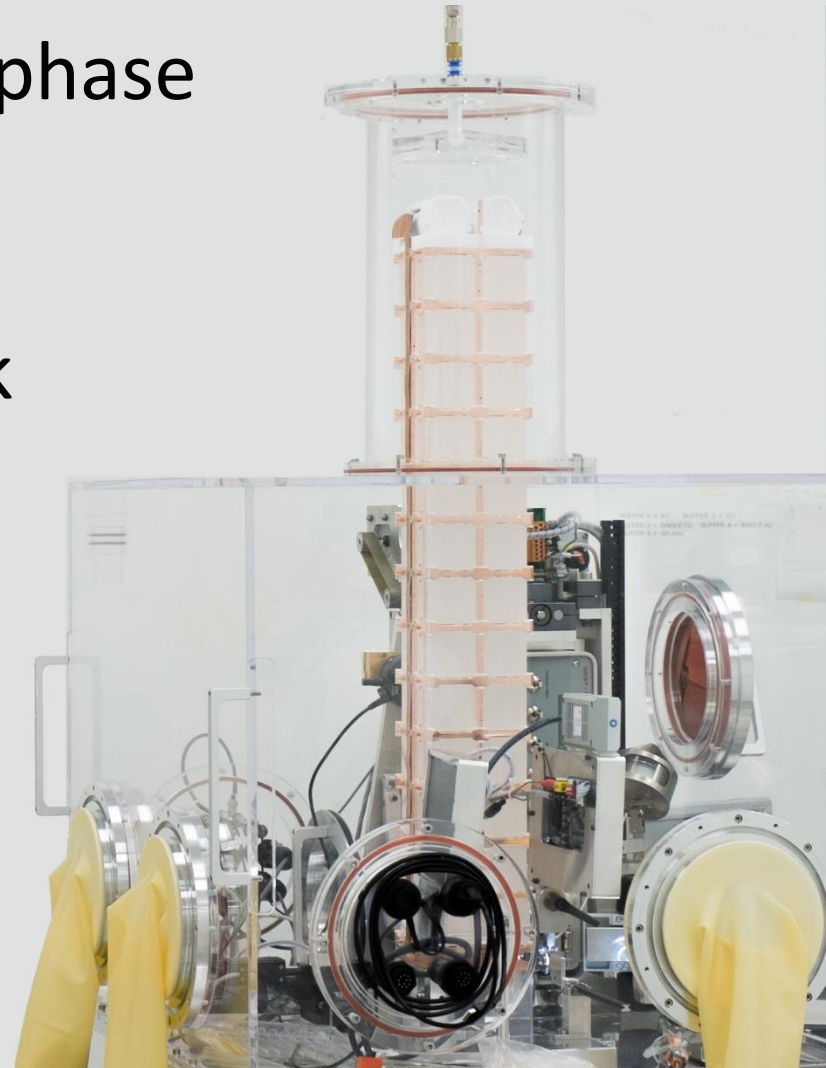


Latest Results from CUORE

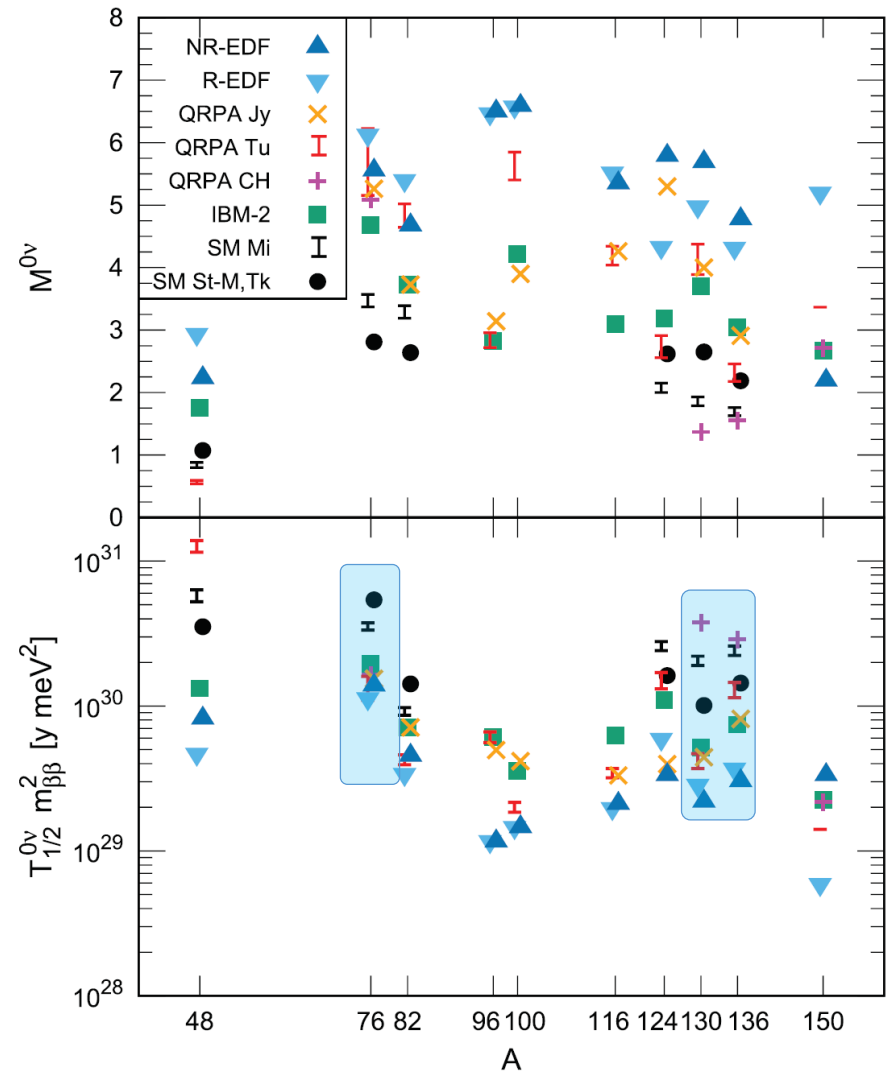
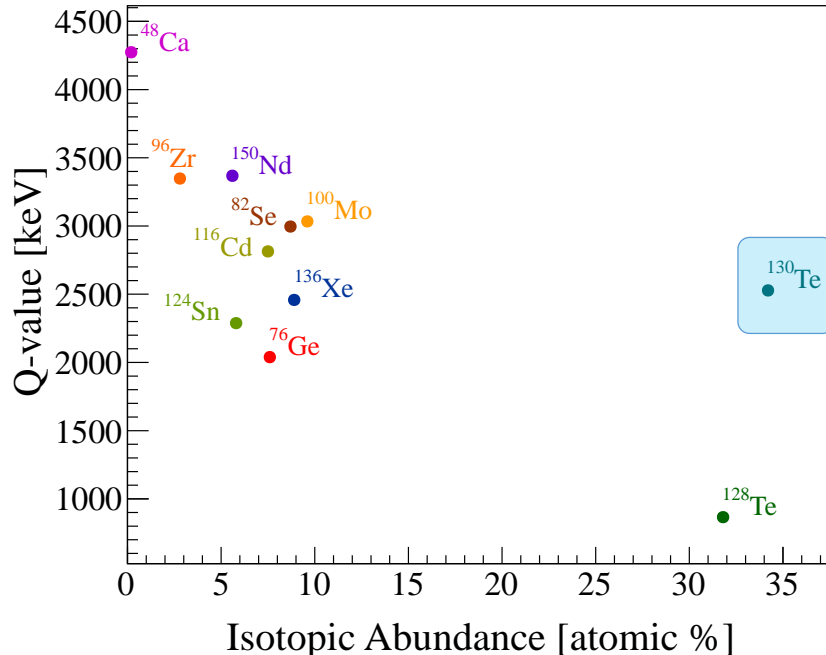
Ke Han
Shanghai Jiao Tong Univ.
June 29, 2018

- CUORE Project overview
- $0\nu\beta\beta$ result from the first phase
- Recent $2\nu\beta\beta$ result
- CUORE status and outlook



Double beta decay in experiments

- Extremely rare events $T_{1/2} > 10^{24}$ year.
- Experimentally try to search for a peak on the spectrum of total energy of electrons.
- ^{130}Te : isotopically abundant and favorable in terms of half-life requirement



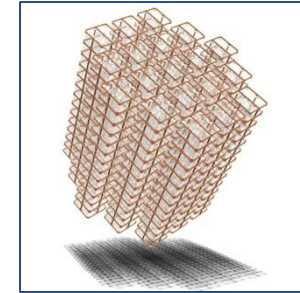
J. Engel and J. Menéndez 2017 Rep. Prog. Phys. 80 046301

Detection channels

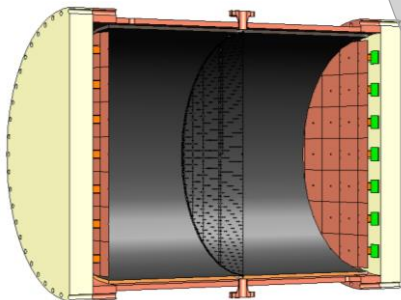
$$(T_{1/2}^{0\nu})^{-1} = F_N \frac{|\langle m_{\beta\beta} \rangle|^2}{m_e^2}$$

Phonons

CUORE



Electrons Holes



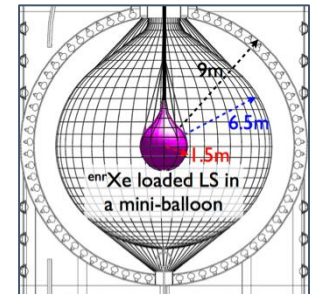
GERDA
Majorana
COBRA
PandaX-III

EXO
NEXT
SuperNEMO

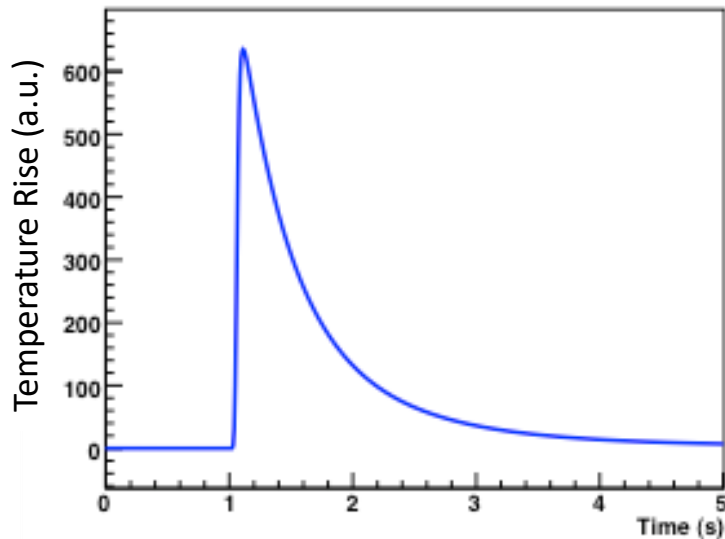
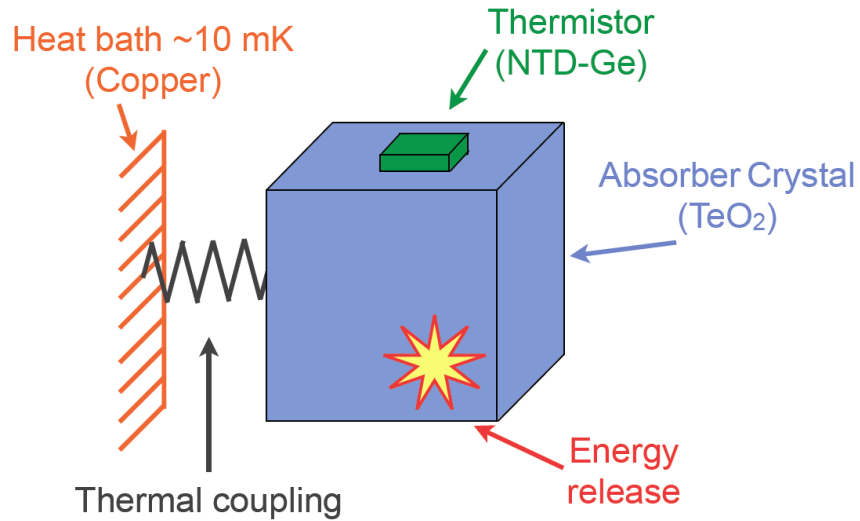
AMoRE
CUPID

