

Interpretation of the diphoton resonance in effective field theory

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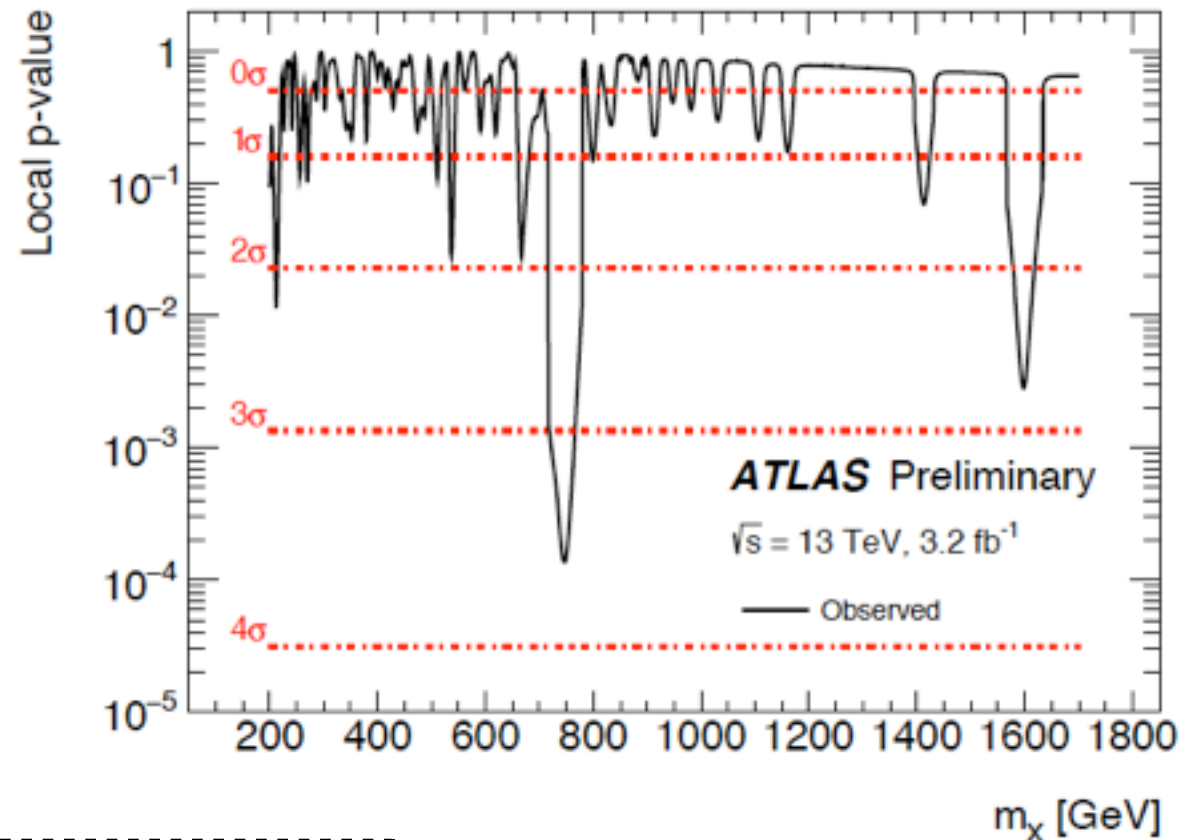
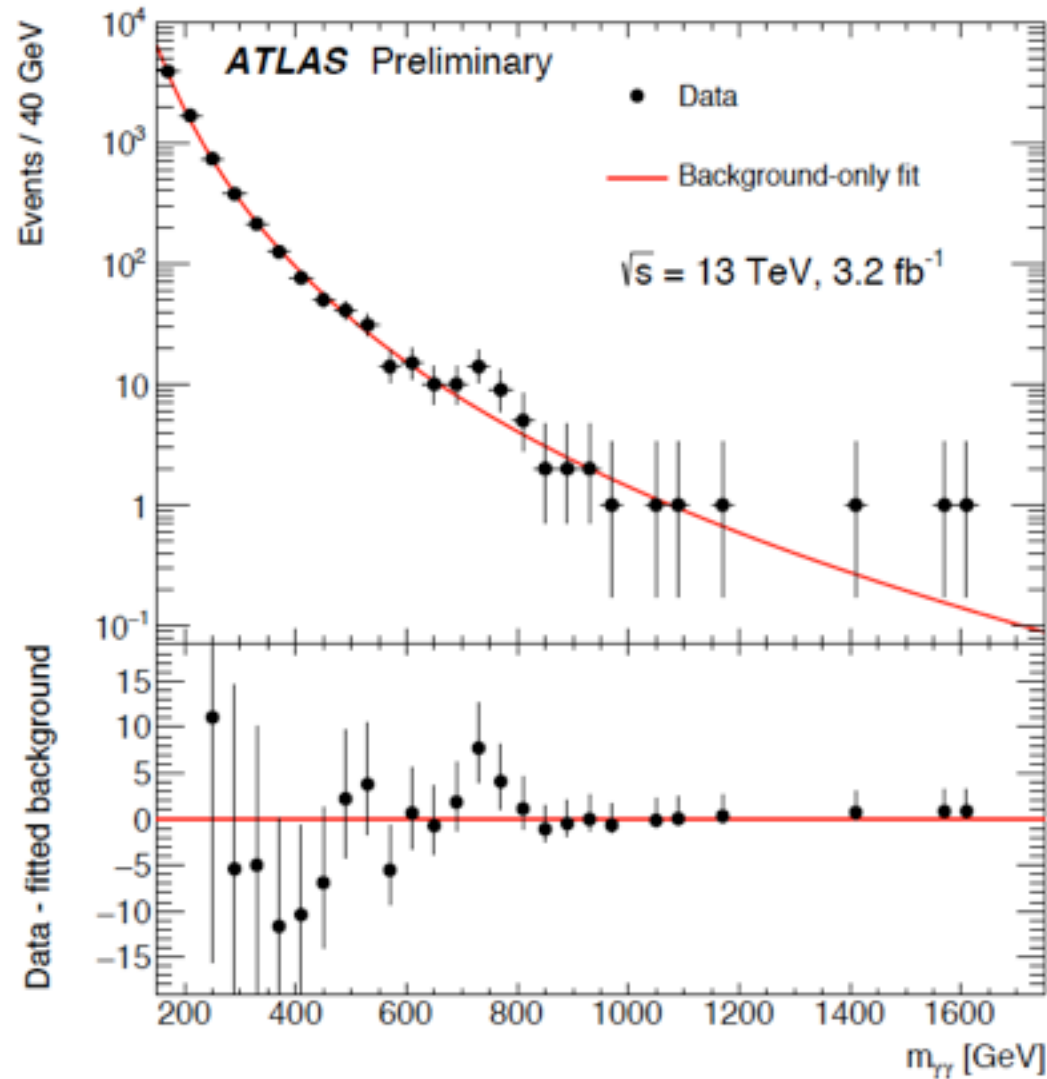
In collaboration with C. Han, M. Park, V. Sanz,
1512.06376 [hep-ph].

Overview on the recent diphoton excess at the LHC Run-2
IBS-CTPU, Jan 8, 2015

Outline

- Diphoton resonance
- Effective field theory for diphoton
- Diphoton as KK graviton
- KK graviton as DM mediator
- Spin determination
- Conclusions

Diphoton excesses



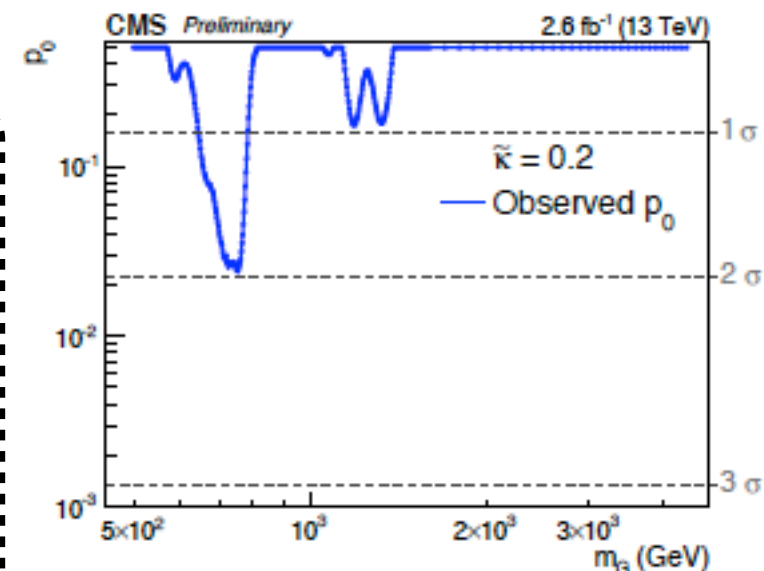
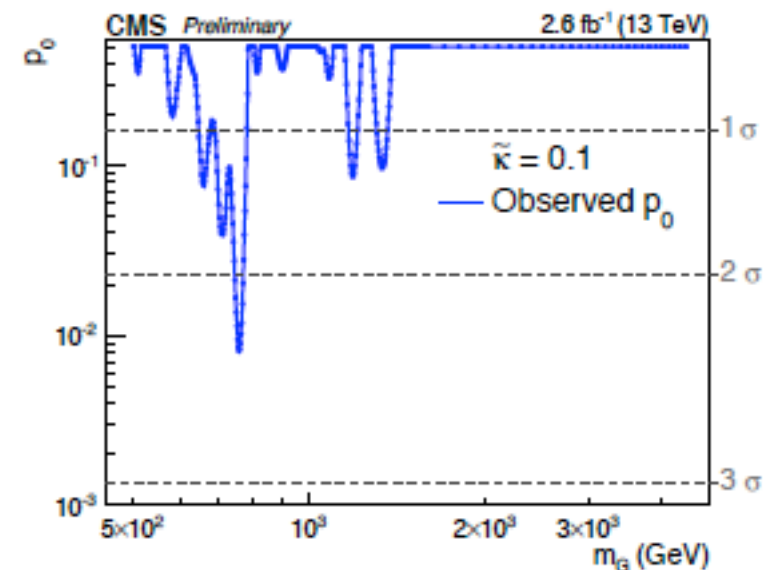
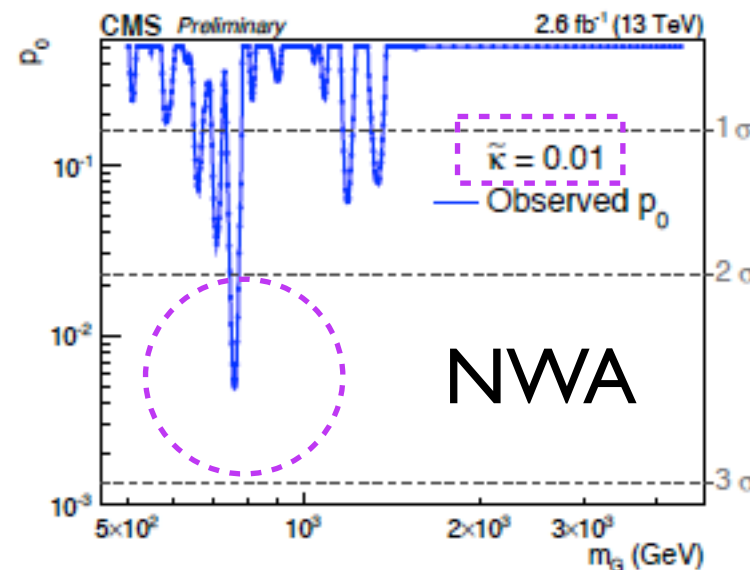
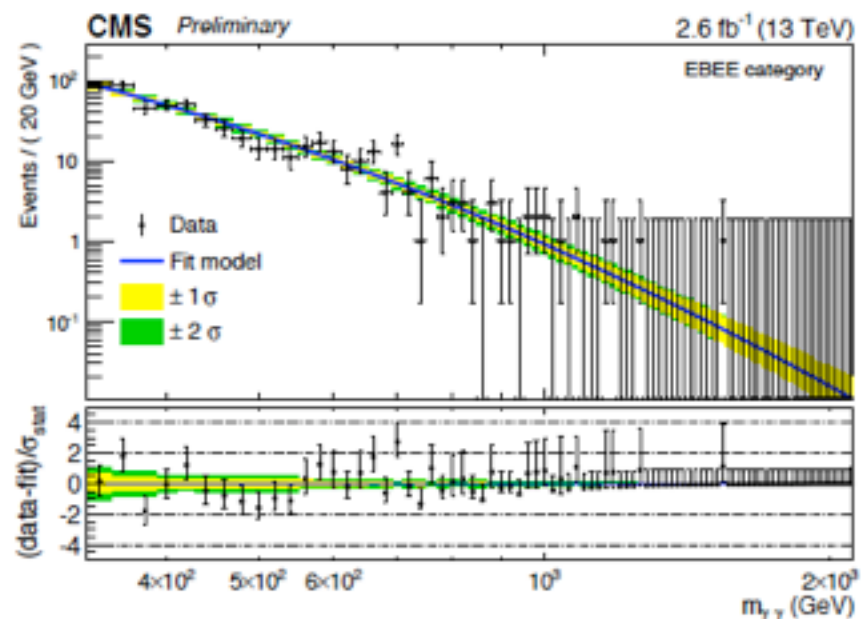
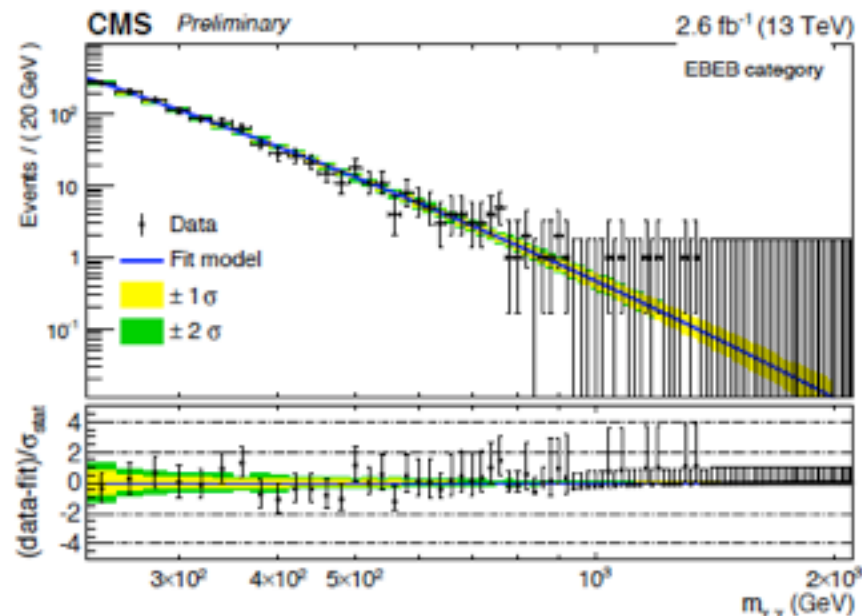
$$\begin{aligned}
 &|\eta^\gamma| < 2.37 \quad \text{"cuts"} \\
 &E_T^\gamma > 25 \text{ GeV} \\
 &E_T^{\gamma^{(1,2)}} / m_{\gamma\gamma} > 0.4(0.3)
 \end{aligned}$$

