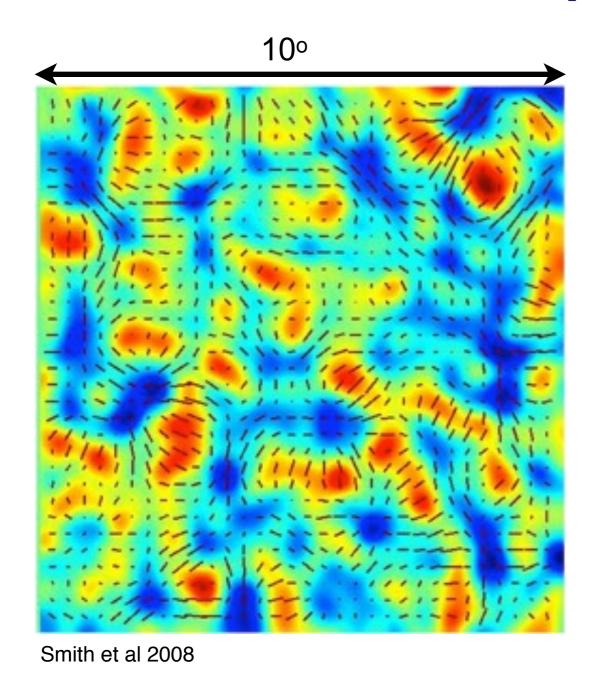
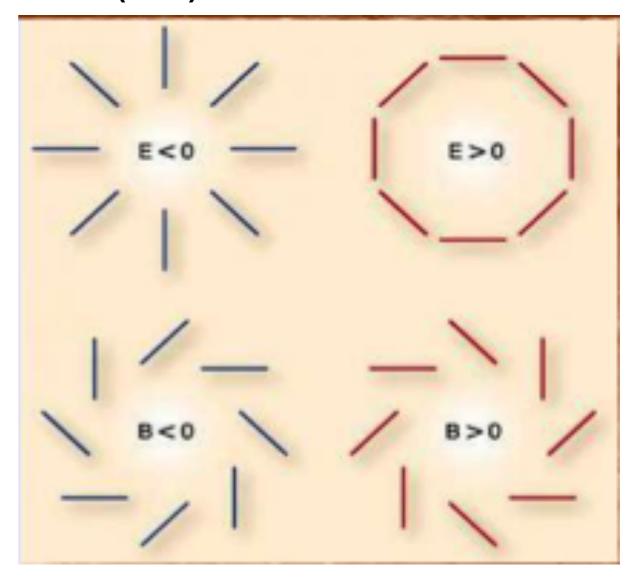


Christian Reichardt University of Melbourne

The CMB is polarised (~10%)

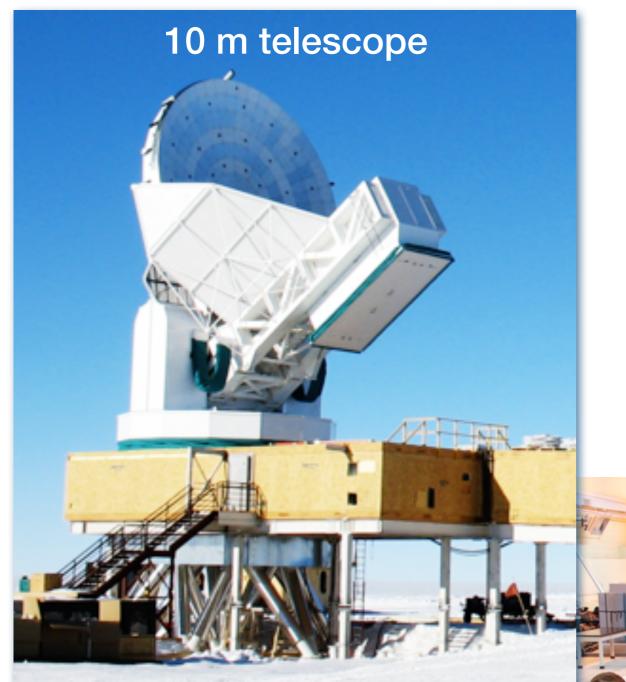


 Any polarisation pattern can be decomposed into "E" (grad) and "B" (curl) modes



 Density fluctuations at LSS do not produce "B" modes!

The South Pole Telescope (SPT)



Sub-millimeter Wavelength Telescope:

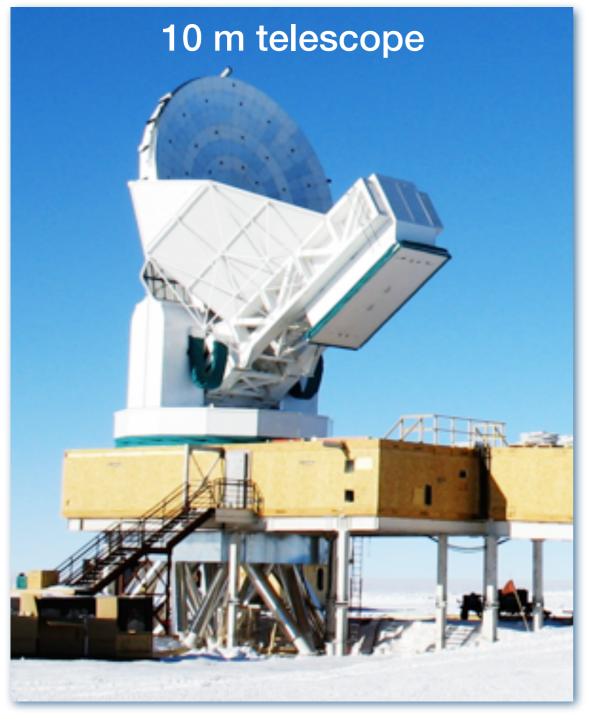
- 10 meter telescope (1.1' FWHM beam)
- Off-axis Gregorian optics design
- Fast scanning (up to 2 deg/sec in azimuth)

