

Neutrino Theory: Mass, Interactions, Unification

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IBS, 6.03.2019

Overview

Neutrino Theory

Introduction

Tree level

Radiative

Dirac case

B-L extension

Unification

Connection to
GW

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- 2 Tree level
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Introduction/Motivation

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- (Naturally) small neutrino masses
- Connection with dark matter existence
- Lepton mixing
- Unification

Standard Model

Neutrino
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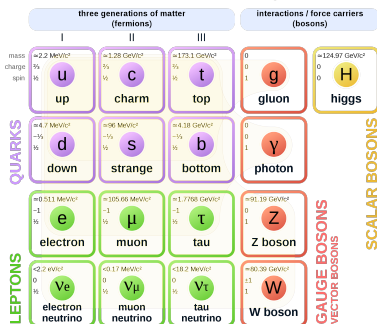
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Standard Model of Elementary Particles



Gauge symmetry group $\mathbb{G}_{SM} = SU(3)_c \times SU(2)_L \times U(1)_Y$

$$\mathcal{L}^Y = \bar{u} Y_u Q H + \bar{Q} Y_d d H + \text{h.c.} \quad (1)$$

$$\langle h^0 \rangle = v \rightarrow m_f = Y_f v \quad (2)$$

Requirements for neutrino mass

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Summary

- Majorana or Dirac type?
- Tree level or radiative?
- New particles? (scalar, fermionic, vector)
- New gauge sectors? ($U(1)$, $SU(2)$, $SU(N)$)

Possible effects/ Ways to test

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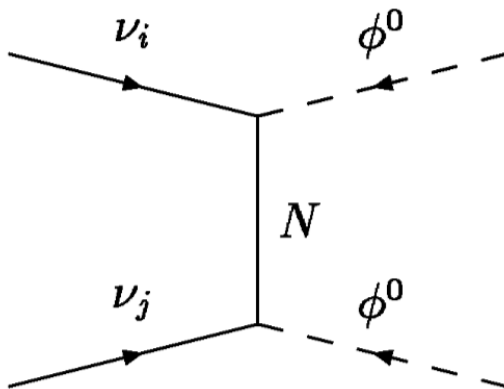
Summary

- Collider (LHC, ILC, FCC)
- Dark matter
- Grand unified theories
- Neutrino mixing
- Leptogenesis (CPV)
- Neutrinoless N-pole beta decay ($0\nu n\beta$)
- GW via scalar sector

Tree level, Majorana

Weinberg Operator 1979: $\frac{LHLH}{\Lambda} = \frac{(l^- H^+ - \nu H^0)(l^- H^+ - \nu H^0)}{\Lambda}$
Add $N_R \sim (1, 1, 0)$ under \mathbb{G}_{SM}

$$\mathcal{L}_{new} = \bar{N} Y_D L H + m_N N_R N_R + \text{h.c.}$$



Seesaw-I

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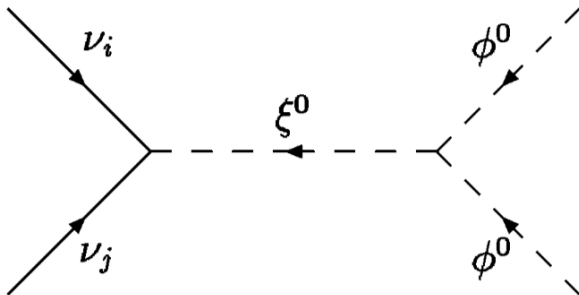
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Add[1980] $\xi = (\xi^{++}, \xi^+, \xi^0) \sim (1, 3, 1)$ under \mathbb{G}_{SM}

$$\mathcal{L}_{new} = YL\xi L + \text{h.c.} - \mu H\xi H \rightarrow m_\nu = Y \langle \xi^0 \rangle = -2 \frac{Y \mu v^2}{M_\xi}$$

