

New physics searches via Axion-like particles

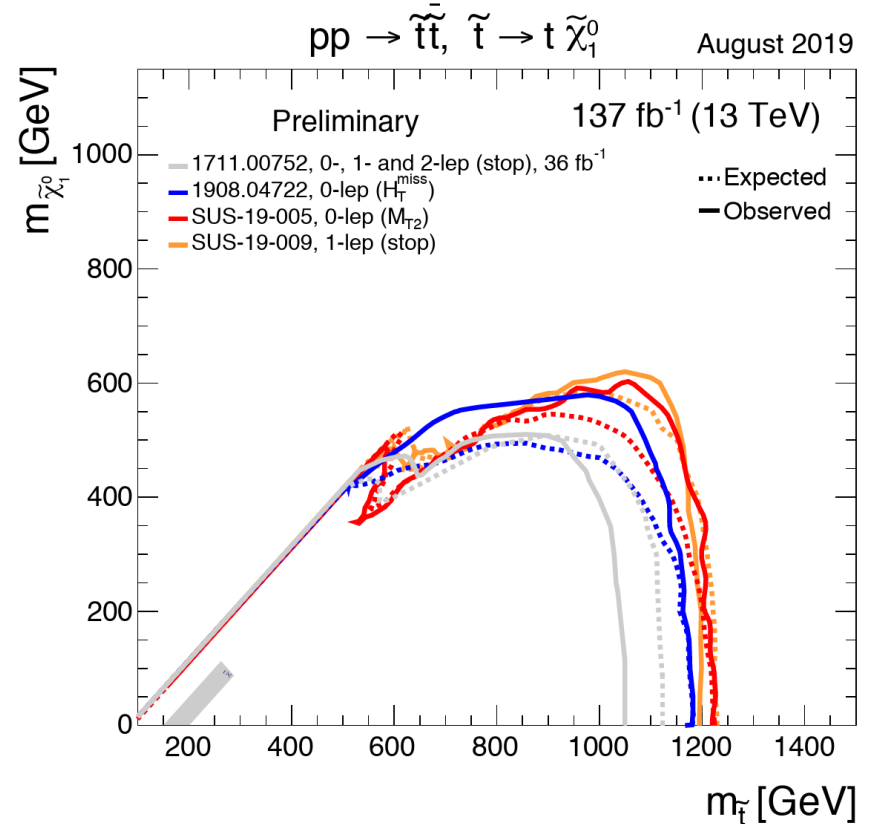
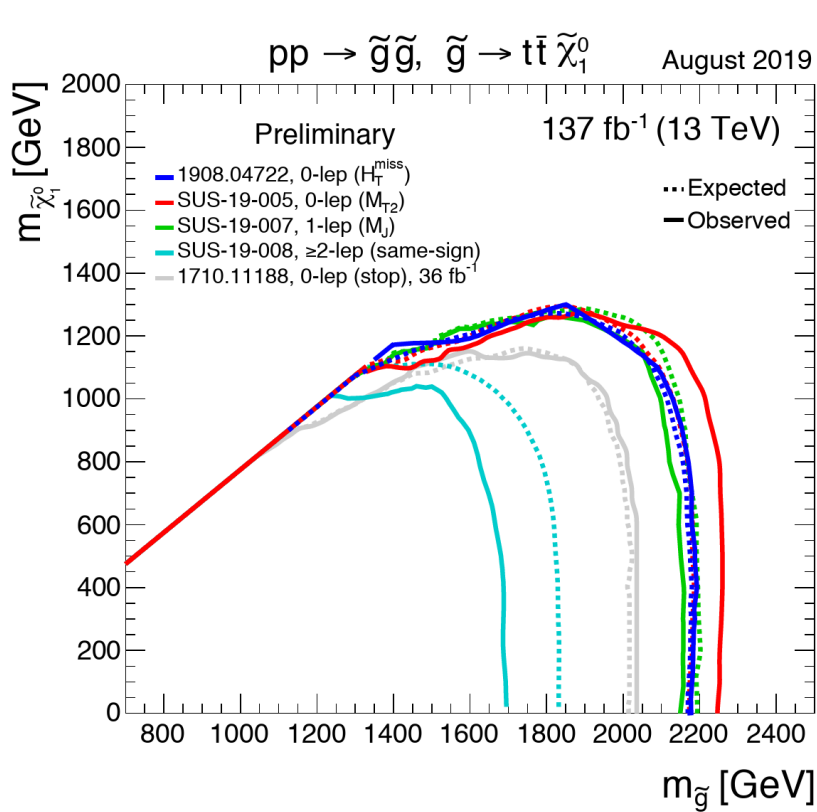
Sang Hui Im (Pusan National Univ.)

Jan 3, 2020, IBS-CTPU

Outline

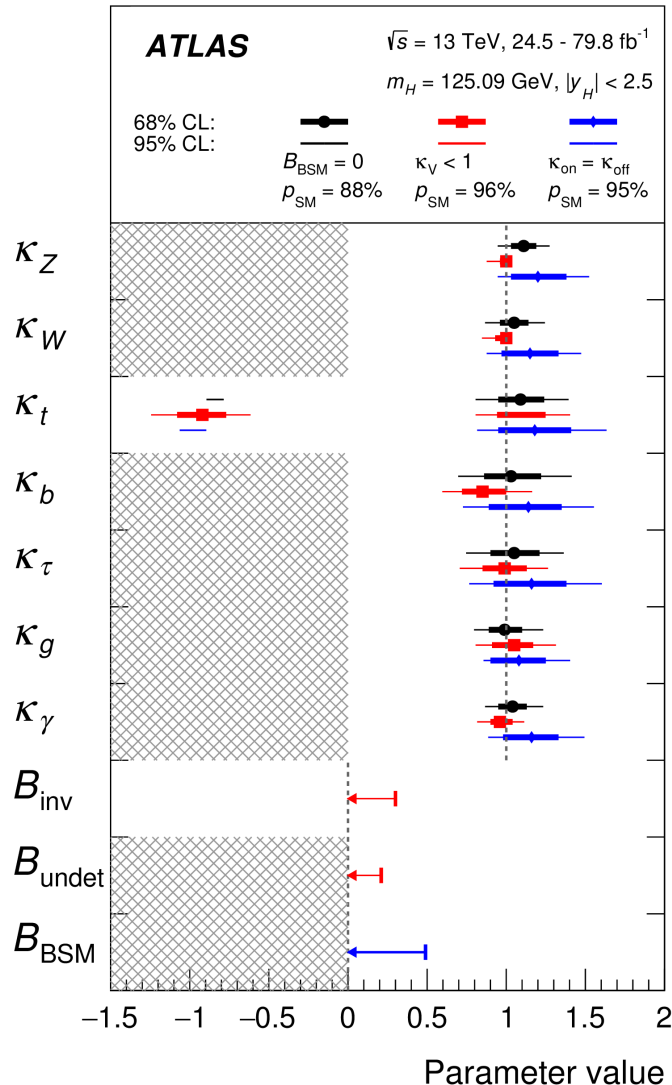
- ALPs as a portal to new physics
: Motivation and structure
- Bottom-up approach : SM problems
- Top-down approach : UV complete theories
- Conclusion

LHC up to 2019



: New physics scale seems higher than conventional expectation.

Higgs couplings

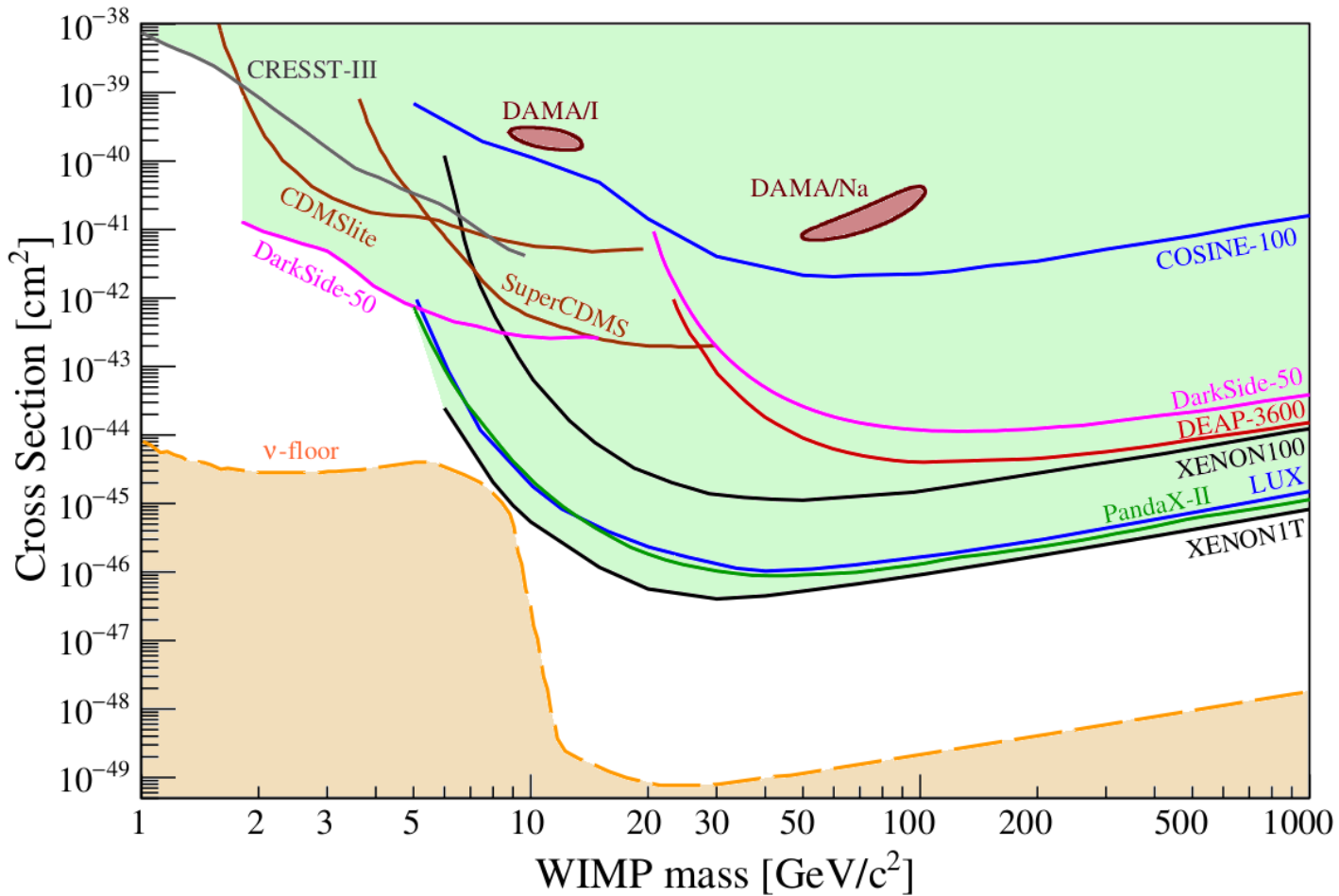


: Good agreement with the SM predictions

: Allow only small mixing with the SM Higgs boson and small radiative corrections to the couplings

WIMP paradigm being challenged

M. Schumann '19



A new physics sector seems coupled to the SM
rather weakly.