KamLAND-Zen
SNO+
CANDLES

Photons



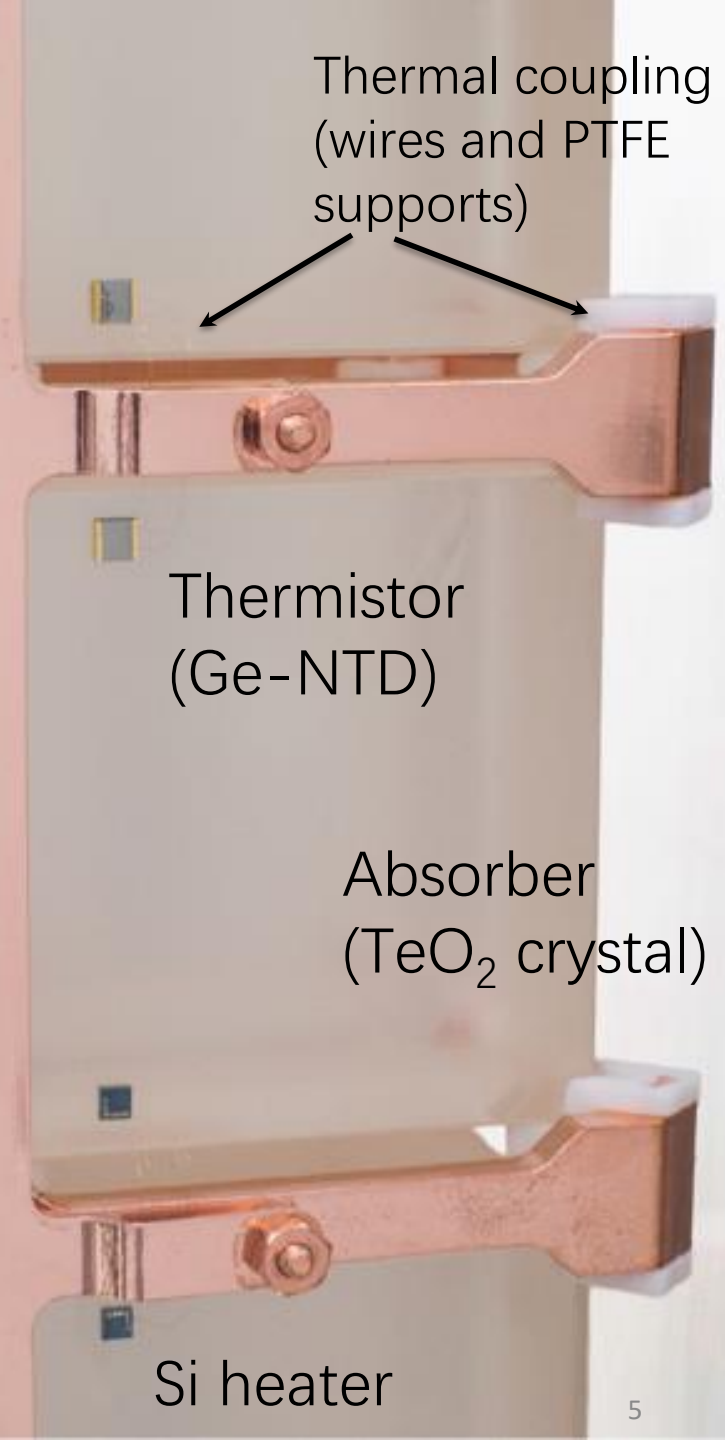
Large mass bolometer



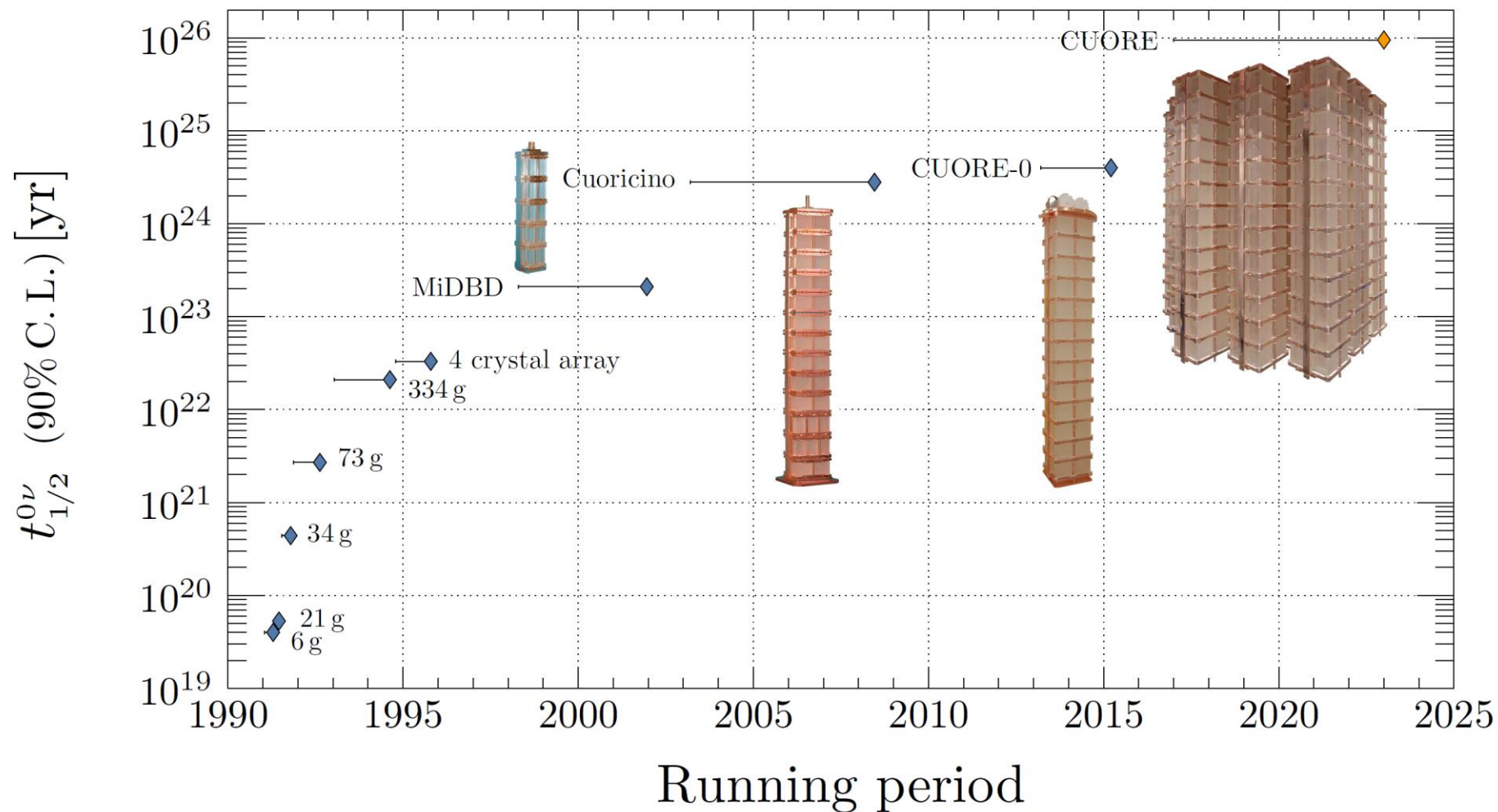
06/29/18

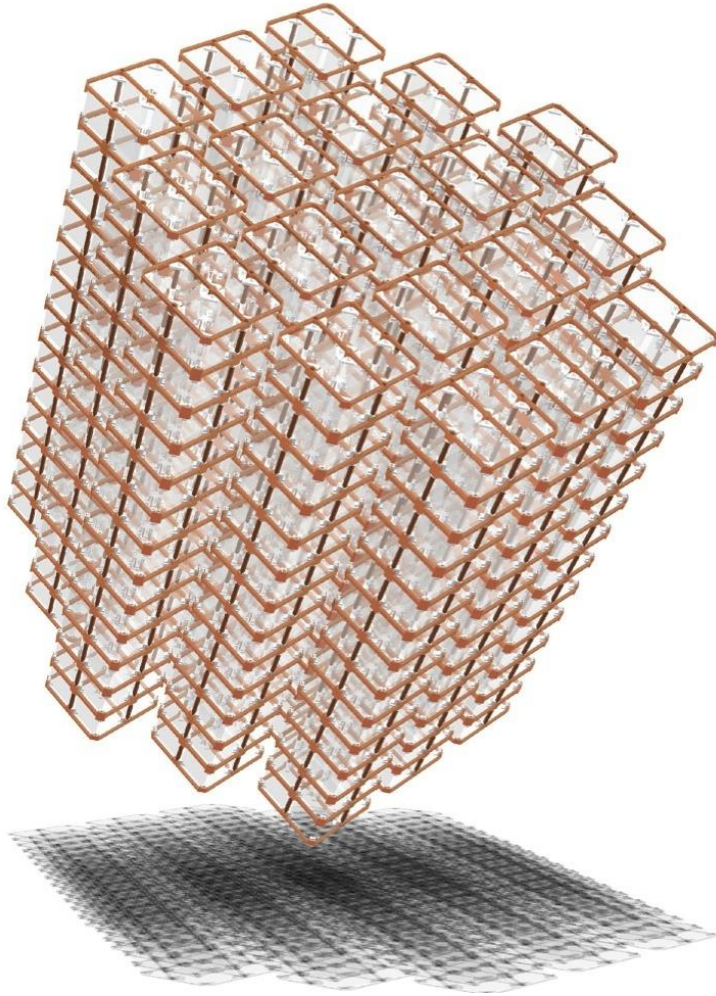
NDM 2018, Ke Han, SJTU

Heat bath ~10mK (Cu frames)



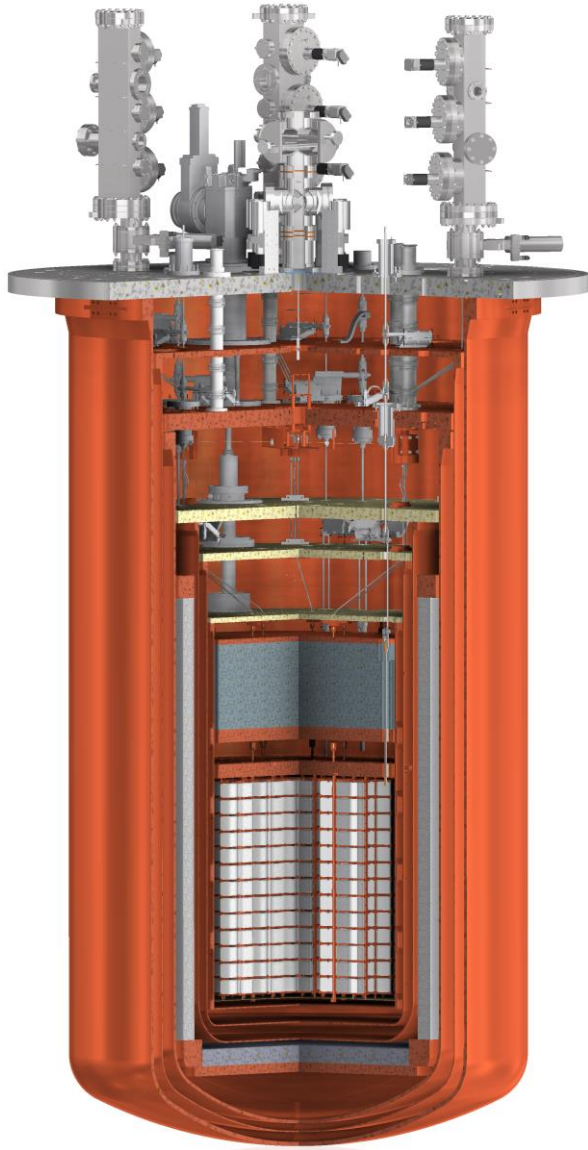
Arrays of TeO_2 bolometers





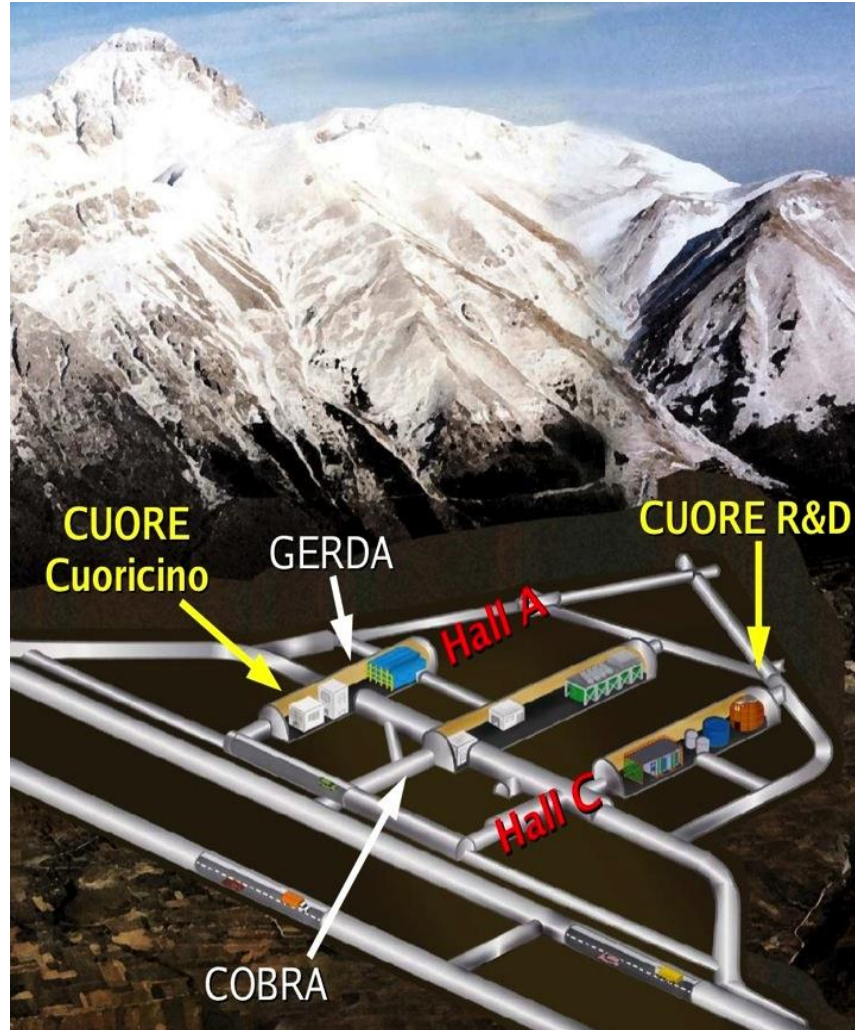
- Search for $0\nu\beta\beta$ of ^{130}Te and other rare events
- 988 TeO_2 crystals run as a bolometer array
 - 19 Towers
 - 13 floors
 - 4 modules per floor
 - 741 kg total; 206 kg ^{130}Te
 - 10^{27} ^{130}Te nuclei
- 10 mK base temperature in a custom dilution refrigerator
- Gran Sasso underground lab (LNGS), Italy