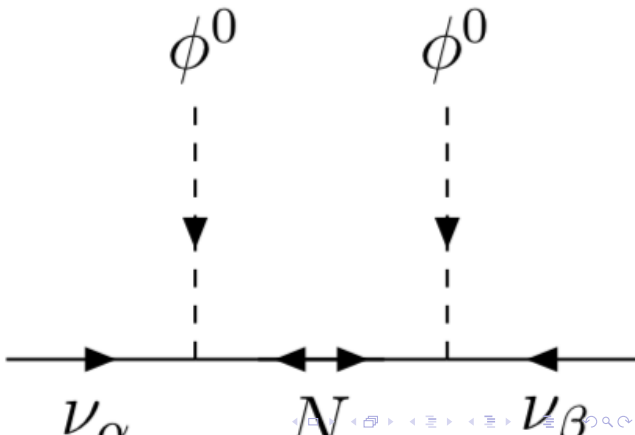


Seesaw-III

Neutrino Theory

Weinberg Operator 1979: $\frac{LHLH}{\Lambda} = \frac{(l^- H^+ - \nu H^0)(l^- H^+ - \nu H^0)}{\Lambda}$
 Add $N_R \sim (1, 1, 0)$ under \mathbb{G}_{SM}

$$\mathcal{L}_{new} = \bar{\Sigma} Y_D L H + m_N \Sigma_R \Sigma_R + \text{h.c.}$$



Seesaw variations

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$$\mathcal{M}_{\nu N} = \begin{pmatrix} 0 & m_D \\ m_D & m_N \end{pmatrix},$$

$$\text{Seesaw[1979]} \quad m_\nu = -m_D^2/m_N$$

Seesaw variations

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Summary

$$\mathcal{M}_{\nu n} = \begin{pmatrix} 0 & m_D & 0 \\ m_D & m_1 & m_N \\ 0 & m_N & m_2 \end{pmatrix}$$

Inverse Seesaw[1986] $m_\nu = m_D^2 m_2 / (m_N^2 - m_1 m_2)$

Seesaw variations

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Summary

$$\mathcal{M}_{\nu N} = \begin{pmatrix} 0 & m_D & m'_D \\ m_D & 0 & m_N \\ m'_D & m_N & 0 \end{pmatrix}$$

Linear Seesaw $m_\nu = -2m_D m_{D'}/m_N$

Radiative neutrino mass, Majorana

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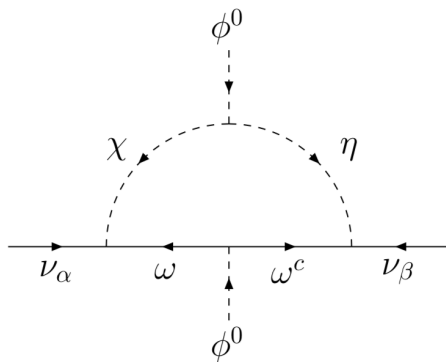
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Zee[1986]

Radiative neutrino mass, Majorana

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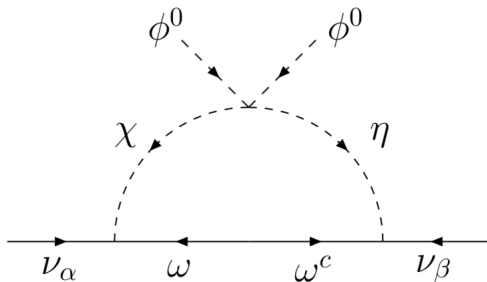
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Ma[2006]

Scotogenic radiative neutrino mass

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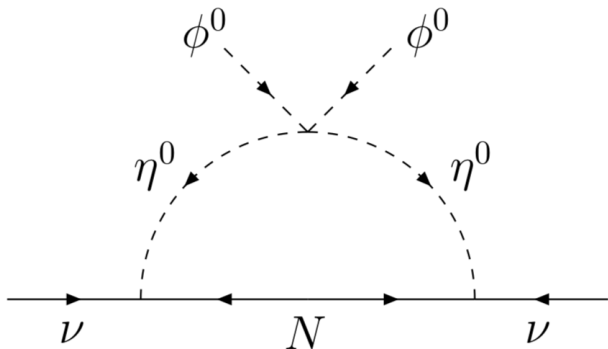
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Add \mathbb{Z}_2 symmetry under which $\eta \sim (1, 2, 1/2)$ and N_R are odd