A good physics candidate for naturally small
couplings while solving the SM problems?

An example : QCD axion

Solves the strong CP problem of the SM :

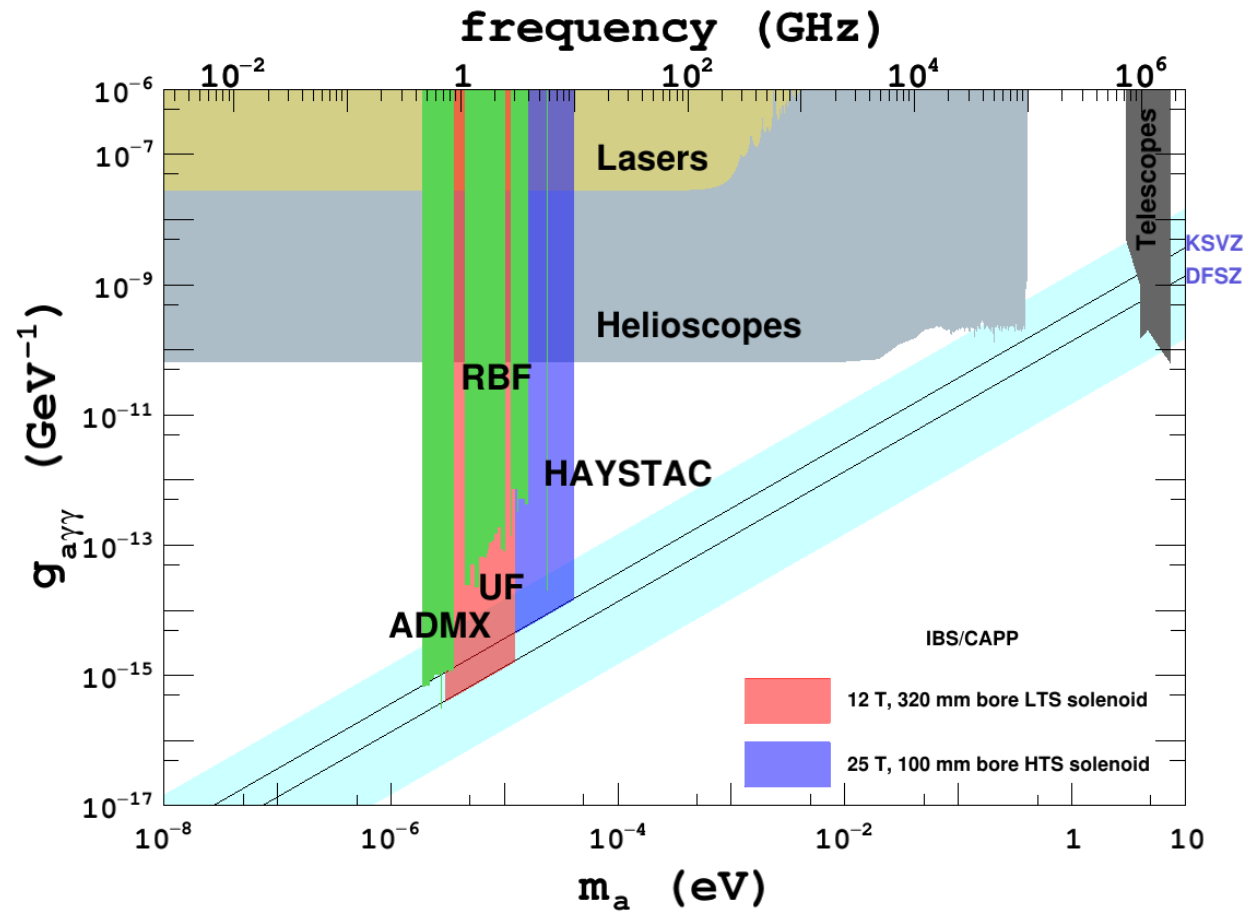
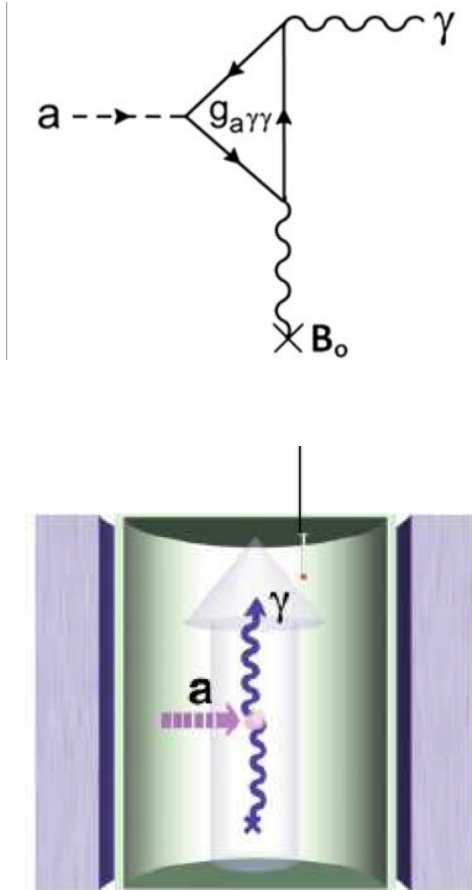
$$\theta_{QCD} < 10^{-10}$$

Dark matter problem as well :

$$f_a \approx 10^9 \text{ or } 10^{12} \text{ GeV}$$

$$\frac{\partial_\mu a}{f_a} \bar{\psi} \gamma^\mu \psi + \frac{1}{32\pi^2} \frac{a}{f_a} G_{\mu\nu} \tilde{G}^{\mu\nu} \quad : \text{tiny couplings to the SM}$$

Still within the experimental reach



with targeted couplings and mass

Axion-like particle (ALP)

- i) Perturbative (approximate) shift symmetry : $a \rightarrow a + c$
- ii) Discrete gauge symmetry : $a \rightarrow a + 2\pi n f$ ($\frac{a}{f} = \theta$; angle)

$$\mathcal{L} = \frac{1}{2}(\partial_\mu a)^2 + \frac{\partial_\mu a}{f} J^\mu + \frac{1}{32\pi^2} \frac{a}{f} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

Natural size of ALP couplings is determined by f .

ALP mass

$$\frac{1}{32\pi^2} \frac{a}{f} G_H \tilde{G}_H \quad \Rightarrow \quad V(a) = \Lambda_{HC}^4 \cos\left(\frac{a}{f}\right)$$

$$m_a = \frac{\Lambda_{HC}^2}{f}$$

m_a and f are free parameters with (technical) naturalness.
What values would we look at experimentally?

ALP effective interactions to the SM

$$\mathcal{L} = \frac{1}{2}(\partial_\mu a)^2 - V(e^{ia/f}, \psi, H)$$

high energy non-perturbative effect

$$+ \underbrace{\frac{\partial_\mu a}{f} \left(\sum_\psi c_\psi \bar{\psi} \gamma^\mu \psi + c_H H^\dagger i \overleftrightarrow{D}^\mu H \right) + \frac{1}{32\pi^2} \frac{a}{f} \sum_F c_F F_{\mu\nu} \tilde{F}^{\mu\nu}}_{\text{Perturbative shift symmetry conserving interactions}}$$

Which interactions are we interested in?

Bottom-up approach

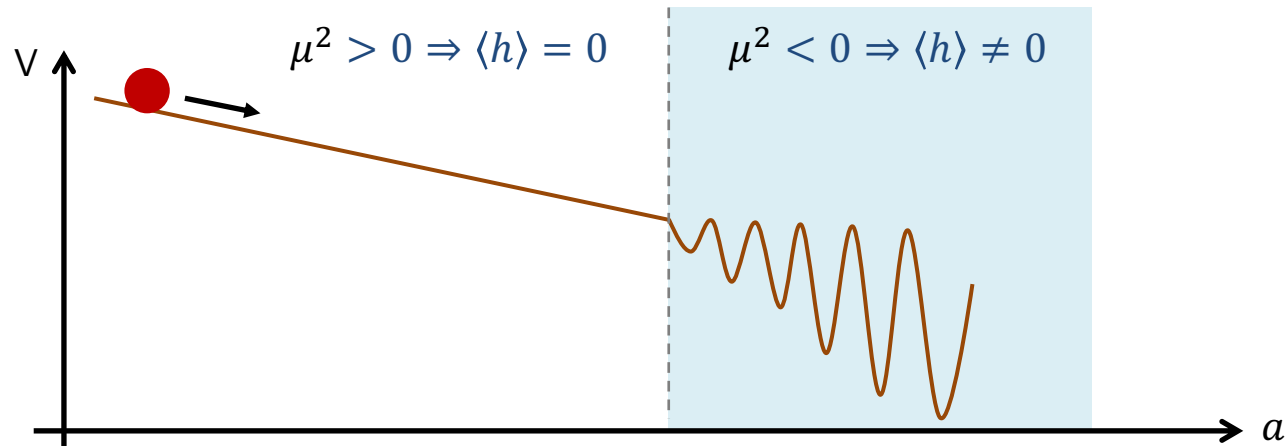
Similarly to the QCD-axion, ALPs may be involved in solving SM problems.

