

Nonlinearities can filter dark energy models

Farbod Hassani⁽¹⁾, Martin Kunz⁽¹⁾, Filippo Vernizzi⁽²⁾, Julian Adamek⁽³⁾, David Daverio⁽⁴⁾, Peter Wittwer⁽¹⁾



⁽¹⁾Department of Theoretical Physics, University of Geneva, Switzerland

⁽²⁾Institut de physique théorique, Université Paris Saclay CEA, France

⁽³⁾School of Physics & Astronomy, Queen Mary University of London, United Kingdom

⁽⁴⁾Centre for Theoretical Cosmology, University of Cambridge, United Kingdom

Action

$$S = \frac{1}{16\pi G} \int \sqrt{-g} R d^4x + \int \sqrt{-g} P(X, \varphi) d^4x \quad X = -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi$$

Equations of motion

$$\zeta \equiv \pi' + \mathcal{H}\pi - \Psi$$

$$\begin{aligned} \zeta' - 3w\mathcal{H}\zeta + 3c_s^2 (\mathcal{H}^2\pi - \mathcal{H}\Psi - \mathcal{H}'\pi - \Phi') - c_s^2 \nabla^2 \pi \\ + \vec{\nabla} [2(c_s^2 - 1)\zeta + c_s^2\Phi - \Psi] \cdot \vec{\nabla} \pi + [(c_s^2 - 1)\zeta + c_s^2\Phi - c_s^2\Psi] \nabla^2 \pi \\ + \frac{\mathcal{H}}{2} \left[(2 + 3w + c_s^2)(\vec{\nabla}\pi)^2 + 6c_s^2(1 + w)\pi \nabla^2 \pi \right] - \frac{c_s^2 - 1}{2} \nabla_i \left(\nabla_i \pi (\vec{\nabla}\pi)^2 \right) = 0 \end{aligned}$$