Radiative neutrino mass, Majorana

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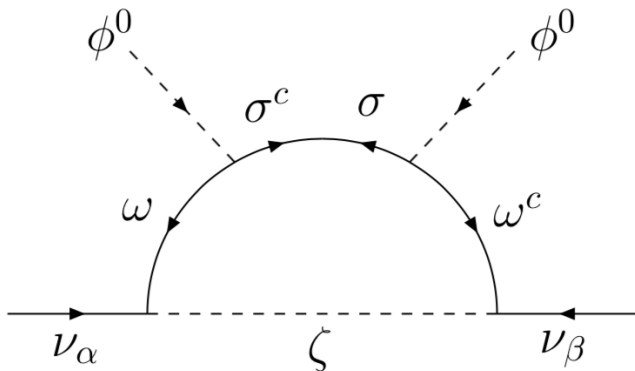
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Fraser, Ma, OP[2014]

Radiative inverse seesaw neutrino mass, Majorana

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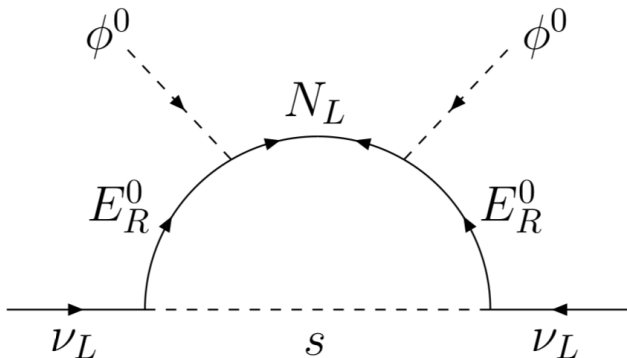
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Add \mathbb{Z}_2 symmetry under which real singlet scalar and $E_{L,R} \sim (1, 2, 1/2)$ and $N_L \sim (1, 1, 0)$ are odd

Dirac neutrinos

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Summary

- Add $N_R \sim (1, 1, 0)$ under G_{SM}
- N_R MUST transform under some other symmetry non-trivially
- New symmetry S is discrete, global, gauged, dark?

Tree Dirac case

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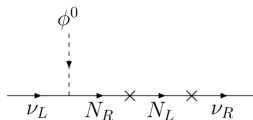
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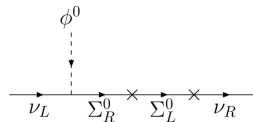
Connection to GW

Summary

- Insert a Dirac fermion singlet N which does not transform under \mathcal{S} , then break \mathcal{S} softly by the dimension-three $\bar{\nu}_R N_L$ term.



- Insert a Dirac fermion triplet $(\Sigma^+, \Sigma^0, \Sigma^-)$ which does not transform under \mathcal{S} , then break \mathcal{S} and $SU(2)_L \times U(1)$ together spontaneously to obtain the dimension-three $\bar{\nu}_R \Sigma_L^0$ term.



Dirac case

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Dirac case

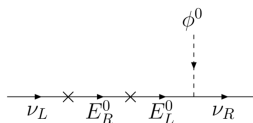
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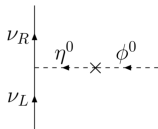
Connection to GW

Summary

- Insert a Dirac fermion doublet (E^0, E^-) which transforms as ν_R under \mathcal{S} , then break \mathcal{S} softly by the dimension-three $(\bar{E}^0 \nu_L + E^+ e^-)$ term.



- Insert a scalar doublet (η^+, η^0) which transforms as ν_R under \mathcal{S} , then break \mathcal{S} softly by the dimension-two $(\eta^- \phi^+ + \bar{\eta}^0 \phi^0)$ term.