Cryogenic Underground Observatory for Rare Events



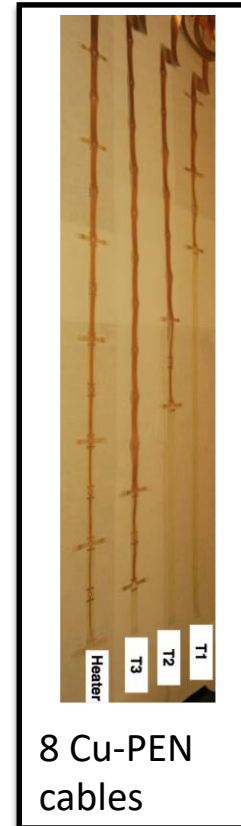
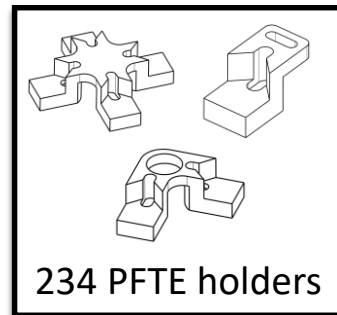
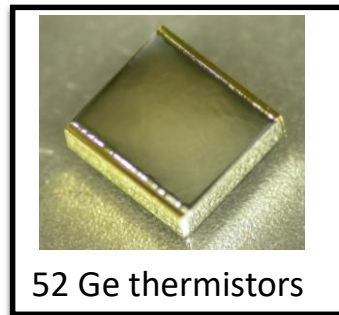
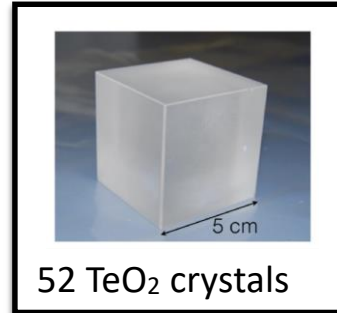
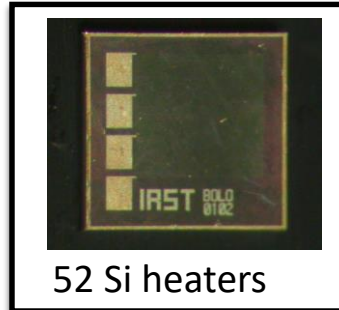
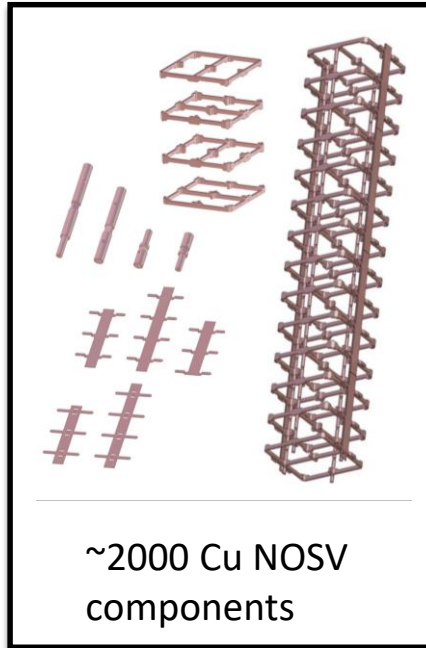
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CUORE-module: a tower



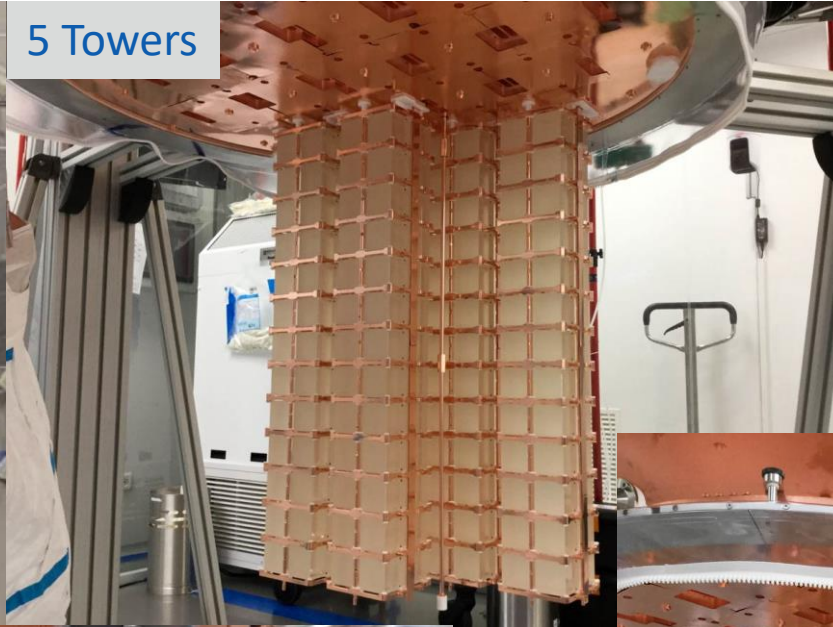
- New detector design structure
- Strict material selection (e.g. raw materials)
- Strict surface cleaning technique for Cu and TeO₂
- Minimization of Rn exposure (Glove Box assembly)

Detector installation (Summer 2016)