Non-linear terms

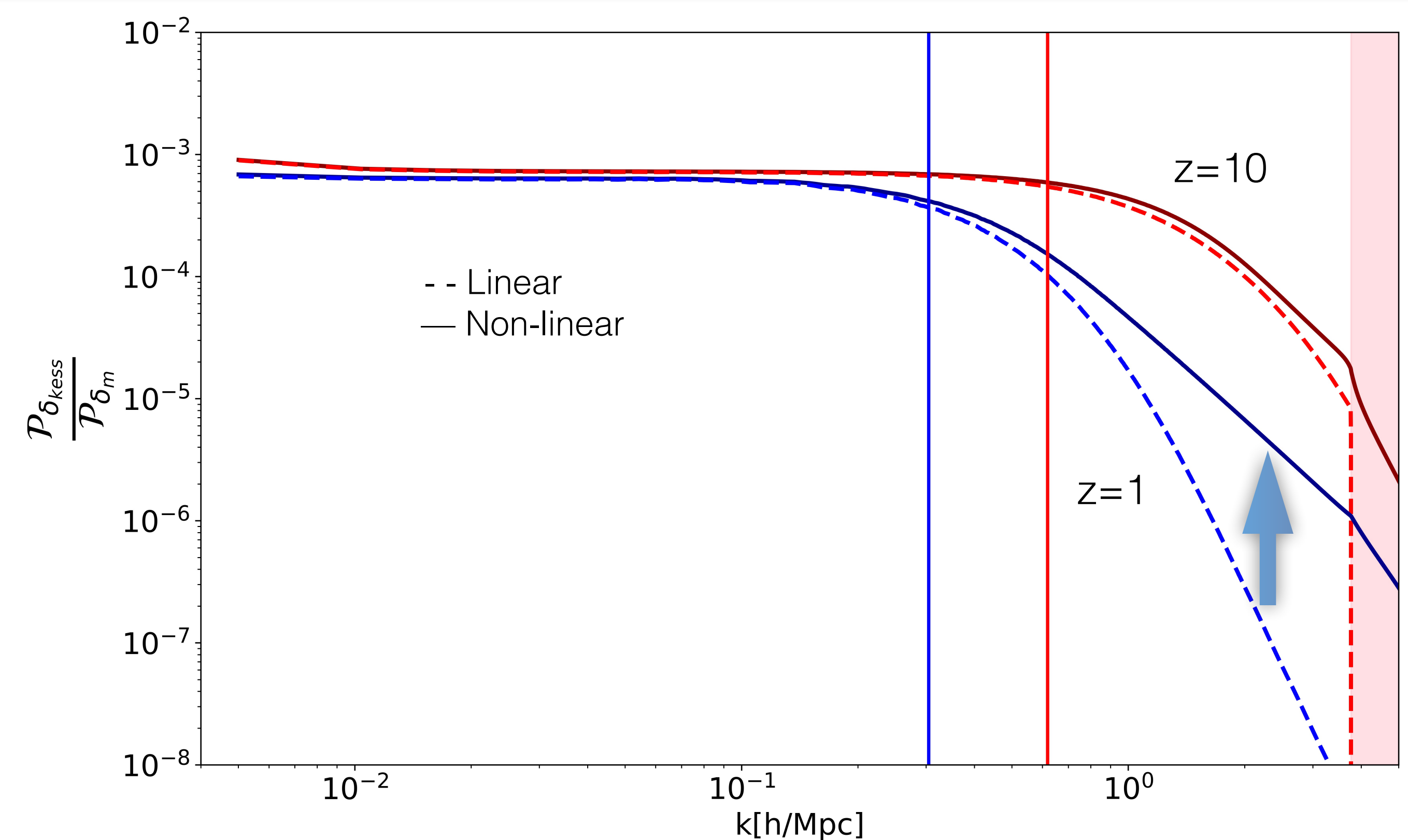
K-essence Stress tensor

$$T_0^0 = -\rho + \frac{\rho + p}{c_s^2} \left[3c_s^2 \mathcal{H}\pi - \zeta - \frac{2c_s^2 - 1}{2} (\vec{\nabla}\pi)^2 \right]$$

$$T_i^0 = -(\rho + p) \left[1 - \frac{1}{c_s^2} \left(3c_s^2(1 + w)\mathcal{H}\pi - \zeta + c_s^2\Psi \right) + \frac{c_s^2 - 1}{2c_s^2} (\vec{\nabla}\pi)^2 \right] \partial_i \pi$$

