

Einstein Double Field Equations

Stephen Angus, Kyoungho Cho, and Jeong-Hyuck Park

Department of Physics, Sogang University, 35 Baekbeom-ro, Mapo-gu, Seoul 04107, KOREA



Core idea: string theory predicts its own gravity rather than GR

In General Relativity the metric $g_{\mu\nu}$ is the only geometric and gravitational field, whereas in string theory the closed-string massless sector comprises a two-form potential $B_{\mu\nu}$ and the string dilaton ϕ in addition to the metric $g_{\mu\nu}$. Furthermore, these three fields transform into each other under T-duality. This hints at a natural augmentation of GR: upon treating the whole closed string massless sector as stringy graviton fields, Double Field Theory [1, 2] may evolve into ‘Stringy Gravity’. Equipped with an $O(D, D)$ covariant differential geometry beyond Riemann [3], we spell out the definitions of the stringy Einstein curvature tensor and the stringy Energy-Momentum tensor. Equating them, all the equations of motion of the closed string massless sector are unified into a single expression [4],

$$G_{AB} = 8\pi GT_{AB} \quad (1)$$

which we dub the **Einstein Double Field Equations**.

Double Field Theory as Stringy Gravity

• Built-in symmetries & Notation:

- $O(D, D)$ T-duality
- DFT diffeomorphisms (ordinary diffeomorphisms plus B -field gauge symmetry)
- Twofold local Lorentz symmetries, $\text{Spin}(1, D-1) \times \text{Spin}(D-1, 1)$
- ⇒ Two locally inertial frames exist separately for the left and the right modes.

Index	Representation	Metric (raising/lowering indices)
A, B, \dots, M, N, \dots	$O(D, D)$ vector	$\mathcal{J}_{AB} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$
p, q, \dots	$\text{Spin}(1, D-1)$ vector	$\eta_{pq} = \text{diag}(- + + \dots +)$
α, β, \dots	$\text{Spin}(1, D-1)$ spinor	$C_{\alpha\beta}, (\gamma^p)^T = C\gamma^p C^{-1}$
\bar{p}, \bar{q}, \dots	$\text{Spin}(D-1, 1)$ vector	$\bar{\eta}_{\bar{p}\bar{q}} = \text{diag}(+ - - \dots -)$
$\bar{\alpha}, \bar{\beta}, \dots$	$\text{Spin}(D-1, 1)$ spinor	$\bar{C}_{\bar{\alpha}\bar{\beta}}, (\bar{\gamma}^{\bar{p}})^T = \bar{C}\bar{\gamma}^{\bar{p}}\bar{C}^{-1}$

The $O(D, D)$ metric \mathcal{J}_{AB} divides doubled coordinates into two: $x^A = (\tilde{x}_\mu, x^\nu)$, $\partial_A = (\tilde{\partial}^\mu, \partial_\nu)$.

• Doubled-yet-gauged spacetime:

The doubled coordinates are ‘gauged’ through a certain equivalence relation, $x^A \sim x^A + \Delta^A$, such that each equivalence class, or gauge orbit in \mathbb{R}^{D+D} , corresponds to a single physical point in \mathbb{R}^D [5]. This implies a section condition, $\partial_A \partial^A = 0$, which can be conveniently solved by setting $\tilde{\partial}^\mu \equiv 0$.

• Stringy graviton fields (closed-string massless sector), $\{d, V_{Mp}, \bar{V}_{N\bar{q}}\}$:

Defining properties of the DFT-metric,

$$\mathcal{H}_{MN} = \mathcal{H}_{NM}, \quad \mathcal{H}_K^L \mathcal{H}_M^N \mathcal{J}_{LN} = \mathcal{J}_{KM}, \quad (2)$$

set a pair of symmetric and orthogonal projectors,

$$P_{MN} = P_{NM} = \frac{1}{2}(\mathcal{J}_{MN} + \mathcal{H}_{MN}), \quad P_L^M P_M^N = P_L^N, \\ \bar{P}_{MN} = \bar{P}_{NM} = \frac{1}{2}(\mathcal{J}_{MN} - \mathcal{H}_{MN}), \quad \bar{P}_L^M \bar{P}_M^N = \bar{P}_L^N, \quad P_L^M \bar{P}_M^N = 0.$$