+NWA hypothesis

- Diphoton invariant mass in ATLAS in Run 2 shows 3.6σ (2.0σ) excesses at about 750 GeV under NWA.

Large width: $\Gamma = 45 \text{ GeV}$ $\longrightarrow$ 3.9σ (2.3σ) excesses

What about CMS ?



$$\frac{\Gamma_G}{m_G} \sim \tilde{\kappa}^2$$

$p_T^\gamma > 75 \text{ GeV}$
 $|\eta^\gamma| < 1.44,$
 or $1.57 < |\eta^\gamma| < 2.5$
 $|\eta^{\gamma^1}| < 1.44$ (no EEEE)
 $m_{\gamma\gamma} > 230 \text{ GeV}$
 (EBEE : 320 GeV)

“cuts”

- CMS in Run 2 shows 2.6σ excess at 760GeV from EBEE category under NWA. (EBEB and EBEE only.)

Diphoton cross section

- Required diphoton cross sections are

$$\sigma(pp \rightarrow \gamma\gamma)_{\text{ATLAS}} = (10 \pm 3) \text{ fb},$$

$$\sigma(pp \rightarrow \gamma\gamma)_{\text{CMS}} = (6 \pm 3) \text{ fb}.$$

- ATLAS data fit favors a large width $\sim 45\text{GeV}$, which leads to smaller significance in CMS data.

cf. Bin size: ATLAS $\sim 50\text{GeV}$, CMS $\sim 20\text{GeV}$)

Width information is not settled!

Diphoton @ 8TeV vs 13TeV

- pp cross section: $\sim \text{pdf}^2/s$ (e.g. ggF)

$$\sigma(pp \rightarrow X \rightarrow \gamma\gamma) = (2J + 1) \frac{C_{gg} K_{gg}}{M_X s} \Gamma(X \rightarrow gg) \cdot \text{BR}(X \rightarrow \gamma\gamma),$$

$$C_{gg} = \frac{\pi^2}{8} \int_{\tau_X}^1 \frac{dx}{x} g(x) g\left(\frac{\tau_X}{x}\right), \quad \tau_X = \frac{M_X^2}{s}; \quad K_{gg} : \text{K-factor.}$$

- Parton-luminosity ratio: $r \equiv \frac{\sigma_{13 \text{ TeV}}}{\sigma_{8 \text{ TeV}}} = \frac{(C_{gg}/s)_{13 \text{ TeV}}}{(C_{gg}/s)_{8 \text{ TeV}}}.$

$\sqrt{s}$	$C_{b\bar{b}}$	$C_{c\bar{c}}$	$C_{s\bar{s}}$	$C_{d\bar{d}}$	$C_{u\bar{u}}$	C_{gg}
8 TeV	1.07	2.7	7.2	89	158	174
13 TeV	15.3	36	83	627	1054	2137

$T_{b\bar{b}}$	$T_{c\bar{c}}$	$T_{s\bar{s}}$	$T_{d\bar{d}}$	$T_{u\bar{u}}$	T_{gg}
5.4	5.1	4.3	2.7	2.5	4.7

[Franceschini et al, 1512.04933]

CMS Run I limit:
0.5~2 fb for 700-800GeV

cf. J. Ellis et al: true with
background increase, 1512.05327

Translated to
2.4~9.4 fb in Run 2:
Consistent with
diphoton excess .

Other searches

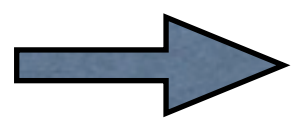
- ggF production necessarily leads to dijet production.
- Diphoton resonance could also decay into other particles in electroweak sector by gauge invariance so it is constrained by Run I searches.

final state f	σ at $\sqrt{s} = 8 \text{ TeV}$			implied bound on $\Gamma(S \rightarrow f)/\Gamma(S \rightarrow \gamma\gamma)_{\text{obs}}$
	observed	expected	ref.	
$\gamma\gamma$	$< 1.5 \text{ fb}$	$< 1.1 \text{ fb}$	[6, 7]	$< 0.8 (r/5)$
$e^+e^- + \mu^+\mu^-$	$< 1.2 \text{ fb}$	$< 1.2 \text{ fb}$	[8]	$< 0.6 (r/5)$
$\tau^+\tau^-$	$< 12 \text{ fb}$	15 fb	[9]	$< 6 (r/5)$
$Z\gamma$	$< 4.0 \text{ fb}$	$< 3.4 \text{ fb}$	[10]	$< 2 (r/5)$
ZZ	$< 12 \text{ fb}$	$< 20 \text{ fb}$	[11]	$< 6 (r/5)$
Zh	$< 19 \text{ fb}$	$< 28 \text{ fb}$	[12]	$< 10 (r/5)$
hh	$< 39 \text{ fb}$	$< 42 \text{ fb}$	[13]	$< 20 (r/5)$
W^+W^-	$< 40 \text{ fb}$	$< 70 \text{ fb}$	[14, 15]	$< 20 (r/5)$
$t\bar{t}$	$< 550 \text{ fb}$	-	[16]	$< 300 (r/5)$
invisible	$< 0.8 \text{ pb}$	-	[17]	$< 400 (r/5)$
$b\bar{b}$	$\lesssim 1 \text{ pb}$	$\lesssim 1 \text{ pb}$	[18]	$< 500 (r/5)$
$j\bar{j}$	$\lesssim 2.5 \text{ pb}$	-	[5]	$< 1300 (r/5)$

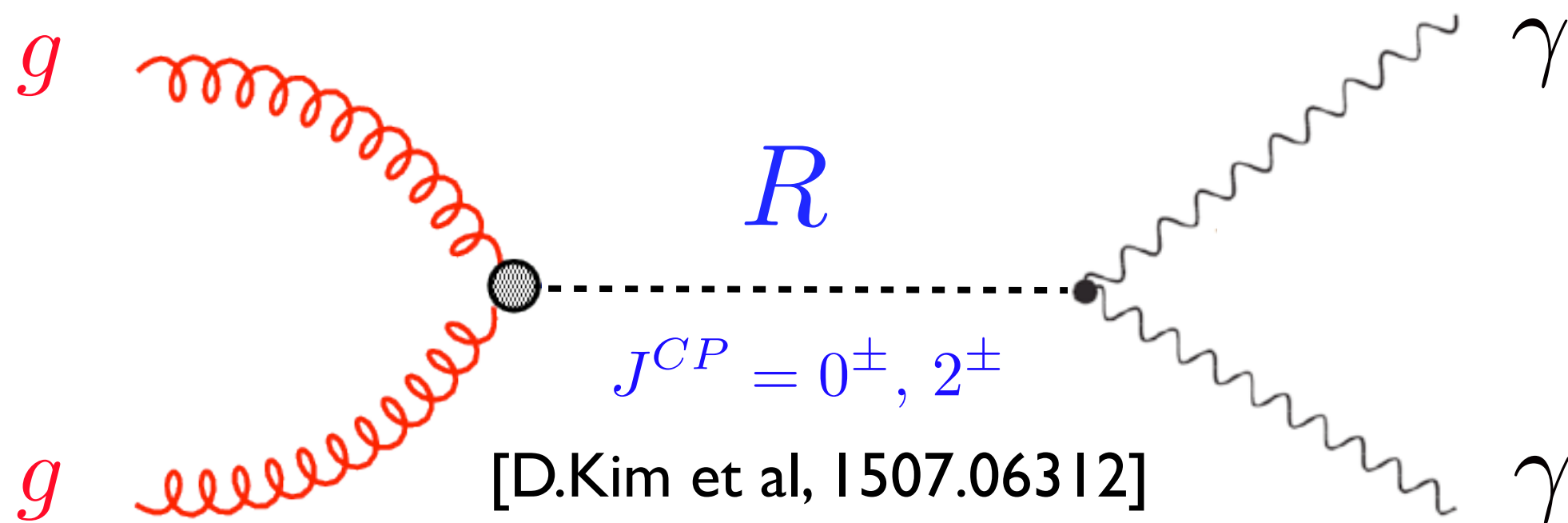
Rescaled Run I limits
[Franceschini et al,
1512.04933]

Diphoton from EFT

- Parton-luminosity ratio favors gluon fusion production.
- Landau-Yang theorem forbids on-shell massive spin-1 field from decaying into a photon pair.



spin-0 or spin-2 resonances with gluon and photon couplings in EFT.



cf. Other interpretation: photon fusion, cascade decays, non-resonant production, MET+photons, etc.