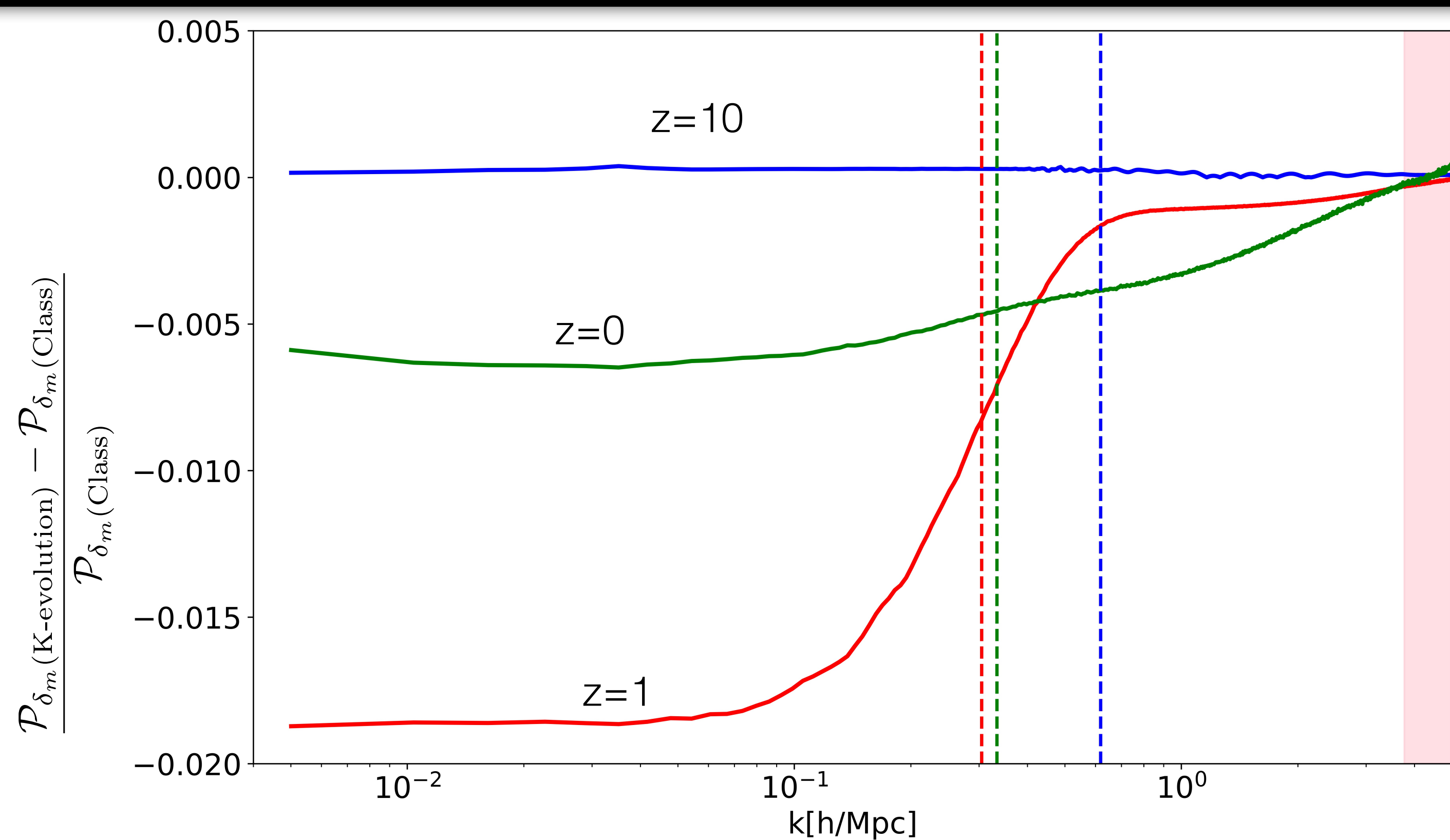
$$T_j^i = p\delta_j^i - (\rho + p) \left[3w\mathcal{H}\pi - \zeta + \frac{1}{2} (\vec{\nabla}\pi)^2 \right] \delta_j^i + (\rho + p) \delta^{ik} \partial_k \pi \partial_j \pi$$