- i) Weak scale hierarchy
- ii) Dark Matter
- iii) Baryogenesis

Weak scale hierarchy : Relaxion

Graham, Kaplan, Rajendran '15

- Promoting the Higgs mass to a dynamical field (ALP) : $\mu^2(a)|H|^2$
- Cosmological ALP evolution to select the Higgs mass



- a slow-rolls while scanning μ^2 from Λ_{cut}^2 to negative, and stops by barriers formed by EWSB.

$$\mu^2(a)|H|^2 \subset V(e^{ia/f}, \psi, H)$$

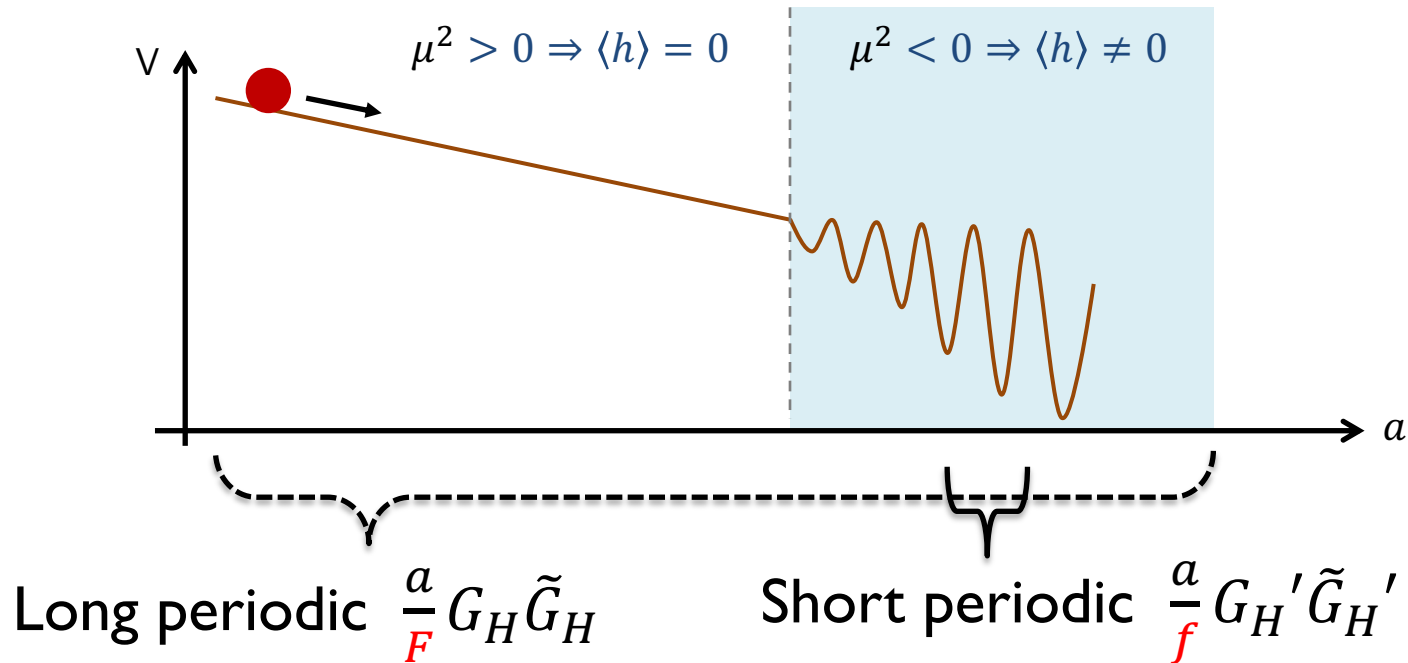
: shift symmetry breaking interaction to the SM

How do we obtain it?

$$\frac{|H|^2}{\Lambda} N_H N_H^c + \frac{1}{32\pi^2} \frac{a}{f} G_H \tilde{G}_H \quad \rightarrow \quad \frac{|H|^2}{\Lambda} \Lambda_{HC}^3 \cos\left(\frac{a}{f}\right)$$

high scale confinement

Trouble : ALP scale hierarchy



To stop relaxation with
the weak scale Higgs VEV

$$\frac{F}{f} \approx \frac{\Lambda_{cut}^4}{v^4} > 10^4$$

Discrete gauge symmetry : $a \rightarrow a + 2\pi n f$

$$\frac{a}{\textcolor{red}{F}} G_H \tilde{G}_H + \frac{a}{\textcolor{red}{f}} G_H' \tilde{G}_H' \quad F = f \times \text{integer}$$



Clockwork scheme

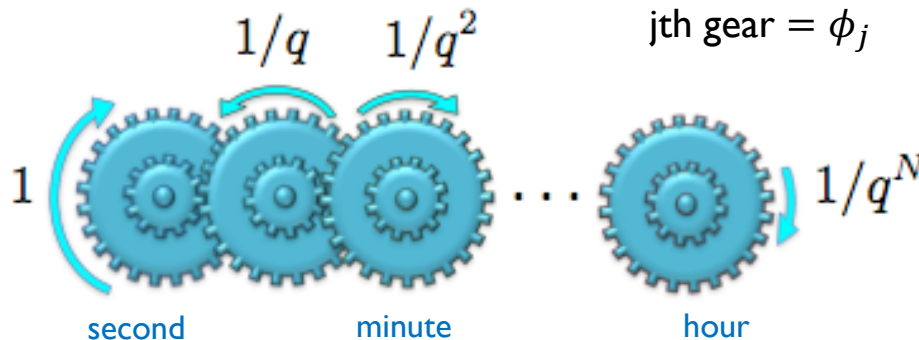
Choi, SHI '15

Kaplan, Rattazzi '15

$$V(\phi_i) = \sum_{i=0}^N \frac{1}{2} m_i^2 (\phi_i - \textcolor{red}{q} \phi_{i+1})^2$$

$$F = f \times q^N$$

N : Number of ALPs



Relaxion interactions to the SM

$$M^2 |H|^2 \cos \left(\frac{a}{f} \right) + c_\gamma \frac{e^2}{16\pi^2} \frac{a}{f} F \tilde{F}$$



$$\langle a \rangle \sim f$$



Relaxion-Higgs mixing

Choi, SHI '16

Flacke, Frugiuale, Fuchs, Gupta, Perez '16

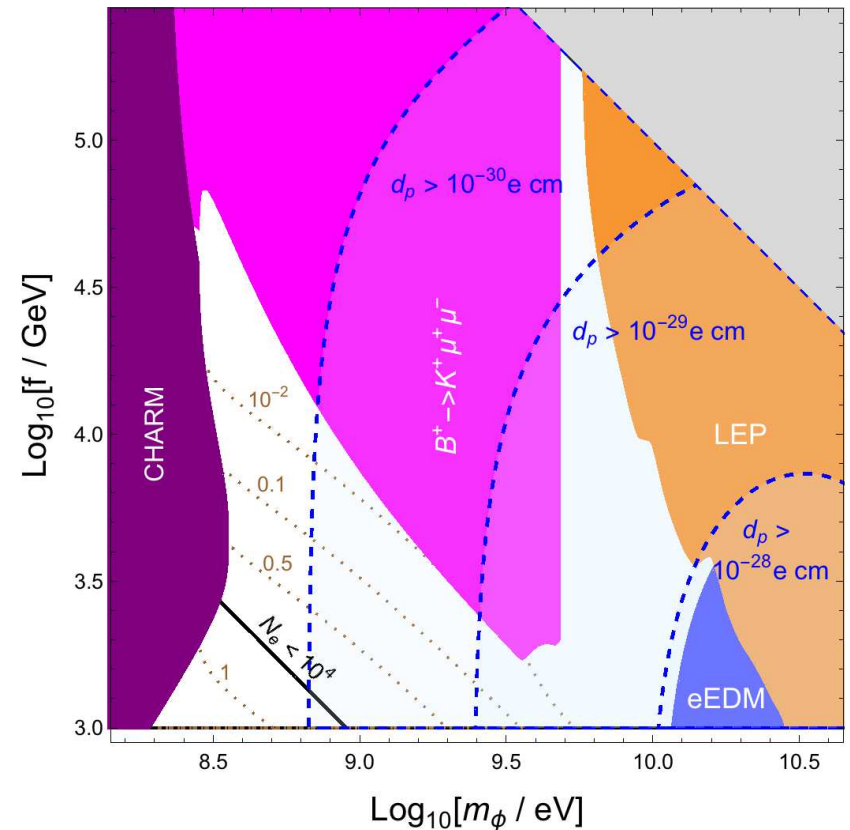
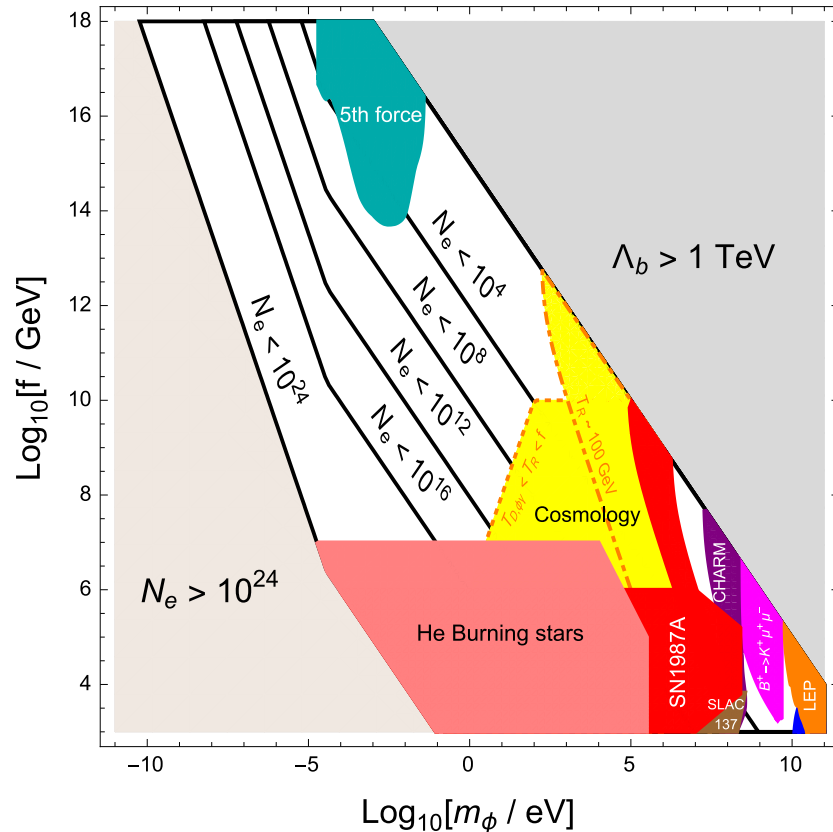
Relaxion-photon coupling

Choi, SHI '16

Simultaneous presence of them breaks CP.

What should we look for?