Collaboration

2" pointing accuracy

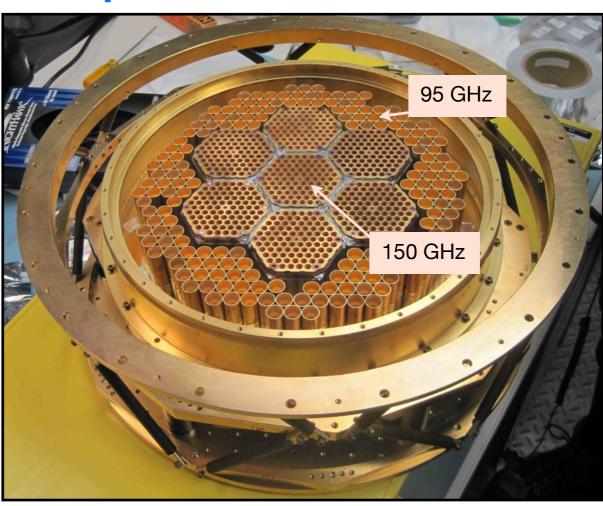


The South Pole Telescope (SPT)





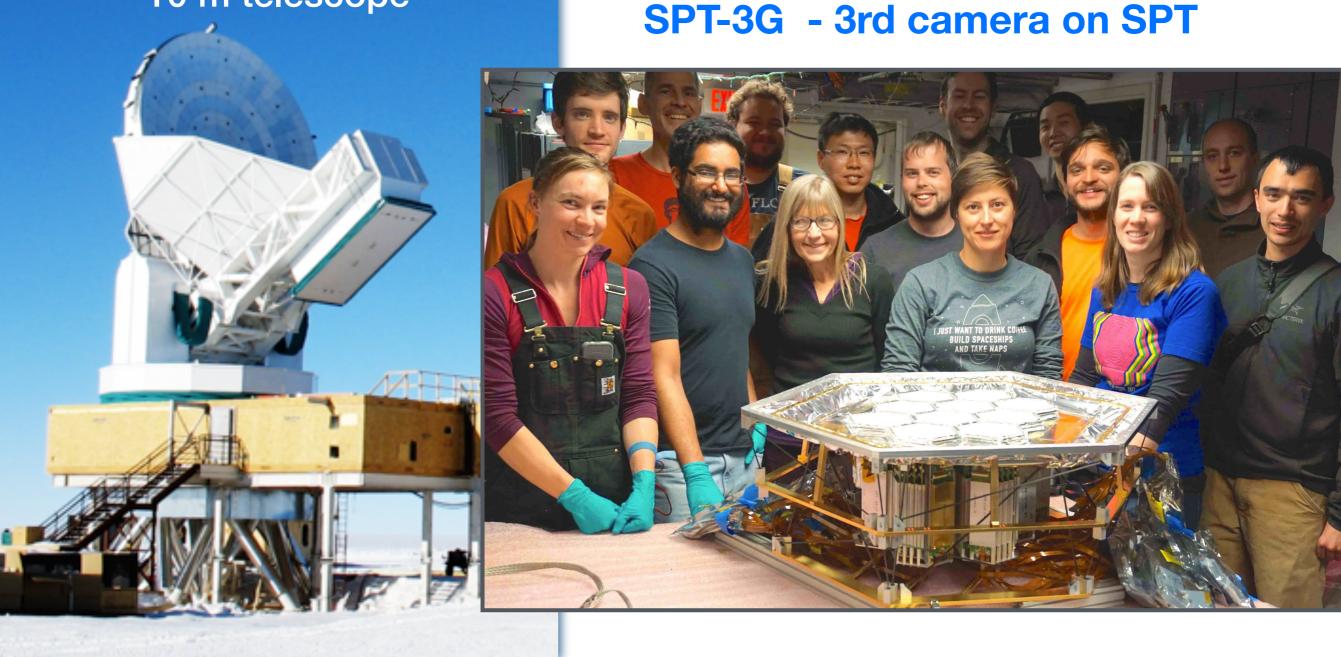
SPTpol - 2nd camera on SPT



- 1536 *polarization-sensitive* bolometers
- 500 deg² survey
 8 μK-arcmin @ 95 GHz
 5 μK-arcmin @ 150 GHz

The South Pole Telescope (SPT)

