

Axion Dark Matter

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✓ Axion(-like) Particle

Radio (SKA-like) survey

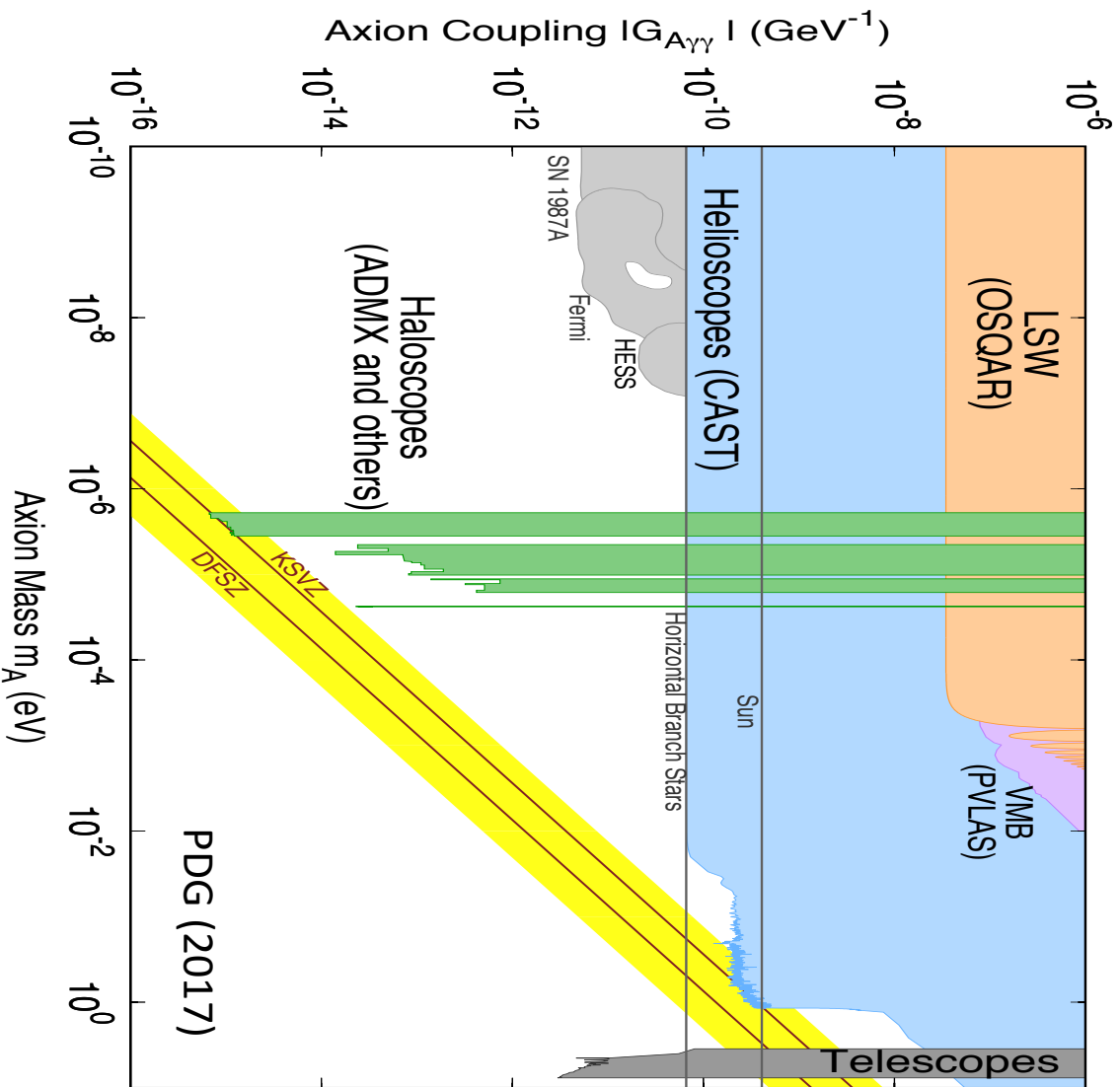
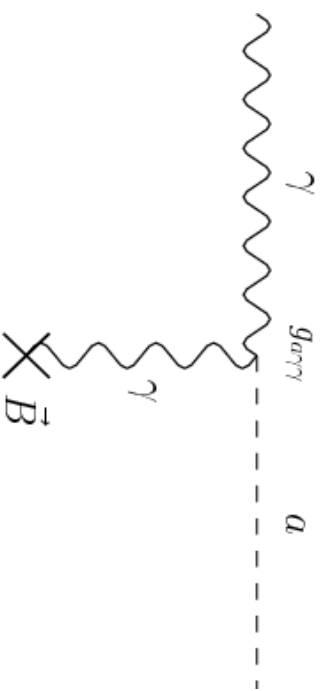
✓ Axino