Choi, SHI '16

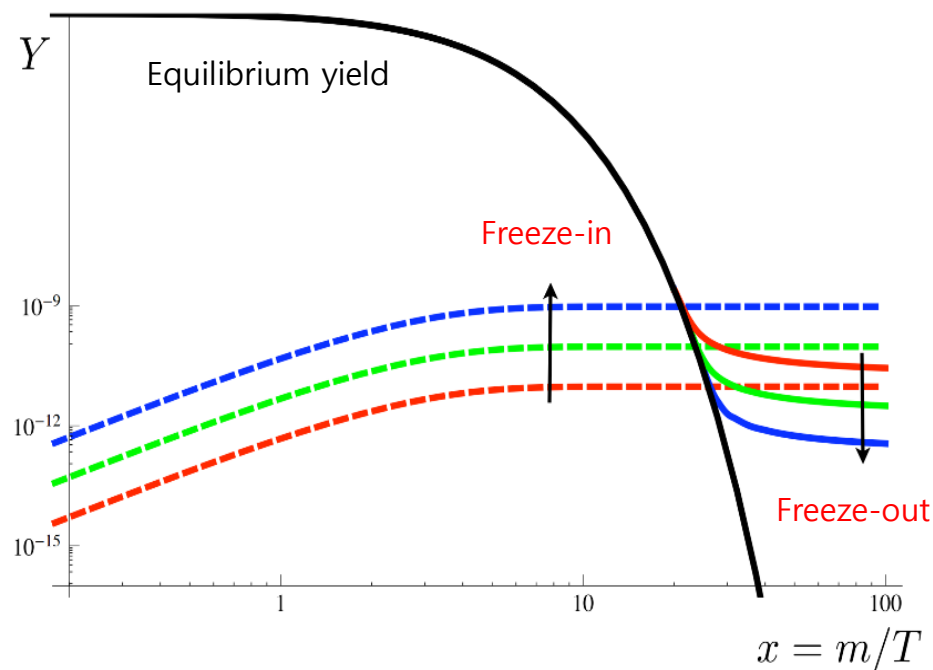


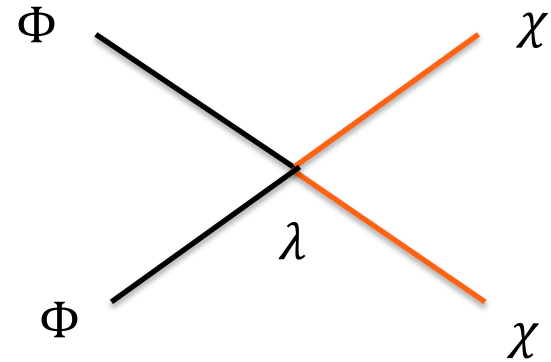
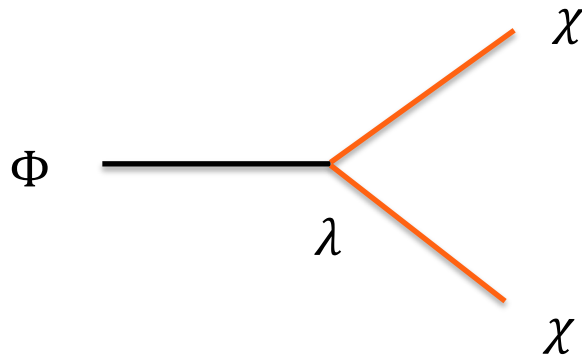
- Beam dump (SHiP), Proton EDM
- HL-LHC, CLIC [Frugiele, Fuchs, Perez, Schlaffer '18]

Dark Matter : Freeze-in paradigm

McDonald 2001, Choi, Roszkowski 2005, Petraki, Kusenko 2007
Hall, Jedamzik, March-Russell, West 2009

- Alternative to freeze-out
- Never in thermal equilibrium: **feebly coupled to SM**
- Produced via decay or annihilation of particles in thermal bath





Φ : particle in thermal bath

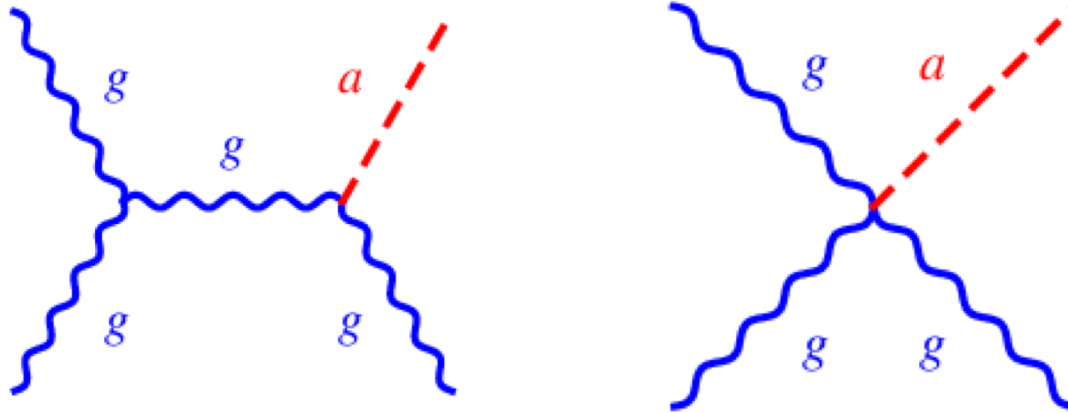
χ : dark matter

$$\Omega_{\chi} \propto m_{\chi} \frac{\lambda^2}{m_{\Phi}}, \quad \lambda^2$$

- Observed DM abundance if $\lambda \sim 10^{-12}$ for $m_{\chi} \sim m_{\Phi}$
- Need an explanation for $\lambda \ll 1$!

QCD-axion freeze-in production

Graf, Steffen '10
Bae, Choi, SHI '11



$$\frac{a}{f_a} G \tilde{G}$$



$$Y_a \propto \frac{M_{\text{Pl}} T_R}{f_a^2}$$

QCD-axion is cosmologically stable due to its small mass (meV~ μ eV) but cannot be freeze-in DM because it shall be hot.

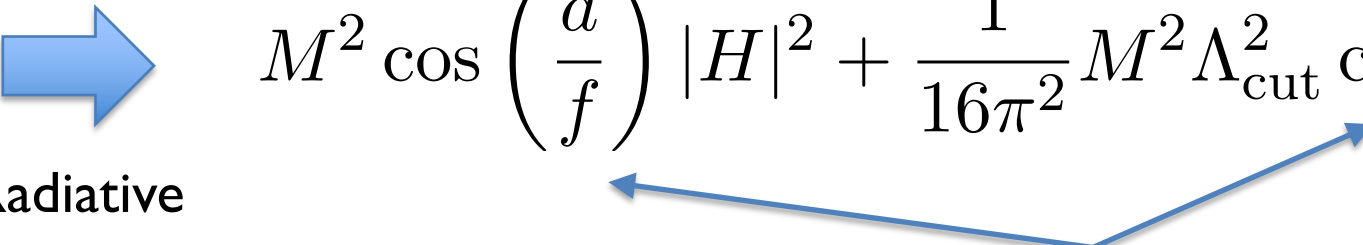
Freeze-in ALP dark matter

SHI, Jeong '19

Cold DM with $m_a > 100$ keV while cosmologically stable

$$M^2 \cos\left(\frac{a}{f}\right) |H|^2$$

Suppose that this is an only operator that breaks ALP shift symmetry.


$$M^2 \cos\left(\frac{a}{f}\right) |H|^2 + \frac{1}{16\pi^2} M^2 \Lambda_{\text{cut}}^2 \cos\left(\frac{a}{f}\right)$$

Radiative correction

No phase difference

$$V = -M^2 \cos\left(\frac{a}{f}\right) |H|^2 - \frac{1}{16\pi^2} M^2 \Lambda_{\text{cut}}^2 \cos\left(\frac{a}{f}\right)$$

ALP is stabilized at $a = 0$.



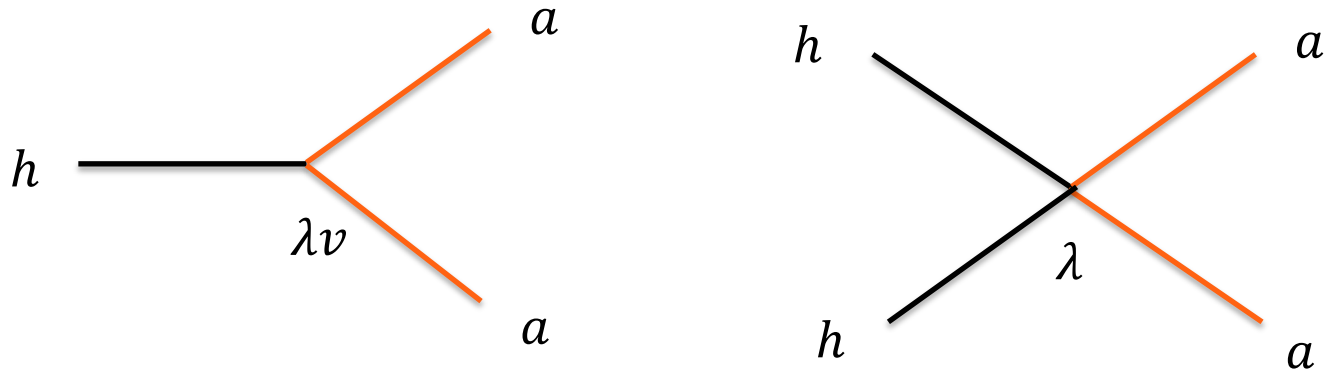
The vacuum preserves
 Z_2 symmetry : $a \rightarrow -a$.
 So ALP does not decay.

Around the vacuum,

$$V = \frac{1}{2} \lambda |H|^2 a^2 + \frac{1}{2} m_a^2 a^2 + \dots$$

where $\lambda = \left(\frac{M}{f}\right)^2$, $m_a \approx \frac{\sqrt{\lambda_{ha}}}{4\pi} \Lambda_{\text{cut}}$

: naturally small λ while having large enough mass !



Fixing λ to obtain the correct relic density,

$$m_a \approx 1 \text{ MeV} \left(\frac{\Lambda_{\text{cut}}}{1 \text{ TeV}} \right)^{4/5}$$

$$f \approx 10^8 \text{ GeV} \left(\frac{M}{1 \text{ TeV}} \right) \left(\frac{\Lambda_{\text{cut}}}{1 \text{ TeV}} \right)^{1/5}$$

Experimental probe?