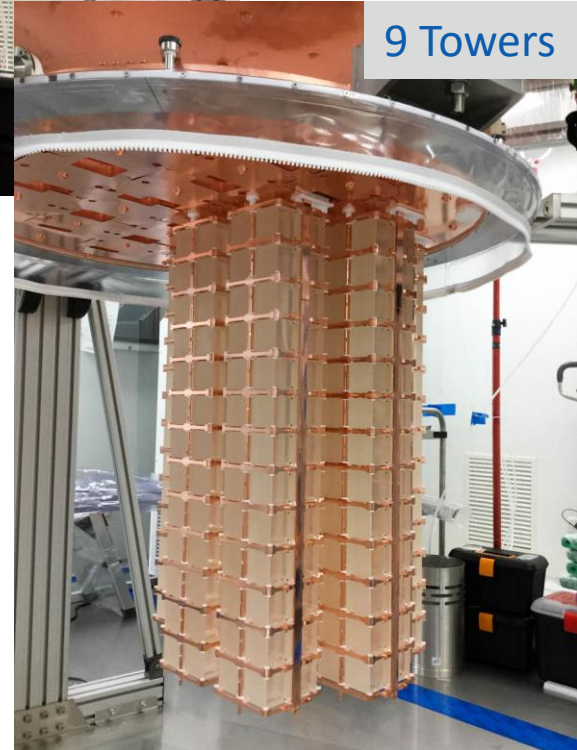
First Tower



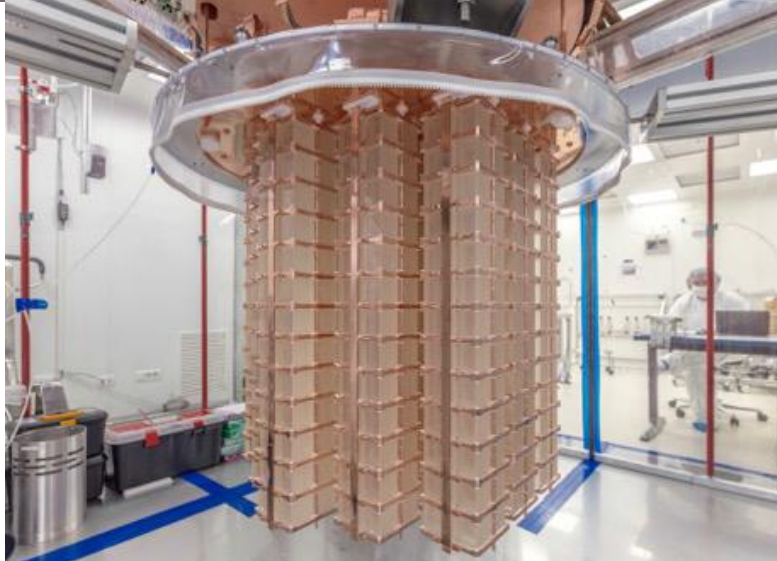
5 Towers



9 Towers

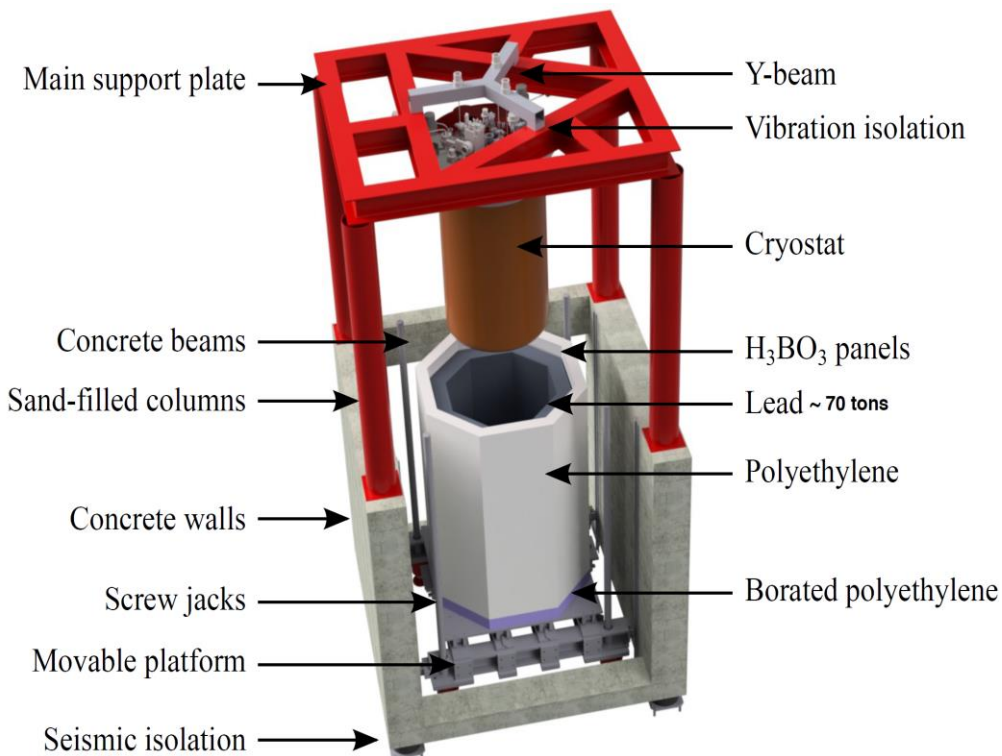


19 Towers



Passive shielding

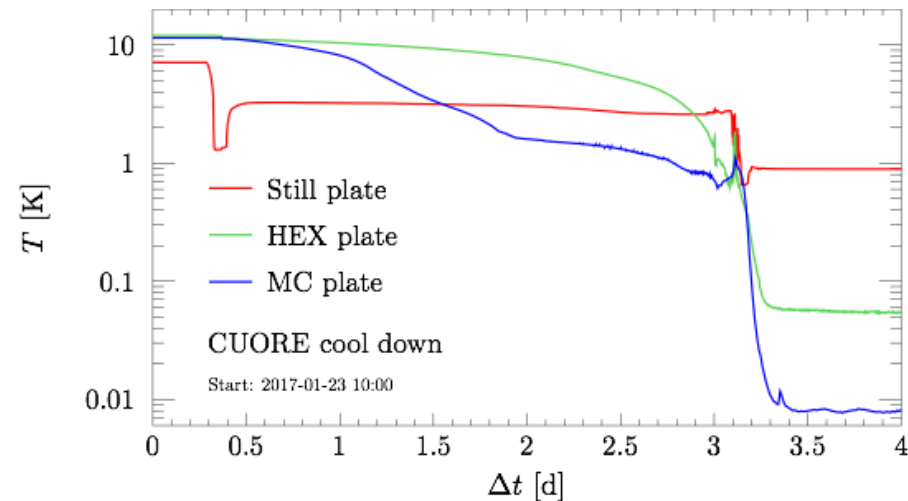
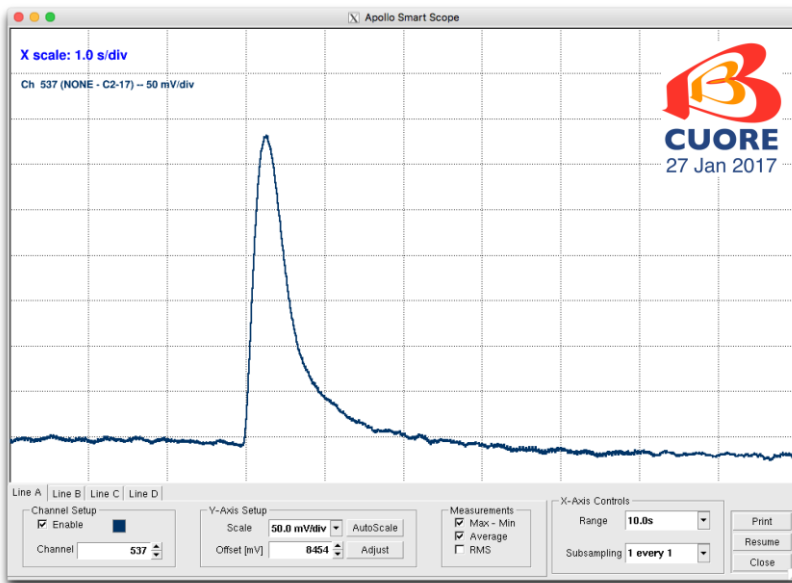
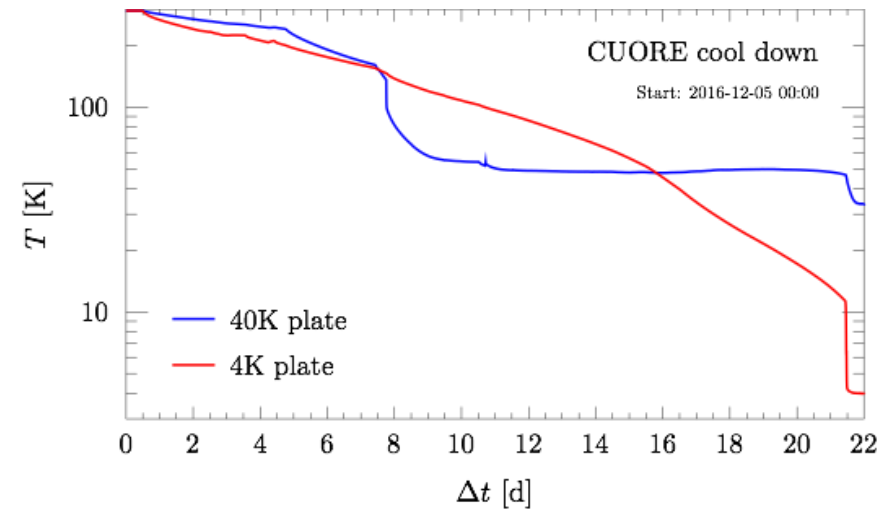
- External gamma and neutron shielding with ~ 70 tons lead + H_3BO_3
- Internal ancient roman lead and selected NOSV copper
- Internal shielding at low temperature: 2.5 tons @ 50mK + 5.5 tons @ 4K



CUORE cooling down

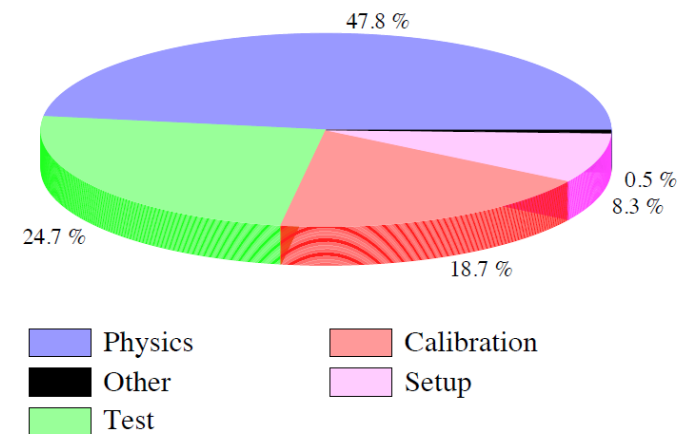
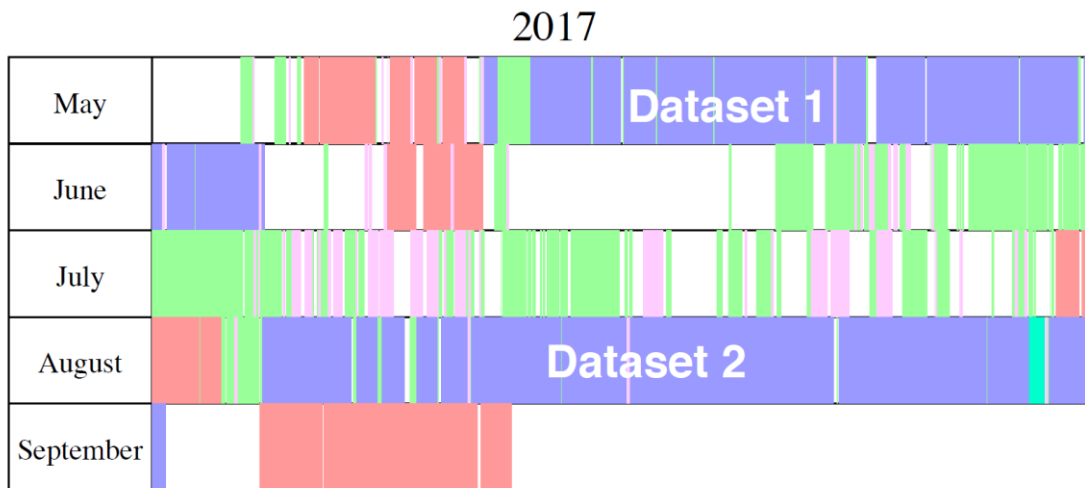


- Fast cooling system (4He gas) allows quick cool down to ~ 40 K
- All 5 PT coolers activated to bring cryostat to ~ 4 K
- DU brings temp down to ~ 6.3 mK
- Residual cooling power at 10 mK: $3 \mu\text{W}$



Science runs

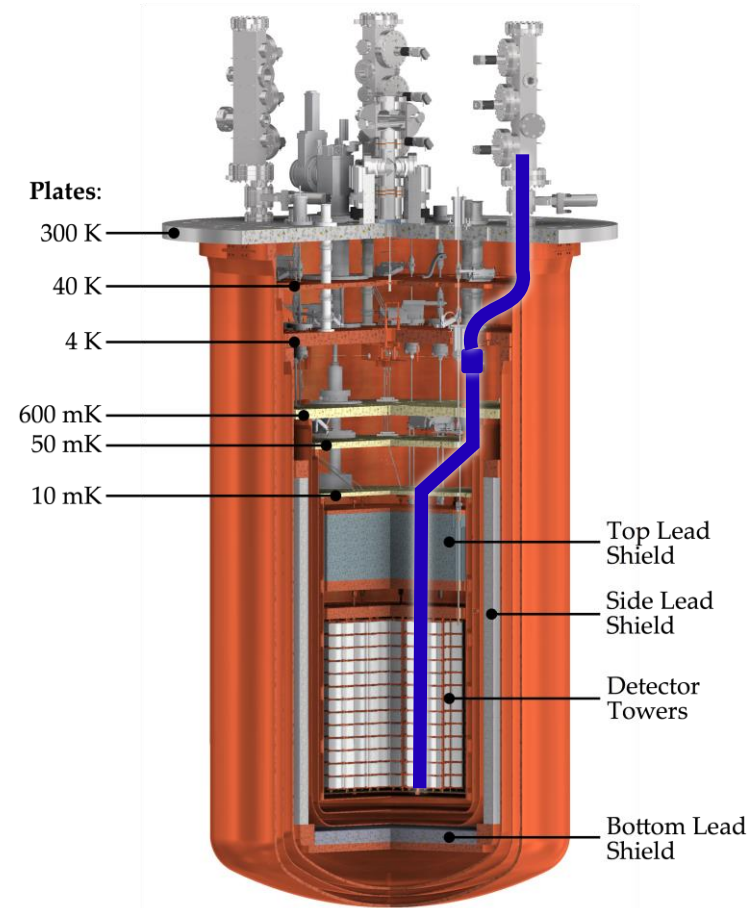
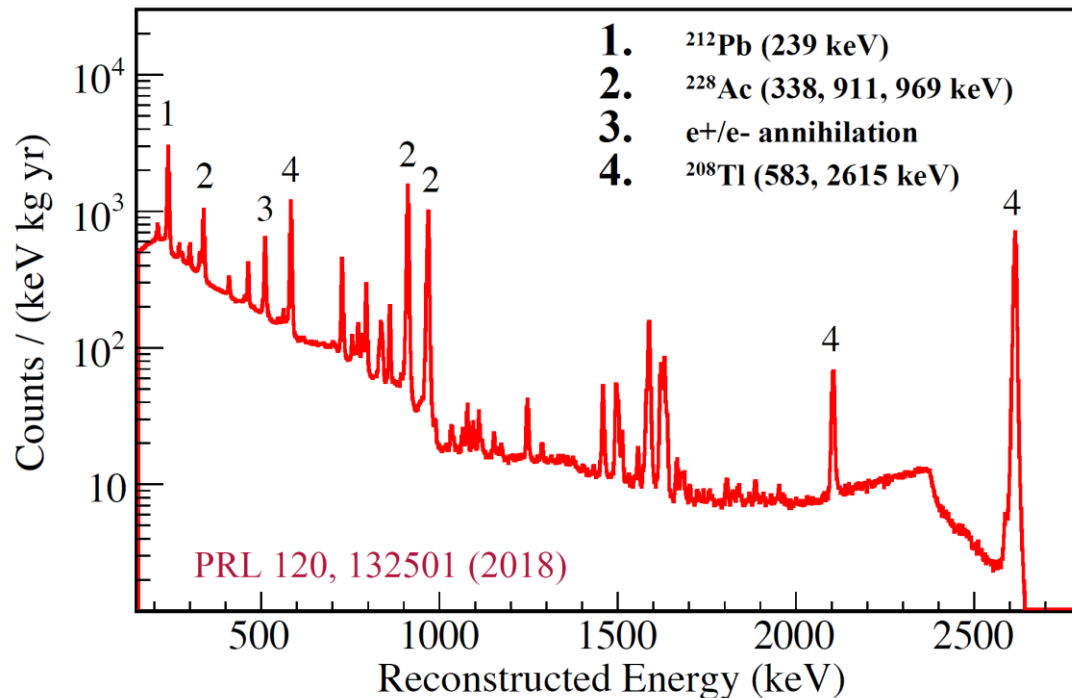
- Science operations started on April 14, 2017
- 984/988 operational channels
- Excellent data-taking efficiency when in operations
- Calibrations/physics data ratio still to be optimized
- $^{\text{nat}}\text{TeO}_2$ exposure: 86.3 kg yr (37.6 + 48.7)
- ^{130}Te exposure: 24.0 kg yr