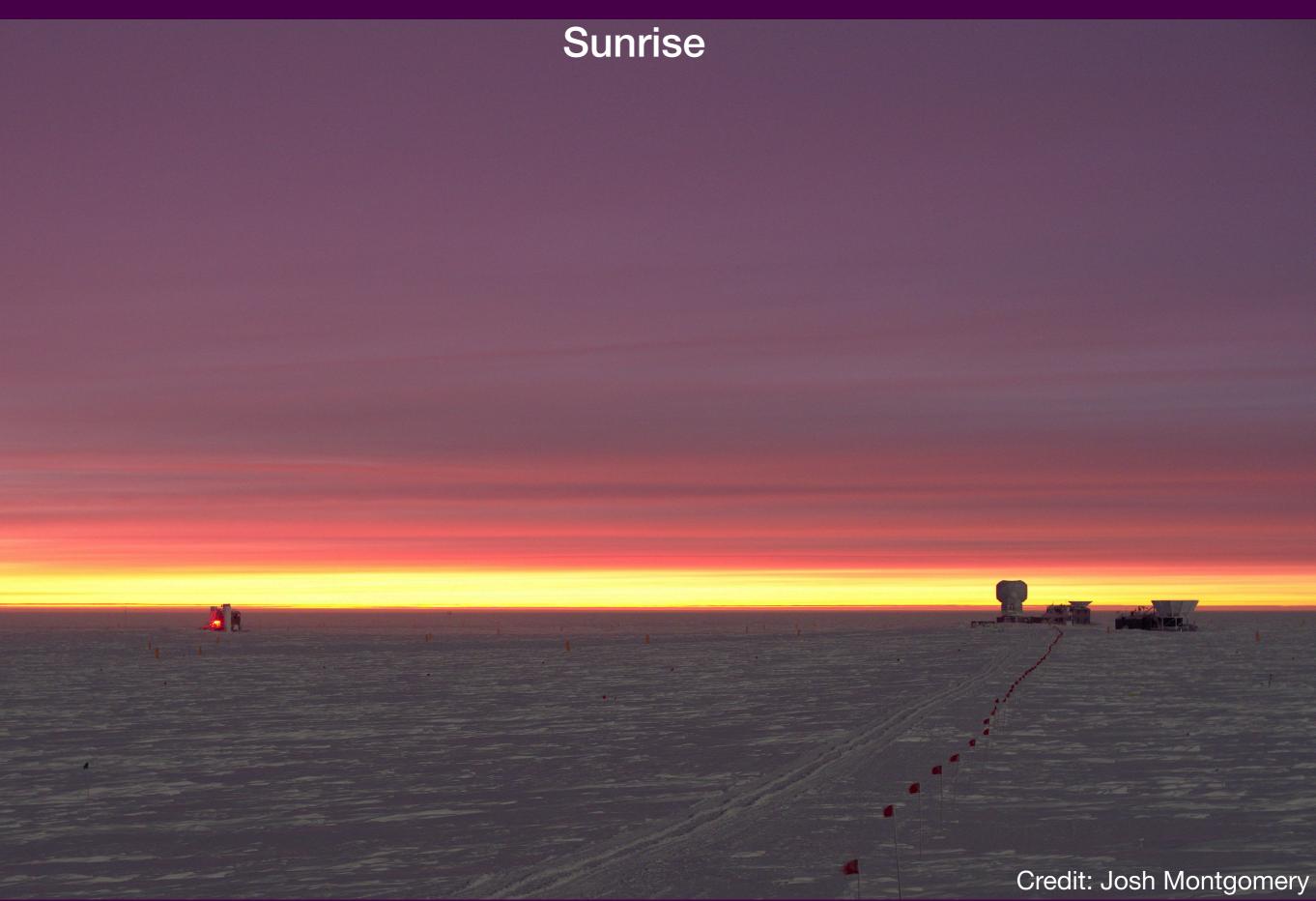




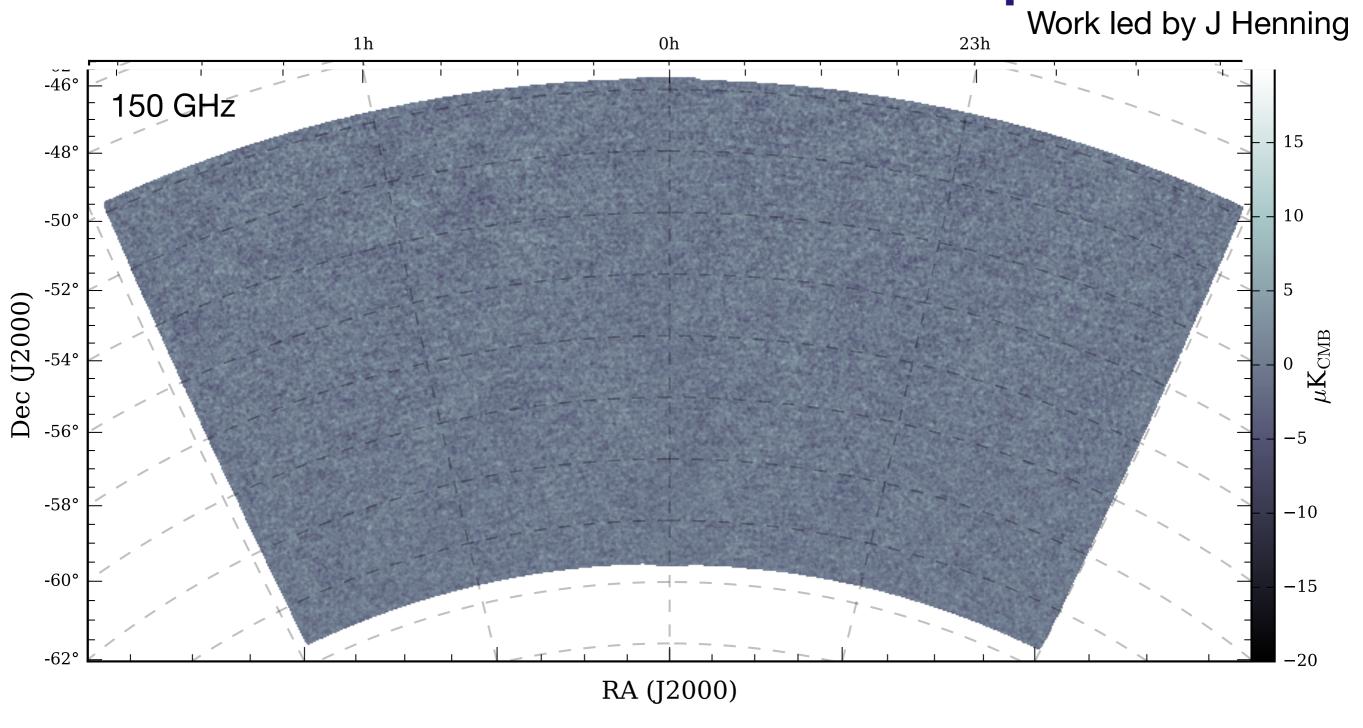


- 10x more bolometers (16,200)
- 1500 deg² survey in 95, 150, 220 GHz
- Currently in 2nd year of observations