Effect of non-linearities on linear k-essence field



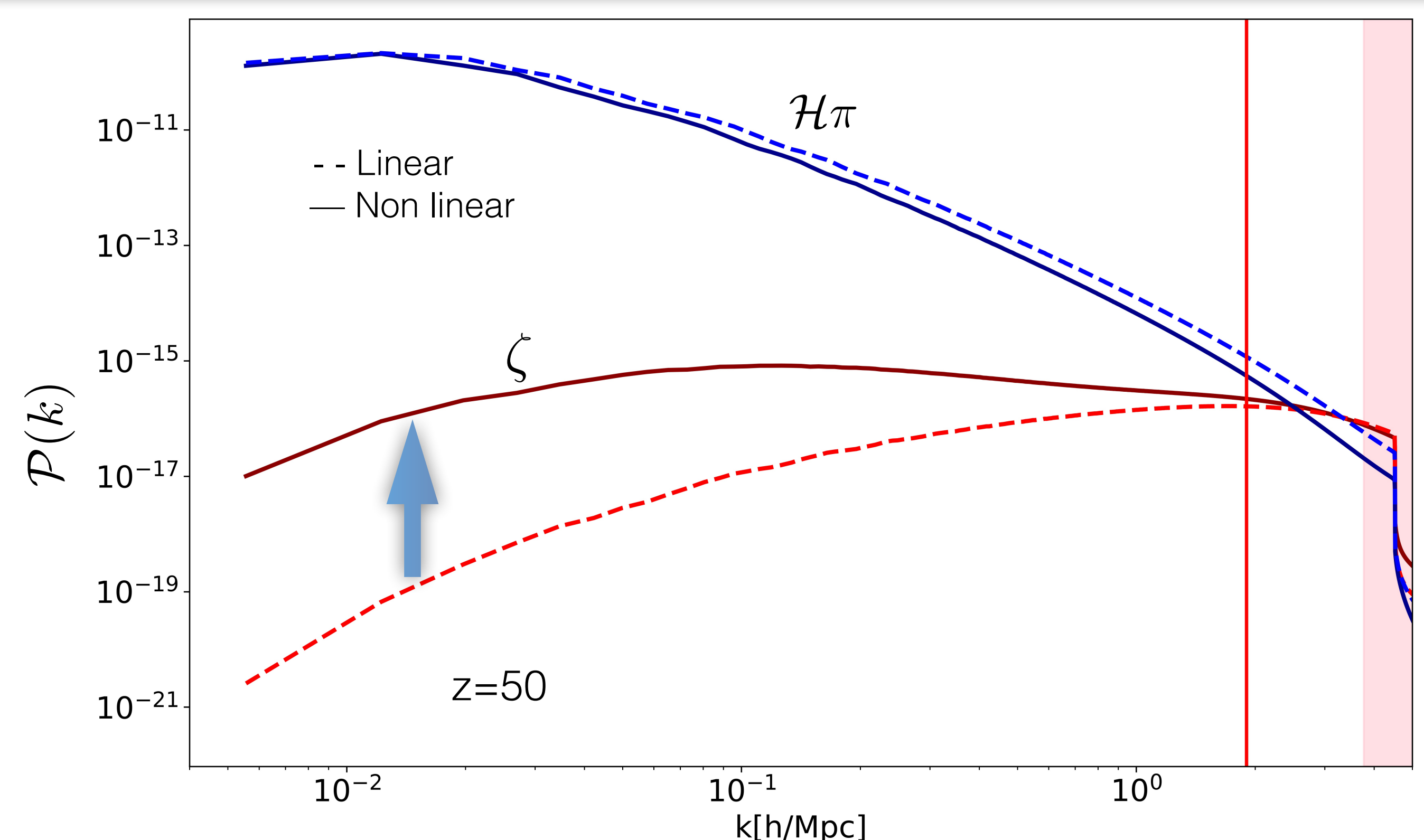
- The non-linear Φ , Ψ and Φ' in k-evolution source linear $\delta_{\text{k-essence}}$
- The equations used for the scalar field in k-evolution are linear