Further, taking the ‘square roots’ of the projectors, we acquire a pair of DFT vielbeins,

$$P_{MN} = V_M^p V_N^q \eta_{pq}, \quad \bar{P}_{MN} = \bar{V}_M^{\bar{p}} \bar{V}_N^{\bar{q}} \bar{\eta}_{\bar{p}\bar{q}},$$

satisfying their own defining properties,

$$V_{Mp} V^M{}_q = \eta_{pq}, \quad \bar{V}_{M\bar{p}} \bar{V}^M{}_{\bar{q}} = \bar{\eta}_{\bar{p}\bar{q}}, \quad V_{Mp} \bar{V}^M{}_{\bar{q}} = 0, \quad V_M^p V_N^q = \mathcal{J}_{MN}.$$

The most general solutions to (2) can be classified by two non-negative integers (n, \bar{n}) [6],

$$\mathcal{H}_{MN} = \begin{pmatrix} H^{\mu\nu} & -H^{\mu\sigma} B_{\sigma\lambda} + Y_i^\mu X_\lambda^i - \bar{Y}_i^\mu \bar{X}_\lambda^i \\ B_{\kappa\rho} H^{\rho\nu} + X_i^\kappa Y_i^\nu - \bar{X}_i^\kappa \bar{Y}_i^\nu & K_{\kappa\lambda} - B_{\kappa\rho} H^{\rho\sigma} B_{\sigma\lambda} + 2X_i^\kappa B_{\lambda\rho} Y_i^\rho - 2\bar{X}_i^\kappa B_{\lambda\rho} \bar{Y}_i^\rho \end{pmatrix}$$

where $1 \leq i \leq n$, $1 \leq \bar{i} \leq \bar{n}$ and

$$H^{\mu\nu} X_\nu^i = 0, \quad H^{\mu\nu} \bar{X}_\nu^i = 0, \quad K_{\mu\nu} Y_i^\nu = 0, \quad K_{\mu\nu} \bar{Y}_i^\nu = 0, \quad H^{\mu\rho} K_{\rho\nu} + Y_i^\mu X_\nu^i + \bar{Y}_i^\mu \bar{X}_\nu^i = \delta_\nu^\mu.$$

Strings become chiral and anti-chiral over n and \bar{n} directions: $X_i^\mu \partial_+ x^\mu = 0$, $\bar{X}_i^\mu \partial_- x^\mu = 0$. Examples include $(0, 0)$ Riemannian geometry as $K_{\mu\nu} = g_{\mu\nu}$, $H^{\mu\nu} = g^{\mu\nu}$, $(1, 1)$ Gomis-Ooguri non-relativistic background, $(1, 0)$ Newton-Cartan gravity, and $(D-1, 0)$ Carroll gravity.

• Covariant derivative:

The ‘master’ covariant derivative, $\mathcal{D}_A = \partial_A + \Gamma_A + \Phi_A + \bar{\Phi}_A$, is characterized by compatibility:

$$\mathcal{D}_A d = \mathcal{D}_A V_{Bp} = \mathcal{D}_A \bar{V}_{B\bar{p}} = 0, \quad \mathcal{D}_A \mathcal{J}_{BC} = \mathcal{D}_A \eta_{pq} = \mathcal{D}_A \bar{\eta}_{\bar{p}\bar{q}} = \mathcal{D}_A C_{\alpha\beta} = \mathcal{D}_A \bar{C}_{\bar{\alpha}\bar{\beta}} = 0.$$