Physics, Calibration, Test, Special Configuration runs

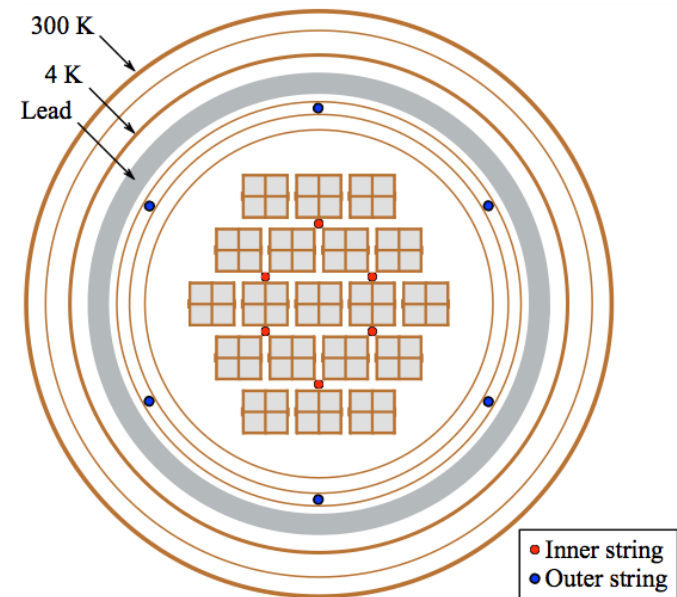
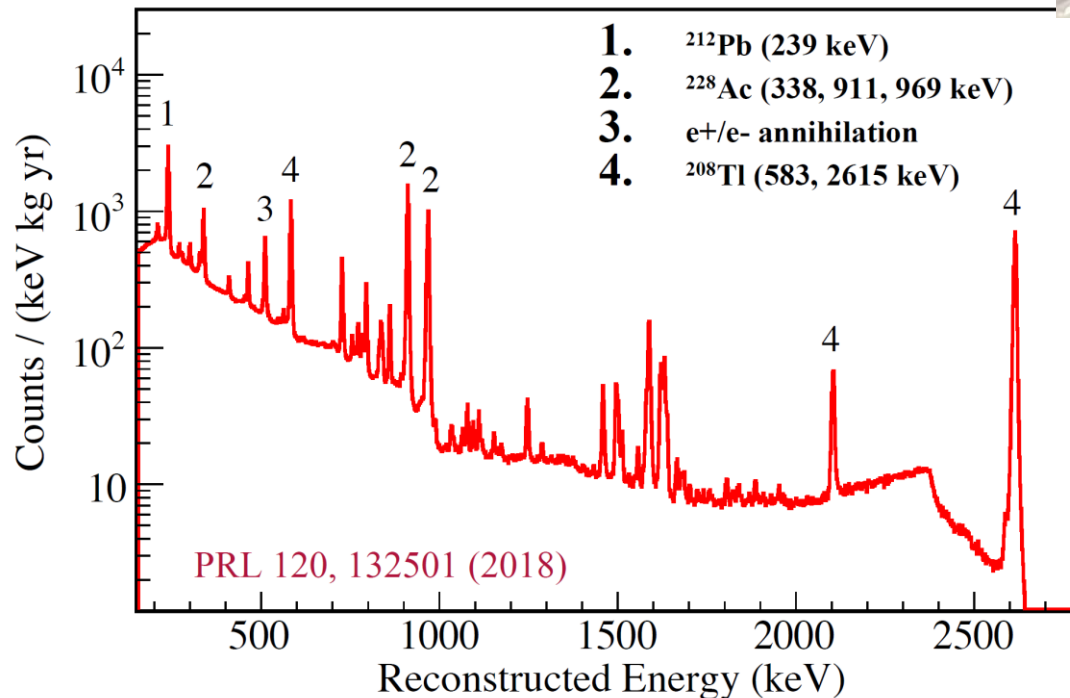
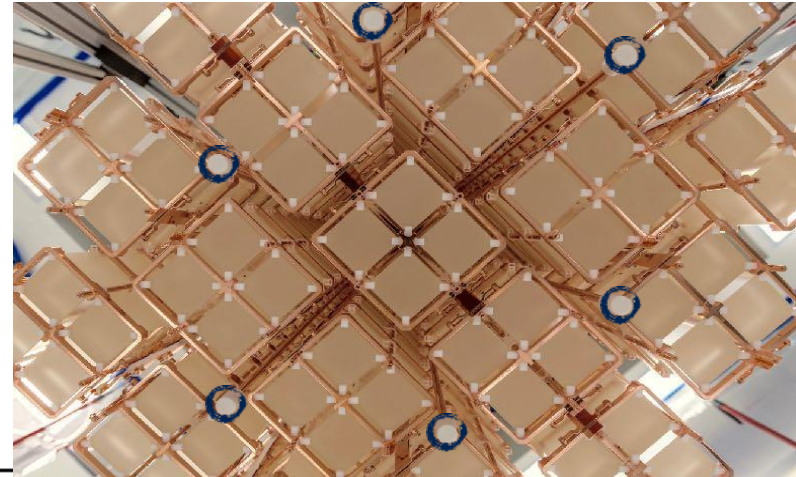
Calibration and gain stabilization

- ^{232}Th source (lower down from 10mK)
 - Energy scale calibration
- Silicon heater (pulser)
 - Thermal gain stabilization
- Detector response (line shape) study

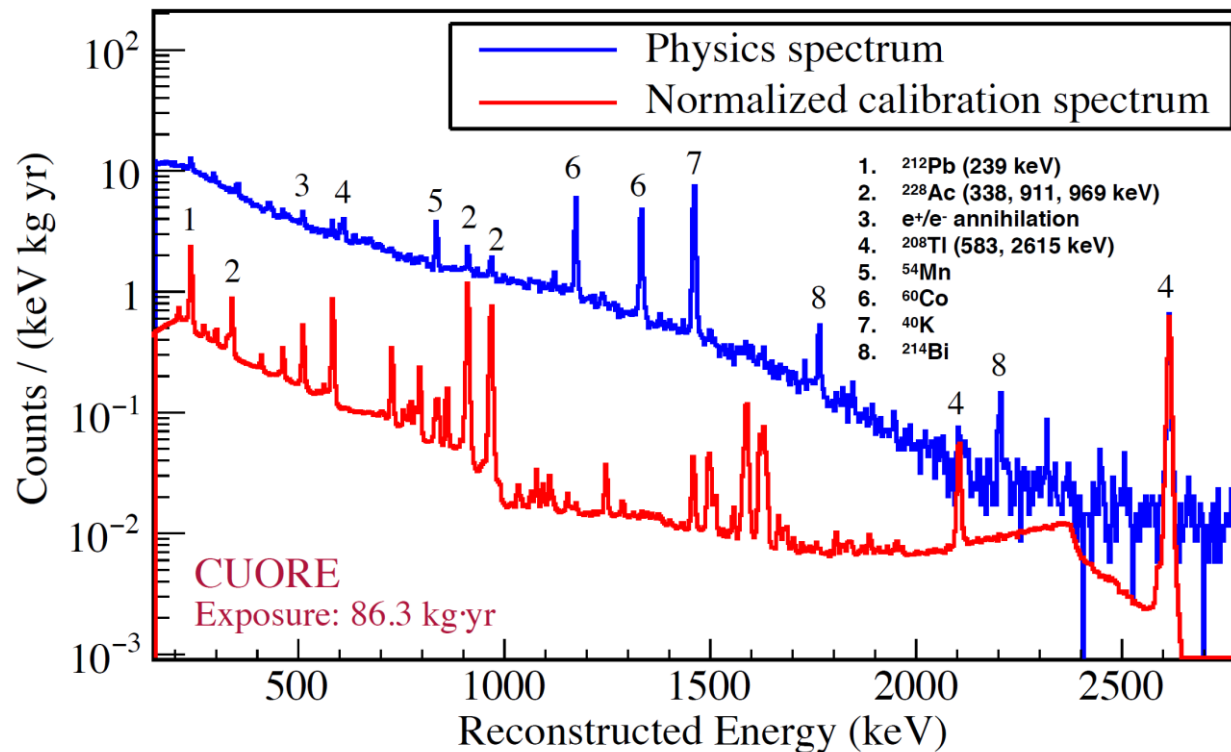


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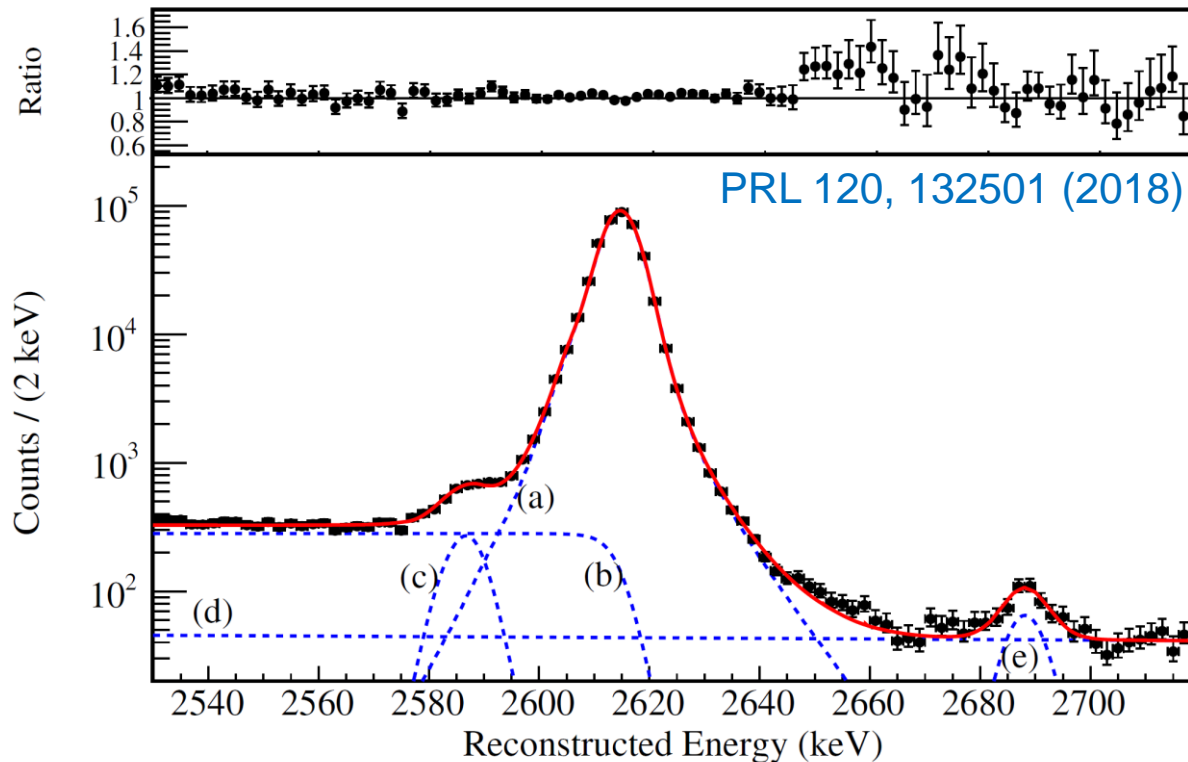
- 1811 channel-dataset pairs utilized
 - 92% of active channels
- Energy spectrum from sum of all active channels used in $0\nu\beta\beta$ search



Line shape studies

Model the ^{208}Tl 2615 keV line

- The most prominent peak from ^{232}Th calibration
- Close to the $0\nu\beta\beta$ Q-value for ^{130}Te in the background spectrum



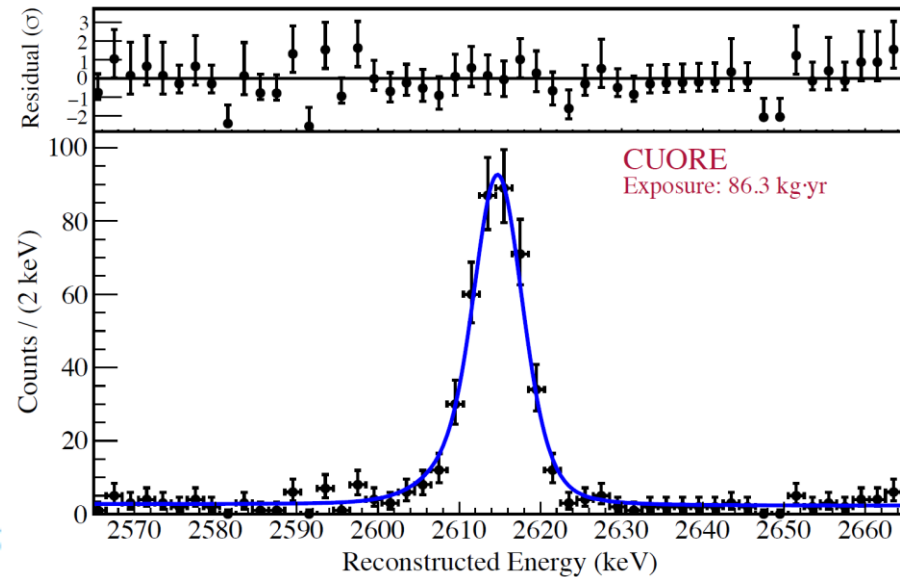
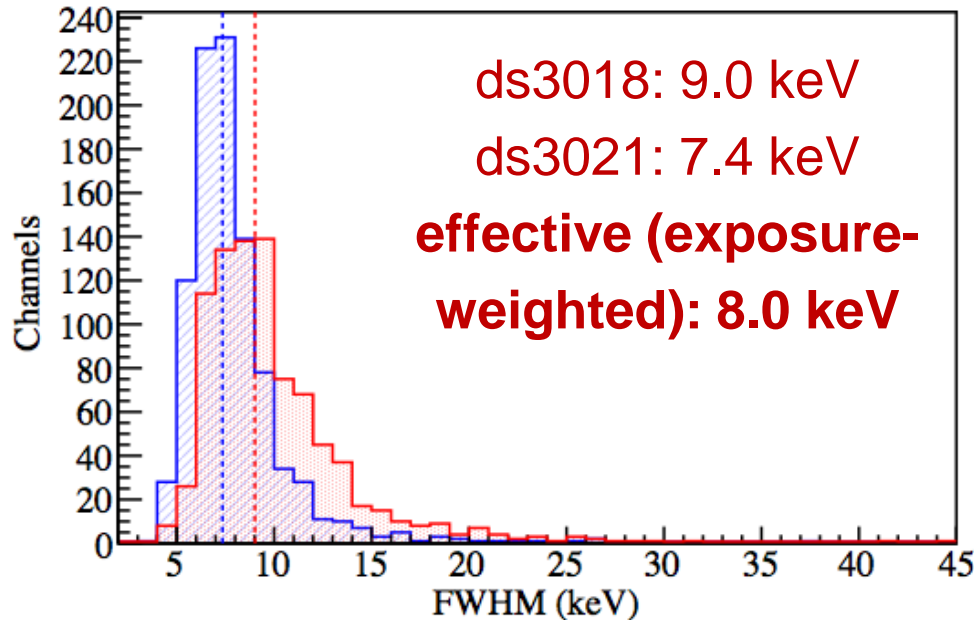
Fit for 2615 keV summed over all channels

- a) Triple Gaussian for main peak
- b) Step-wise multi-Compton background
- c) X-ray escape peak (~30 keV)
- d) Linear background
- e) 2687 (2615 + 583 - 511) keV peak