Conditions on resonances

- Both CP states for spin-0 are possible; only CP-even state for spin-2 is possible.

cf. On-shell CP-odd spin-2 can be produced only by VBF.

[D.Kim et al, 1507.06312]

- Diphoton production cross section leads to a model-independent constraint on decay rates:

$$(2J + 1)\Gamma(X \rightarrow gg) \times \text{BR}(X \rightarrow \gamma\gamma) = 6.2 \times 10^{-4} \text{ GeV} \left(\frac{\sigma_{pp \rightarrow \gamma\gamma}}{6 \text{ fb}} \right)$$

or

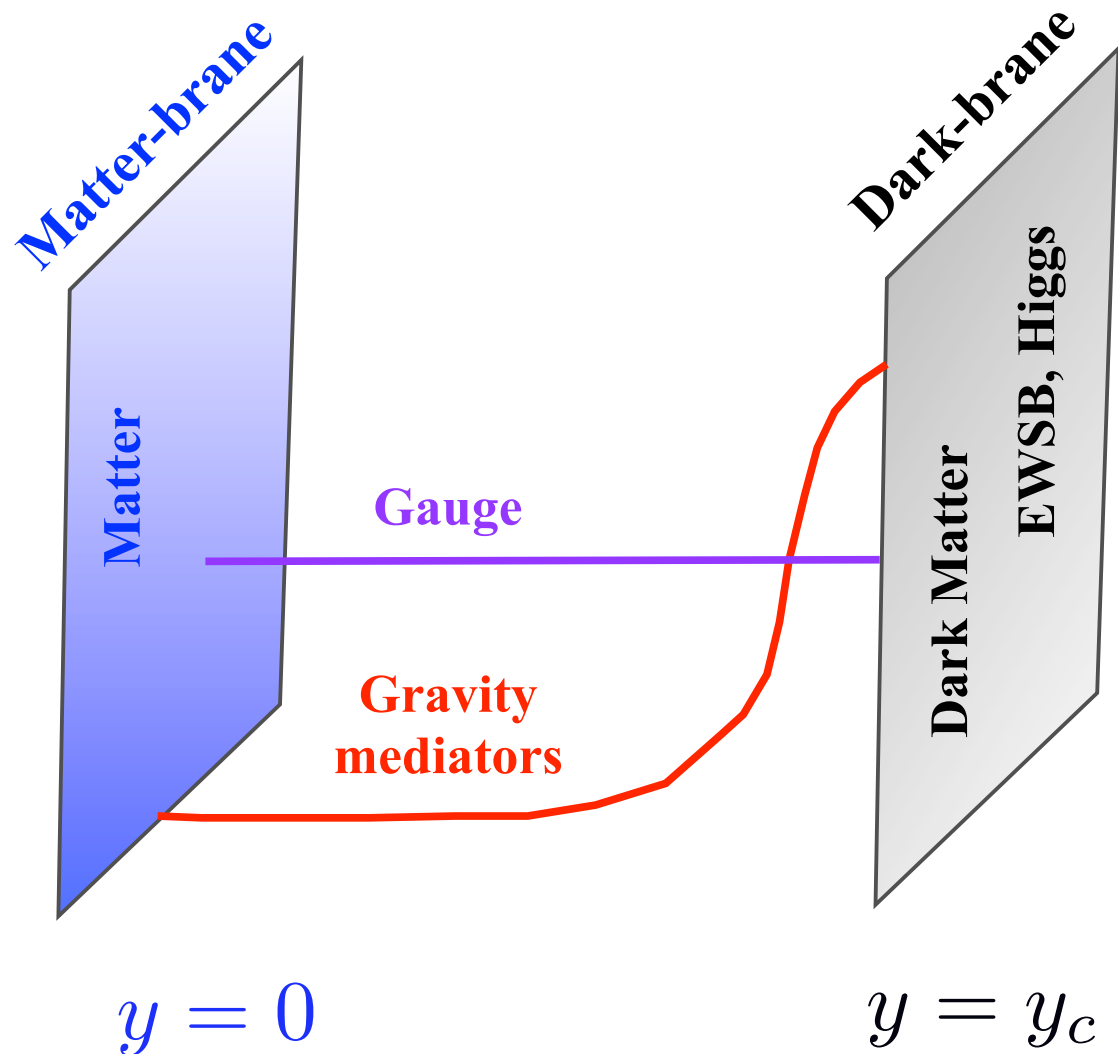
$$(2J + 1)\Gamma(X \rightarrow gg) \times \Gamma(X \rightarrow \gamma\gamma) = 2.8 \times 10^{-2} \text{ GeV}^2 \left(\frac{\Gamma_X / M_X}{0.06} \right) \left(\frac{\sigma_{pp \rightarrow \gamma\gamma}}{6 \text{ fb}} \right).$$

CP-even spin-0 in P. Ko's talk, etc;
CP-odd spin-0 in KS Jeong, JC Park, etc;
CP-even spin-2 is the topic of this talk.

Diphoton as KK graviton

Bulk-RS model (A)

- Higgs and dark matter on IR brane;
- SM fermions on UV brane (for flavor problem/proton stability).



$$ds^2 = e^{-2k|y|} \eta_{\mu\nu} dx^\mu dx^\nu + dy^2$$

$$M_P^2 = \frac{M^3}{k} (1 - e^{-2ky_c}),$$

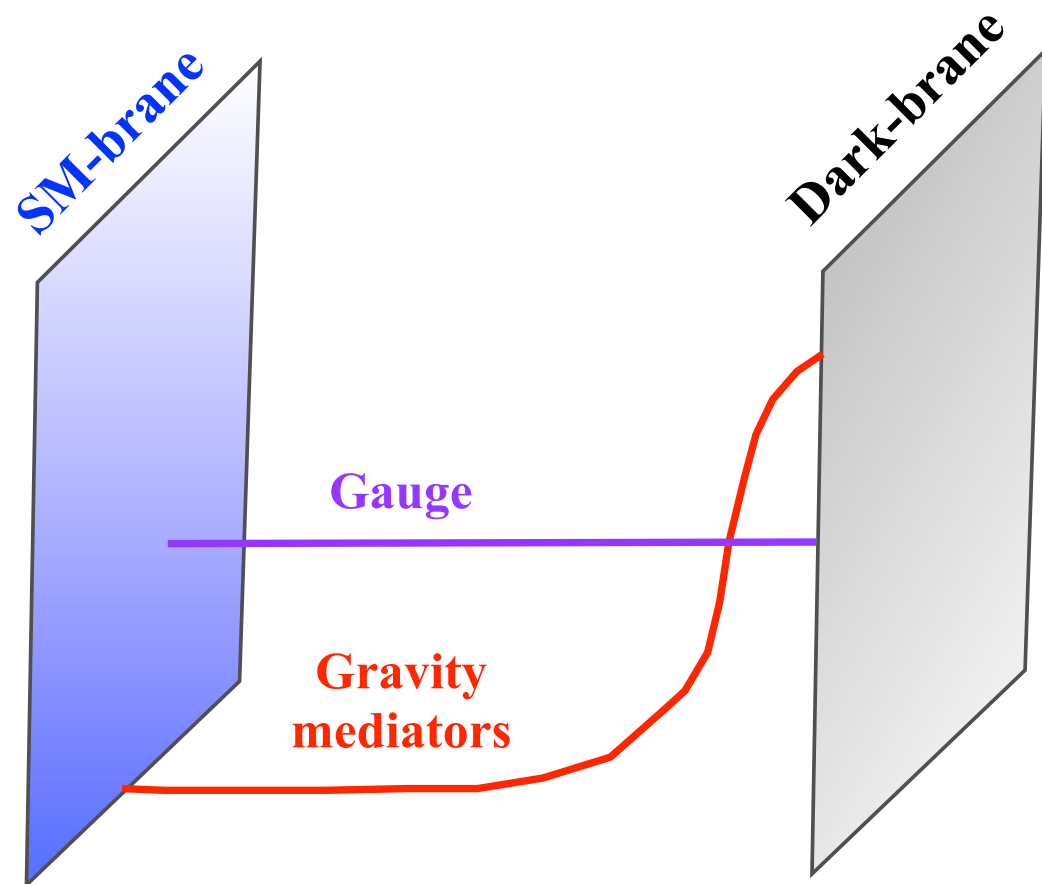
$$\frac{m_H}{M_P}, \frac{m_X}{M_P} \sim e^{-ky_c} \ll 1.$$

KK graviton strongly couples to Higgs doublet and dark matter, as well as bulk gauge bosons.

Dark-brane model (B)

[HML, M.Park, V. Sanz, 2013, 2014]

- DM on IR brane;
- All the SM particles on UV brane, except gauge bosons in bulk.



$$\frac{m_X}{M_P} \sim e^{-ky_c} \ll 1.$$

KK graviton strongly couples to dark matter and bulk (transverse) gauge bosons, depending on localization.

Focus for diphoton resonance!

Couplings of KK graviton

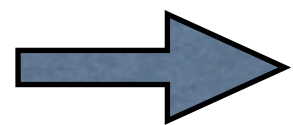
- KK graviton mass controlled by warp factor:

$$\frac{m_G}{\Lambda} = x_G \tilde{\kappa} : \Gamma(G \rightarrow gg) = \frac{x_G^2 \tilde{\kappa}^2}{10\pi} m_G.$$

e.g. in RS, $\tilde{\kappa} \equiv \frac{k}{M_P}$, $\Lambda = e^{-ky_c} M_P$, $x_G = 3.83$.