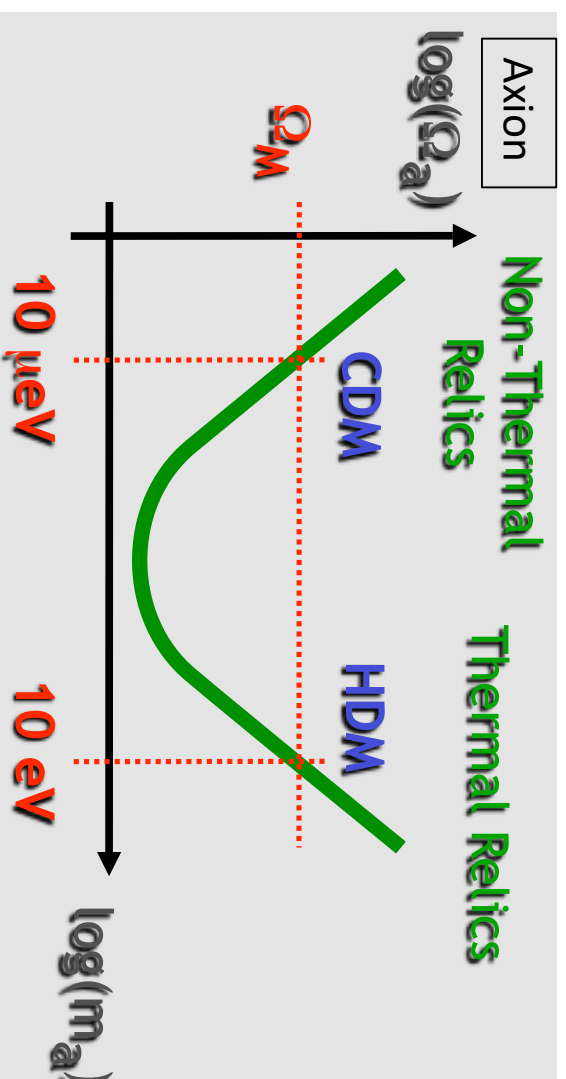
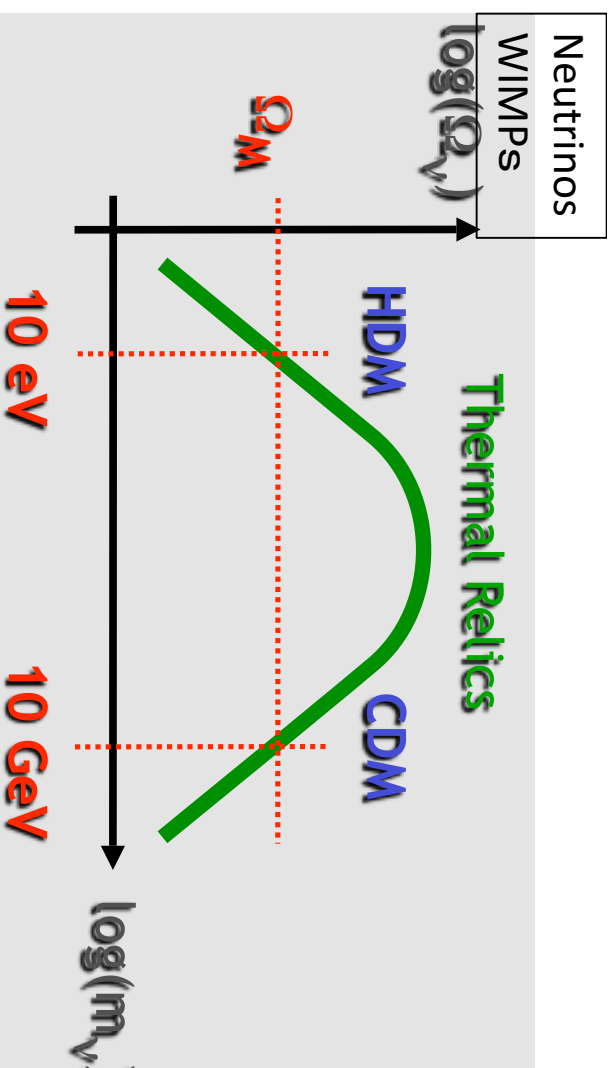
Fixed target experiment

➤ Conclusion

Primakoff effect

$$\frac{g_{a\gamma\gamma}}{4} a F \tilde{F} = - g_{a\gamma\gamma} a \mathbf{E} \cdot \mathbf{B}$$