Unfortunately very challenging as much as other freeze-in models

$$V = \frac{1}{2}\lambda|H|^2a^2 + \frac{1}{2}m_a^2a^2 + \dots \quad \lambda < 10^{-10}$$

For a light ALP around MeV, there should be a new degree of freedom around TeV which couples to the Higgs.

$$m_a \approx 1 \text{ MeV} \left(\frac{\Lambda_{\text{cut}}}{1 \text{ TeV}} \right)^{4/5}$$

Hidden-color confinement $\Lambda_{cut} = \Lambda_{HC}$

$$M^2 \cos\left(\frac{a}{f}\right) |H|^2 \quad \leftarrow \quad \underbrace{\frac{|H|^2}{\Lambda} N_H N_H^c}_{H L_H N_H^c + H^\dagger L_H^c N_H} + \frac{1}{32\pi^2} \frac{a}{f} G_H \tilde{G}_H$$

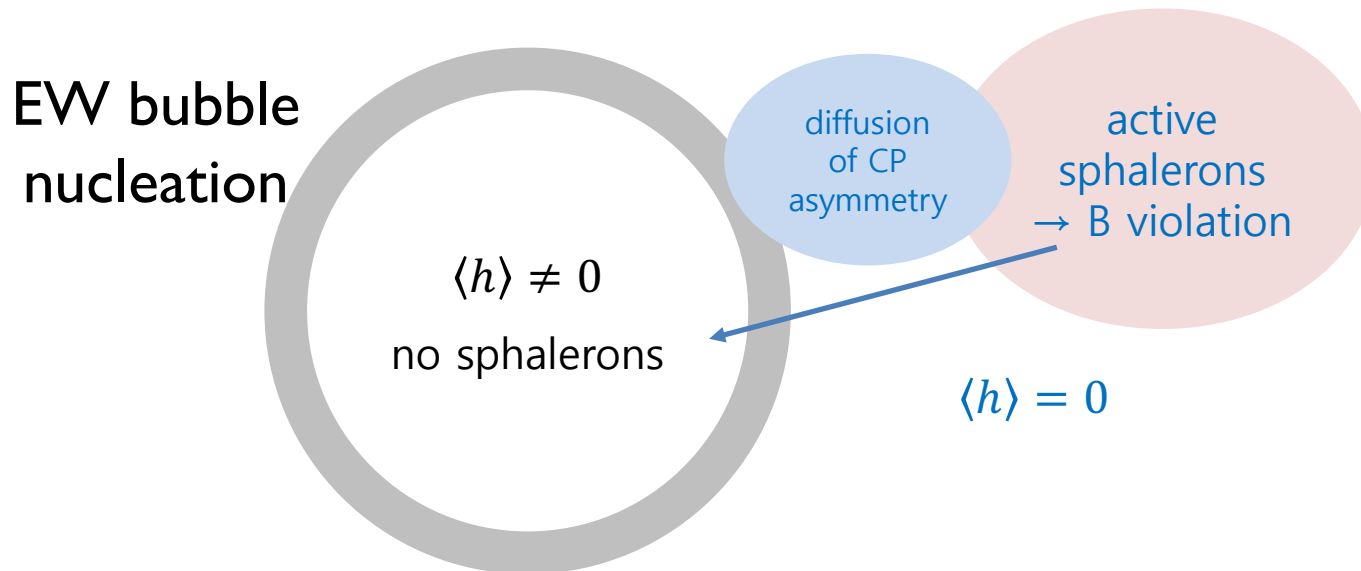
$$H L_H N_H^c + H^\dagger L_H^c N_H + m_L L_H L_H^c$$

$$(\Lambda = m_L)$$

Higgs can sizably mix with hidden-color η' meson of mass Λ_{cut} .

Electroweak Baryogenesis

Appealing framework for experimental testability



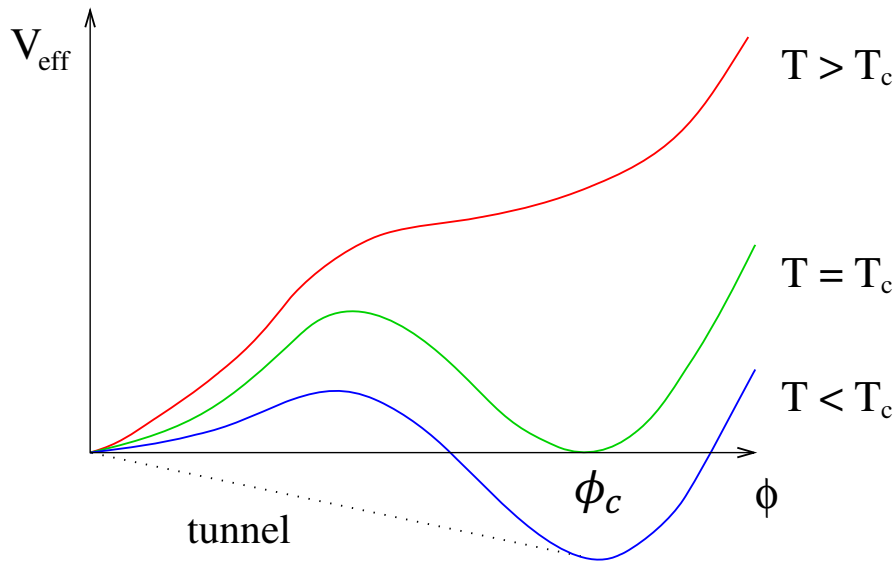
- i) B violation : sphalerons ✓
- ii) C & CP violation : CKM not enough ✗
- iii) Out of equilibrium : strongly first-order EWPT ✗



NP

Finite temperature Higgs potential

$$V_{\text{eff}}(\phi, T) \simeq D(T^2 - T_0^2)\phi^2 - \boxed{ET\phi^3} + \frac{\bar{\lambda}}{4}\phi^4$$



$$\frac{\phi_c}{T_c} = \frac{2E}{\bar{\lambda}} > 1$$

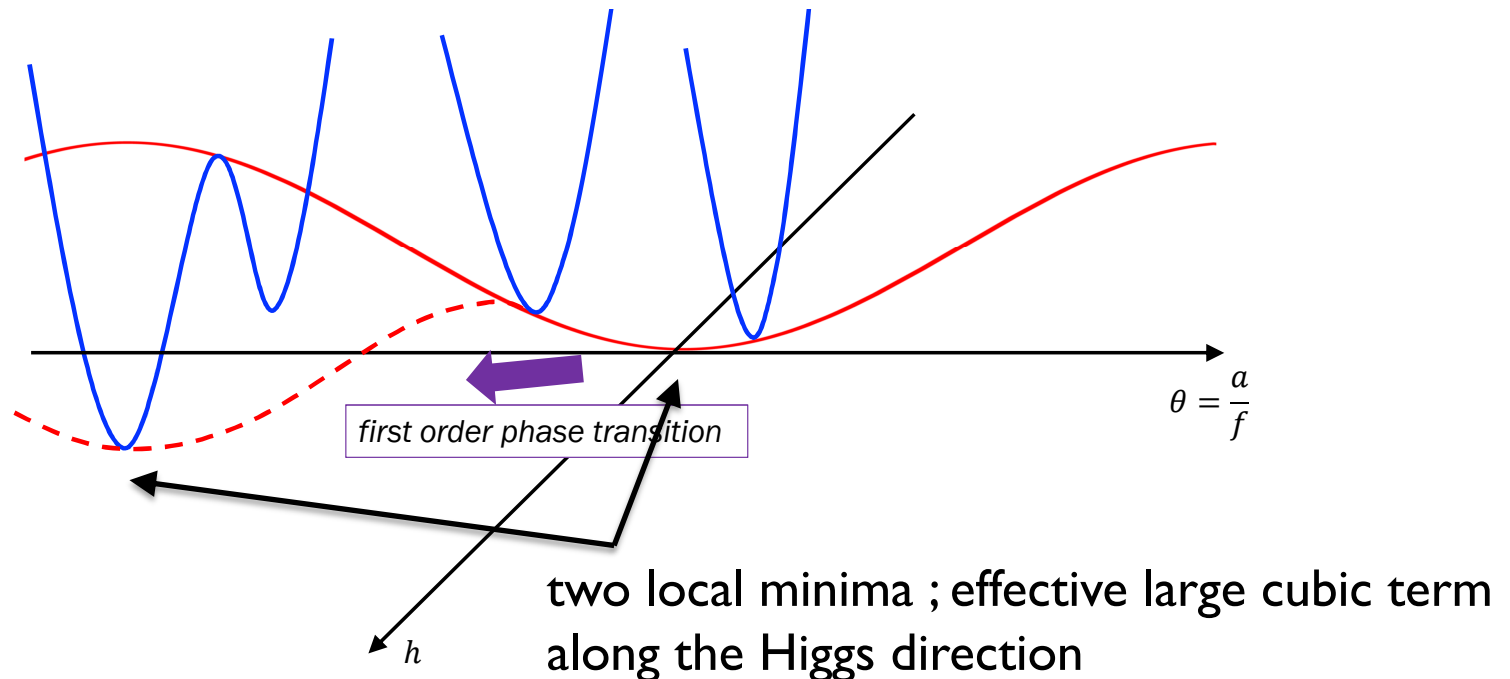
strongly first-order EWPT

- A large cubic term is needed.
- It makes two local minima in the Higgs potential.

Axionic electroweak baryogenesis

K S Jeong, T H Jung, C S Shin '18

$$V = \left[\mu^2 - M^2 \cos \left(\frac{a}{f} + \alpha \right) \right] |H|^2 - \Lambda^4 \cos \left(\frac{a}{f} \right) + \lambda |H|^4$$



$$V = \left[\mu^2 - M^2 \cos \left(\frac{a}{f} + \alpha \right) \right] |H|^2 - \Lambda^4 \cos \left(\frac{a}{f} \right) + \lambda |H|^4$$

The potential only depends on $\theta = \frac{a}{f}$, insensitive to f .