- Couplings to dark matter of spin 0, 1/2 or 1, and SM gauge bosons, are parametrized by c's:

$$\mathcal{L}_{\text{KK}} = -\frac{1}{\Lambda} G^{\mu\nu} \left[T_{\mu\nu}^{\text{DM}} + \sum_{a=1}^3 c_a \left(\frac{1}{4} g_{\mu\nu} F_a^{\lambda\rho} F_{\lambda\rho,a} - F_{\mu\lambda,a} F^{\lambda}_{\nu,a} \right) \right]$$



couplings to physical gauge bosons:

$$\begin{aligned} \mathcal{L}_{\text{KK}}^V = & -\frac{1}{\Lambda} G^{\mu\nu} \left[c_{\gamma\gamma} \left(\frac{1}{4} g_{\mu\nu} A^{\lambda\rho} A_{\lambda\rho} - A_{\mu\lambda} A^{\lambda}_{\nu} \right) + c_{Z\gamma} \left(\frac{1}{4} g_{\mu\nu} A^{\lambda\rho} Z_{\lambda\rho} - A_{\mu\lambda} Z^{\lambda}_{\nu} \right) \right. \\ & + c_{ZZ} \left(\frac{1}{4} g_{\mu\nu} Z^{\lambda\rho} Z_{\lambda\rho} - Z_{\mu\lambda} Z^{\lambda}_{\nu} \right) + c_{WW} \left(\frac{1}{4} g_{\mu\nu} W^{\lambda\rho} W_{\lambda\rho} - W_{\mu\lambda} W^{\lambda}_{\nu} \right) \\ & \left. + c_{gg} \left(\frac{1}{4} g_{\mu\nu} G^{\lambda\rho} G_{\lambda\rho} - G_{\mu\lambda} G^{\lambda}_{\nu} \right) \right] \end{aligned}$$

$$\begin{aligned} c_{\gamma\gamma} &= c_1 \cos^2 \theta_W + c_2 \sin^2 \theta_W, \\ c_{Z\gamma} &= (c_2 - c_1) \sin(2\theta_W), \\ c_{ZZ} &= c_1 \sin^2 \theta_W + c_2 \cos^2 \theta_W, \\ c_{WW} &= 2c_2, \\ c_{gg} &= c_3. \end{aligned}$$

Production/decay of KK graviton

- Production cross section for $gg \rightarrow G$

$\sqrt{s}=8$ TeV	$\sqrt{s}=13$ TeV	$\sigma_{13\text{TeV}}/\sigma_{8\text{TeV}}$	$\Gamma_{G \rightarrow gg}$
105 fb	465 fb	4.4	0.015 GeV

$$\Lambda=3 \text{ TeV and } c_3 = 0.1$$

- Visible decay rates

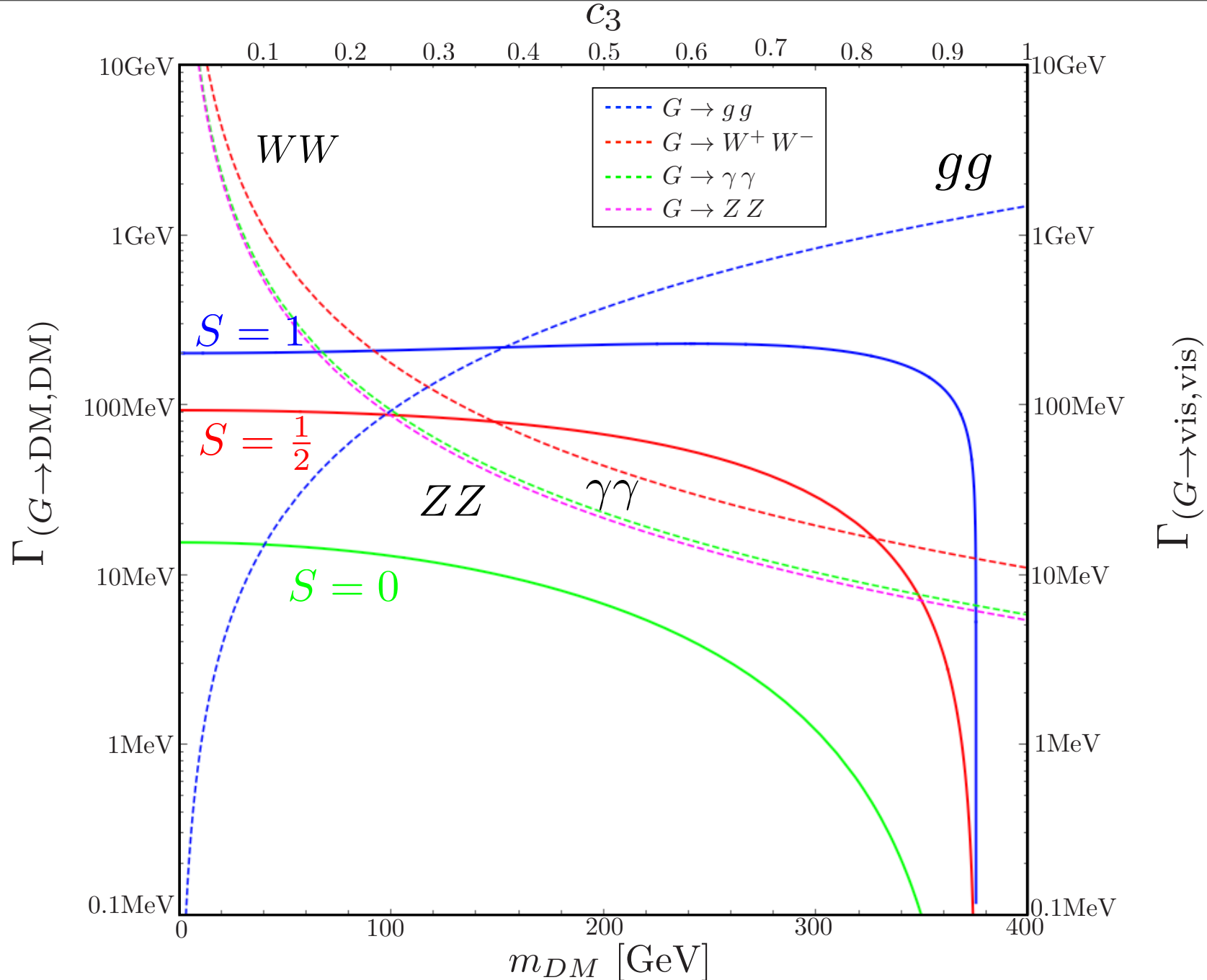
$$\begin{aligned}\Gamma(\gamma\gamma) &= \frac{c_{\gamma\gamma}^2 m_G^3}{80\pi\Lambda^2}, \\ \Gamma(Z\gamma) &= \frac{c_{Z\gamma}^2 m_G^3}{160\pi\Lambda^2} \left(1 - \frac{m_Z^2}{m_G^2}\right)^3 \left(1 + \frac{m_Z^2}{2m_G^2} + \frac{m_Z^4}{6m_G^4}\right), \\ \Gamma(ZZ) &= \frac{c_{ZZ}^2 m_G^3}{80\pi\Lambda^2} \left(1 - \frac{4m_Z^2}{m_G^2}\right)^{\frac{1}{2}} \left(1 - \frac{3m_Z^2}{m_G^2} + \frac{6m_Z^4}{m_G^4}\right), \\ \Gamma(WW) &= \frac{c_{WW}^2 m_G^3}{160\pi\Lambda^2} \left(1 - \frac{4m_W^2}{m_G^2}\right)^{\frac{1}{2}} \left(1 - \frac{3m_W^2}{m_G^2} + \frac{6m_W^4}{m_G^4}\right), \\ \Gamma(gg) &= \frac{c_{gg}^2 m_G^3}{10\pi\Lambda^2}.\end{aligned}$$

- Invisible decay rates

$$\begin{aligned}\Gamma(SS) &= \frac{N_f c_S^2 m_G^3}{960\pi\Lambda^2} \left(1 - \frac{4m_S^2}{m_G^2}\right)^{\frac{5}{2}}, \\ \Gamma(\chi\bar{\chi}) &= \frac{N_f c_X^2 m_G^3}{160\pi\Lambda^2} \left(1 - \frac{4m_X^2}{m_G^2}\right)^{\frac{3}{2}} \left(1 + \frac{8m_X^2}{3m_G^2}\right), \\ \Gamma(XX) &= \frac{N_f c_X^2 m_G^3}{960\pi\Lambda^2} \left(1 - \frac{4m_X^2}{m_G^2}\right)^{\frac{1}{2}} \left(13 + \frac{56m_X^2}{m_G^2} + \frac{48m_X^4}{m_G^4}\right).\end{aligned}$$

phase space suppressed

Henceforth, $c_1 = c_2$: no $Z\gamma$ decay. (discussion generalized otherwise.)



$\Gamma_G = 45 \text{ GeV}$, diphoton signal rates imposed;

$$\Lambda = 3 \text{ TeV}, m_G = 750 \text{ GeV}$$

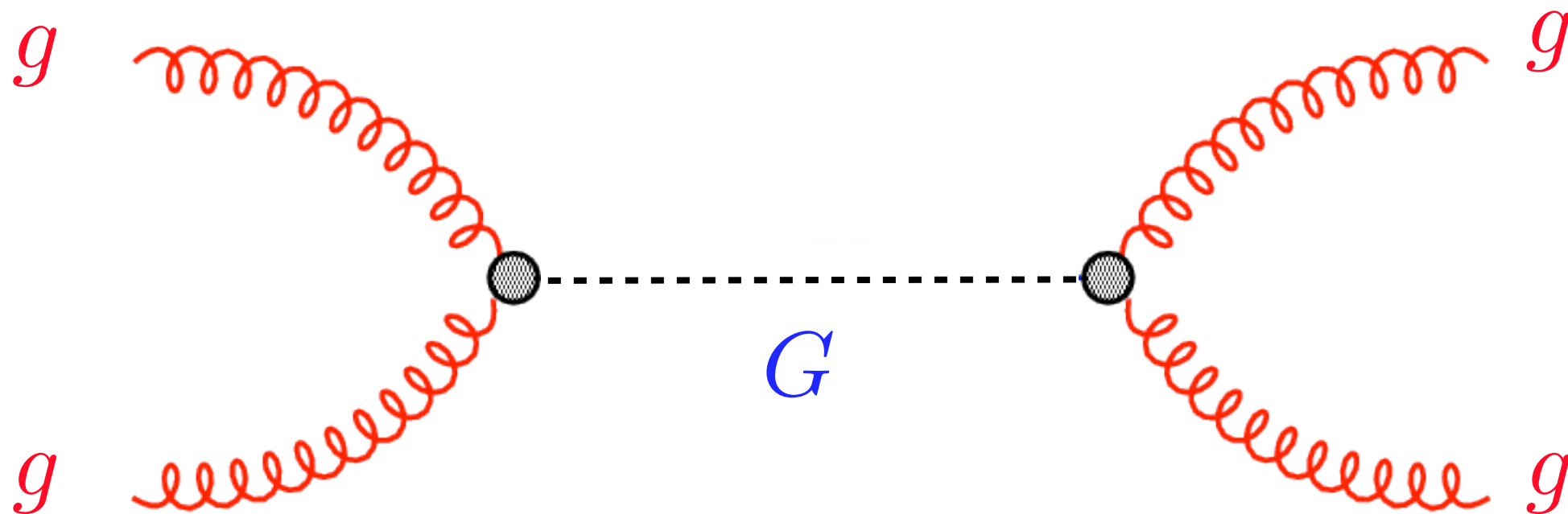
$c_X = 1$: Invisible decay rate is subdominant.

Vector DM is the largest.

Bounds on KK graviton

- Null results in searches for KK graviton in Run I & II set limits on other channels.