Effect of linear k-essence field on matter power spectrum



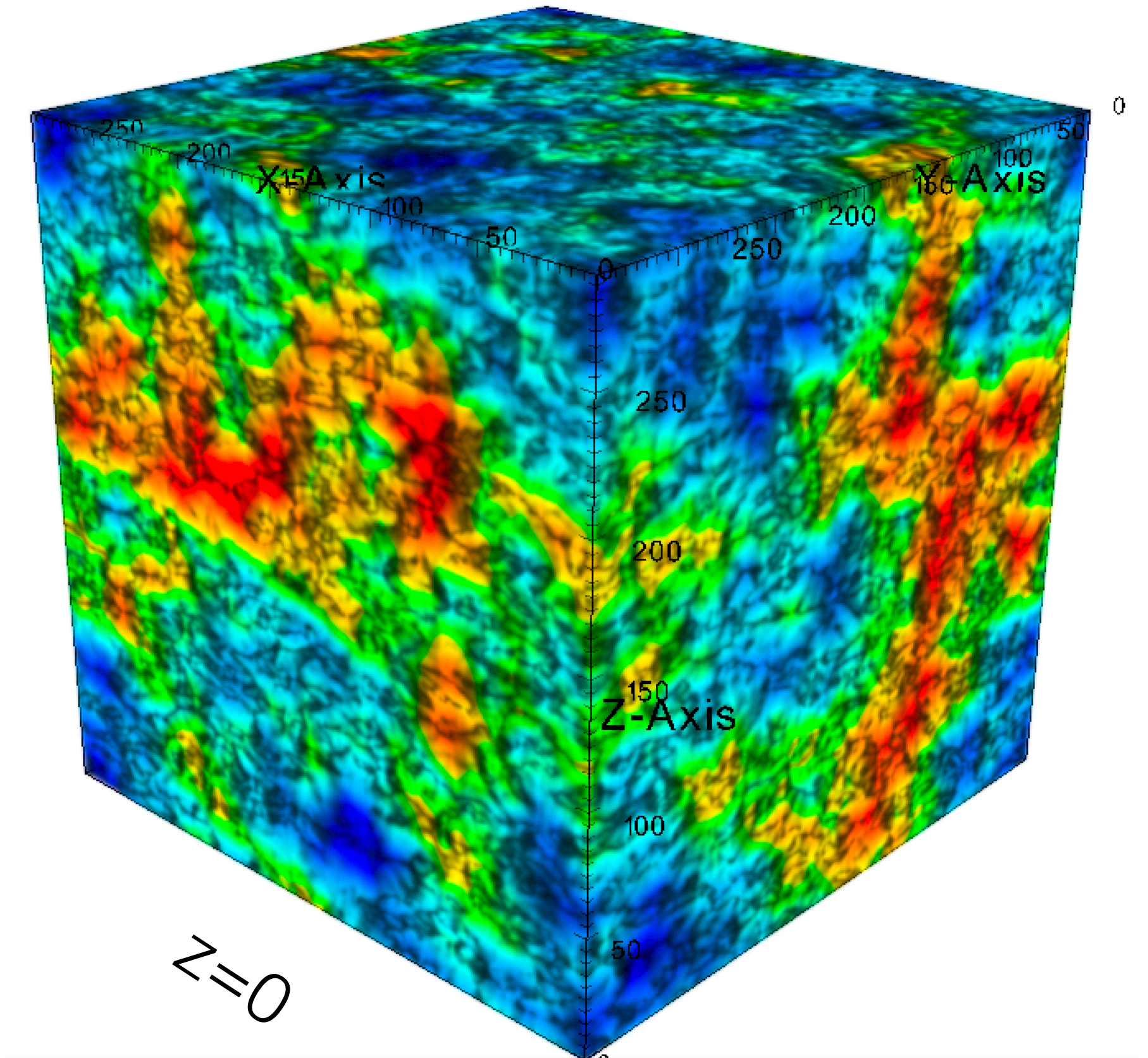
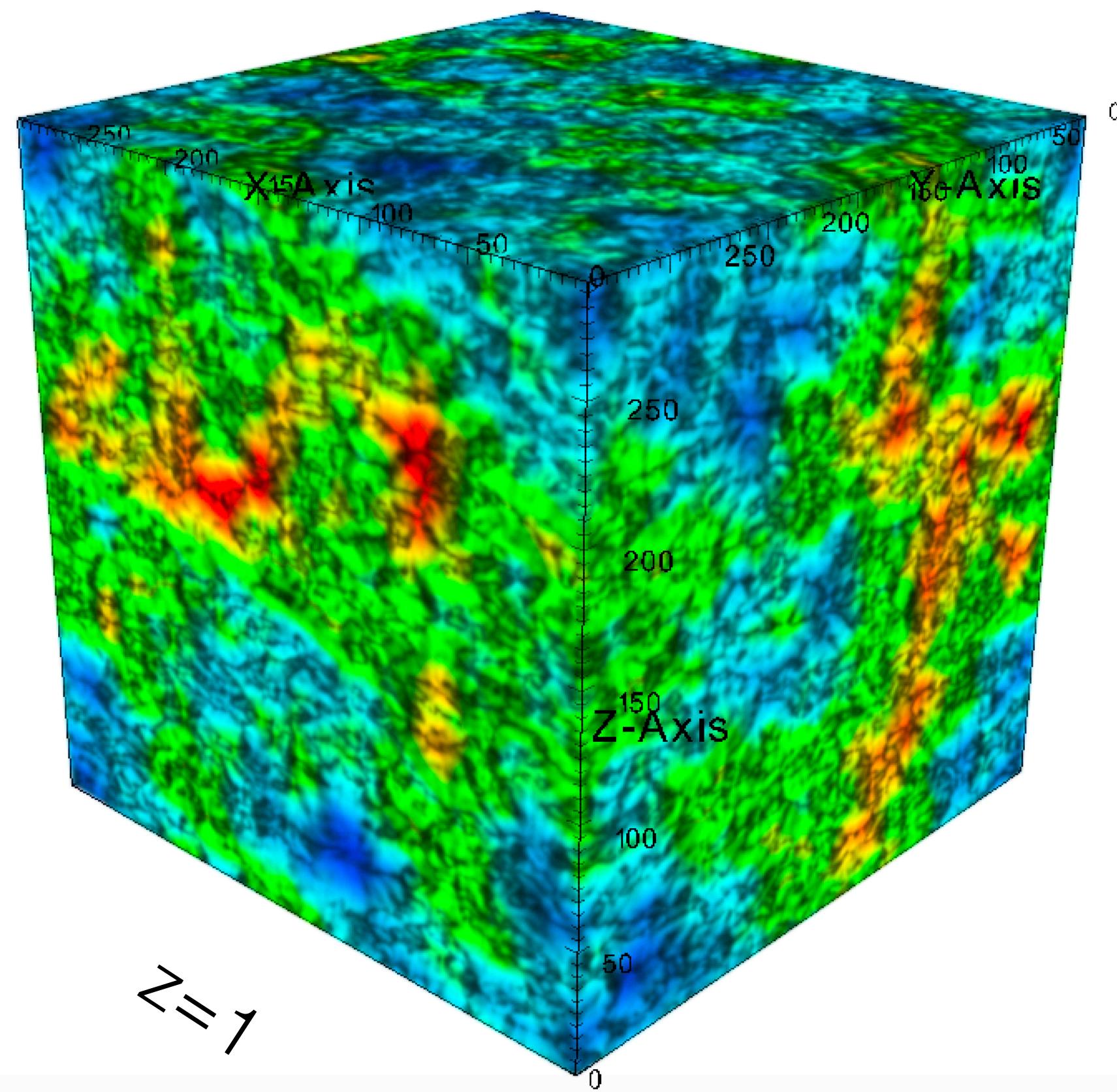
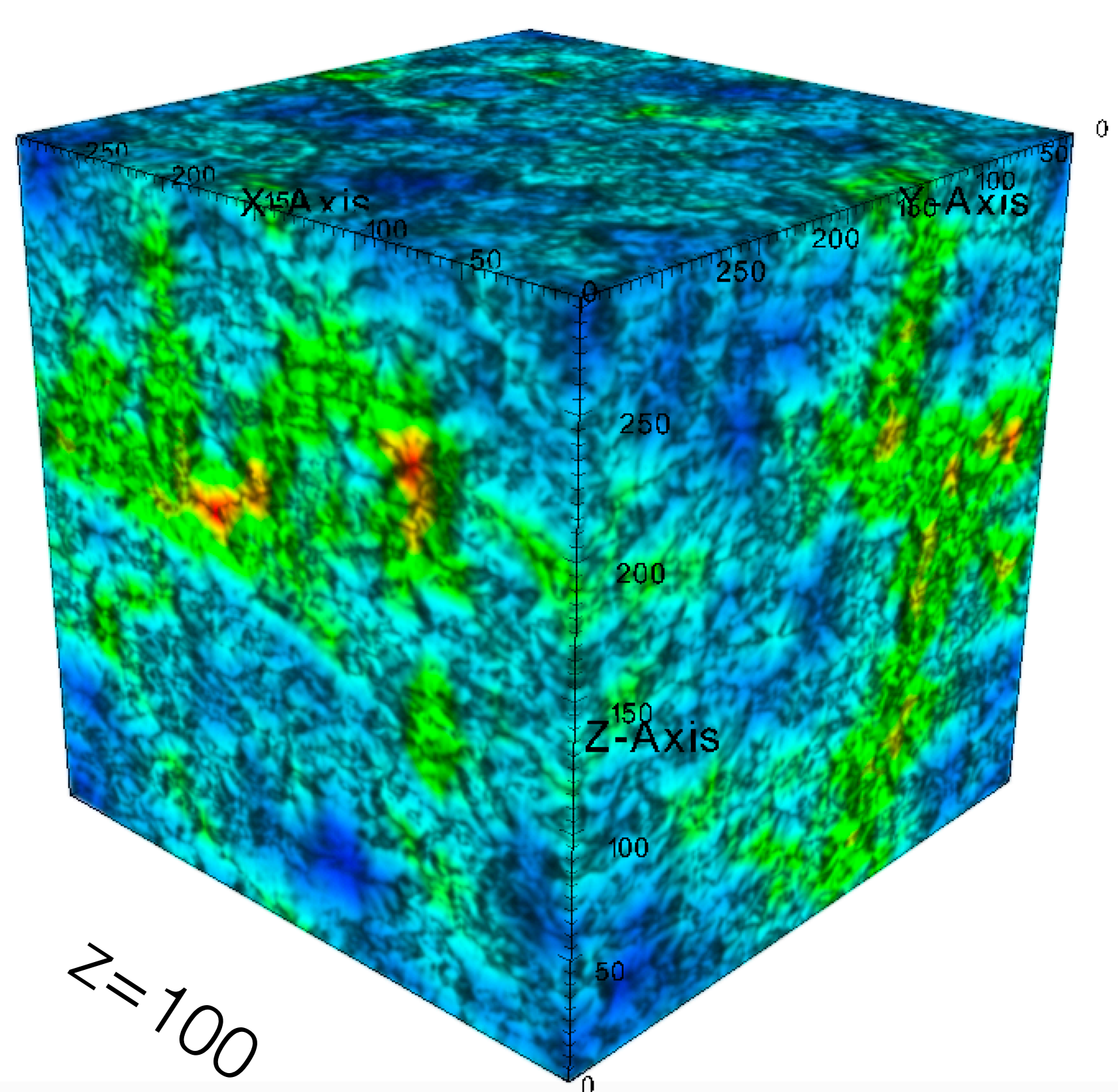
- The non-linearities affect k-essence field clustering and therefore resulting in different matter power spectrum
- Φ' which is more important at large scales is the source of ζ which changes total stress tensor and reduces the matter clustering at large scales

Can non-linearities filter dark energy models?



- Adding non-linearities can make the system unstable
- $\mathcal{H}\pi$ decays but ζ grows and $\delta_{\text{k-essence}}$ becomes large which causes the particles speed up

K-essence field clustering



Acknowledgement:

I would like to thank Jean-Pierre Eckmann, Mona Jalilvand, Ruth Durrer, Miguel Zumalcarregui, Joyce Byun, Yanchuan Cai and Cyril Pitrou. This work is done with financial support from the Swiss National Science Foundation and was supported by a grant from the Swiss National Supercomputing Centre (CSCS) under project IDs 710. The computations were performed on Piz Daint at CSCS.



UNIVERSITÉ
DE GENÈVE

FACULTÉ DES SCIENCES
Département de physique théorique