Latest from the Pole



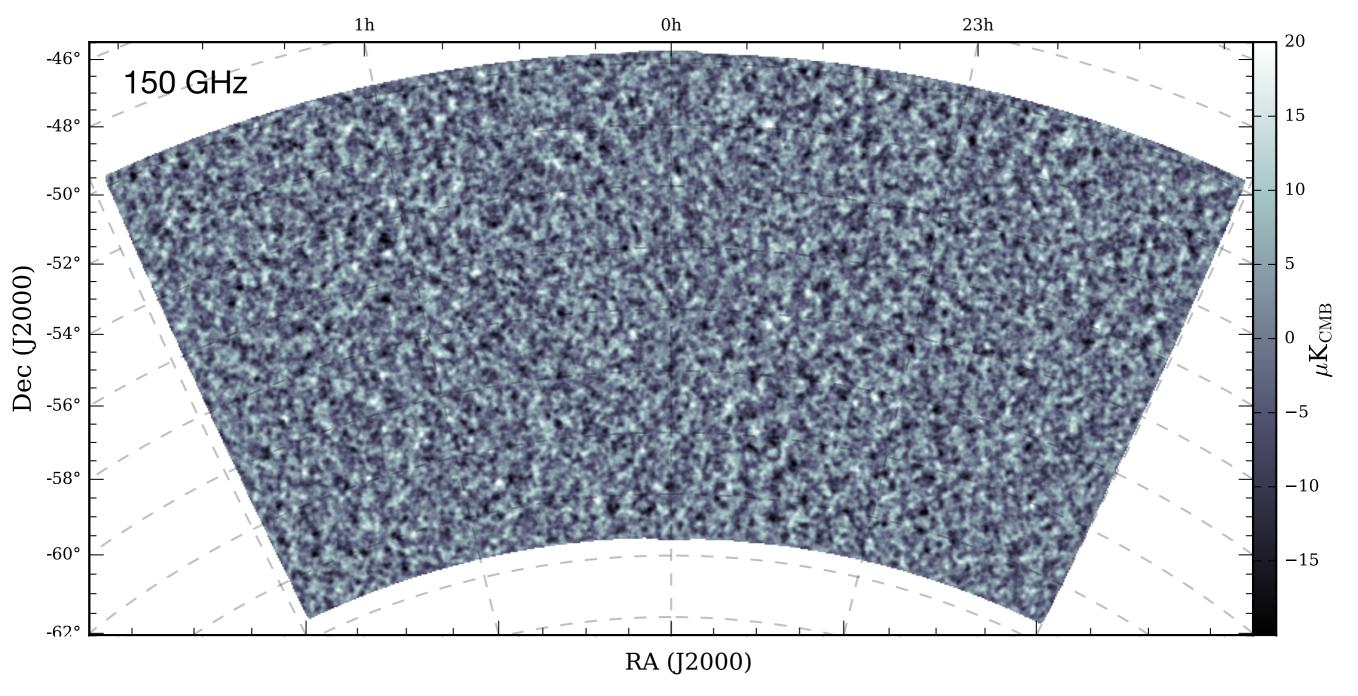
E-modes - A noise map



- Map from the first half of the season minus the second half
- Noise: 9.4 μK-arcmin between 1000 < ℓ < 3000.