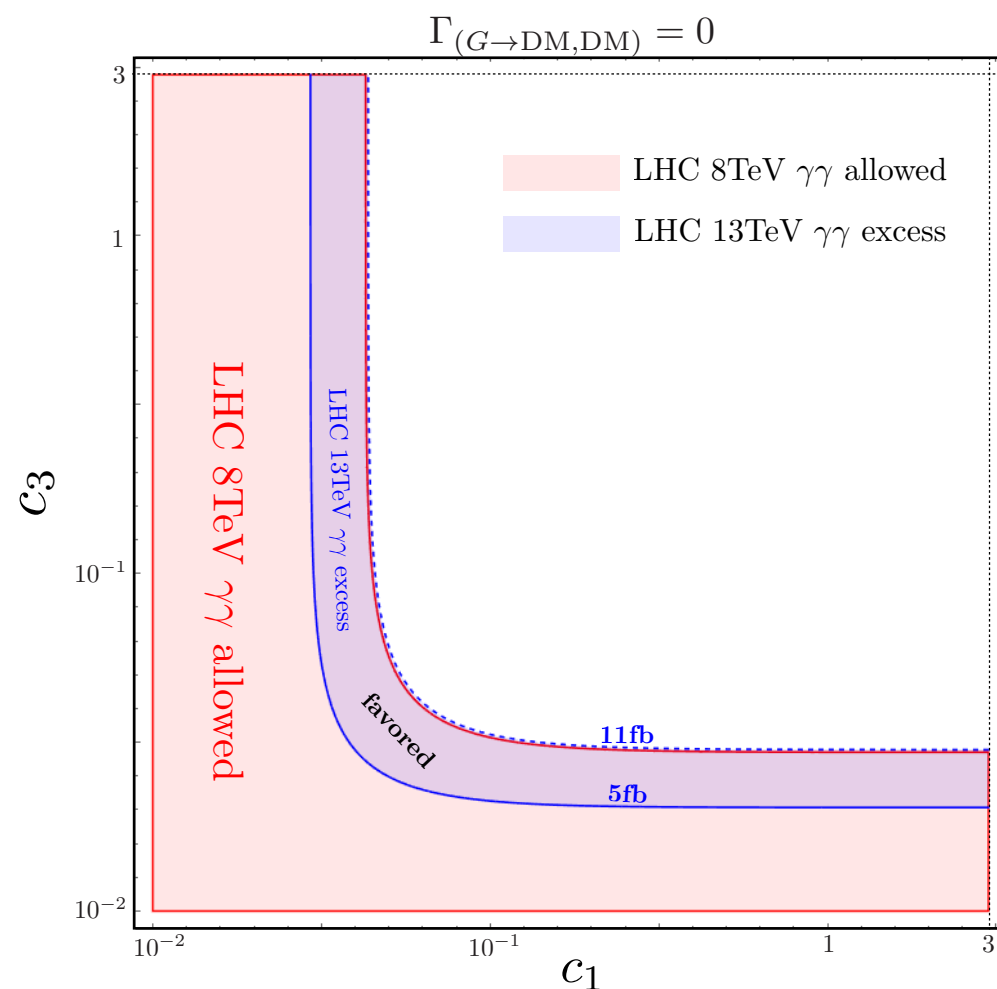
Channels	$\sqrt{s}=8$ TeV	$\sqrt{s}=13$ TeV
$WW(l\nu jj)$	$\lesssim 68$ fb [10]	$\lesssim 259$ fb [11]
$ZZ(ll jj)$	$\lesssim 37$ fb [12]	$\lesssim 151$ fb [13]
$\gamma\gamma$	$\lesssim 2.4$ fb [14]	$\lesssim 11$ fb
dijet	14 pb^1 [15]	-
Monojet	$\lesssim 270$ fb [16]	-



Diphoton = KK graviton

- Diphoton production cross section constrains KK graviton couplings to

$$c_{gg} \times c_{\gamma\gamma} = 0.16 \left(\frac{\sigma_{pp \rightarrow \gamma\gamma}}{8 \text{ fb}} \right)^{1/2} \left(\frac{\Lambda}{3 \text{ TeV}} \right)^2 \left(\frac{45 \text{ GeV}}{\Gamma_G} \right)^{1/2}.$$



No invisible decay

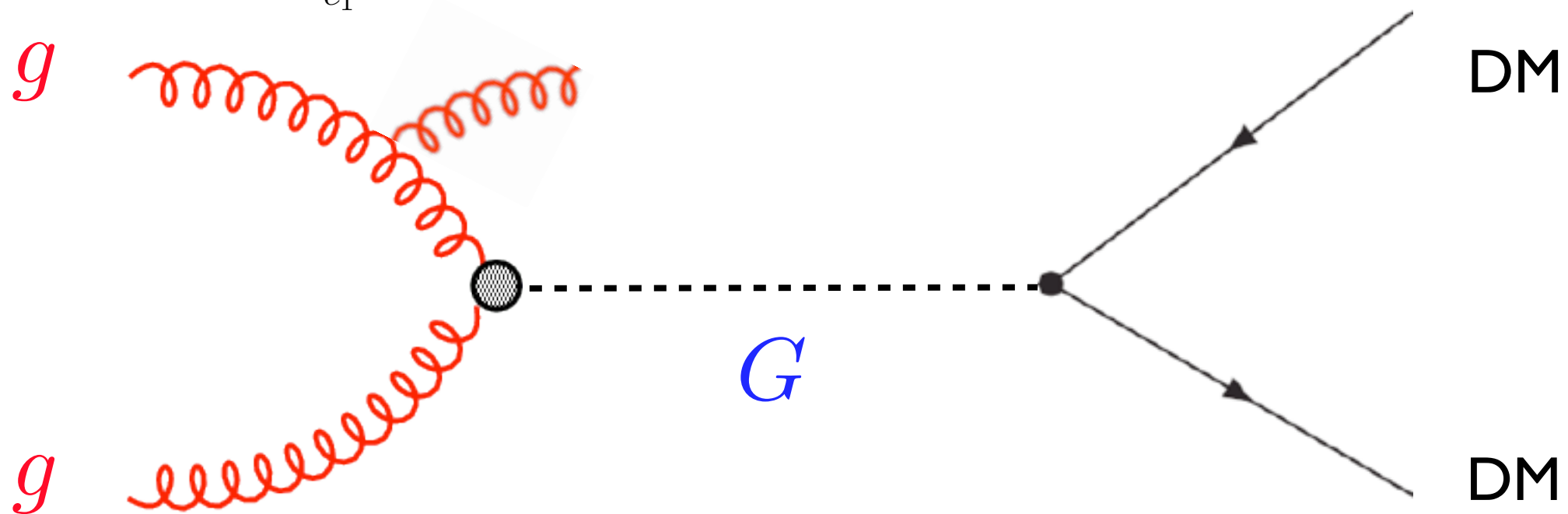
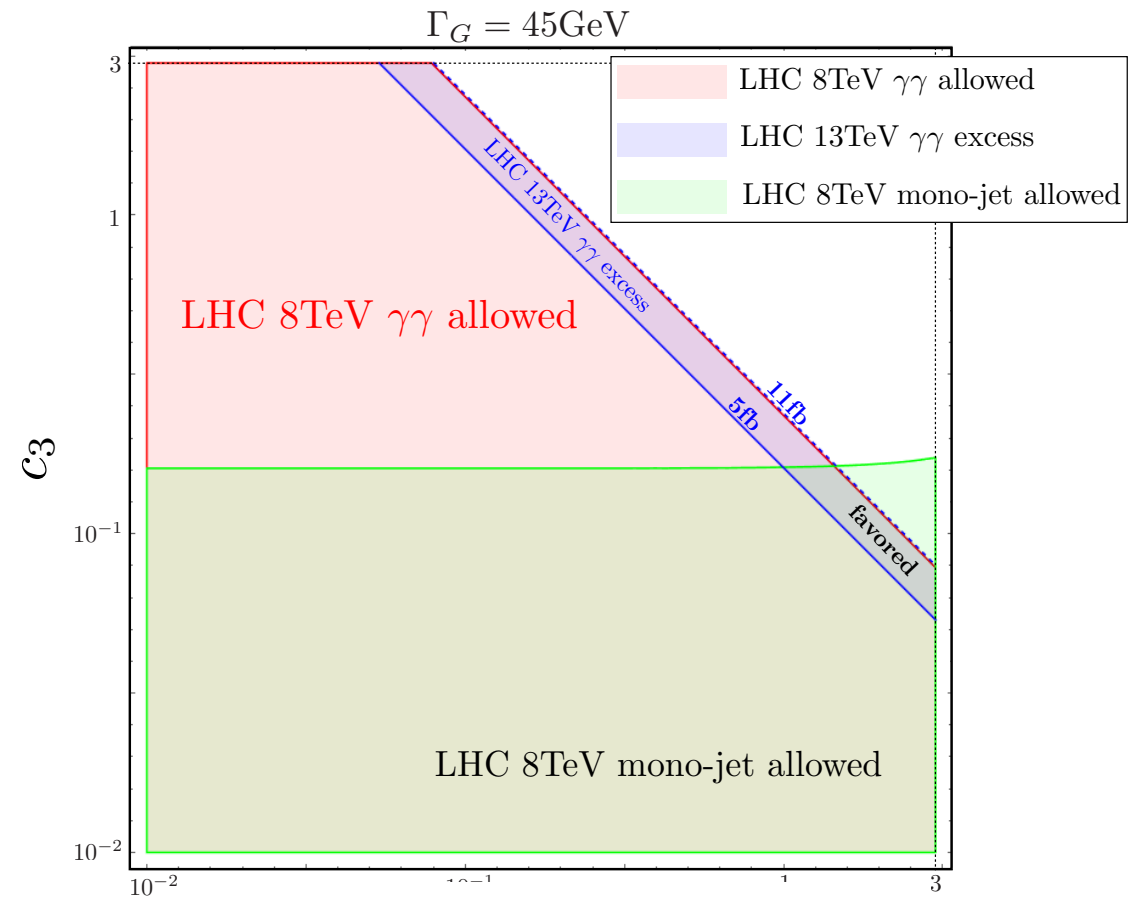
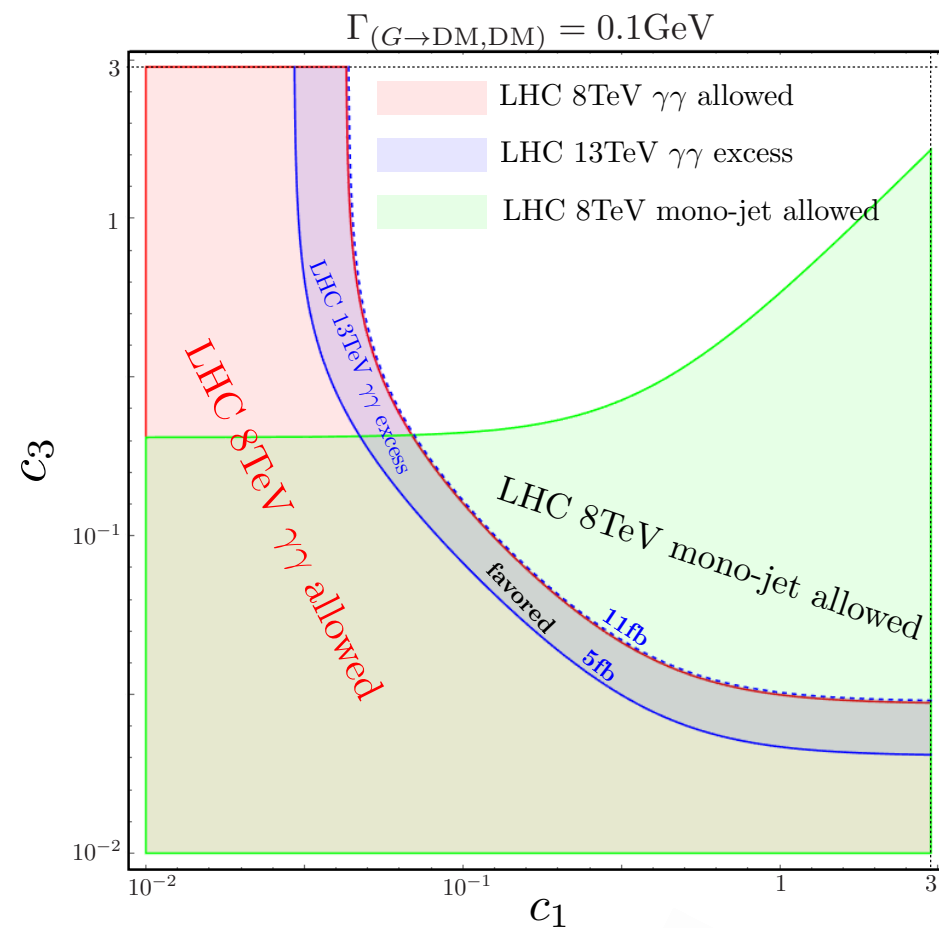
$$\Lambda = 3 \text{ TeV}, m_G = 750 \text{ GeV}$$