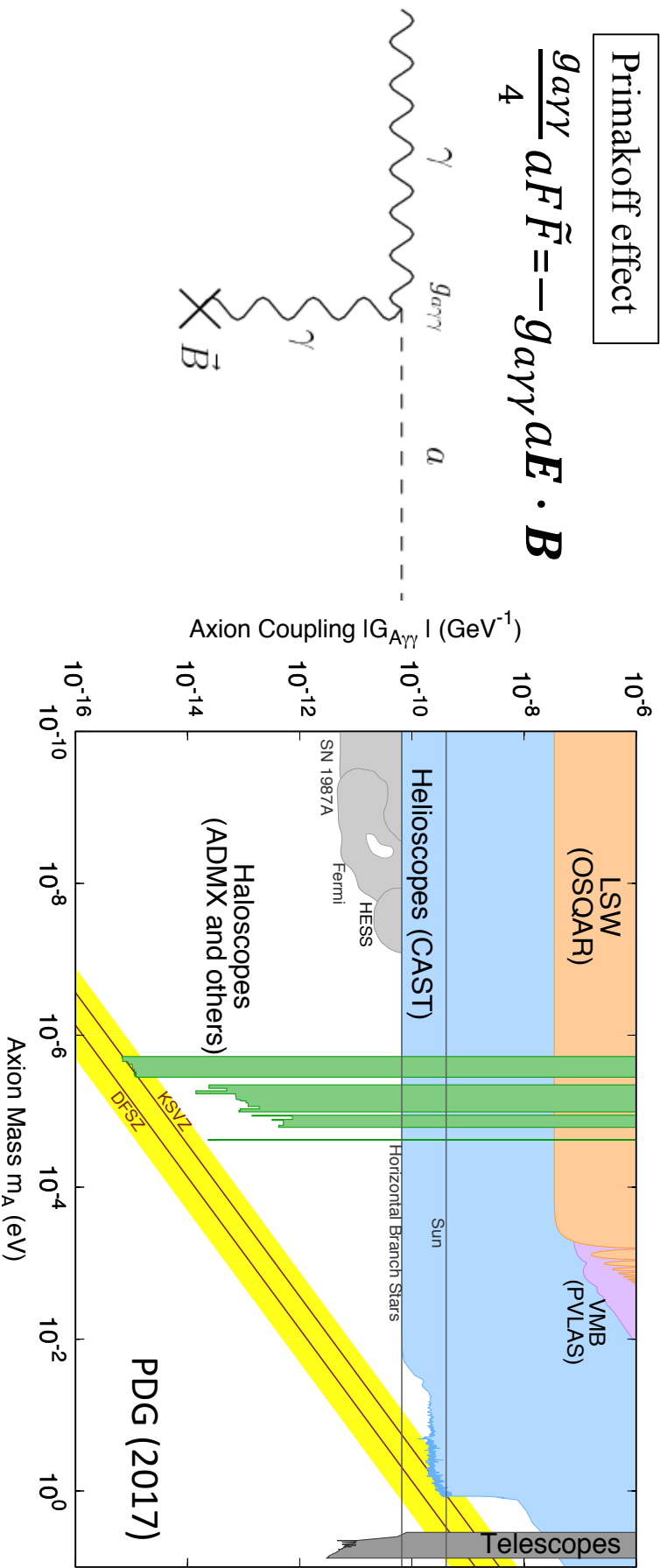




Taken from G. Raffelt

Primakoff effect

$$\frac{g_{a\gamma\gamma}}{4} a F \tilde{F} = -g_{a\gamma\gamma} a \mathbf{E} \cdot \mathbf{B}$$



QCD axion as a CDM candidate : mass range $\mu\text{eV} \sim \text{meV}$ ($0.1\text{GHz} \sim 100\text{GHz}$)

Previous works: CDM axions converted into photons in the labs.