Energy resolution

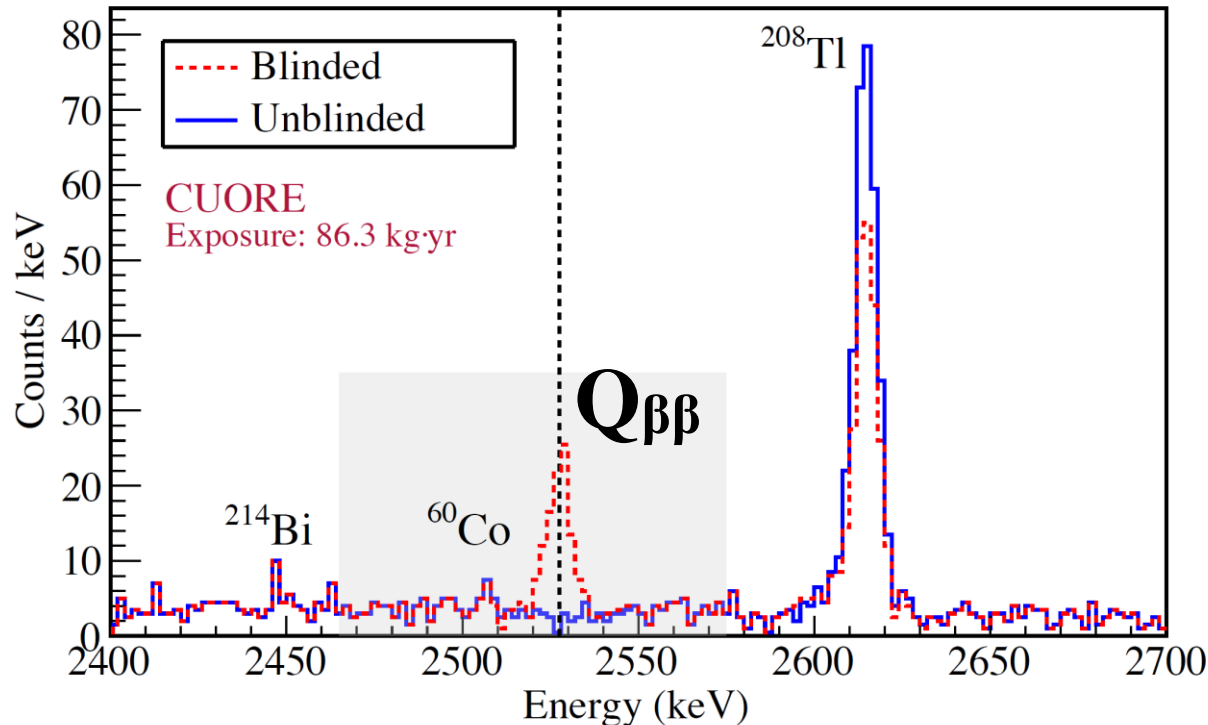
- 1811 (92% of live channels) channels-dataset
- Obvious improvements in ds3021
- Exposure weighted average resolution at Q-value of 7.7 ± 0.5 keV FWHM

Resolution at 2615 keV gamma peak



Blinding procedure

- To blind our data we randomly move a fraction of events from ± 20 keV of 2615 keV to the Q-value and vice versa
- The blinding algorithm produces an artificial peak around the $0\nu\beta\beta$ Q-value hiding the real $0\nu\beta\beta$ rate of ^{130}Te
- When all data analysis procedures are fixed the data are eventually unblinded

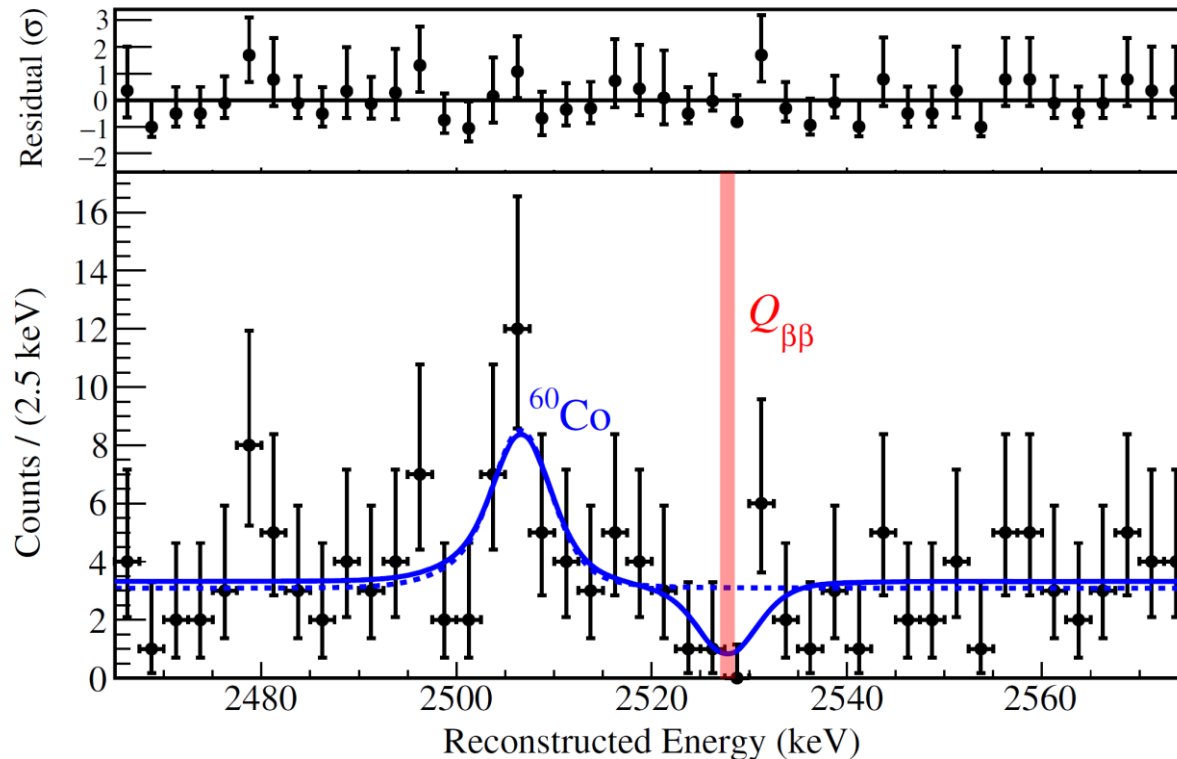


Event selection

- 155 (65+90) events survive cuts in the unblinded region of interest

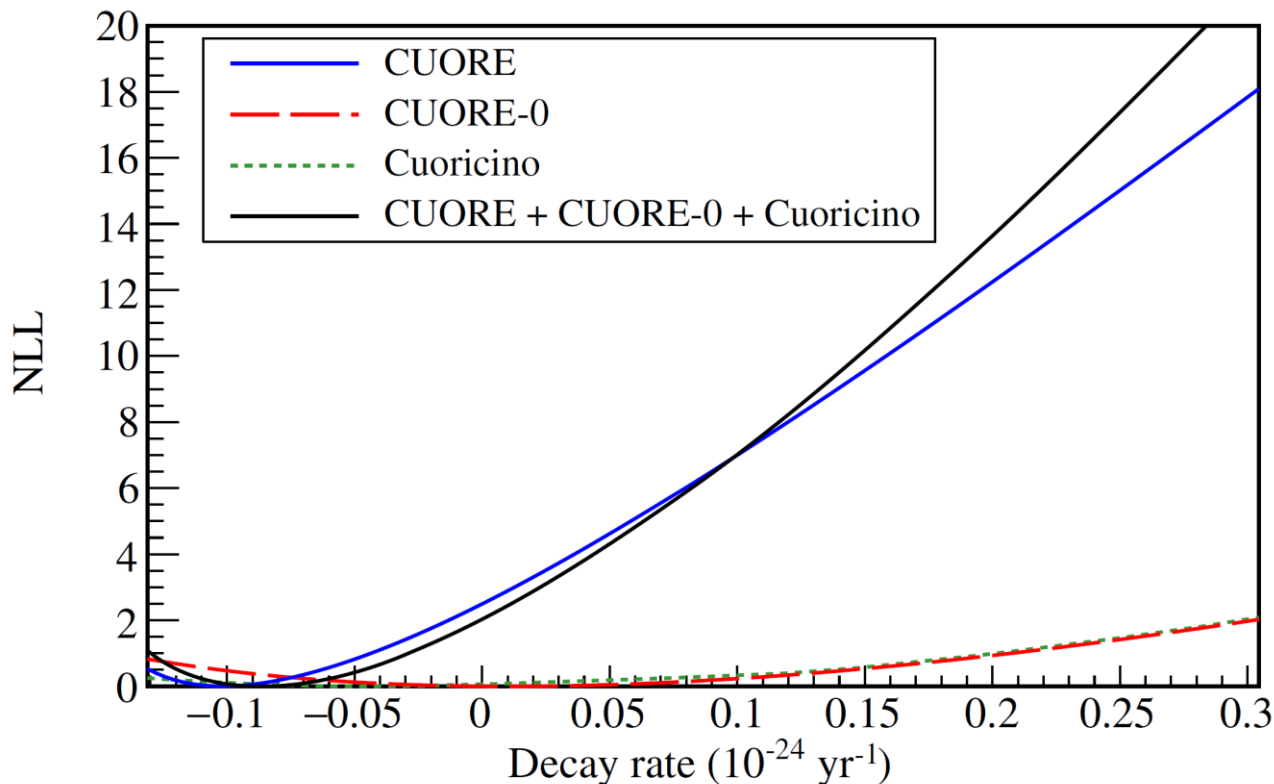


- Region of interest: 2465 to 2575 keV
- Signal efficiency $(75.7 \pm 3.0)\%$ - ds3018 $[(83.0 \pm 2.6)\% - \text{ds3021}]$
- Background index (c/(keV·kg·yr)): $(1.49_{-0.17}^{+0.18}) \times 10^{-2}$ $[(1.35_{-0.18}^{+0.20}) \times 10^{-2}]$
- **Best Fit Signal Decay Rate: $(-1.0_{-0.3}^{+0.4} \text{ (stat.)} \pm 0.1 \text{ (syst.)}) \times 10^{-25} / \text{yr}$**



$0\nu\beta\beta$ results

- No $0\nu\beta\beta$ signals found
- Half-life limit established
- Profile likelihood integrated on the physical region ($\Gamma^{0\nu} > 0$)



Decay rate limit (90% CL, including systematics):
 $0.51 \times 10^{-25} / \text{yr}$

Half-life limit (90% CL, including systematics):
 $1.3 \times 10^{25} \text{ yr}$

Median expected sensitivity:
 $7.0 \times 10^{24} \text{ yr}$

Combining with previous ^{130}Te results

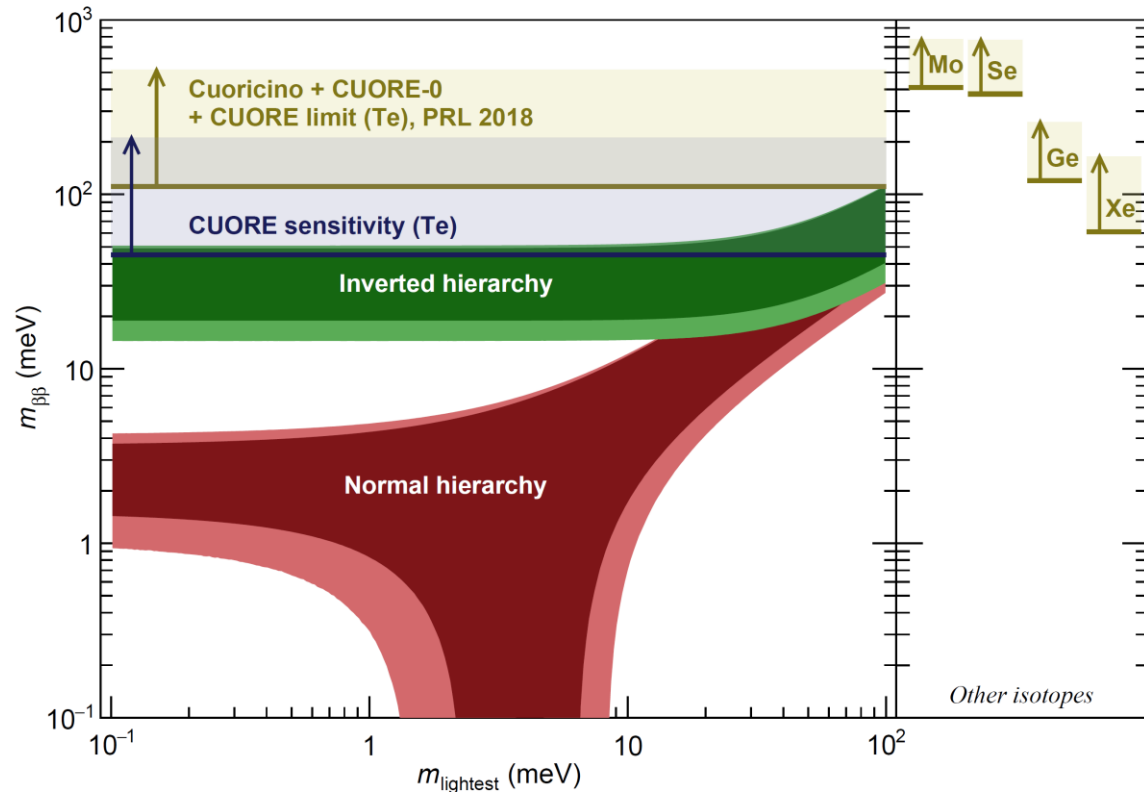
We combined the CUORE result with the existing ^{130}Te