The stringy Christoffel symbols are [3]

$$\Gamma_{CAB} = 2(P\partial_C P\bar{P})_{[AB]} + 2(\bar{P}_{[A}^D \bar{P}_{B]}^E - P_{[A}^D P_{B]}^E) \partial_D P_{EC} \\ - 4\left(\frac{1}{P_{[M}^N - 1} P_{C[A} P_{B]}^D + \frac{1}{\bar{P}_{[M}^N - 1} \bar{P}_{C[A} \bar{P}_{B]}^D\right) (\partial_D d + (P\partial^E P\bar{P})_{[ED]}),$$

and the spin connections are $\Phi_{Apq} = V^B{}_p (\partial_A V_{Bq} + \Gamma_{AB}^C V_{Cq})$, $\bar{\Phi}_{A\bar{p}\bar{q}} = \bar{V}^B{}_{\bar{p}} (\partial_A \bar{V}_{B\bar{q}} + \Gamma_{AB}^C \bar{V}_{C\bar{q}})$. In Stringy Gravity, there are no normal coordinates where Γ_{CAB} would vanish point-wise: the Equivalence Principle holds for point particles but is generically broken for strings (i.e. extended objects).

• Scalar and ‘Ricci’ curvatures:

The semi-covariant Riemann curvature in Stringy Gravity is defined by

$$S_{ABCD} := \frac{1}{2} (R_{ABCD} + R_{CDAB} - \Gamma_{AB}^E \Gamma_{ECD}),$$

where $R_{CDAB} = \partial_A \Gamma_{BCD} - \partial_B \Gamma_{ACD} + \Gamma_{ACE} \Gamma_B^E D - \Gamma_{BCE} \Gamma_A^E D$ (the ‘field strength’ of Γ_{CAB}). The completely covariant ‘Ricci’ and scalar curvatures are, with $S_{AB} = S_{ACB}^C$,

$$S_{p\bar{q}} := V^A{}_p \bar{V}^B{}_{\bar{q}} S_{AB}, \quad S_{(0)} := (P^A C P^B D - \bar{P}^A \bar{C} \bar{P}^B D) S_{ABCD}.$$

While $e^{-2d} S_{(0)}$ corresponds to the original DFT Lagrangian density [1, 2], or the ‘pure’ Stringy Gravity, the master covariant derivative fixes its minimal coupling to extra matter fields, e.g. type II maximally supersymmetric DFT [7] or the Standard Model [8].

Derivation of Einstein Double Field Equations

Variation of the action for Stringy Gravity coupled to generic matter fields, Υ_a , gives

$$\delta \int e^{-2d} \left[\frac{1}{16\pi G} S_{(0)} + L_{\text{matter}} \right] \\ = \int e^{-2d} \left[\frac{1}{4\pi G} \bar{V}^A \bar{q} \delta V_A^p (S_{p\bar{q}} - 8\pi G K_{p\bar{q}}) - \frac{1}{8\pi G} \delta d (S_{(0)} - 8\pi G T_{(0)}) + \delta \Upsilon_a \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} \right] \\ = \int e^{-2d} \left[\frac{1}{8\pi G} \xi^B \mathcal{D}^A \{G_{AB} - 8\pi G T_{AB}\} + (\hat{\mathcal{L}} \xi \Upsilon_a) \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} \right],$$

where the second line is for generic variations and the third line is specifically for diffeomorphic transformations. We are naturally led to define

$$K_{p\bar{q}} := \frac{1}{2} \left(V_{Ap} \frac{\delta L_{\text{matter}}}{\delta V_A^q} - \bar{V}_{A\bar{q}} \frac{\delta L_{\text{matter}}}{\delta V_A^p} \right), \quad T_{(0)} := e^{2d} \times \frac{\delta (e^{-2d} L_{\text{matter}})}{\delta d},$$

and subsequently the stringy Einstein curvature, G_{AB} , and Energy Momentum tensor, T_{AB} ,

$$G_{AB} = 4V_{[A}^p \bar{V}_{B]}^{\bar{q}} S_{p\bar{q}} - \frac{1}{2} \mathcal{J}_{AB} S_{(0)}, \quad \mathcal{D}_A G^{AB} = 0 \quad (\text{off-shell}), \\ T_{AB} := 4V_{[A}^p \bar{V}_{B]}^{\bar{q}} K_{p\bar{q}} - \frac{1}{2} \mathcal{J}_{AB} T_{(0)}, \quad \mathcal{D}_A T^{AB} = 0 \quad (\text{on-shell}).$$

The equations of motion of the stringy graviton fields are thus unified into a single expression, the Einstein Double Field Equations (1). Note that $G_A^A = -DS_{(0)}$, $T_A^A = -DT_{(0)}$.