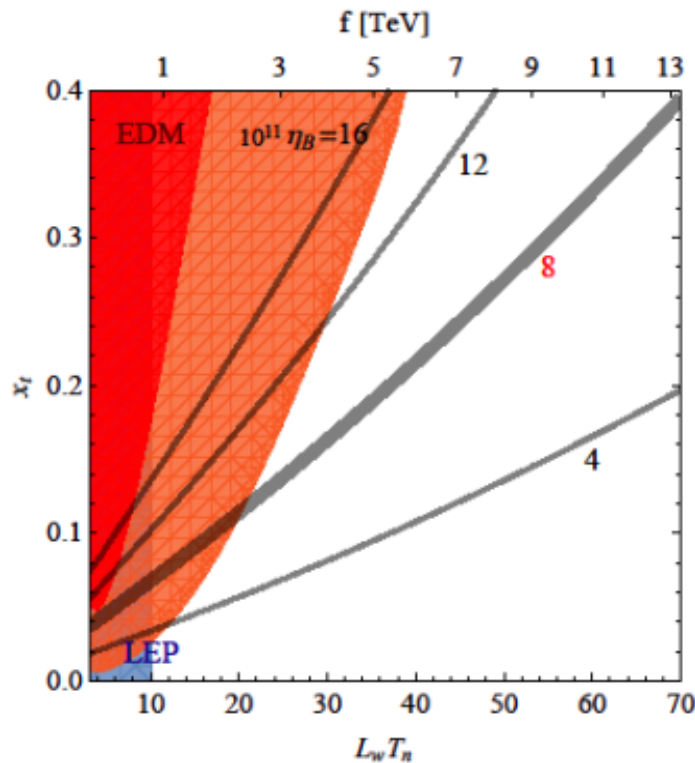
→ Weak coupling limit $f \gg v$.

→ Free from EDM and LHC constraints while allowing testability.

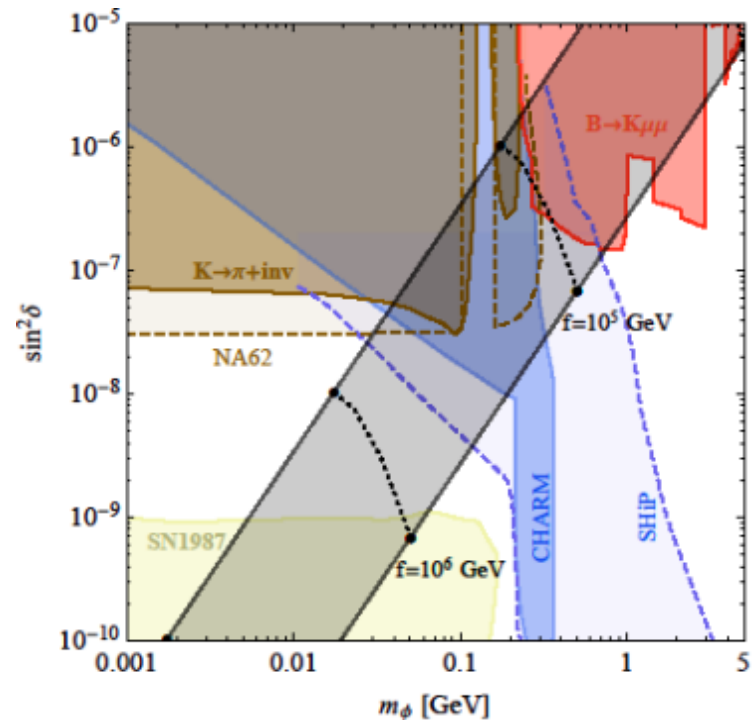
CP violation from top Yukawa : $Y_t = y_t + x_t e^{i\phi/f}$

For the correct baryon asymmetry, $3 \text{ TeV} \leq f \leq 10 \text{ TeV}$
 $\text{GeV} \leq m_a \leq 20 \text{ GeV}$

Electron EDM



Beam dump



Spontaneous EWBG with 2nd order EWPT

SHI, K.S. Jeong under progress

$$\theta_{EW} W_L \widetilde{W}_L = \partial_\mu \theta_{EW} J_B^\mu = \underbrace{\dot{\theta}_{EW} (n_B - n_{\bar{B}})}_{\text{B wash-out by sphalerons}}$$

$\dot{\theta}_{EW} \neq 0$: energy preference of baryons over anti-baryons

➔

$$\frac{dn_B}{dt} \approx \underbrace{\frac{3}{2} \frac{\Gamma_{\text{sph}}}{T}}_{\text{B violation}} \underbrace{\dot{\theta}_{EW}}_{\text{C \& CP and Lorentz violation}} - \frac{39}{4} \frac{\Gamma_{\text{sph}}}{T^3} n_B \quad \left. \vphantom{\frac{dn_B}{dt}} \right\} \text{B wash-out by sphalerons}$$



Baryogenesis has to happen just before/during the sphaleron process turns off (i.e. EWPT) to avoid a strong wash-out.

$$V = V_{\text{SM}} - \underbrace{M^2 |H|^2}_{M^2 < v^2} \cos \left(\underbrace{\frac{a}{F} + \alpha}_{\theta} \right) - \Lambda_{\text{HC}}^4 \cos \left(\frac{a}{F} \right) + \underbrace{\frac{a}{f} W_L \widetilde{W}_L}_{\theta_{EW}}$$

As EWPT happens, a starts to move due to the Higgs VEV dependent potential, while not significantly affecting the SM phase transition structure.

$$\frac{dn_B}{dt} \approx \underbrace{\frac{3}{2} \frac{\Gamma_{\text{sph}}}{T} \dot{\theta}_{EW}}_{\text{persistent time scale} \sim \text{Hubble}} - \underbrace{\frac{39}{4} \frac{\Gamma_{\text{sph}}}{T^3} n_B}_{\text{persistent time scale} \sim \text{Hubble}}$$

$$\dot{\theta}_{EW} = \frac{\dot{a}}{f} = \frac{F}{f} \dot{\theta} \sim \frac{F}{f} \times \text{Hubble}$$


For $F \sim f$, the produced baryon asymmetry is negligible.


For the correct baryon asymmetry, $F \sim 10^9 f$ (clockwork)

Top-down approach

Origin of defining properties of ALP ?

i) Discrete gauge symmetry : $a \rightarrow a + 2\pi n f$

 $\frac{a(x)}{f} = \theta(x)$ angle variable

 $\left\{ \begin{array}{l} \Phi(x) = (S(x) + f)e^{i\theta(x)} \\ B_2(x) = \frac{1}{2\pi}\omega_i^{(1,1)}\theta^i(x) \\ \star H_3(x) = d\theta(x) \end{array} \right\}$
axial degree of freedom of
(elementary or composite)
complex scalar field

string theoretic form
fields

ii) Perturbative shift symmetry : $a \rightarrow a + c$

; Global $U(1)$ \rightarrow incompatible with quantum gravity (black hole argument)

\rightarrow cannot be a fundamental symmetry

\rightarrow To be identified as accidental symmetry from gauge symmetry or discrete symmetry

Example in the SM

: $U(1)_B$ and $U(1)_L$ accidentally originated from the SM gauge symmetries

ALP physics often requires a very high quality of global $U(1)$.

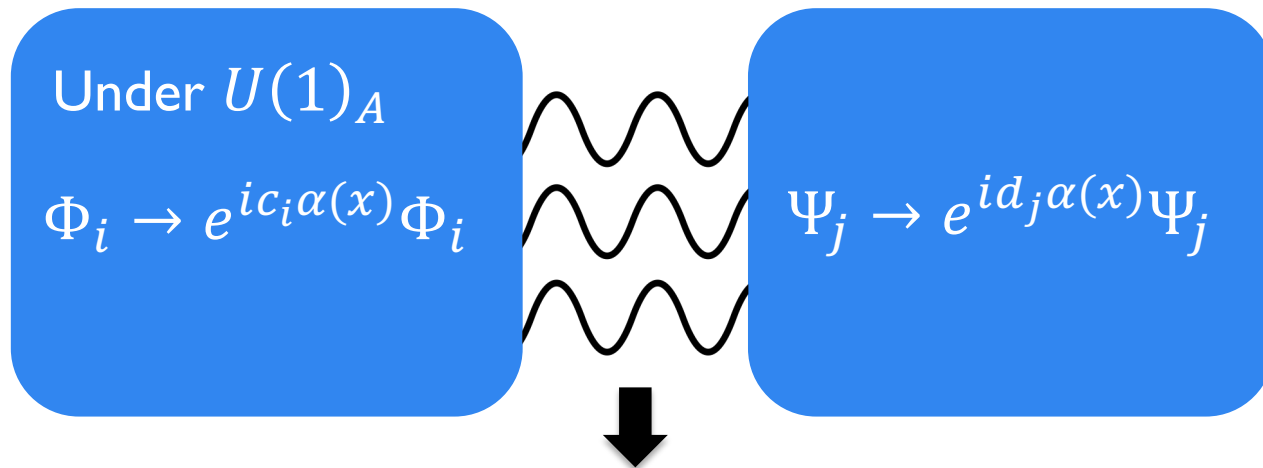
eg) QCD axion, Freeze-in ALP DM, ...

How to obtain it specifically?