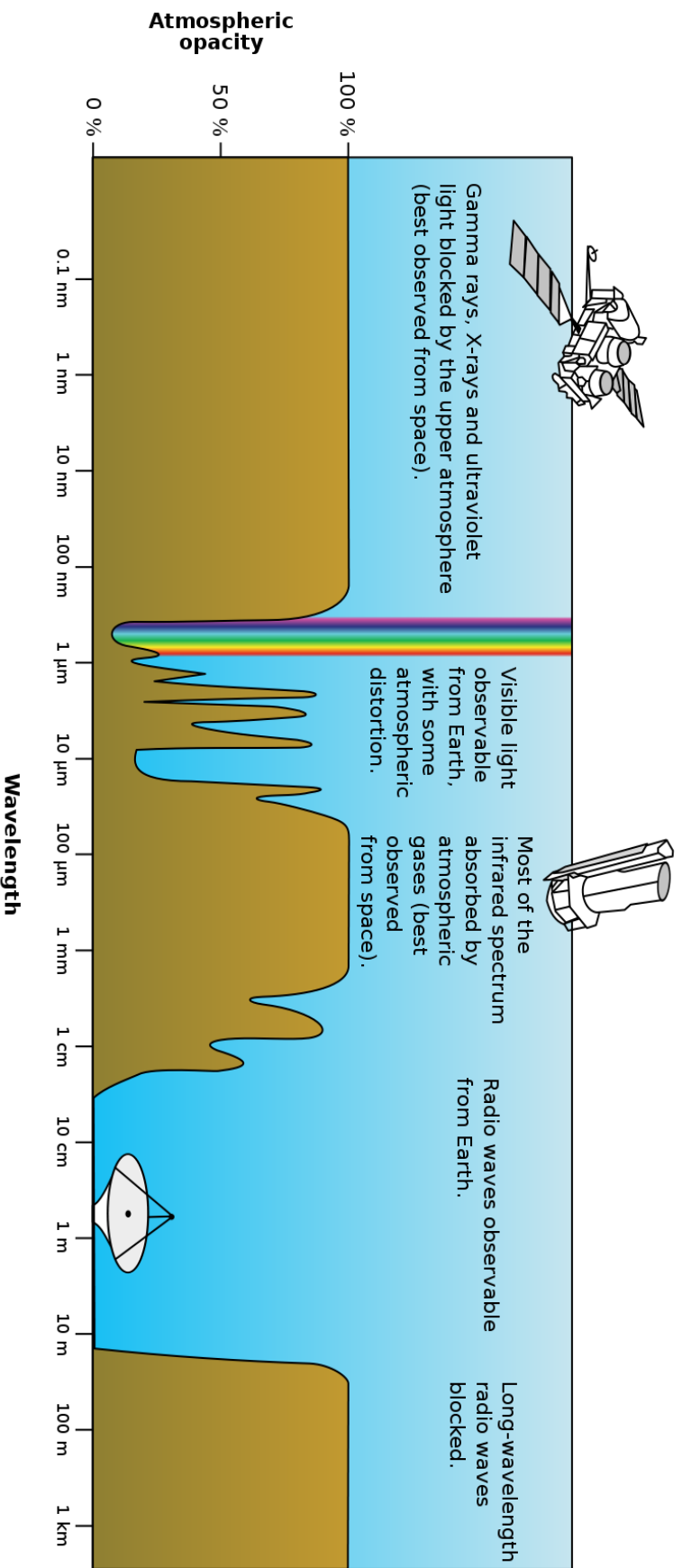
New works: How about the astrophysically sourced magnetic fields?

Non-resonant conversion: Kelley and Quinn (2017), Sigl (2017)

Resonant conversion: Huang, KK, Sekiguchi and Tashiro (2018), Hook, Kahn, Safdi and Sun (2018)

Line-like radio signal for non-relativistic axion conversion:

$$f \sim \frac{m_a}{2\pi} \sim 240 \left(\frac{m_a}{\mu\text{eV}} \right) \text{MHz}$$