Dijet bound:

$$|c_3| \lesssim 1.5$$

Invisible decay & mono-jet

Allow large coupling to DM: Large invisible decay

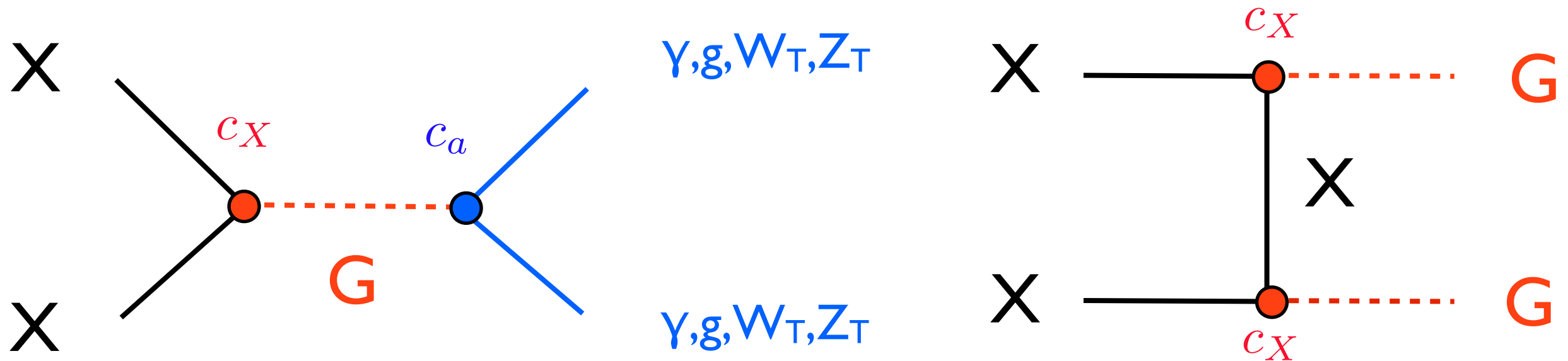


KK graviton as DM mediator

DM annihilation

[HML, M.Park, V. Sanz, 2013, 2014]

- DM annihilates via KK graviton exchanges, depending on the spin of dark matter.

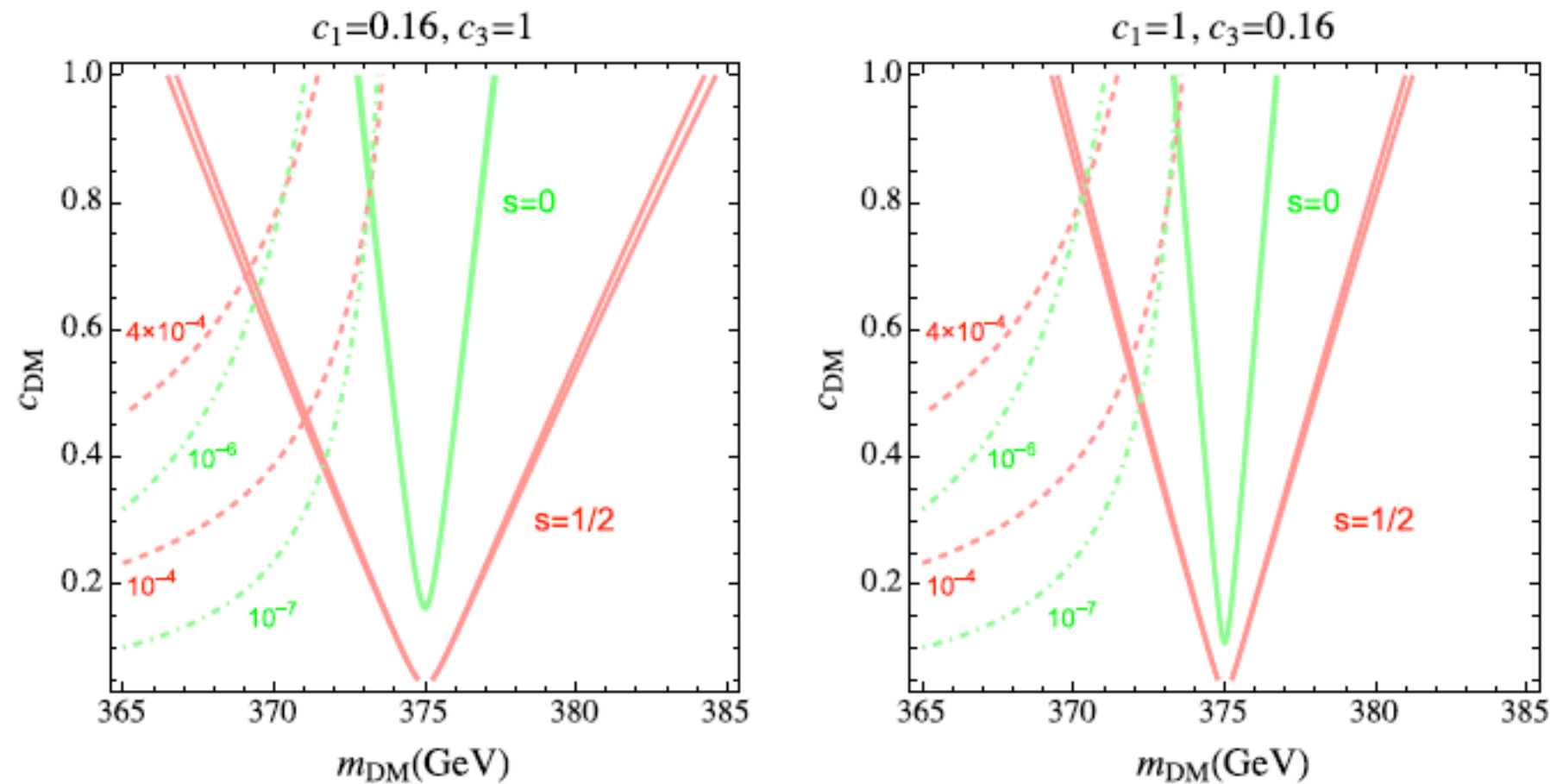


$$(\sigma v)_s \sim v^n \frac{c_X^2 c_a^2 m_X^2}{\Lambda^4} \left(\frac{m_X}{m_G} \right)^4$$

$$(\sigma v)_t \sim \frac{c_X^4 m_X^2}{\Lambda^4} \left(\frac{m_X}{m_G} \right)^8$$

channels	DM mass	X (s=0)	X (s=1/2)	X (s=1)
s-channel	$m_{\text{DM}} < m_W$	d-wave	p-wave	s-wave
s-channel	$m_{\text{DM}} > m_W$	s-wave	p-wave	s-wave
t-channel	$m_{\text{DM}} > m_G$	s-wave	s-wave	s-wave

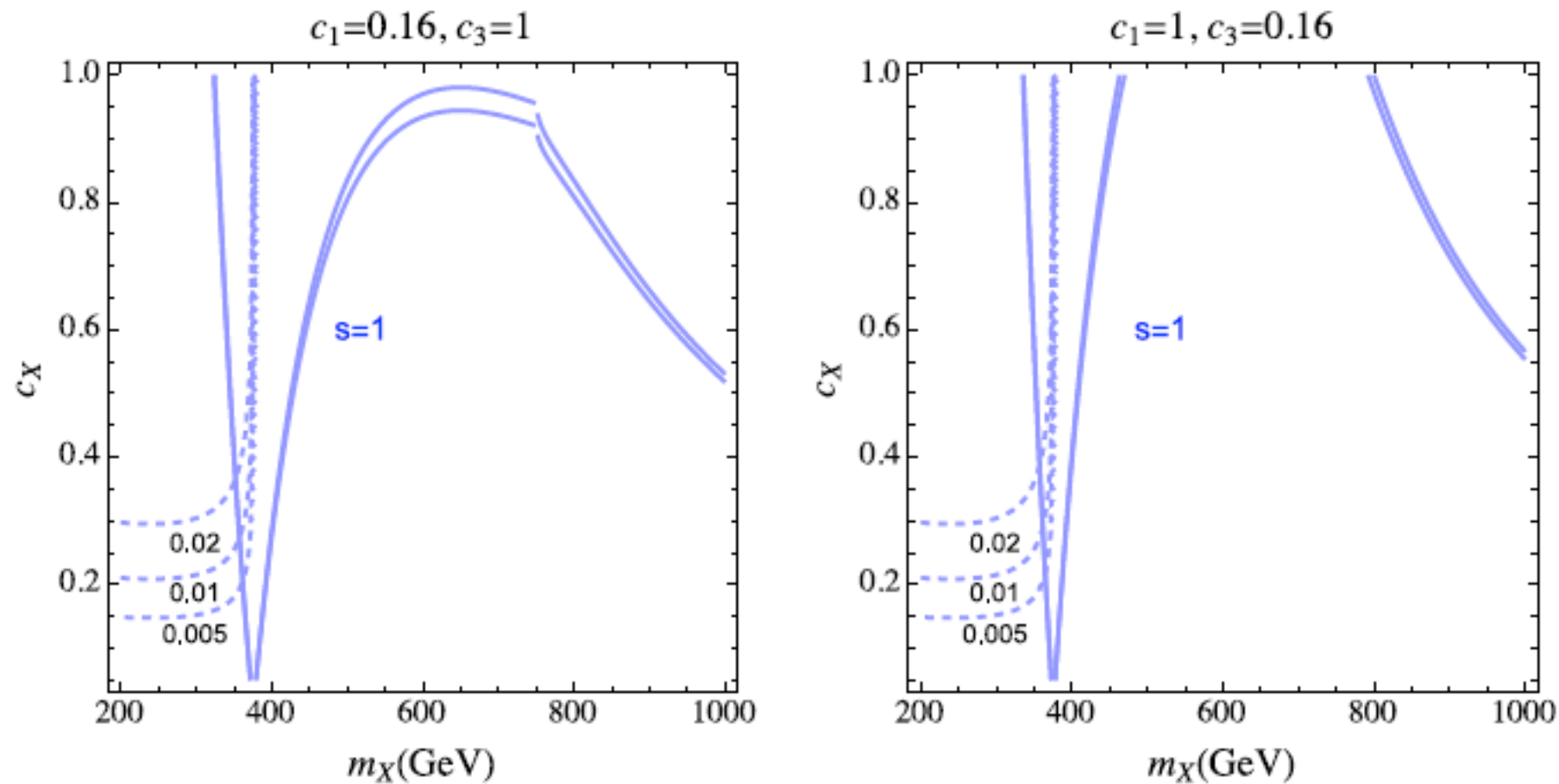
Scalar and fermion DM



$$\Lambda = 3 \text{ TeV and } m_G = 750 \text{ GeV}$$

- Correct relic density is obtained near resonance, due to velocity-suppressed annihilation.
- Invisible decay rates (given in unit of GeV) are small in the region of correct relic density.