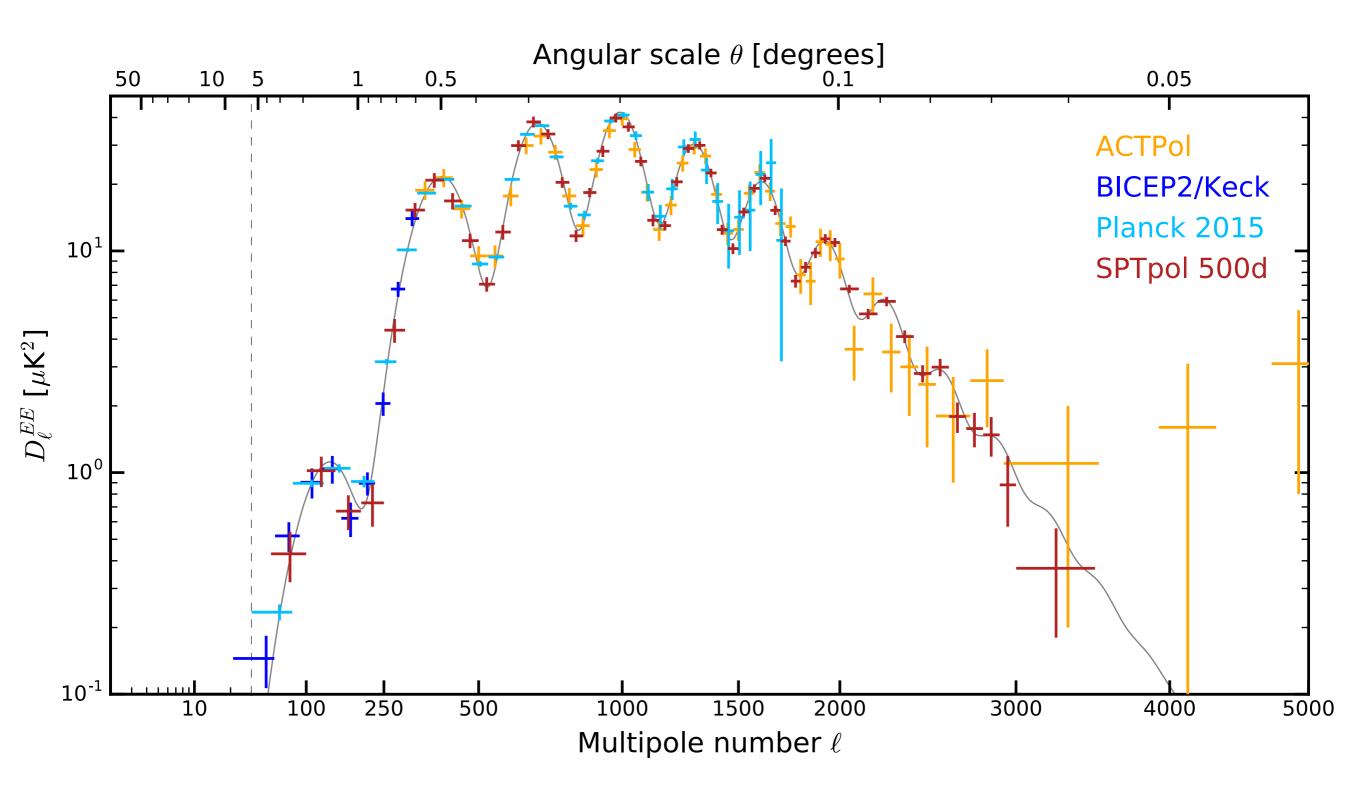
E-modes measured with SPTpol

Work led by J Henning



- High signal-to-noise!

E-mode power spectra



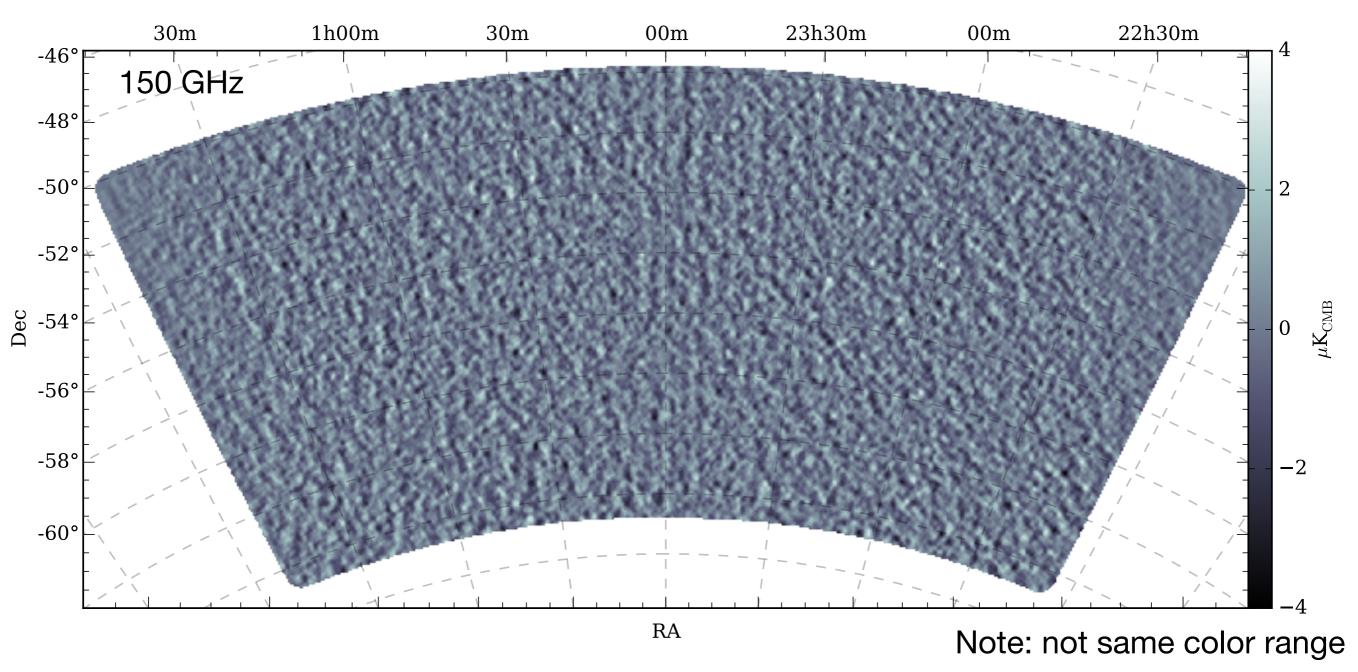
Best measurements at ell > 1050!

