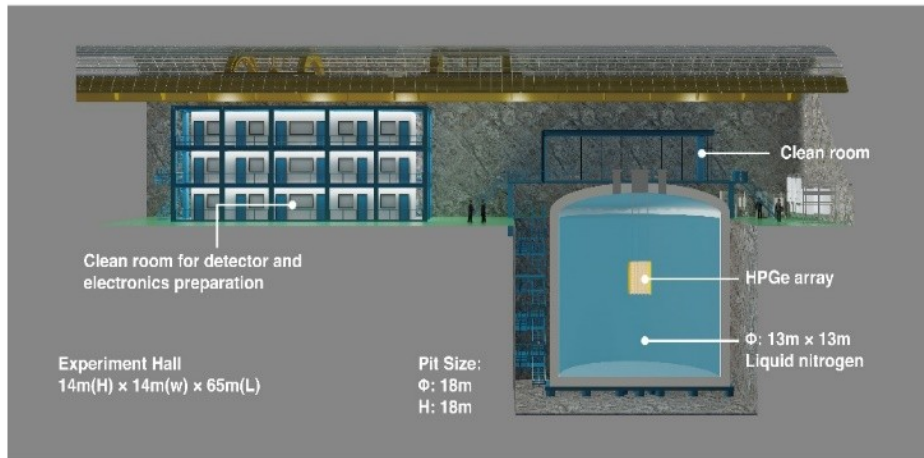
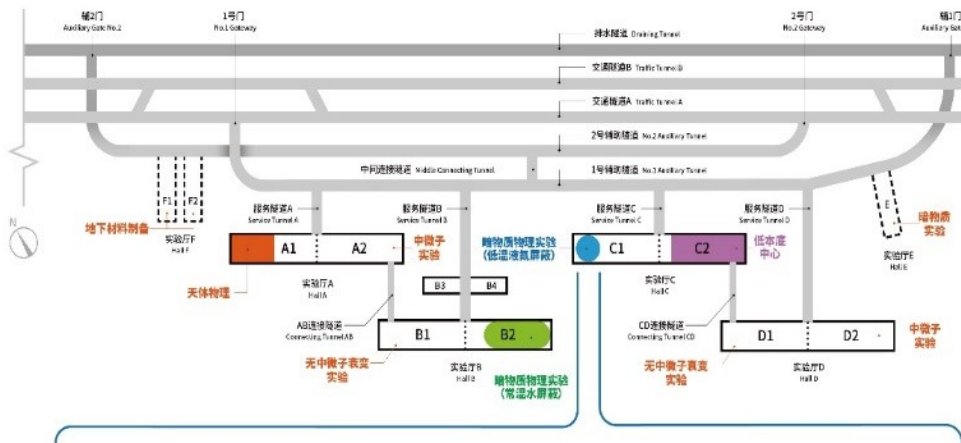
C1-- $\phi 18\text{m} \times 32\text{m(H)}$ →CDEX-1T

B2-- $27\text{m(L)} \times 14\text{m(W)} \times 30\text{m(H)}$

	CJPL-I	CJPL-II
Rock Work	4100 m ³	210000+151000m ³
Electric Power	70x2 kVA	10x2 MVA
Fresh Air	2400 m ³ /h	15000x3 m ³ /h



CDEX-1T@CJPL-II



Summary

- CDEX has achieved competitive DM physical results since 2013
- The first physical results from CDEX-10 published, competitive SI/SD sensitivities at light WIMP mass region.
- CDEX is developing the key technologies of low background Ge detector towards **CDEX-100** → **CDEX-1T** experiment for DM (+DBD+Solar Neutrino).



Thanks for your attention!