Vector DM



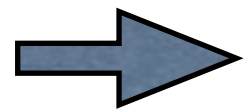
$$\Lambda = 3 \text{ TeV and } m_G = 750 \text{ GeV}$$

- Correct relic density can be obtained even away from resonance, due to s-wave annihilation.
- Invisible decay rate is larger than the cases with other spins, but is still small.

Direct detection

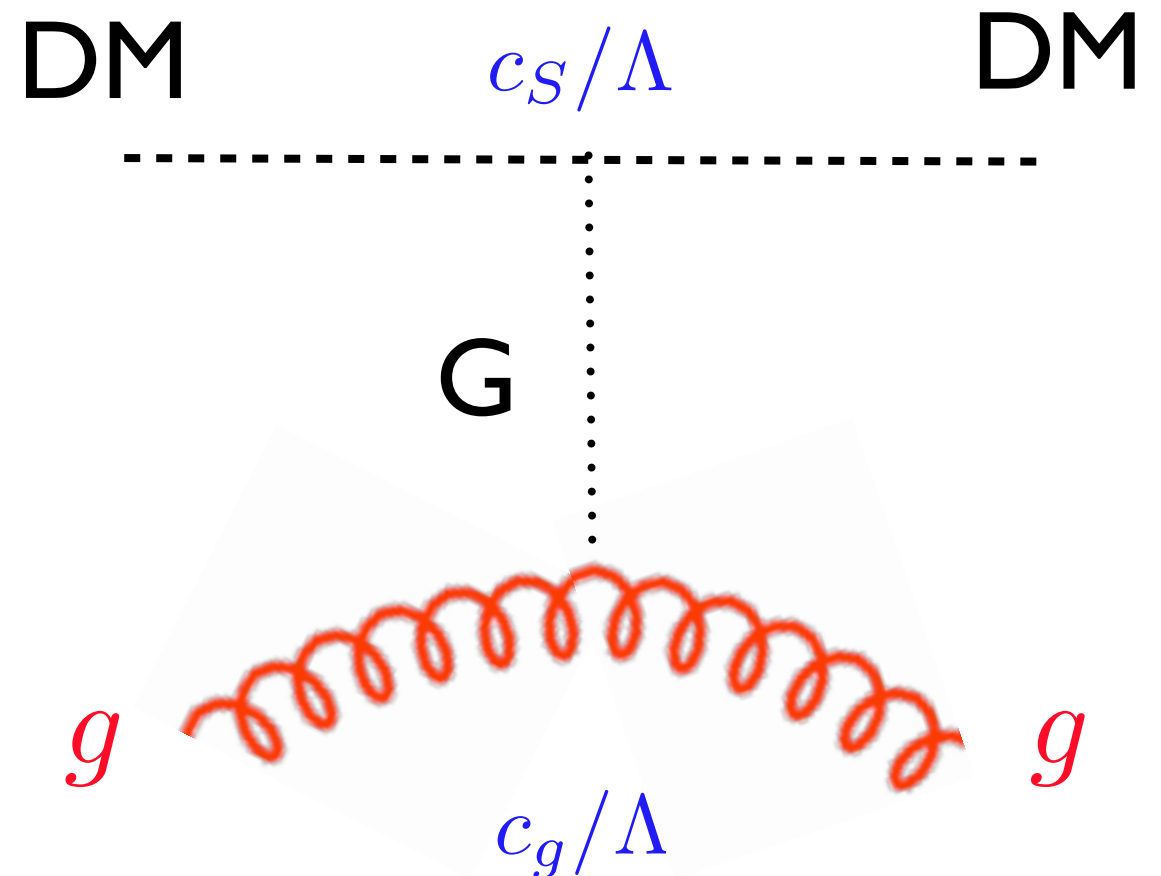
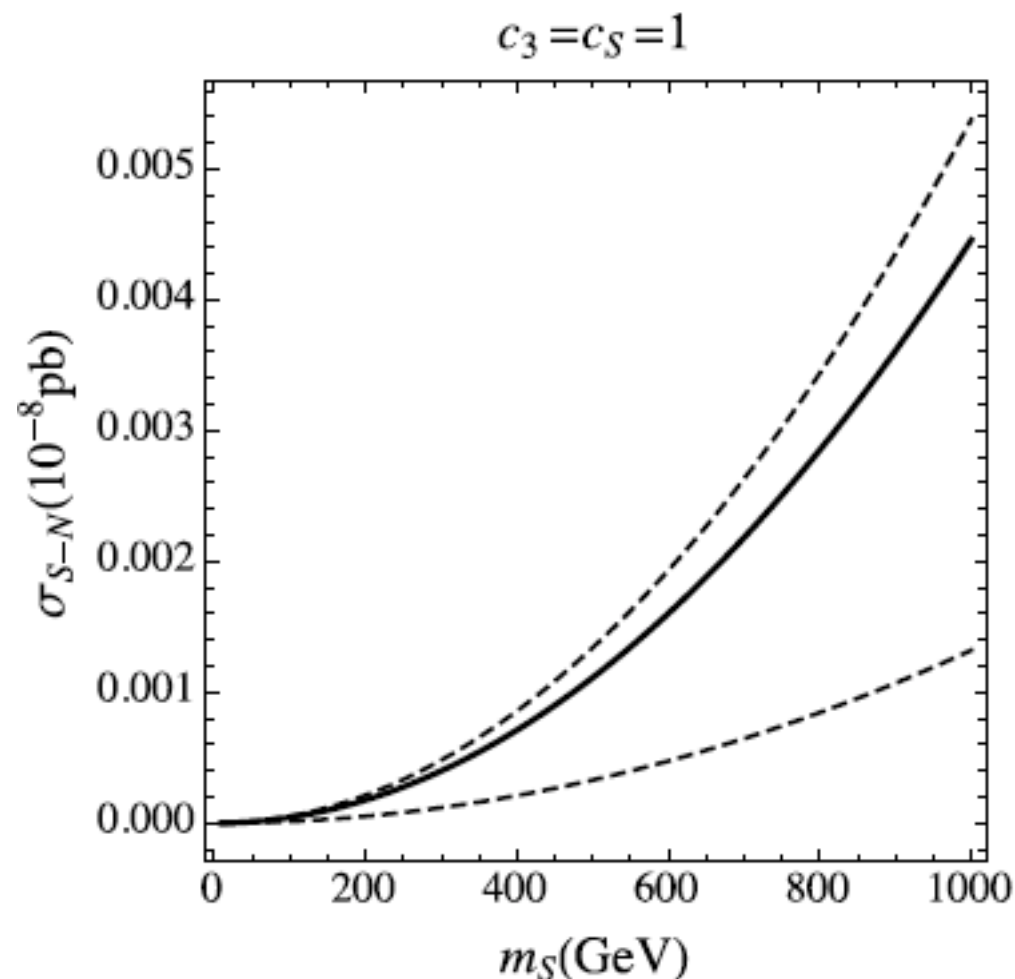
- **Gluon coupling** is unconstrained by direct detection.

$$\mathcal{L}_{S-N} = \xi_g S^2 G_{\mu\nu} G^{\mu\nu}, \quad \xi_g = \frac{c_S c_g}{6\Lambda^2} \frac{m_S^2}{m_G^2},$$



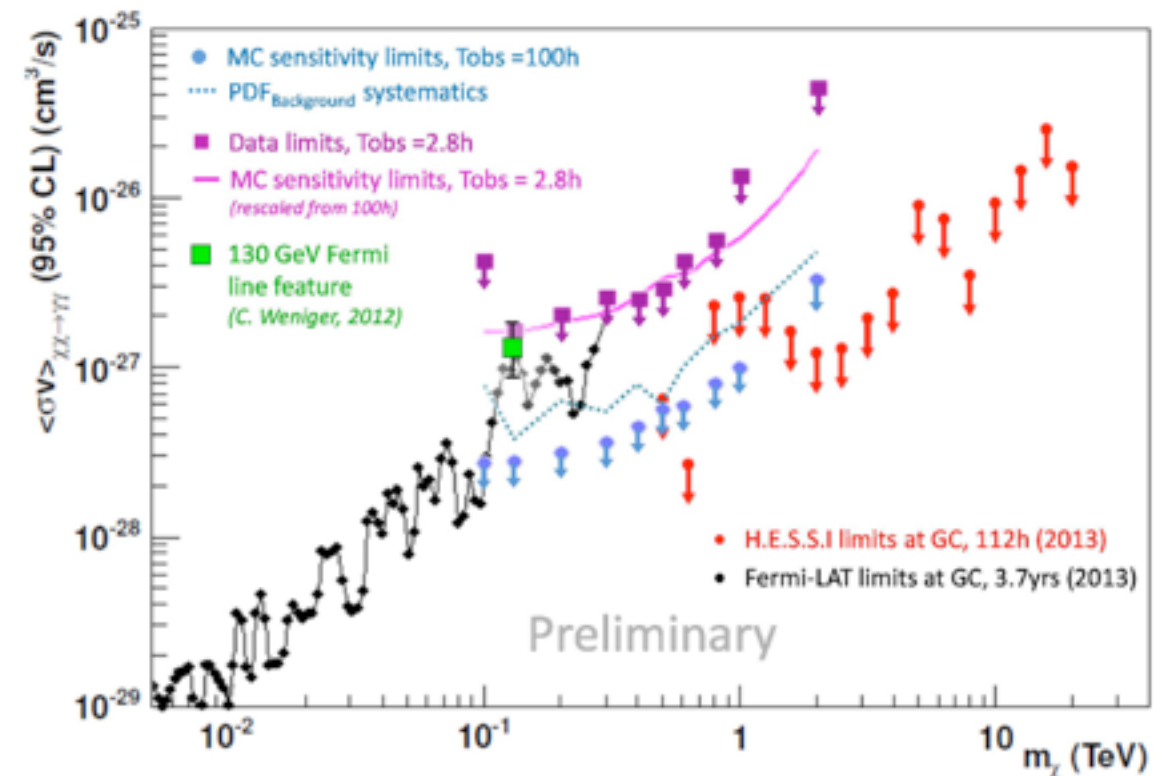
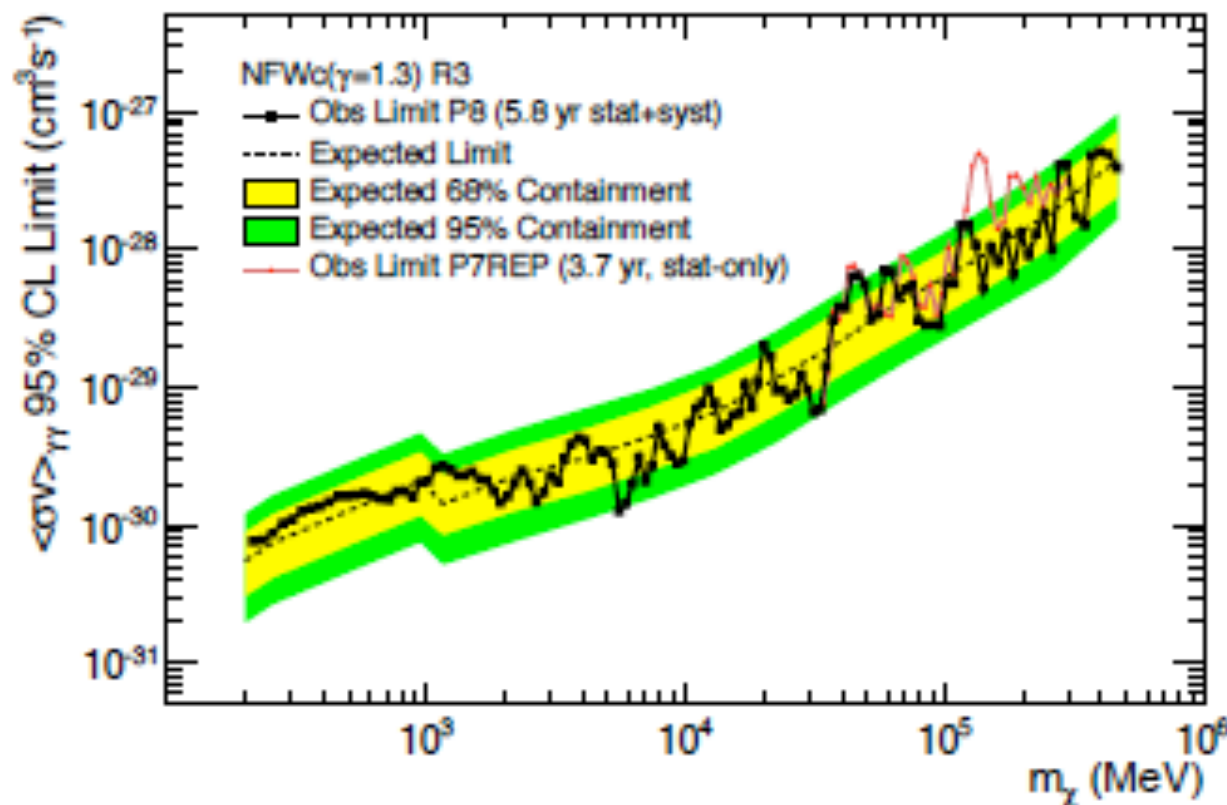
$$\sigma_{S-N} = \frac{\mu^2}{\pi m_S^2} \left(\frac{8\pi}{9\alpha_s} \right)^2 m_N^2 \xi_g^2 f_{TG}^2, \quad f_{TG} = \frac{1}{m_N} \langle N | \frac{-9\alpha_s}{8\pi} G_{\mu\nu} G^{\mu\nu} | N \rangle$$