But what about B? (with an focus on r)

- Direct measurements
- Delensing for inflation

B-modes measured with SPTpol

Work led by JT Sayre

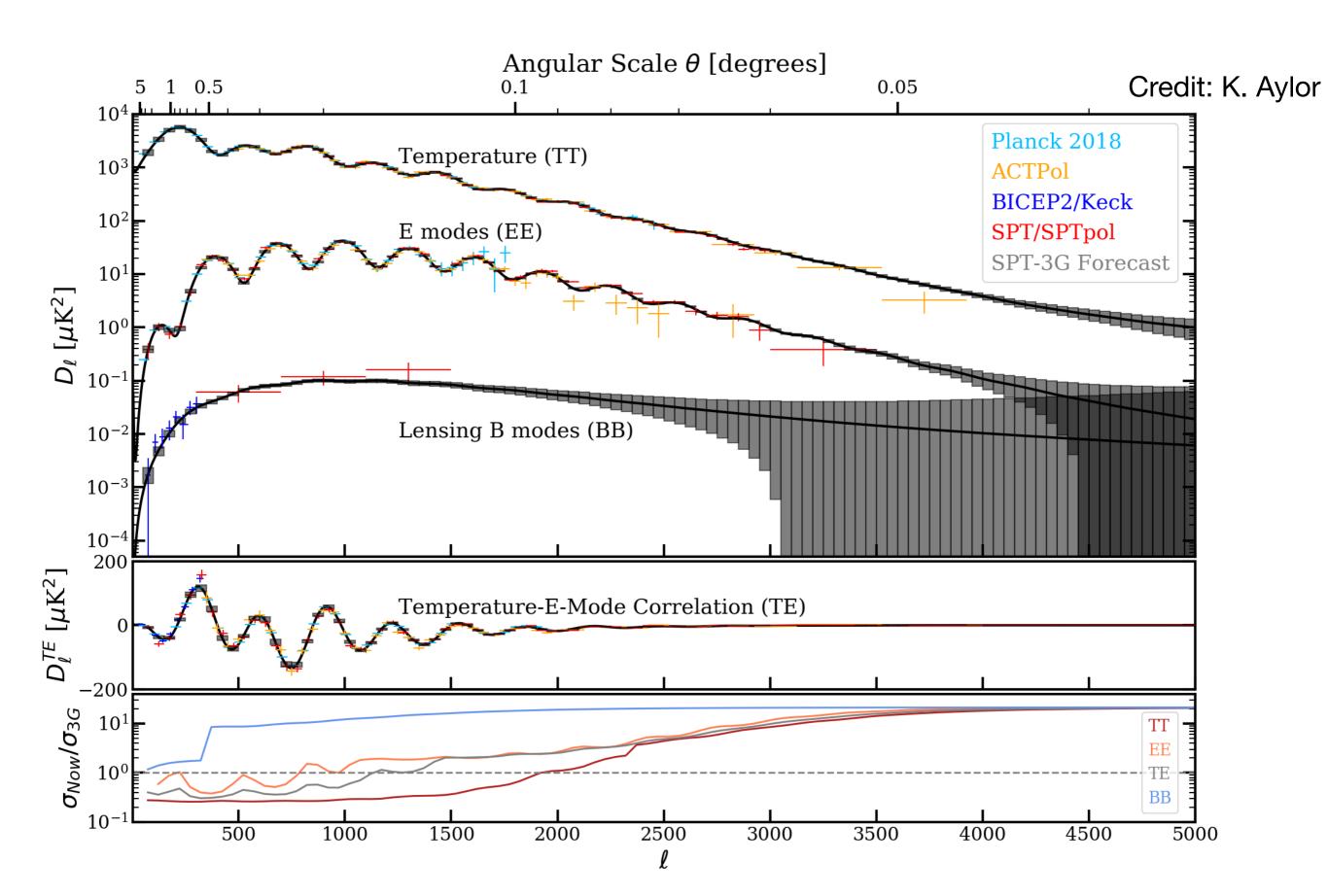


Signal-to-noise < 1 per mode (but lots of modes)

Also have 95 GHz maps

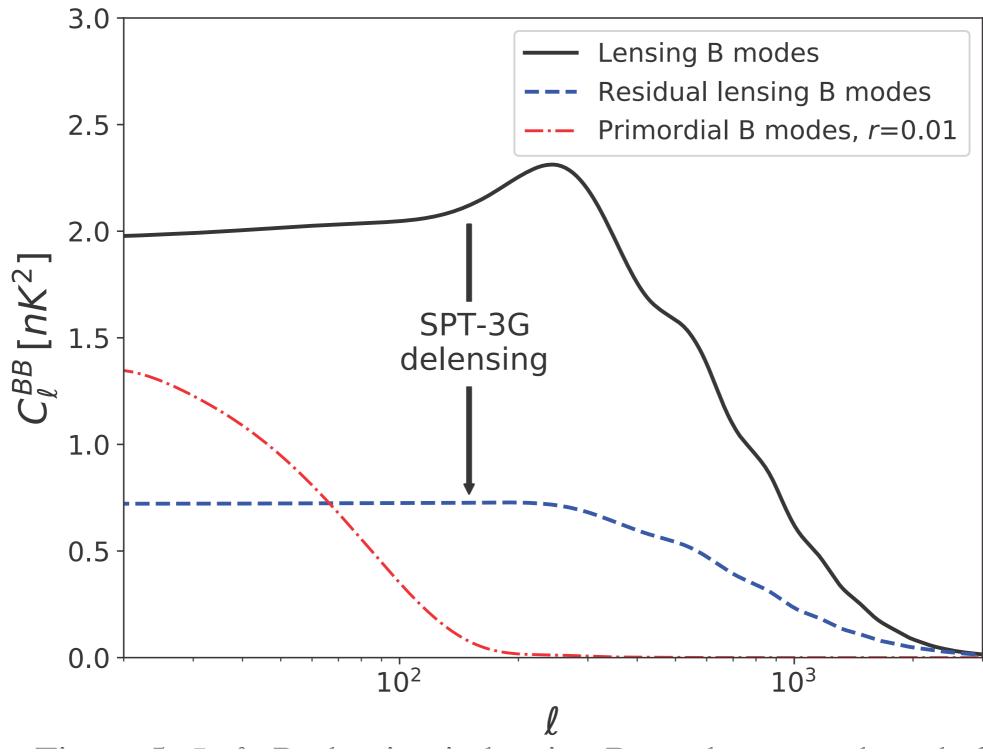
New SPTpol B-mode power spectrum coming soon!