- 19.75 kg·yr of Cuoricino
- 9.8 kg·yr of CUORE-0

- The combined 90% C.L. limit is

$$T_{0\nu} > 1.5 \times 10^{25} \text{ yr}$$

$$m_{\beta\beta} < 110\text{--}520 \text{ meV}$$



^{130}Te : 1.5×10^{25} yr from this analysis PRL 120, 132501 (2018)

^{76}Ge : 8.0×10^{25} yr from PRL 120, 132503 (2018)

^{136}Xe : 1.1×10^{26} yr from Phys. Rev. Lett. 117, 082503 (2016)

^{100}Mo : 1.1×10^{24} yr from Phys. Rev. D 89, 111101 (2014)

^{82}Se : 2.4×10^{24} yr from Phys. Rev. Lett. 120, 232502 (2018)

CUORE sensitivity: 9.0×10^{25} yr

CUPID sensitivity: 5.1×10^{27} yr

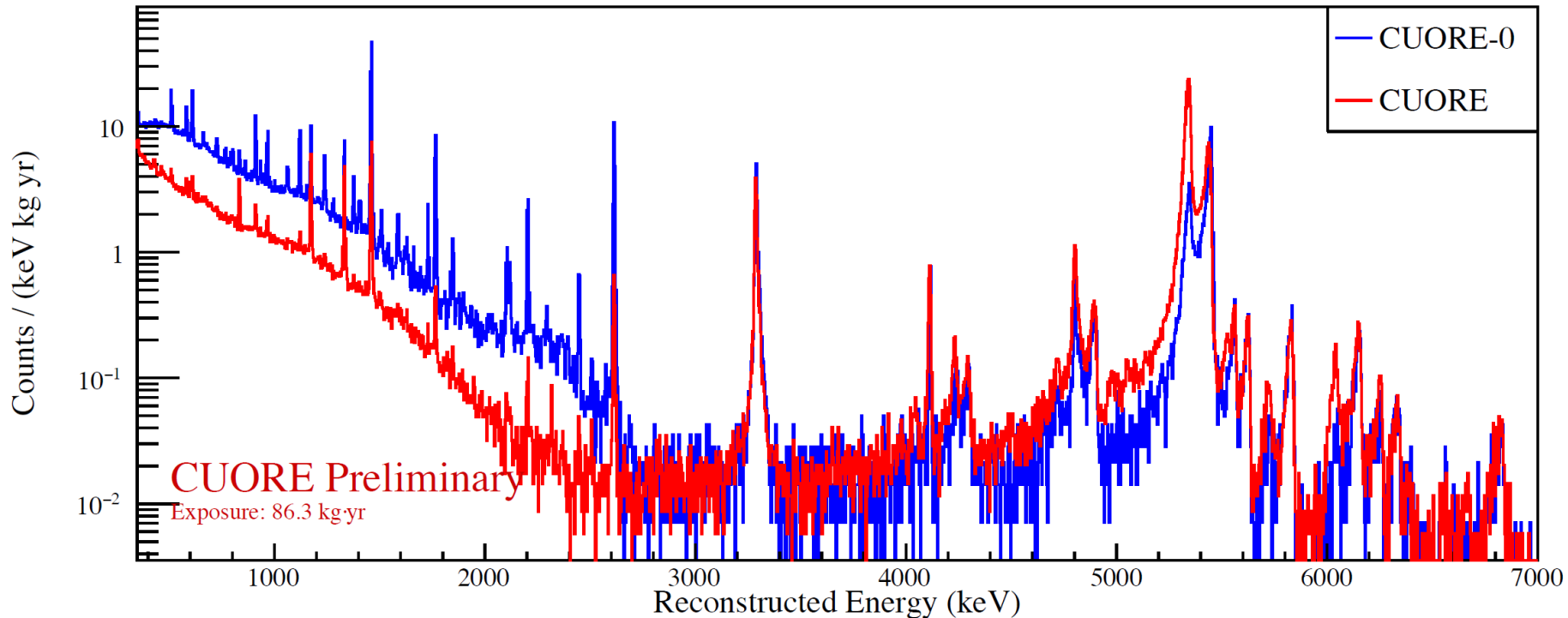
NME:

JHEP02 (2013) 025
 Nucl. Phys. A 818, 139 (2009)
 Phys. Rev. C 87, 045501 (2013)
 Phys. Rev. C 87, 064302 (2014)
 Phys. Rev. C 91, 034304 (2015)
 Phys. Rev. C 91, 024613 (2015)
 Phys. Rev. C 91, 024309 (2015)
 Phys. Rev. C 91, 024316 (2015)
 Phys. Rev. Lett. 105, 252503 (2010)
 Phys. Rev. Lett. 111, 142501 (2013)

CUORE background

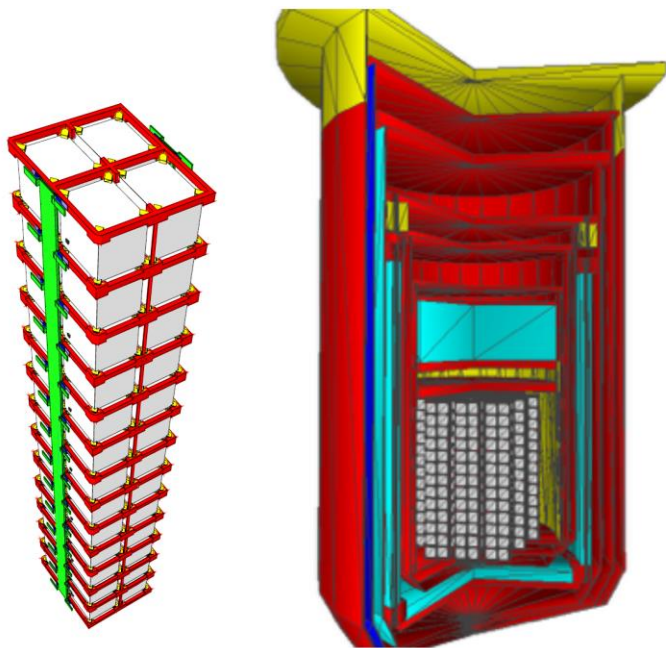


- Significant reduction in the gamma region with respect to CUORE-0
- ^{210}Po excess appears to be from shallow contamination in copper around the detectors
 - Estimated contribution to ROI at the level of $\sim 10^{-4}$ c/(keV kg yr)



Fitting the CUORE Background Model

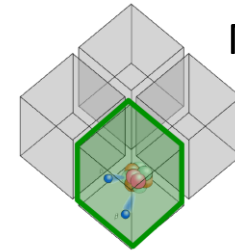
- Simulate the contaminations coming from different cryostat components using a detailed Geant4 MC simulation
- ~60 independent parameters representing various contaminations that could contribute to the CUORE background model
- Bayesian fit to data with flat priors except for muons which come from a cosmogenic analysis



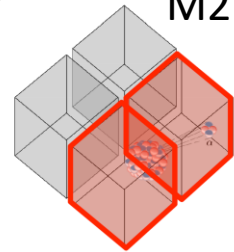
Volume	Type	Components
TeO ₂	Bulk	$2\nu\beta\beta$, ^{210}Pb , ^{232}Th , ^{228}Ra - ^{208}Pb , ^{238}U - ^{230}Th , ^{230}Th , ^{226}Ra - ^{210}Pb , ^{40}K , ^{60}Co , ^{125}Sb , ^{190}Pt
TeO ₂	Surface (0.01 μm)	^{232}Th , ^{228}Ra - ^{208}Pb , ^{238}U - ^{230}Th , ^{226}Ra - ^{210}Pb , ^{210}Pb
TeO ₂	Surface (1 μm)	^{210}Pb
TeO ₂	Surface (10 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Bulk	^{232}Th , ^{238}U , ^{40}K , ^{60}Co , ^{54}Mn
CuNOSV	Surface (0.01 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Surface (1 μm)	^{210}Pb , ^{232}Th , ^{238}U
CuNOSV	Surface (10 μm)	^{210}Pb , ^{232}Th , ^{238}U
Roman lead	Bulk	^{232}Th , ^{238}U , ^{108m}Ag
Top lead	Bulk	^{232}Th , ^{238}U , ^{210}Bi
Ext. lead	Bulk	^{210}Bi
CuOFE	Bulk	^{232}Th , ^{238}U , ^{60}Co
External	-	Cosmic muons

More recent runs

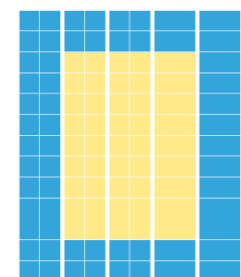
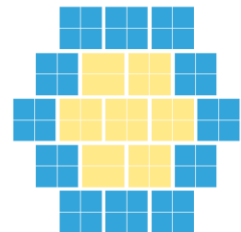
- Multiplicity 1 spectra very sensitive to signal events
- Many contaminations constrained by the higher multiplicity spectra



M1

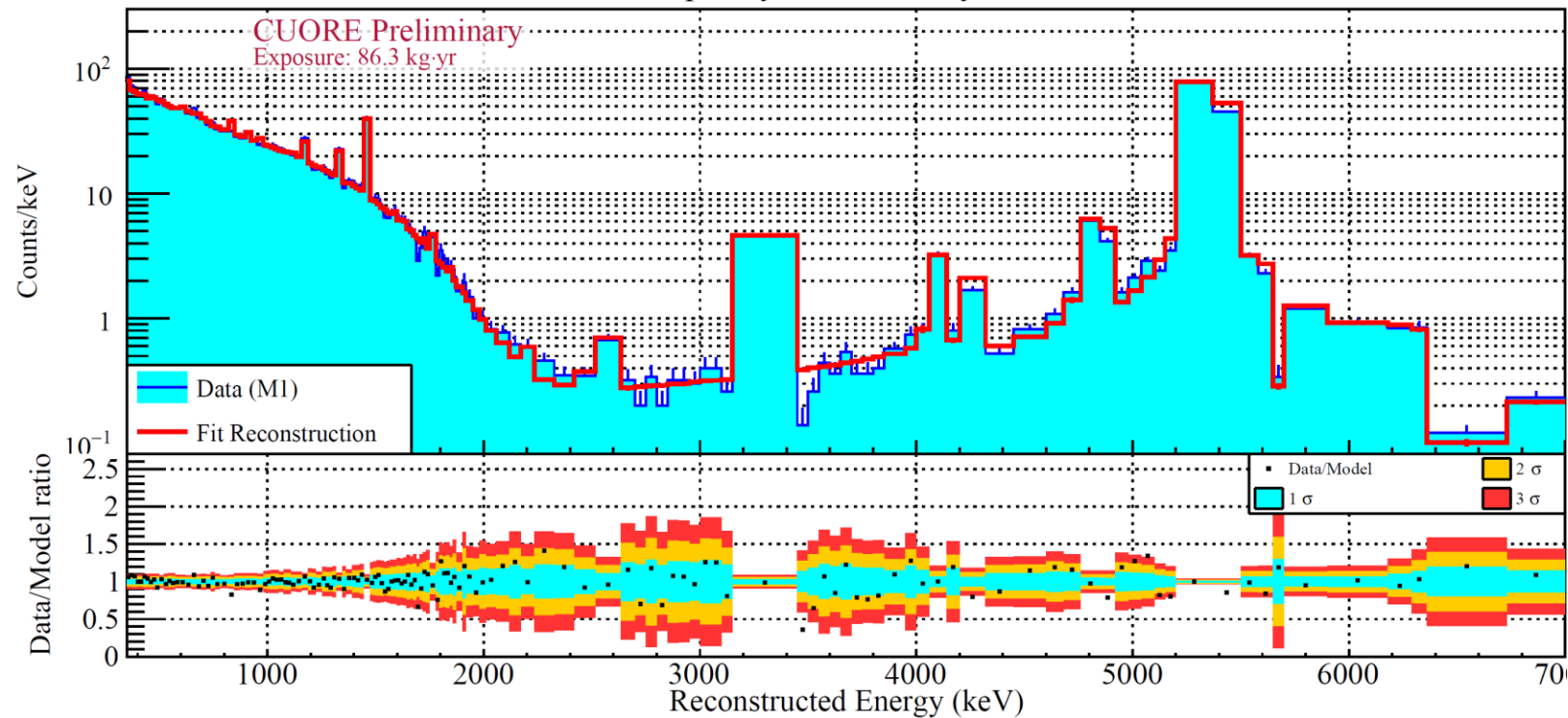


M2



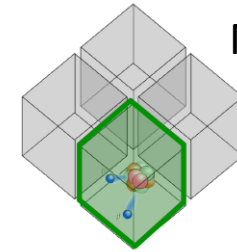
Inner Layer
Outer Layer

Multiplicity 1 - Inner Layer

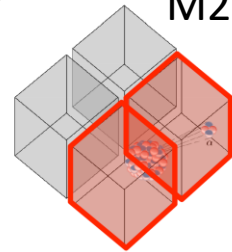


More recent runs

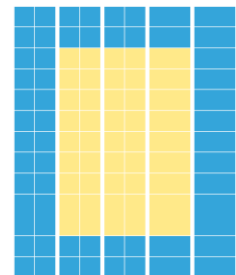
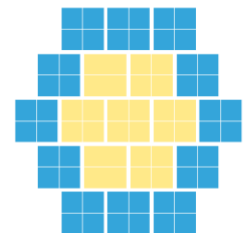
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M1