Restricting to the $(0, 0)$ Riemannian background, the Einstein Double Field Equations reduce to

$$R_{\mu\nu} + 2\nabla_\mu (\partial_\nu \phi) - \frac{1}{4} H_{\mu\rho\sigma} H_\nu{}^{\rho\sigma} = 8\pi G K_{(\mu\nu)}, \\ \nabla^\rho (e^{-2\phi} H_{\rho\mu\nu}) = 16\pi G e^{-2\phi} K_{[\mu\nu]},$$

$$R + 4\Box\phi - 4\partial_\mu \phi \partial^\mu \phi - \frac{1}{12} H_{\lambda\mu\nu} H^{\lambda\mu\nu} = 8\pi G T_{(0)},$$

which imply the conservation law, $\mathcal{D}_A T^{AB} = 0$, given explicitly by

$$\nabla^\mu K_{(\mu\nu)} - 2\partial^\mu \phi K_{(\mu\nu)} + \frac{1}{2} H_\nu{}^{\lambda\mu} K_{[\lambda\mu]} - \frac{1}{2} \partial_\nu T_{(0)} = 0, \quad \nabla^\mu (e^{-2\phi} K_{[\mu\nu]}) = 0.$$

The Einstein Double Field Equations also govern the dynamics of other non-Riemannian cases, $(n, \bar{n}) \neq (0, 0)$, where the Riemannian metric, $g_{\mu\nu}$, cannot be defined.

Examples

– Pure Stringy Gravity with cosmological constant:

$$\frac{1}{16\pi G} e^{-2d} (S_{(0)} - 2\Lambda_{\text{DFT}}), \quad K_{p\bar{q}} = 0, \quad T_{(0)} = \frac{1}{4\pi G} \Lambda_{\text{DFT}}.$$

– RR sector, given by a $\text{Spin}(1, 9) \times \text{Spin}(9, 1)$ bi-spinorial potential, $C^\alpha_{\bar{\alpha}}$:

$$L_{\text{RR}} = \frac{1}{2} \text{Tr}(\mathcal{F}\bar{\mathcal{F}}), \quad K_{p\bar{q}} = -\frac{1}{4} \text{Tr}(\gamma_p \mathcal{F} \bar{\gamma}_{\bar{q}} \bar{\mathcal{F}}), \quad T_{(0)} = 0,$$

where $\mathcal{F} = \mathcal{D}_+ C + \gamma^p \mathcal{D}_p C + \gamma^{(11)} \mathcal{D}_{\bar{p}} C \bar{\gamma}^{\bar{p}}$ is the RR flux set by an $O(D, D)$ covariant ‘ H -twisted’ cohomology, $(\mathcal{D}_+)^2 = 0$, and $\bar{\mathcal{F}} = \bar{C}^{-1} \mathcal{F}^T C$ is its charge conjugate [7].

– Spinor field: $L_\psi = \bar{\psi} \gamma^p \mathcal{D}_p \psi + m_\psi \bar{\psi} \psi$, $K_{p\bar{q}} = -\frac{1}{4} (\bar{\psi} \gamma_p \mathcal{D}_{\bar{q}} \psi - \mathcal{D}_{\bar{q}} \bar{\psi} \gamma_p \psi)$, $T_{(0)} = 0$.