High quality global $U(1)_{PQ}$: gauge symmetry origin

Fukuda, Ibe, Suzuki, Yanagida '17

Two **sequestered** sectors interacting with each other only via gauge interactions ($U(1)_A$ and $SU(N)_C$)

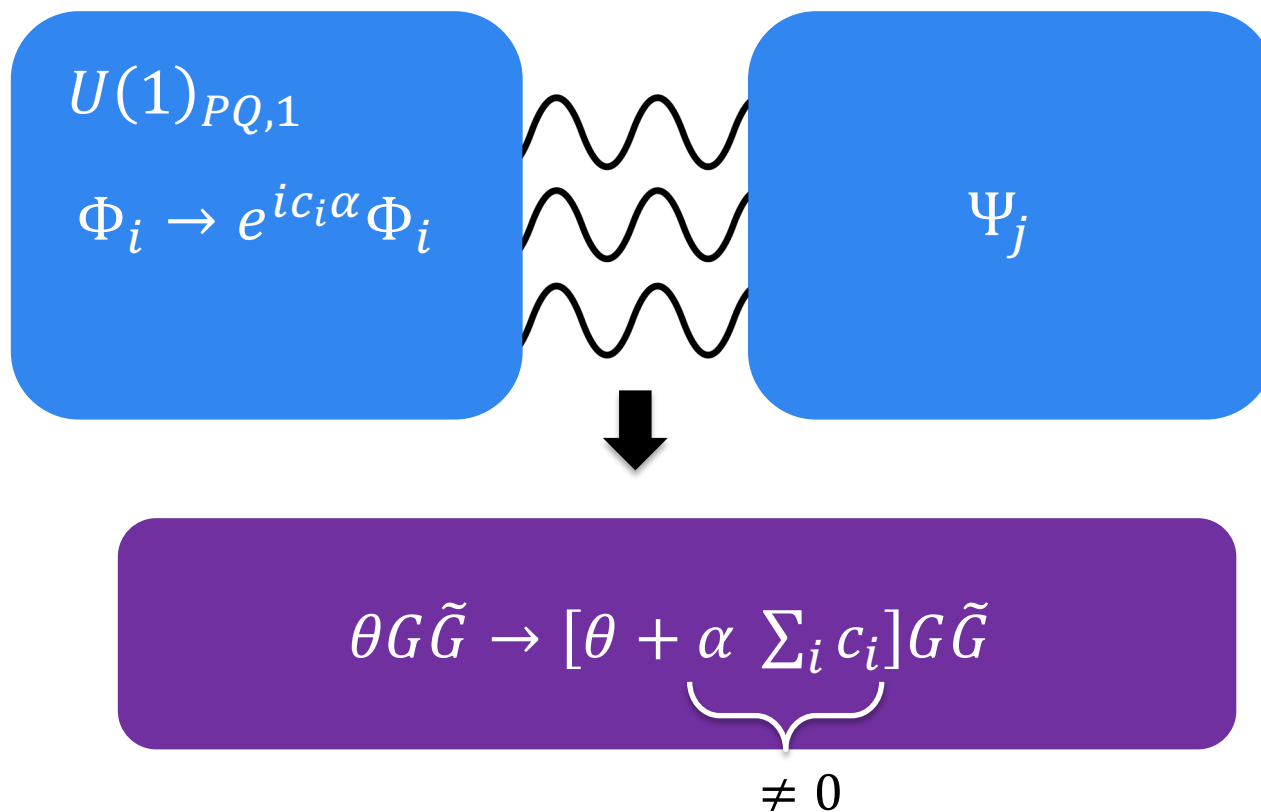


$$\theta G \tilde{G} \rightarrow \left[\theta + \alpha(x) \underbrace{\left(\sum_i c_i + \sum_j d_j \right)} \right] G \tilde{G}$$

0 for $U(1)_A \times [SU(N)_C]^2$
anomaly cancellation

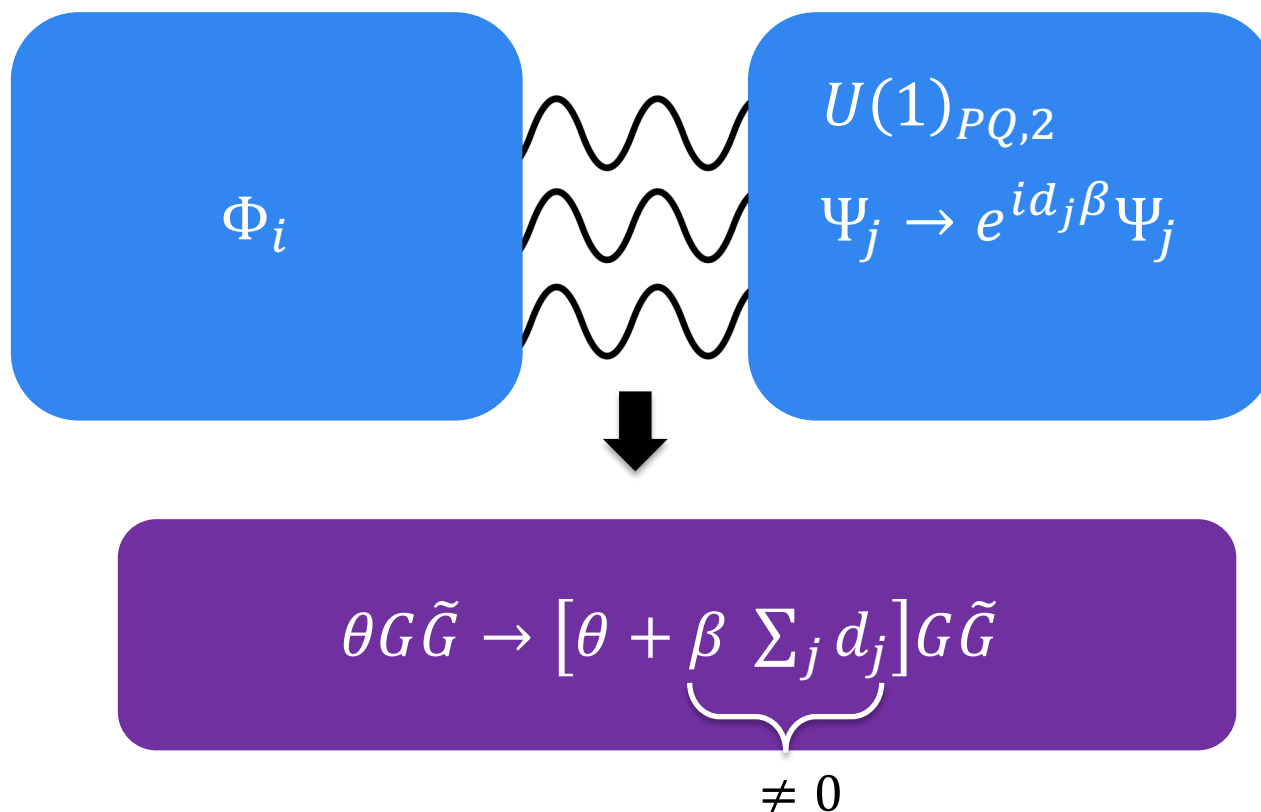
High quality global $U(1)_{PQ}$: gauge symmetry origin

We then have two accidental $U(1)_{PQ}$ broken only by anomalies.



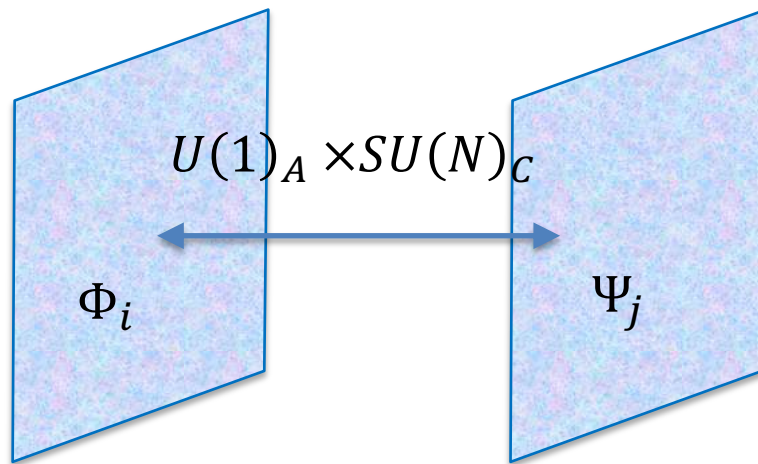
High quality global $U(1)_{PQ}$: gauge symmetry origin

We then have two accidental $U(1)_{PQ}$ broken only by anomalies.



The previous construction needs a good sequestering of the two sectors. How could it be?

Extra dimensional locality



H C Cheng and D E Kaplan '01
Izawa, Watari, Yanagida '02

...

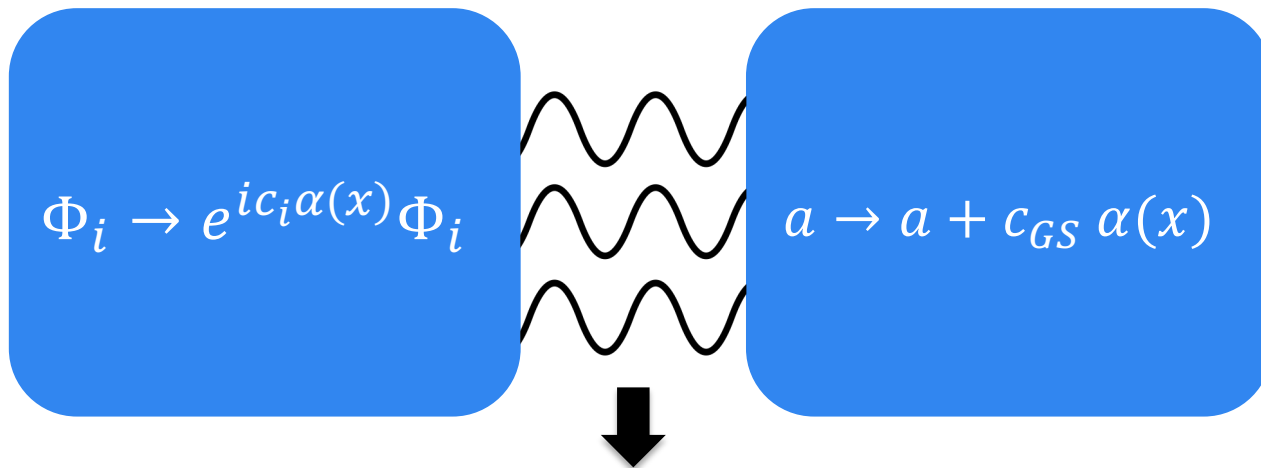
Implications for bottom-up ALPs? To be studied

$U(1)_A$ in string theory

Witten '84

Matter sector

Green-Schwarz modulus



$$\theta G \tilde{G} \rightarrow [\theta + \alpha(x) (\underbrace{\sum_i c_i + c_{GS}})] G \tilde{G}$$

0 for $U(1)_A \times [SU(N)_C]^2$
anomaly cancellation

$U(1)_A$ gauge boson mostly gains its mass by eating GS modulus remaining accidental $U(1)_A$ for the matter sector.

Matter sector

$$\Phi_i \rightarrow e^{ic'_i \alpha} \Phi_i$$



$$\theta G \tilde{G} \rightarrow [\theta + \underbrace{\alpha \sum_i c'_i}_{\neq 0}] G \tilde{G}$$

In order to preserve supersymmetry, however,

$$D_A = \sum_i q_i^2 |\Phi_i|^2 - \xi_{\text{FI}}^2 = 0$$

$U(1)_A$ Fayet-Iliopoulos term

$$\xi_{\text{FI}} \sim \frac{\alpha_{\text{YM}}}{4\pi} M_{\text{Pl}} \quad \text{in most of compactifications}$$



$$f \sim \langle \Phi_i \rangle \sim M_{\text{GUT}}$$

It seems hard to get a low scale ALP decay constant.

