Potentially interesting aspects of the effective potential

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27.03.2019, IBS-CTPU

Ref.

JhP, 1902.04559

Essential object, e.g. for studying:

- symmetry breaking
- vacuum stability
- phase transitions at zero and finite temperatures
- cosmic inflation
- radiative corrections to Higgs mass

Generating functionals

"Partition function (Zustandssumme)"

$$Z[J] = \int \mathcal{D}\phi \exp\left[i \int d^4x \left(\mathcal{L}(x) + J(x)\phi(x)\right)\right]$$
$$= Z[0] \langle \Omega | \Omega \rangle_J = \exp\left(iW[J]\right)$$

= generating functional for correlation functions

Phase of vacuum to vacuum amplitude

$$W[J] = -i \ln Z[J] = -TE_{\Omega}[J]$$

= generating functional for connected correlation functions

1PI effective action

Legendre transform of vacuum energy functional Jona-Lasinio, Nuovo Cim.34(1964)1790

$$\Gamma[\phi_{cl}] = W[J] - \int d^4x J(x)\phi_{cl}(x)$$

= generating functional for 1PI correlation functions

$$\frac{\delta W[J]}{\delta J(x)} = \frac{-i}{Z[J]} \frac{\delta Z[J]}{\delta J(x)} = \frac{\int \mathcal{D}\phi \, e^{i \int d^4 x \, (\mathcal{L} + J\phi)} \phi(x)}{\int \mathcal{D}\phi \, e^{i \int d^4 x \, (\mathcal{L} + J\phi)}}$$
$$= \frac{\langle \Omega | \phi(x) | \Omega \rangle_J}{\langle \Omega | \Omega \rangle_J} = \phi_{\text{cl}}(x)$$
$$= \text{Expectation value of } \phi(x) \text{ in the presence of } J$$
$$\frac{\delta \Gamma[\phi_{\text{cl}}]}{\delta \phi_{\text{cl}}(x)} = -J(x) = 0 \text{ if } \phi_{\text{cl}} = \langle \phi \rangle$$

(X)

Effective potential or "free energy" Effective action for *x*-independent ϕ_{cl}

$$V_{\rm np}(\phi_{\rm cl}) = \left. -\frac{1}{VT} \Gamma[\phi_{\rm cl}] \right|_{\phi_{\rm cl}={\rm const.}}_{\rm Jackiw. \ PRD9(1974)1686}$$

- real-valued by definition
- depends on gauge and renormalization scale



Loop expansion of effective potential

$$\begin{split} V_{np} &\simeq V_{pert} = V_0 + V_1 + V_2 + \cdots \\ V_0(\phi_{cl}) &= \text{tree-level potential in } \mathcal{L} \\ V_1(\phi_{cl}) &= \frac{M^4(\phi_{cl})}{64\pi^2} \left(\ln \frac{M^2(\phi_{cl})}{\mu^2} - \frac{3}{2} \right) \quad \text{in } \overline{\text{MS}} \text{ scheme} \end{split}$$

develops imaginary part where V₀ is concave
 In this talk, allow V_{pert} to include nonperturbative contributions as in axion or meson potential

Energy interpretation

Kurt Symanzik, CMP16(1970)48

$$V_{\rm np}(\phi) = \frac{1}{V} \min_{\Omega} \frac{\langle \Omega | H | \Omega \rangle}{\langle \Omega | \Omega \rangle} \quad \text{s.t.} \quad \frac{\langle \Omega | \Phi | \Omega \rangle}{\langle \Omega | \Omega \rangle} = \phi$$

 $V_{\rm np}(\phi)$ is the minimum of the expectation value of the energy density for all states constrained by the condition that the scalar fields Φ have expectation values ϕ

Convexity of effective potential

Kurt Symanzik, CMP16(1970)48

$$V_{np}(x\phi_1 + (1-x)\phi_2) \le xV_{np}(\phi_1) + (1-x)V_{np}(\phi_2)$$

for $0 \le x \le 1$

 Can be understood by taking linear combination of states

$$\sqrt{x} |\Omega_1\rangle + \sqrt{1-x} |\Omega_2\rangle \quad \text{with} \quad \langle \Omega_{1,2} | \Omega_{1,2} \rangle = 1 \langle \Omega_{1,2} | \Phi | \Omega_{1,2} \rangle = \phi_{1,2}, \quad \langle \Omega_{1,2} | H | \Omega_{1,2} \rangle / V = V_{\text{np}}(\phi_{1,2})$$

or by imagining a state with volumes in different phases

 Convexity of Gibbs free energy is also known in statistical mechanics

"Maxwell construction"