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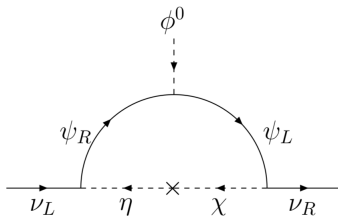
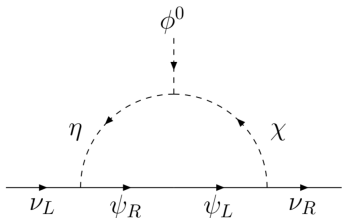
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$U(1)_{B-L}$ case

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- Add 3 $N_R \sim (1, 1, 0)$ which carry $L = (1, 1, 1)$
- Add 3 $N_R \sim (1, 1, 0)$ which carry $L = (4, 4, -5)$
- Other variations are possible
- Makes $U(1)_{B-L}$ anomaly free.
- $U(1)_{B-L}$ can global or gauged
- Global: softly or spontaneously broken(Majorana, Dirac)
- Gauged: spontaneously broken(Majorana, Dirac)

Scotogenic model in $SU(6)$ GUT*

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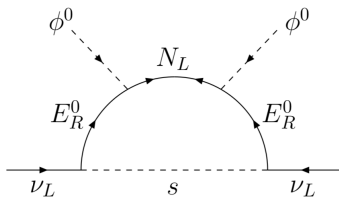
Simple case: Extend $SU(5)$ to $SU(6)$ to include BSM particles needed

$\underline{5}_F^* \times \underline{10}_F \times \underline{5}_S^*$, $\underline{10}_F \times \underline{10}_F \times \underline{5}_S$ $SU(5)$ Yukawa terms are extended to

$\underline{6}_F^* \times \underline{15}_F \times \underline{6}_S^*$, $\underline{15}_F \times \underline{15}_F \times \underline{15}_S$ $SU(6)$ Yukawas

Anomaly free combinations: $\underline{5}_F^* + \underline{10}_F$ for $SU(5)$,
 $\underline{6}_F^* + \underline{6}_F^* + \underline{15}_F$ for $SU(6)$. New $SU(6)$ $\underline{21}_S$ scalar is added to obtain 2'nd Higgs doublet ($\mathbb{Z}_2 \sim -$) with new interactions

$\underline{6}_F^* \times \underline{6}_F^* \times \underline{21}_S$, $\underline{15}_S^* \times \underline{15}_S^* \times \underline{21}_S \times \underline{21}_S$



Scotogenic model in $SU(7)$ GUT[†]

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Less simple case: Extend $SU(5)$ to $SU(7)$ to include BSM particles needed

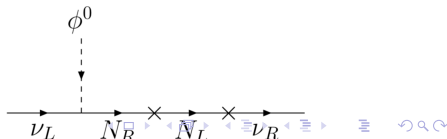
$\underline{5}_F^* \times \underline{10}_F \times \underline{5}_S^*$, $\underline{10}_F \times \underline{10}_F \times \underline{5}_S$ $SU(5)$ Yukawa terms are extended to

$\underline{7}_F^* \times \underline{21}_F \times \underline{7}_S^*$, $\underline{21}_F \times \underline{21}_F \times \underline{35}_S$ $SU(7)$ Yukawas

Anomaly free combination for $SU(7)$: $\underline{7}_F^* + \underline{7}_F^* + \underline{7}_F^* + \underline{21}_F$.

New $\underline{28}_S$ scalar needed to accomodate $SU(2)_N$ doublet, with new interactions: $\underline{7}_F^* \times \underline{7}_F^* \times \underline{28}_S$, $\underline{21}_S^* \times \underline{21}_S^* \times \underline{28}_S \times \underline{28}_S$

- Insert a Dirac fermion singlet N which does not transform under \mathcal{S} , then break \mathcal{S} by the dimension-three $\bar{\nu}_R N_L$ term.



Connection to GW signals

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Summary

- Neutrino mass requires new scalars or gauge bosons
- Strong first order phase transition (BSM gauge symmetry)
- Strong first order phase transition gives GW
- Some possible models: LR models, $SU(N)$ scotogenic symmetry

Leptogenesis

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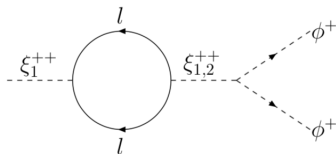
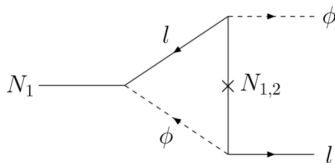
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Summary

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Summary

- Neutrino masses can be Majorana or Dirac
- Tree or radiative
- Require BSM fields and symmetries (global or gauged)
- Connection with DM
- Connection with GW
- Leptogenesis

Thank you!