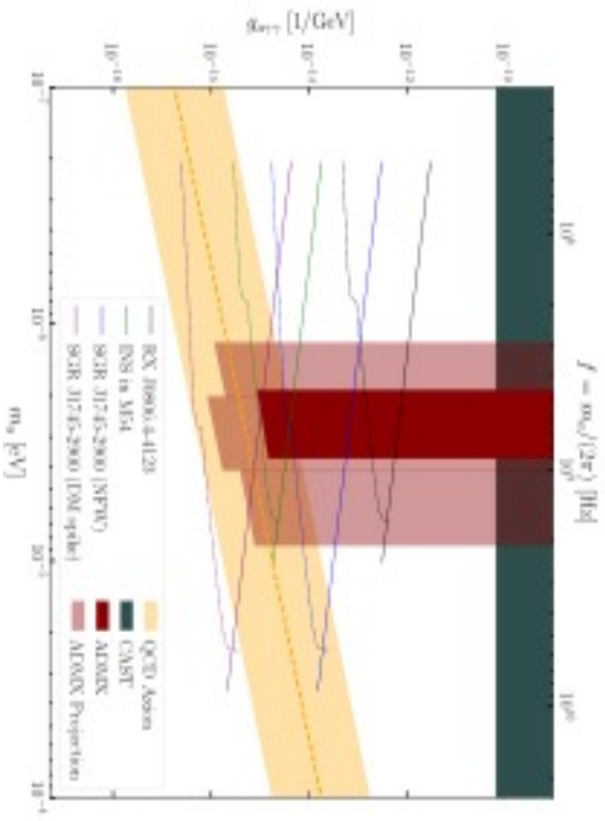
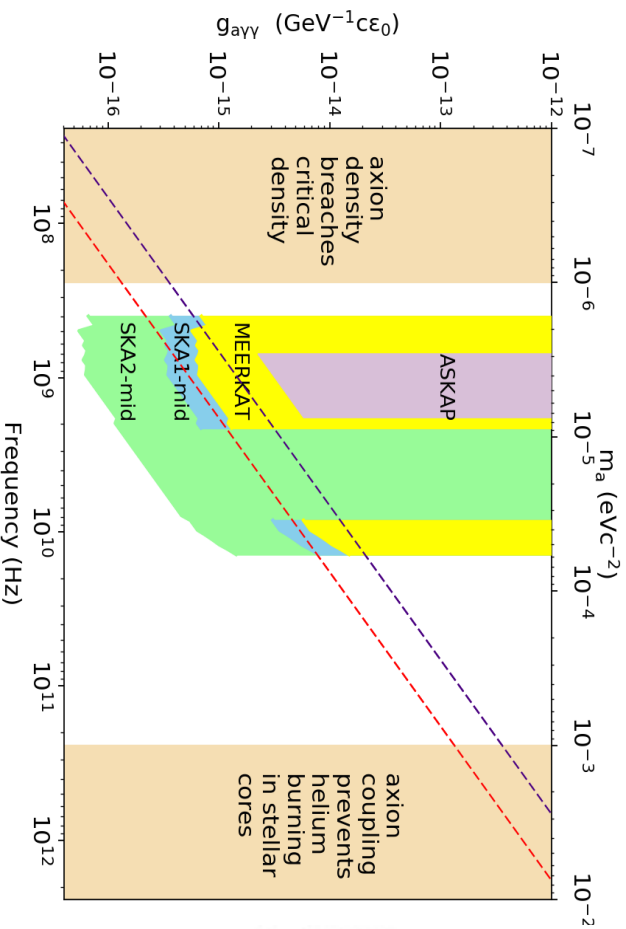
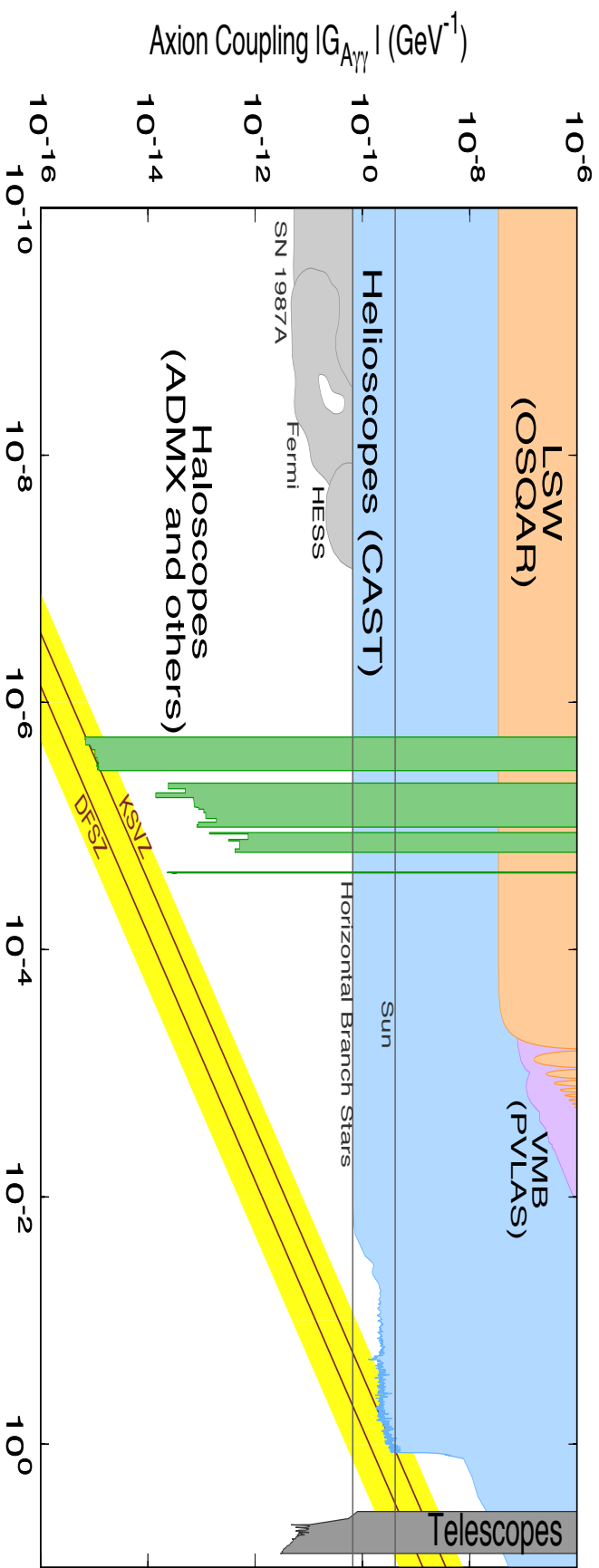
Australia: SKA low: 50-350 MHz

S. Africa: SKA mid: 350 MHz-14GHz

Axion mass: $0.2 \sim 60 \mu\text{eV}$

QCD axion as a CDM candidate :

Mass $\mu\text{eV} \sim \text{meV}$ ($0.1\text{GHz} \sim 100\text{GHz}$)



Kelley and Quinn (2017)

Hook, Kahn, Safdi and Sun(2018)

Model: ALP (Axion-like particles) i.e. Ultra-light scalars

- Ultra-light mass :

$$m_u \sim H_0 \sim 10^{-33} \text{ eV}$$

DE (Barbieri et al (2005),...)

$$m_u \sim 10^{-22} \text{ eV}$$

Fuzzy DM (Hu (2000),...)