– Green-Schwarz superstring (κ -symmetric):

$$e^{-2d} L_{\text{string}} = \frac{1}{4\pi\alpha'} \int d^2\sigma \left[-\frac{1}{2} \sqrt{-h} h^{ij} \Pi_i^M \Pi_j^N \mathcal{H}_{MN} - \epsilon^{ij} D_i y^M (\mathcal{A}_{jM} - i\Sigma_{jM}) \right] \delta^D(x - y(\sigma)),$$

$$K_{p\bar{q}}(x) = \frac{1}{4\pi\alpha'} \int d^2\sigma \sqrt{-h} h^{ij} (\Pi_i^M V_{Mp}) (\Pi_j^N \bar{V}_{N\bar{q}}) e^{2d} \delta^D(x - y(\sigma)), \quad T_{(0)} = 0,$$

where $\Sigma_i^M = \bar{\theta} \gamma^M \partial_i \theta + \bar{\theta}' \gamma^M \partial_i \theta'$ and $\Pi_i^M = \partial_i y^M - \mathcal{A}_i^M - i\Sigma_i^M$ (doubled-yet-gauged) [9].

Gravitational effect

The regular spherical solution to the $D = 4$ Einstein Double Field Equations shows that Stringy Gravity modifies GR (Schwarzschild geometry), in particular at ‘short’ dimensionless scales, R/MG , i.e. distance normalized by mass times Newton constant. This might shed new light upon the dark matter/energy problems, as they arise essentially from ‘short distance’ observations. Furthermore, it would be intriguing to view the B -field and DFT dilaton d as ‘dark gravitons’, since they decouple from the geodesic motion of point particles, which should be defined in string frame [10].

	Electron ($R \approx 0$)	Proton	Hydrogen Atom	Billiard Ball	Earth	Solar System (1AU/ $M_\odot G$)	Milky Way (visible)	Galaxy Cluster	Universe ($M \propto R^3$)
$R/(MG)$	0^+	7.1×10^{38}	2.0×10^{43}	2.4×10^{26}	1.4×10^9	1.0×10^8	1.5×10^6	$\sim 10^5$	0^+

References

- [1] W. Siegel, ‘Two vierbein formalism for string inspired gravity,’ Phys. Rev. D **47** (1993) 5453.
- [2] C. Hull and B. Zwiebach, ‘Double Field Theory,’ JHEP **0909** (2009) 099 [arXiv:0904.4664].
- [3] I. Jeon, K. Lee and J. H. Park, ‘Stringy differential geometry, beyond Riemann,’ Phys. Rev. D **84** (2011) 044022 [arXiv:1105.6294 [hep-th]].
- [4] S. Angus, K. Cho and J. H. Park, ‘Einstein Double Field Equations,’ arXiv:1804.00964.
- [5] J. H. Park, ‘Comments on double field theory and diffeomorphisms,’ JHEP **1306** (2013) 098 [arXiv:1304.5946 [hep-th]].
- [6] K. Morand and J. H. Park, ‘Classification of non-Riemannian doubled-yet-gauged spacetime,’ Eur. Phys. J. C **77** (2017) no.10, 685 [arXiv:1707.03713 [hep-th]].
- [7] I. Jeon, K. Lee, J. H. Park and Y. Suh, ‘Stringy Unification of Type IIA and IIB Supergravities under $N = 2D = 10$ Supersymmetric Double Field Theory,’ Phys. Lett. B **723** (2013) 245 [arXiv:1210.5078 [hep-th]]. Twofold spin group, $\text{Spin}(1, 9) \times \text{Spin}(9, 1)$, unifies IIA and IIB.
- [8] K. S. Choi and J. H. Park, ‘Standard Model as a Double Field Theory,’ Phys. Rev. Lett. **115** (2015) no.17, 171603 [arXiv:1506.05277 [hep-th]].
- [9] J. H. Park, ‘Green-Schwarz superstring on doubled-yet-gauged spacetime,’ JHEP **1611** (2016) 005 [arXiv:1609.04265 [hep-th]].
- [10] S. M. Ko, J. H. Park and M. Suh, ‘The rotation curve of a point particle in stringy gravity,’ JCAP **1706** (2017) no.06, 002 [arXiv:1606.09307 [hep-th]].