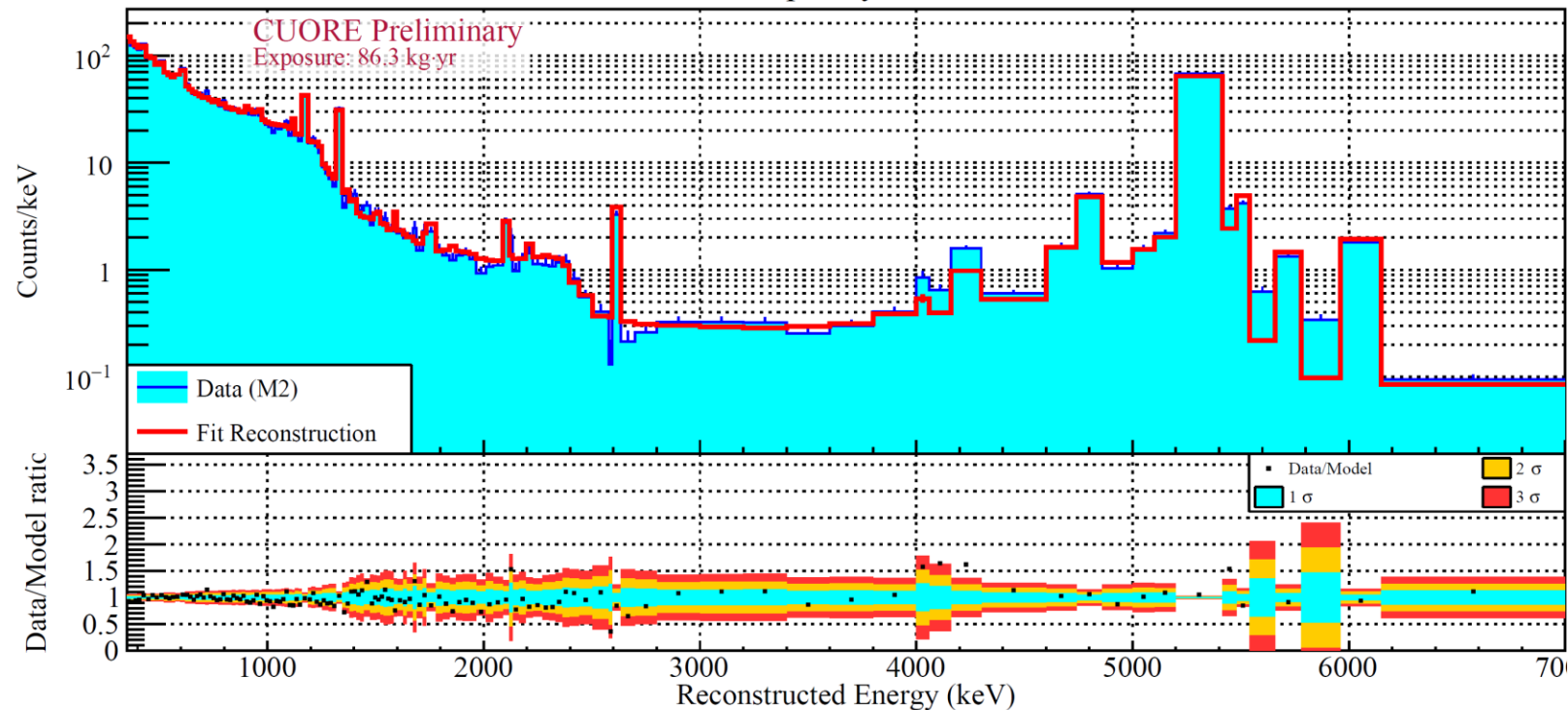


M2



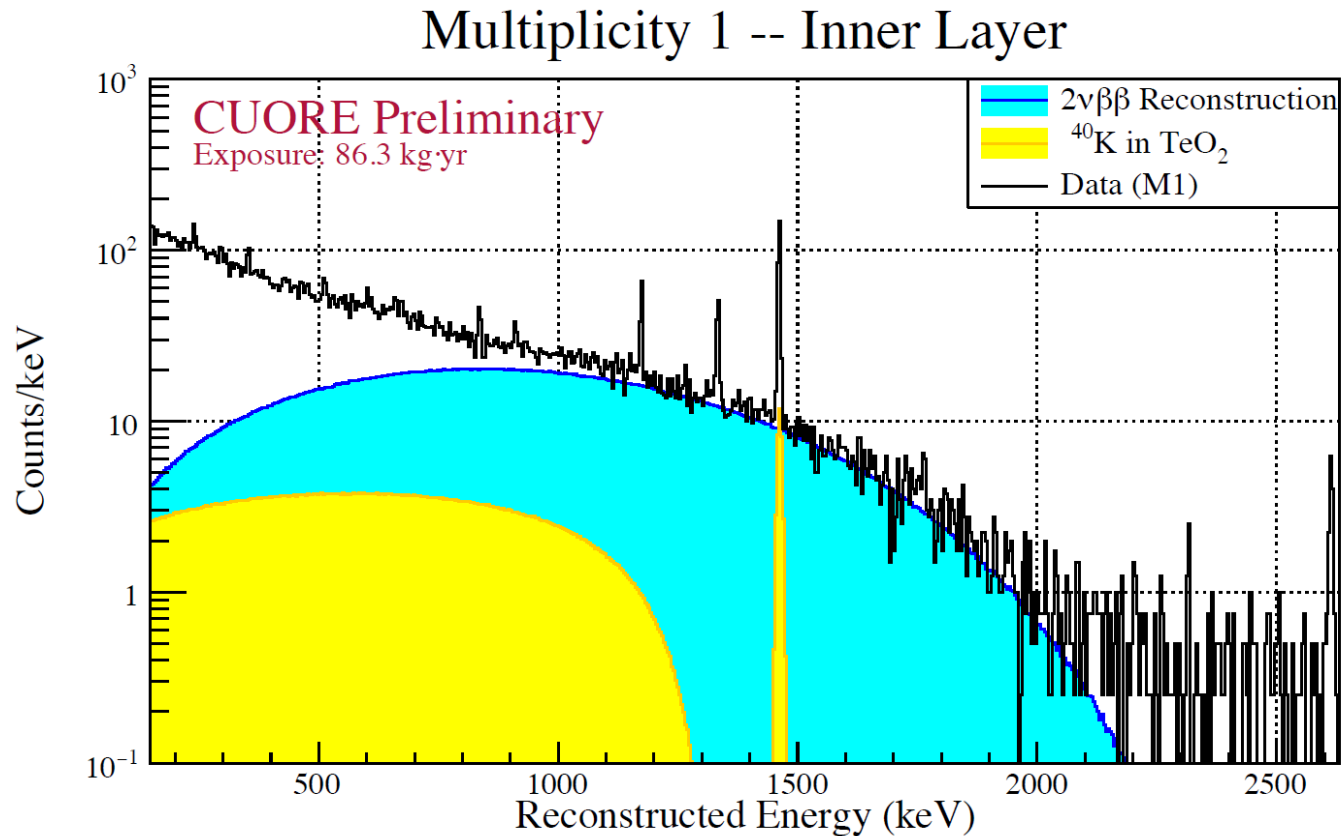
Inner Layer
Outer Layer

Multiplicity 2



Measuring precisely $2\nu\beta\beta$ half-life

- $2\nu\beta\beta$ decay rate: $[8.7 \pm 0.1 \text{ (stat.)} \pm 0.2 \text{ (syst.)}] \times 10^{-21} \text{ yr}^{-1}$
- Half-life: $[7.9 \pm 0.1 \text{ (stat.)} \pm 0.2 \text{ (syst.)}] \times 10^{20} \text{ yr}$

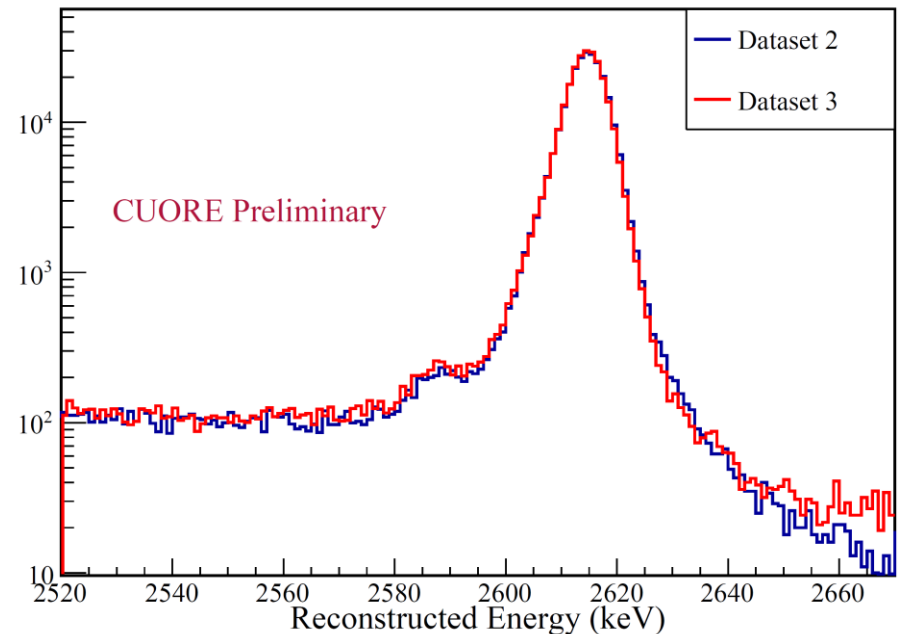
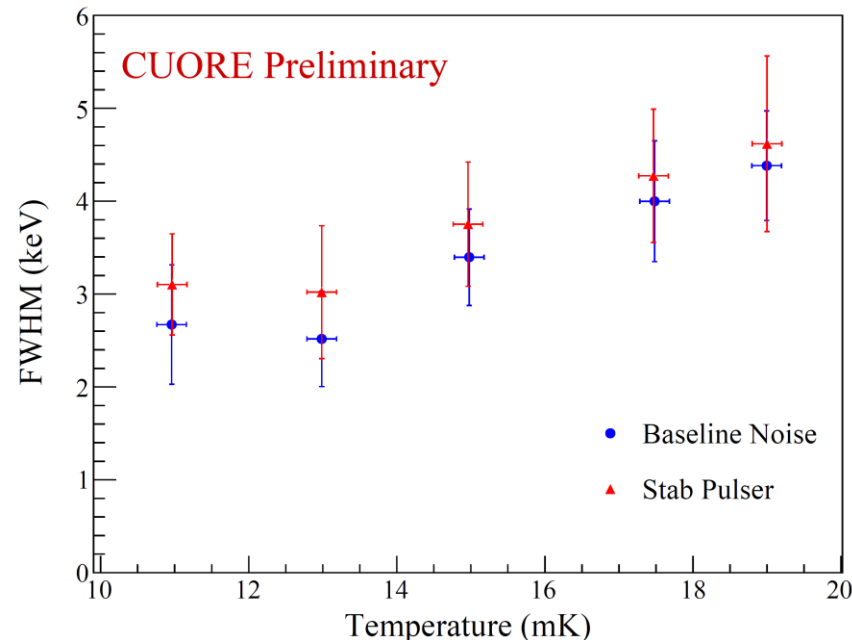


CUORE-0 : $[8.2 \pm 0.2 \text{ (stat.)} \pm 0.6 \text{ (syst.)}] \times 10^{20} \text{ yr}$; NEMO-3 : $[7.0 \pm 0.9 \text{ (stat.)} \pm 1.1 \text{ (syst.)}] \times 10^{20} \text{ yr}$

More recent data taking

- Detector temperature scan and cryostat optimization after the first PRL result
- Stable physics data taking since May 2018 at 11 mK.
- Data quality on par with previous datasets.

Median FWHM vs Temperature - October 2017 Temperature Scan



What's next for CUORE



- We set the most stringent limit on the $0\nu\beta\beta$ half-life of ^{130}Te and had the most precise measurement of the $2\nu\beta\beta$ half-life of ^{130}Te
- More physics analysis in the pipeline:
 - Double beta decay with Majoron emission
 - Dark Matter WIMP and axion searches with low energy analysis
 - $\beta\beta$ decays to excited states, $\beta^+/\text{E.C.}$ decays
- CUORE will continue to be one of the most sensitive searches for $0\nu\beta\beta$ for the next few years.
- Sensitivity to $0\nu\beta\beta$ half-life with 5 year live time reaches 9×10^{25} yr
- CUORE Upgrade with Particle ID (CUPID) R&D is on-going.