• $V_{\rm np}(\phi)$ is linear between two local minima of $V_{\rm pert}(\phi)$

Fujimoto, O'Raifeartaigh, Parravicini, NPB212(1983)268

Approximation by linear interpolation

$$V_{np}(x\phi_1 + (1-x)\phi_2) \simeq xV_{pert}(\phi_1) + (1-x)V_{pert}(\phi_2)$$

for $0 < x < 1$

guaranteed to be real

 QFT analogue of Maxwell construction for free energies in thermodynamics

dS swampland conjecture

- A low energy effective theory belongs to landscape if it has string theory as its UV completion
- Otherwise it belongs to swampland
- Scalar potential of a low energy effective theory in landscape satisfies

Obied, Ooguri, Spodyneiko, Vafa (2018)

$$M_{\mathsf{Pl}} |\nabla V| > c V, \quad 0 < c \sim \mathcal{O}(1)$$

Excludes de Sitter extrema
 Refined to include alternative condition

Ooguri, Palti, Shiu, Vafa (2018)

$$M_{\mathsf{Pl}}^2 \min(\nabla_i \nabla_j V) \leq -c' V, \quad 0 < c' \sim \mathcal{O}(1)$$

Potentially interesting aspects of the effective potential

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Old dS swampland conjecture vs HEP models

 Quintessence could accommodate observed cosmological constant and local maximum of Higgs potential

Denef, Hebecker, Wrase (2018)

 Severely constrained by a long-range force and time dependence of proton-to-electron mass ratio

Hamaguchi, Ibe, Moroi (2018)

- Quintessence + pion extremum requires $c < 1.4 \times 10^{-2}$ K. Choi, D. Chway, C. S. Shin (2018)
- QCD axion becomes difficult

Murayama, Yamazaki, Yanagida (2018)

 Root of problems is exclusion of positive local maxima of V

QFT questions about dS conjectures

- Which V? V_{np}? V₀? V_{pert}?
- How to interpret inequalities

$$\begin{split} M_{\mathsf{Pl}} \left| \nabla V \right| &> c \, V, \quad 0 < c \sim \mathcal{O}(1) \\ M_{\mathsf{Pl}}^2 \min(\nabla_i \nabla_j V) \leq -c' \, V, \quad 0 < c' \sim \mathcal{O}(1) \end{split}$$

- if $\text{Im } V \neq 0$?
- Gauge and scale dependence of V

Could free energy interpretation save old dS conjecture?

Kobakhidze, 1901.08137; JhP, 1902.04559

Suppose *V* is V_{np} in dS swampland conjecture $M_{Pl} |\nabla V_{np}| > c V_{np}, \quad 0 < c \sim O(1)$

- V_{np} is real-valued, as quantum as possible
- More permissive than old and refined dS criteria as local maxima of V_{pert} are flattened in V_{np}
- Global minima still need to be AdS
- We are not living in a true vacuum, i.e. we are living in
- !(true && vacuum) == (false || !vacuum)

(1)

Option ! vacuum: living on a slope

- Quintessence solution to cosmological constant with $|\nabla V_{pert}| > 0$
- Enough to add quintessence terms to Higgs potential as V_{np} has no local maximum
- No problem with a long-range force or time dependence of proton-to-electron mass ratio







Potentially interesting aspects of the effective potential

Disallowed

Use Maxwell construction to represent condition (1) by



$$V_{\text{pert}}(\phi_+) < rac{M_{\text{Pl}}}{c} \left| rac{V_{\text{pert}}(\phi_+) - V_{\text{pert}}(\phi_-)}{\phi_+ - \phi_-}
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$$V_{
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Implications for V_{pert}

Our EW vacuum is required to be metastableMaybe due to an extra field such as in:

supersymmetry with CCB minima

Claudson, Hall, Hinchliffe (1983)

metastable supersymmetry breaking sectors

Intriligator, Seiberg, Shih (2006)

relaxion mechanism

Graham, Kaplan, Rajendran (2015)

scalar extensions of the Higgs sector

Otherwise metastability is due to SM Higgs

Near-criticality of Higgs potential

 RG-improved effective potential

$$V_{\text{pert}}(h) = rac{\lambda(\mu = |h|)}{4} h^4$$

• Condition (1) + distance conjecture + UV cutoff of SM suggest $\lambda(\mu)$ should turn negative at $\mu \lesssim M_{\rm Pl}$

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Buttazzo, Degrassi, Giardino, Giudice, Sala,

Salvio, Strumia (2013)

Connection to string theory?

- Hard to judge if condition (1) has something to do with string theory
- Admits false dS vacua contrary to original motivation for dS criteria
- Conjectures have not been proved

Summary

Yet another refinement of dS swampland conjecture

- Form is identical to original but effective potential integrates all (non)perturbative quantum effects
- V_{np} is real \rightarrow no inequality on complex numbers
- *V*_{np} is convex ~ allows local maxima and false dS vacua in *V*_{pert} if slope of *V*_{np} is everywhere steep enough ~ compatible with HEP models
- True vacua must still be AdS ~> quintessence or metastable EW vacuum
- Reason for near-criticality of SM Higgs potential?