$$m_u \sim 10^{-22} \text{ eV} - 10^{-10} \text{ eV}$$

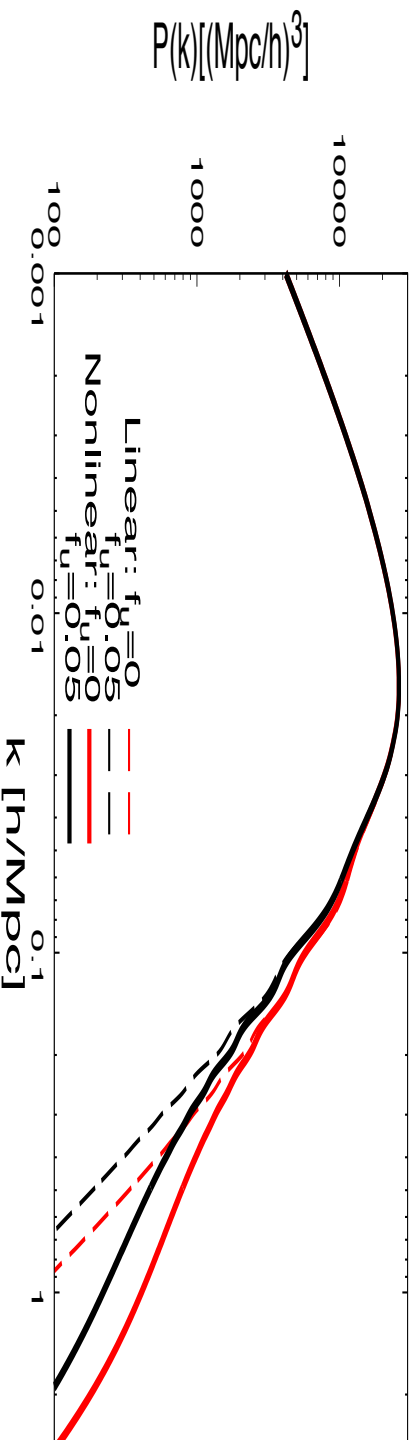
String axiverse (Arvanitaki et al (2009),...)

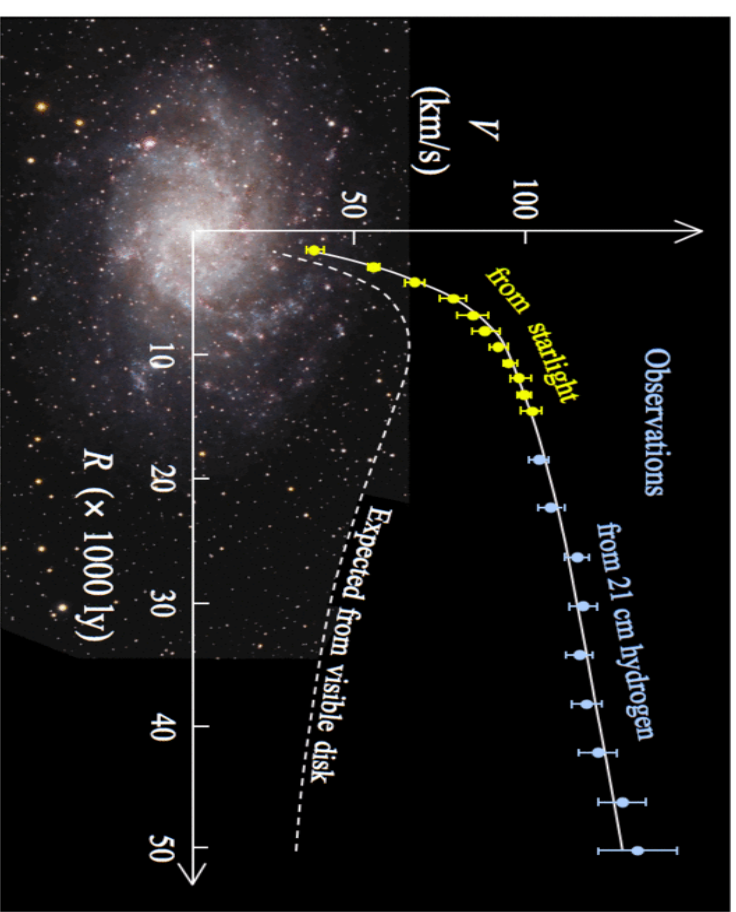
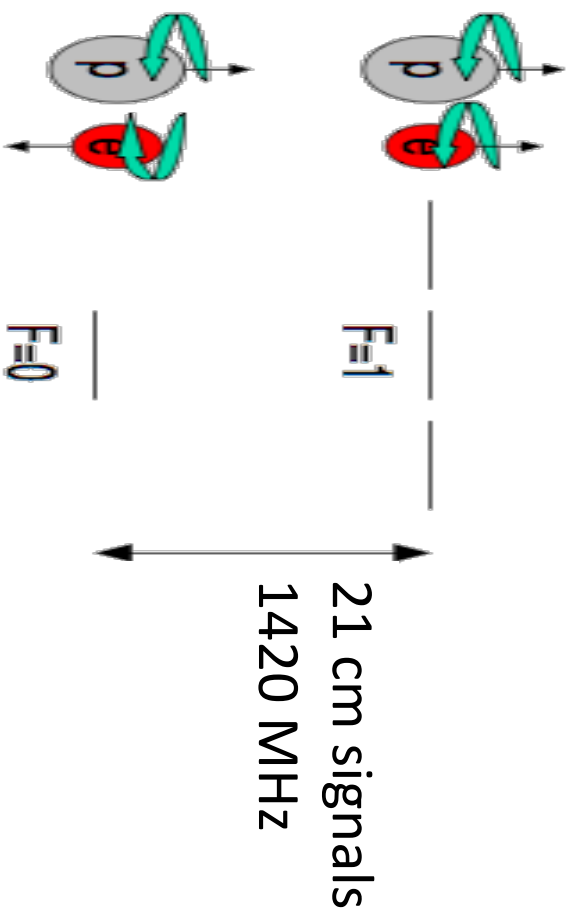
$$m_u, f_u = \Omega_u / \Omega_m \sim \mathcal{O}(0.01)$$