SPT-3G



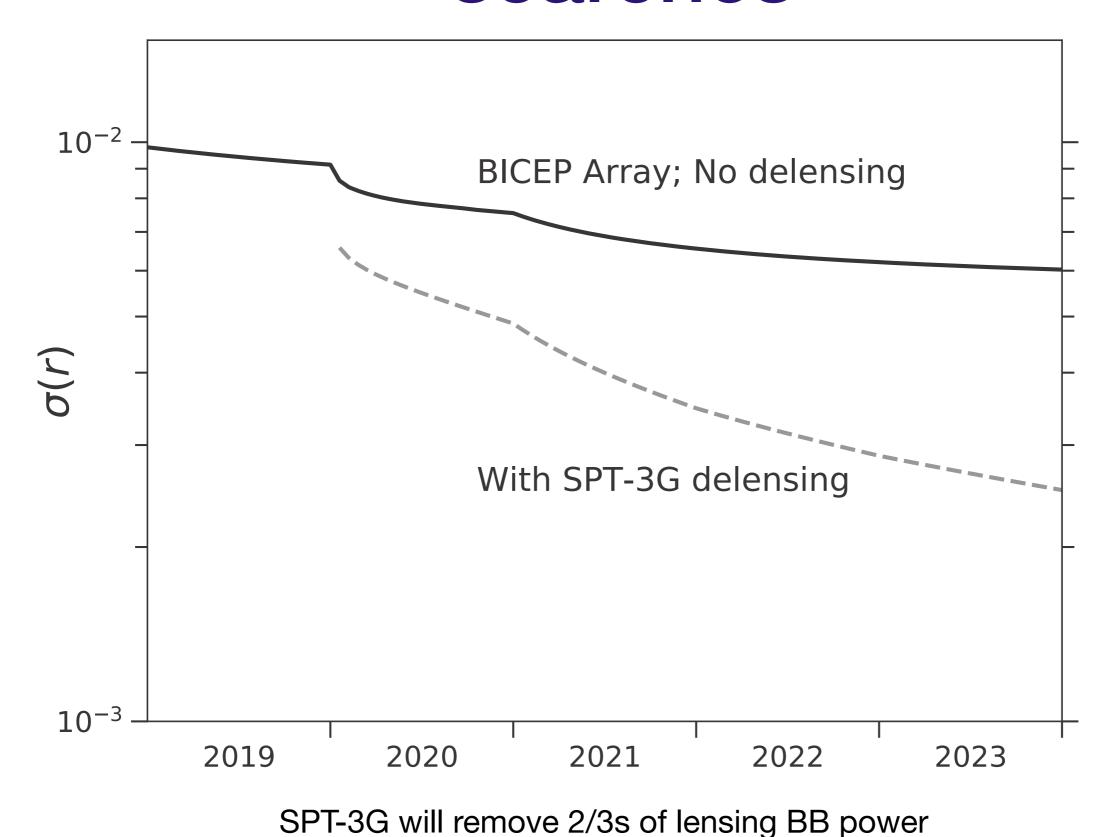
Delensing

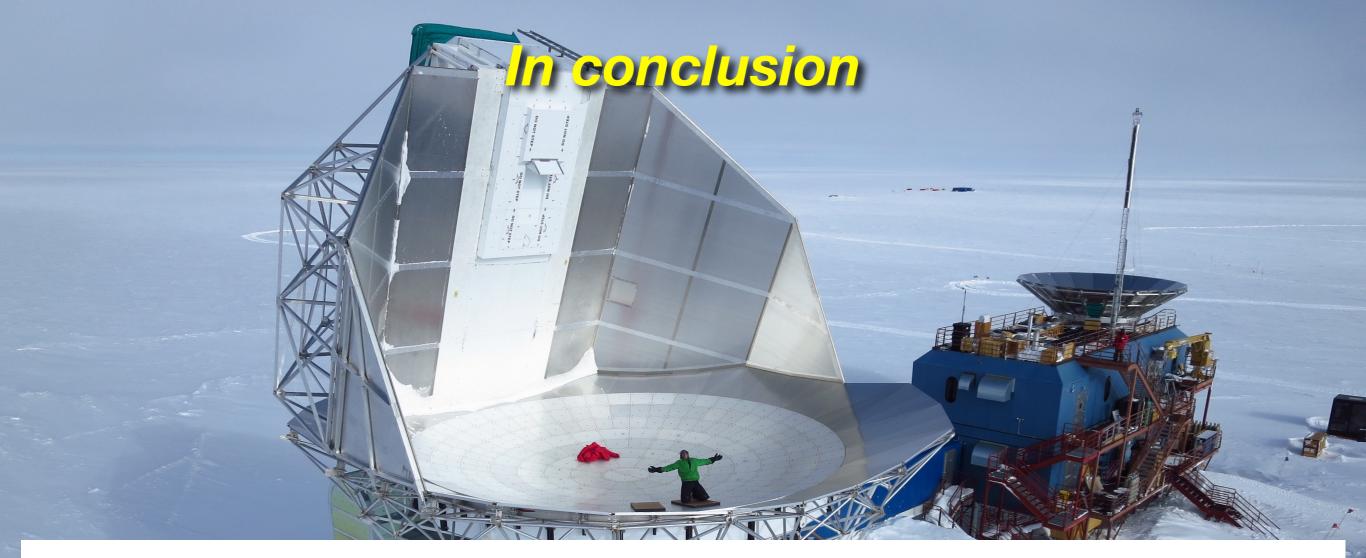
Last COSMO: Alessandro showed delensing of 24% on the SPTpol 100d field (Manzotti et al. ApJ 2017)