The situation is similar for stringy form field ALPs.

Low scale ALP decay constant from large internal volume

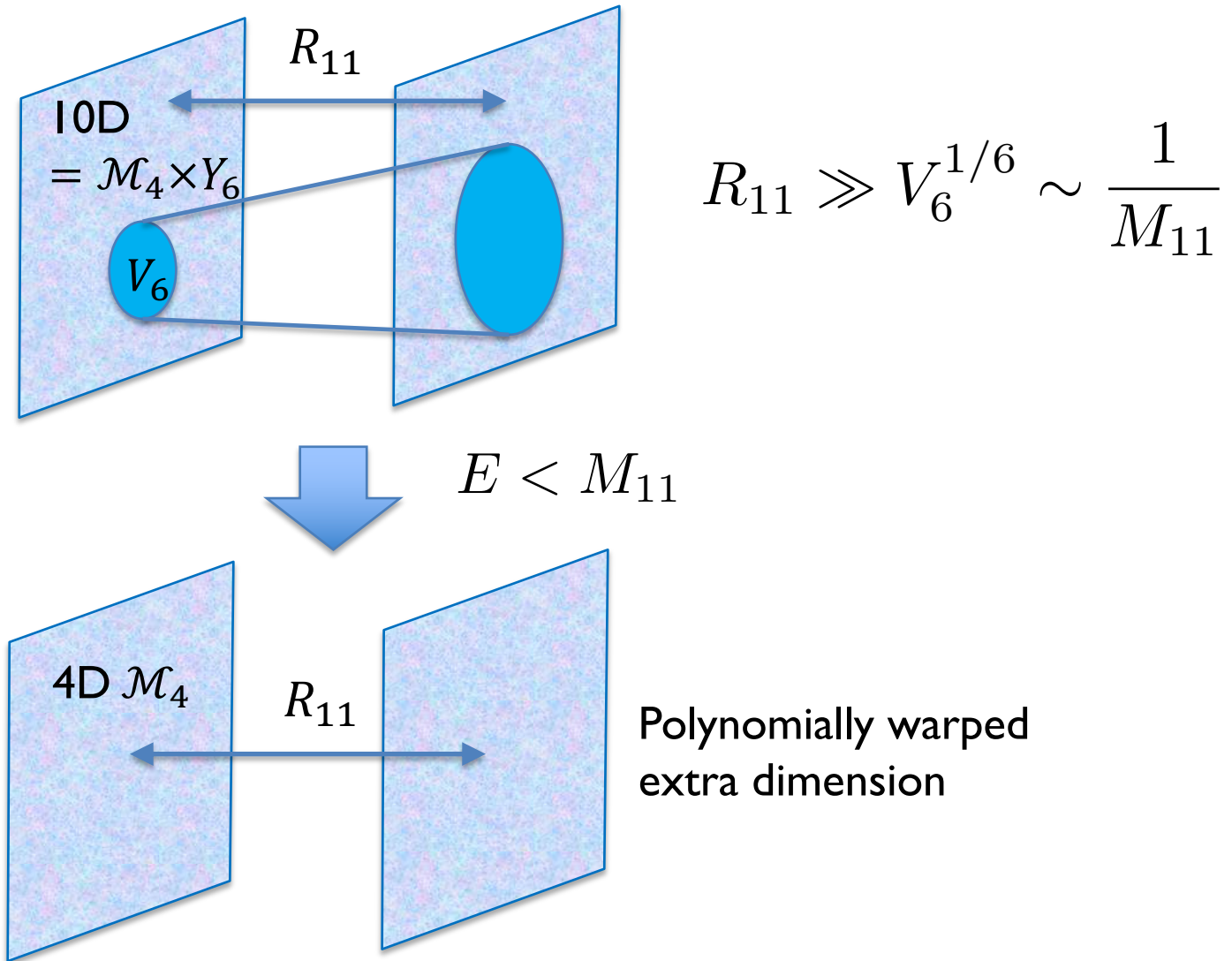
- Large Volume Scenario in Type IIB

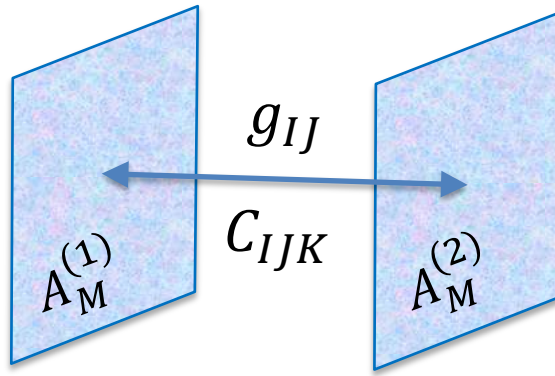
Balasubramanian, Berglund,
Conlon, Quevedo '05

- Warped large IID in Heterotic M-theory

Svrcek, Witten '06
SHI, Nilles, Olechowski '19

Heterotic M-theory : 5D effective description





4D axions : C_{mn11} , $C_{\mu\nu11}$

gauge fixing

$$S_{\sigma-A_{11}} = - \int d^5x \frac{1}{2} \eta^{\mu\nu} \partial_\mu \phi_L \partial_\nu \phi_L + \frac{1}{2} \eta^{\mu\nu} \partial_\mu \phi_R \partial_\nu \phi_R \\ + e^{2px_{11}} \frac{1}{2} [(\partial_{11} \phi_L + m_{\mathcal{A}} \phi_L)^2 + (\partial_{11} \phi_R - m_{\mathcal{A}} \phi_R)^2]$$

: two continuum clockwork axions

Choi, SHI, C S Shin '17

zero modes

$$\mathcal{L}_{4D} = -\frac{1}{2}(\partial_\mu a_1)^2 - \frac{1}{2}(\partial_\mu a_2)^2 - \epsilon_k \partial_\mu a_1 \partial^\mu a_2 + \frac{1}{16\pi^2} \left[\frac{a_1}{f_1} \text{tr} F_{(1)} \tilde{F}_{(1)} + \frac{a_2}{f_2} \text{tr} F_{(2)} \tilde{F}_{(2)} \right]$$

: two axions with kinetic mixing

$$\mathcal{L}_{4D} = -\frac{1}{2}(\partial_\mu a_1)^2 - \frac{1}{2}(\partial_\mu a_2)^2 - \epsilon_k \partial_\mu a_1 \partial^\mu a_2 + \frac{1}{16\pi^2} \left[\frac{a_1}{f_1} \text{tr} F_{(1)} \tilde{F}_{(1)} + \frac{a_2}{f_2} \text{tr} F_{(2)} \tilde{F}_{(2)} \right]$$

$$\epsilon_k \sim \left(\frac{M_{11}}{M_{\text{Pl}}} \right)^5$$

$$f_1 \sim M_{11}$$

$$f_2 \sim M_{11} \sqrt{\frac{M_{11}}{M_{\text{Pl}}}}$$

Visible sector



M_{11} can be lowered up to TeV scale. SHI, Nilles, Olechowski '18

→ Successful QCD axion with M_{11} lying in the axion window

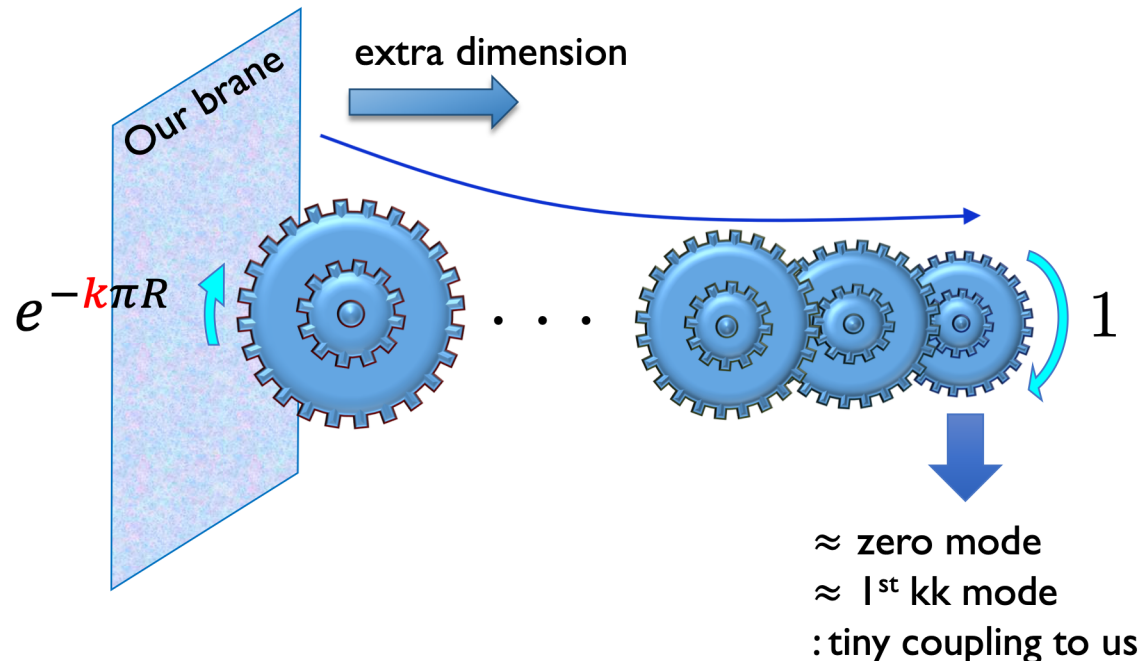
SHI, Nilles, Olechowski '19

Kaluza-Klein axions as ALPs from heterotic M-theory

SHI, Nilles, Olechowski '19

$$M_n \approx 9n \text{ GeV} \left(\frac{M_{11}}{10^{10} \text{ GeV}} \right)^{9/4}$$

$$f_{L1}^{(n)} \sim f_{R1}^{(n)} \sim \frac{M_P}{n^{3/10}} > 2.4 \times 10^{15} \text{ GeV} \left(\frac{M_{11}}{10^{10} \text{ GeV}} \right)^{3/8}$$



Conclusion

- A new physics sector seems to be quite weakly coupled to the SM.
- ALP provides a natural candidate as a weakly coupled portal for a new physics sector to interact with the SM.
- It has good potential to resolve the SM problems having experimental testability.
- From UV perspective, the origin of ALP shift symmetry needs a non-trivial structure in the theories, which may have implications for low energy physics.