$$m_u \leq H(t) : \rho_u = \text{const}$$

$$m_u > H(t) : \rho_u \propto 1/a^3$$

KK, Mao, Ichiki, Silk (2014)



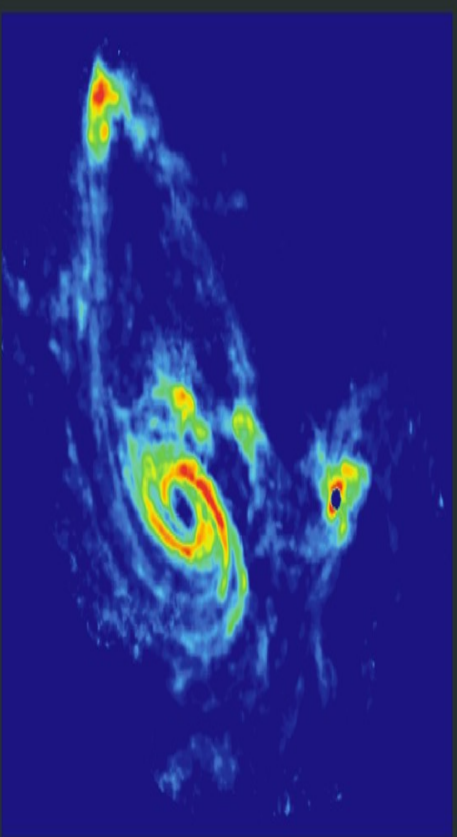


TIDAL INTERACTIONS IN M81 GROUP

Stellar Light Distribution



21 cm HI Distribution



Brief History of Universe

Years since
the Big Bang

~300000
($z \sim 1000$)



← Big Bang:

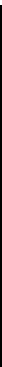
the Universe is filled with ionized gas

← Recombination: The gas cools and becomes neutral

Dark Ages



~100 million
($z \sim 20-40$)



← The first structures begin to form.

Reionization starts ($z \sim 12$)

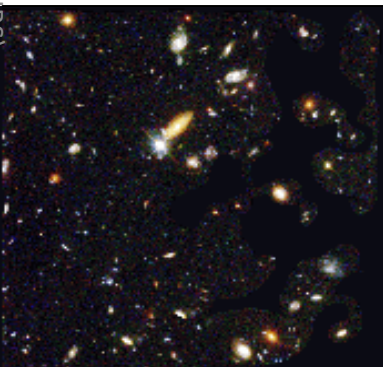
Reionization



~1 billion
($z \sim 6$)

← Reionization is complete

~13 billion
($z=0$)