SPT-3G will remove 2/3s of lensing BB power

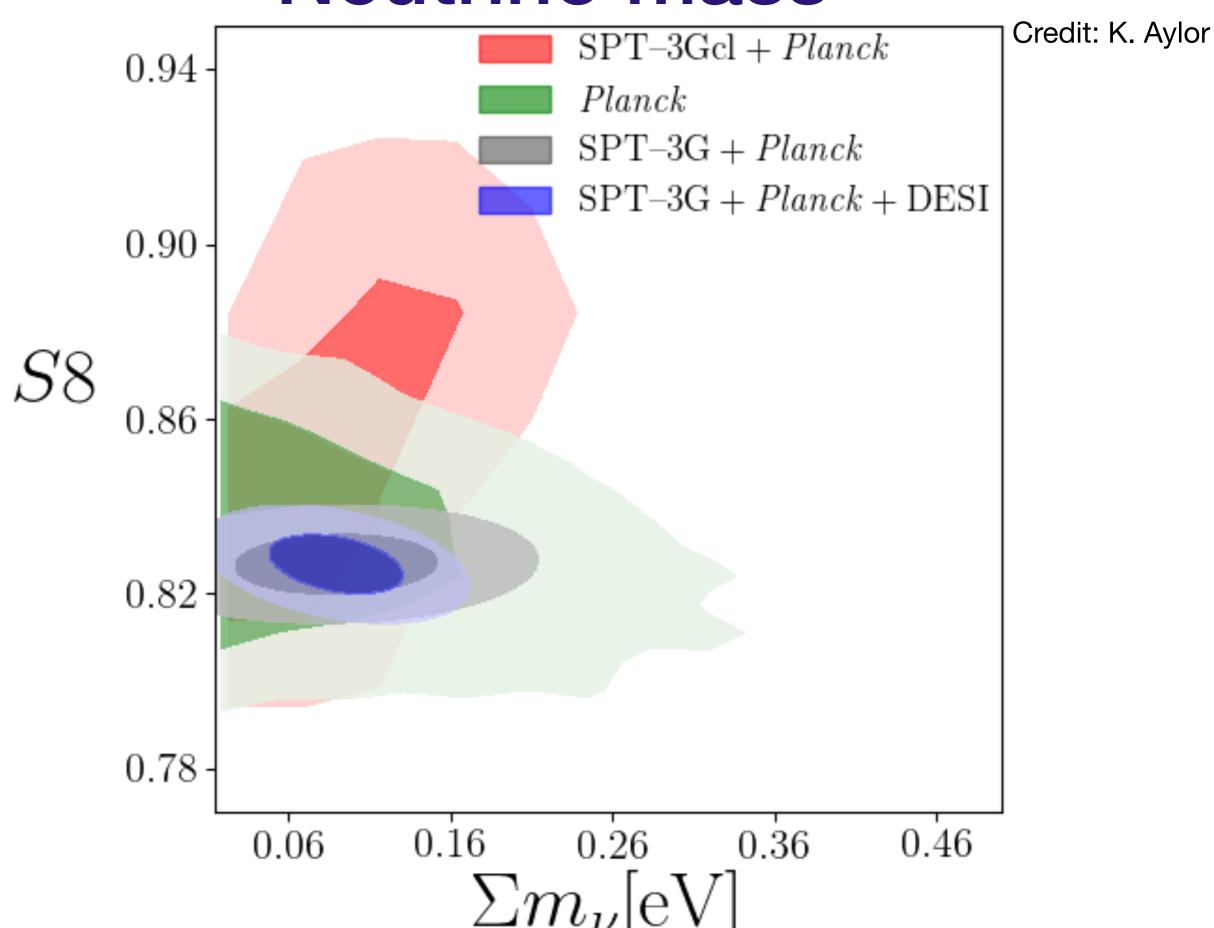
New era: delensing crucial to IGW searches



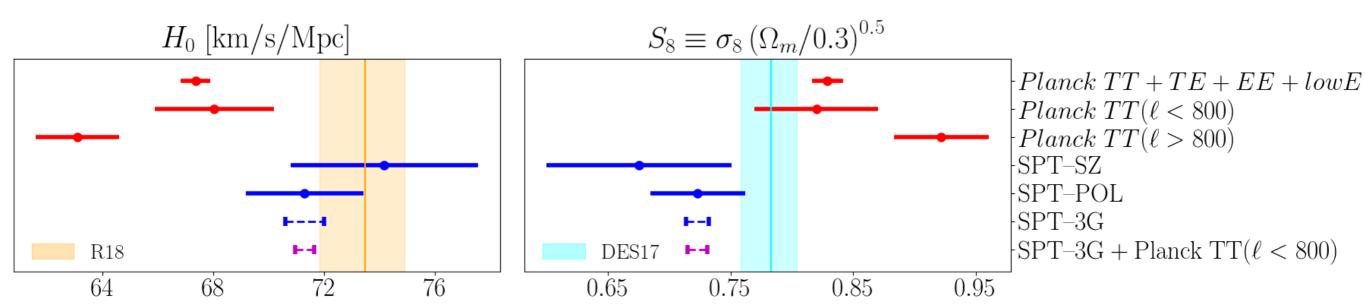


- SPTpol has made the best measurements of the E-mode damping tail
 - New B-mode spectra from SPTpol coming soon!
- SPT3G is on sky with x10 more detectors
 - significant upgrade in sensitivity!

Neutrino mass



Hubble constant tension



Credit: K. Aylor