← Today's structures

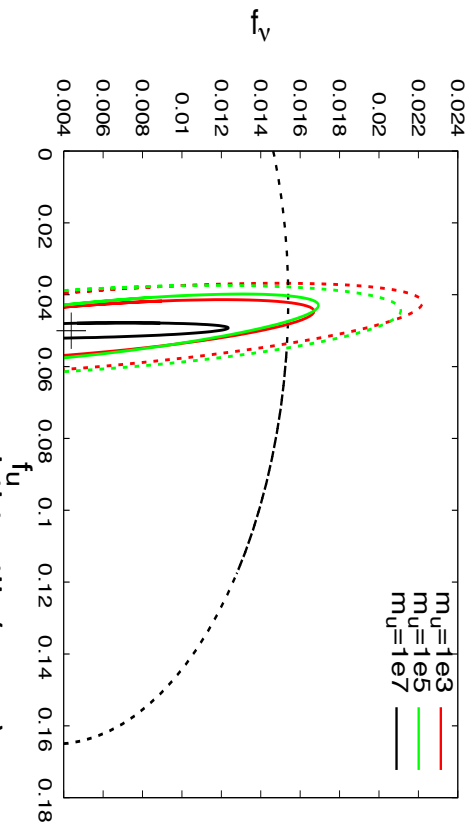
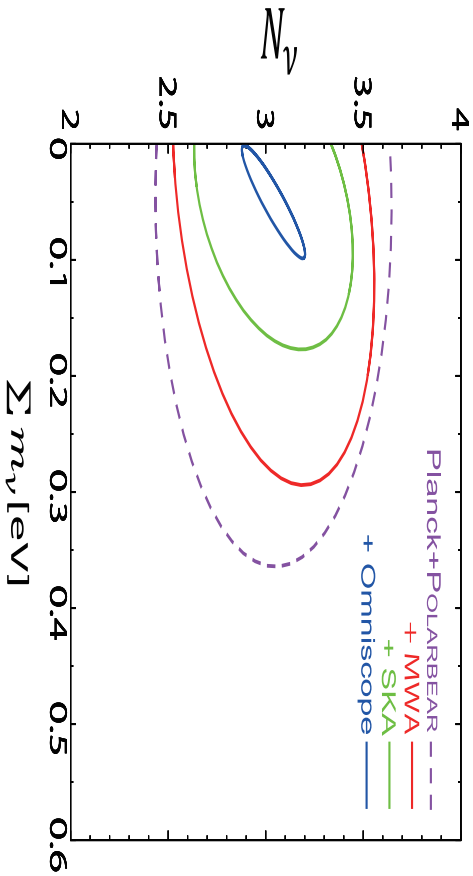
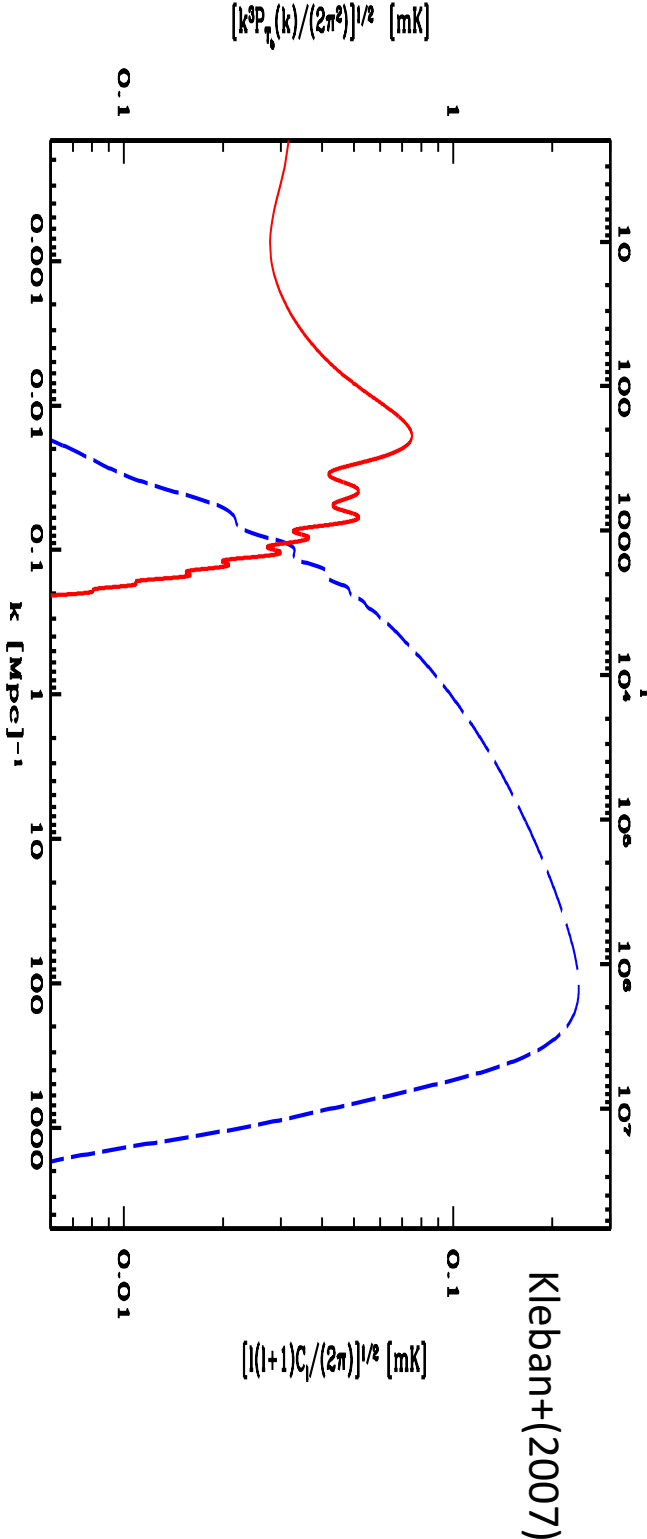
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CTPU Summer School July 2018

What can we do with 21cm?

High precision on small-scale power spectrum

$$\Delta P / P \sim 1 / \sqrt{N}$$



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Oyama+(2013)^{CTPU Summer School July 2018}

KK, Mao, Ichiki, Silk (2014)