$$= 0.472 - 0.952(\text{MILC}).$$



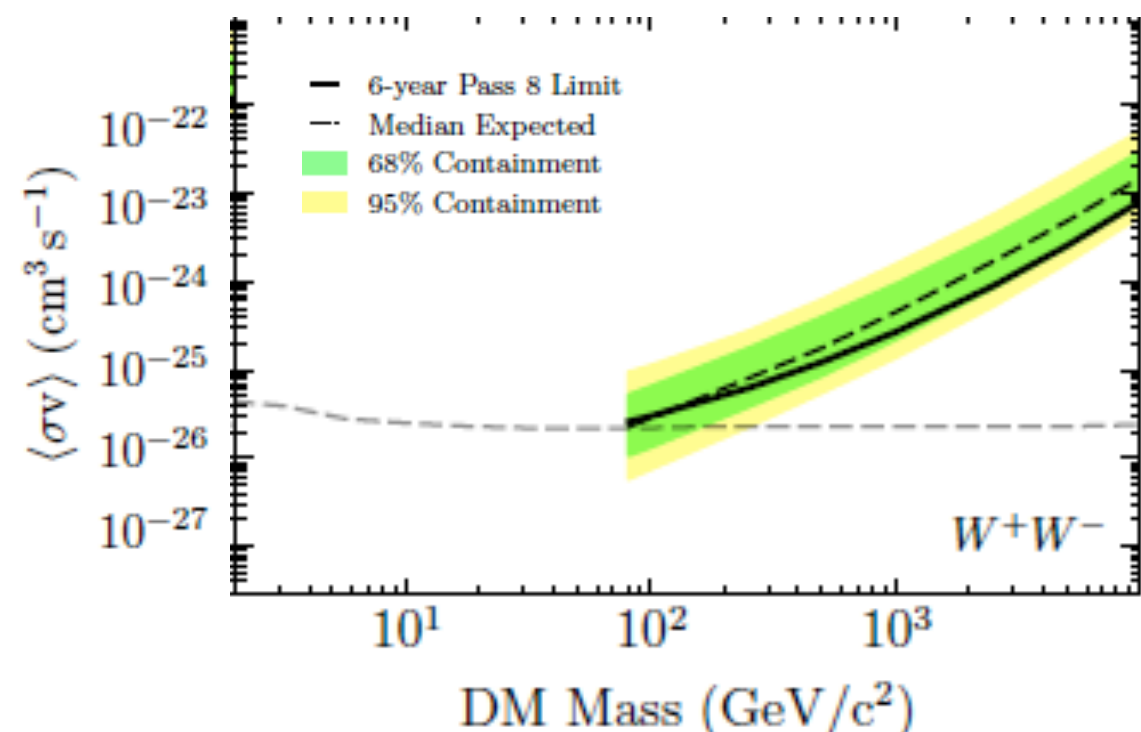
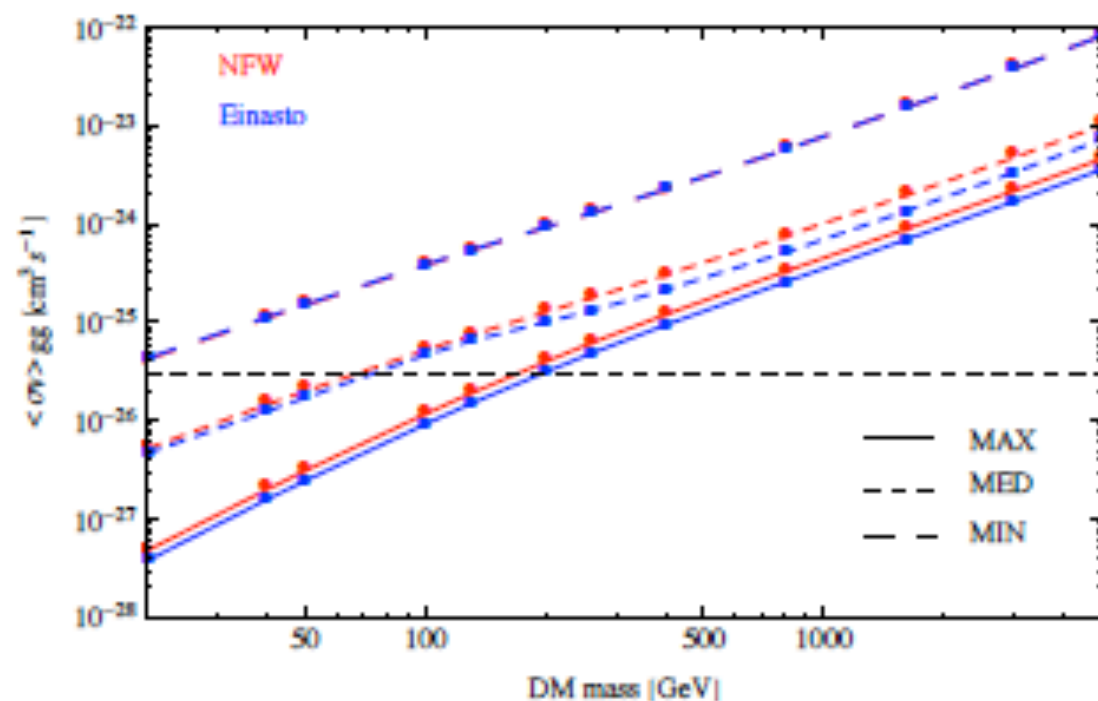
Indirect detection - lines

- Spectral gamma-ray line (Fermi-LAT, HESS, CTA, etc):
 $\gamma\gamma, GG \rightarrow \gamma\gamma\gamma\gamma$ channels (vector DM)
- BR of DM annihilation into a photon pair less than 1% of thermal cross section for DM mass $\sim$ a few 100GeV.



Indirect detection - continuum

- Continuum gamma-ray (Fermi-LAT dwarf galaxies):
WW, ZZ channels (scalar & vector DM)
- Anti-proton (PAMELA, AMS-02): gg channel (vector DM)
- Bounds from Anti-proton & Fermi dwarf galaxies constrain thermal cross section for gg & WW.



[Chu, Hambye, Scarna, Tytgat, 2012]

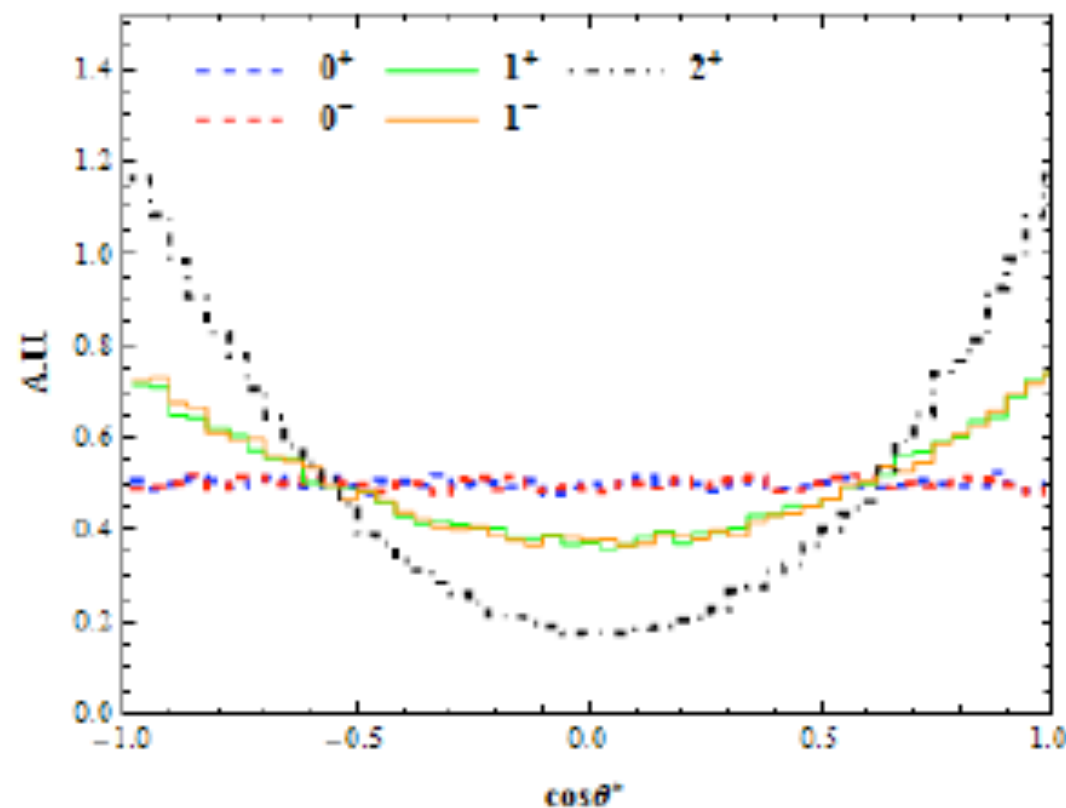
Spin determination

Spin/CP of resonances

- Angular correlations between decay products depend on spin/CP of the resonance particle.

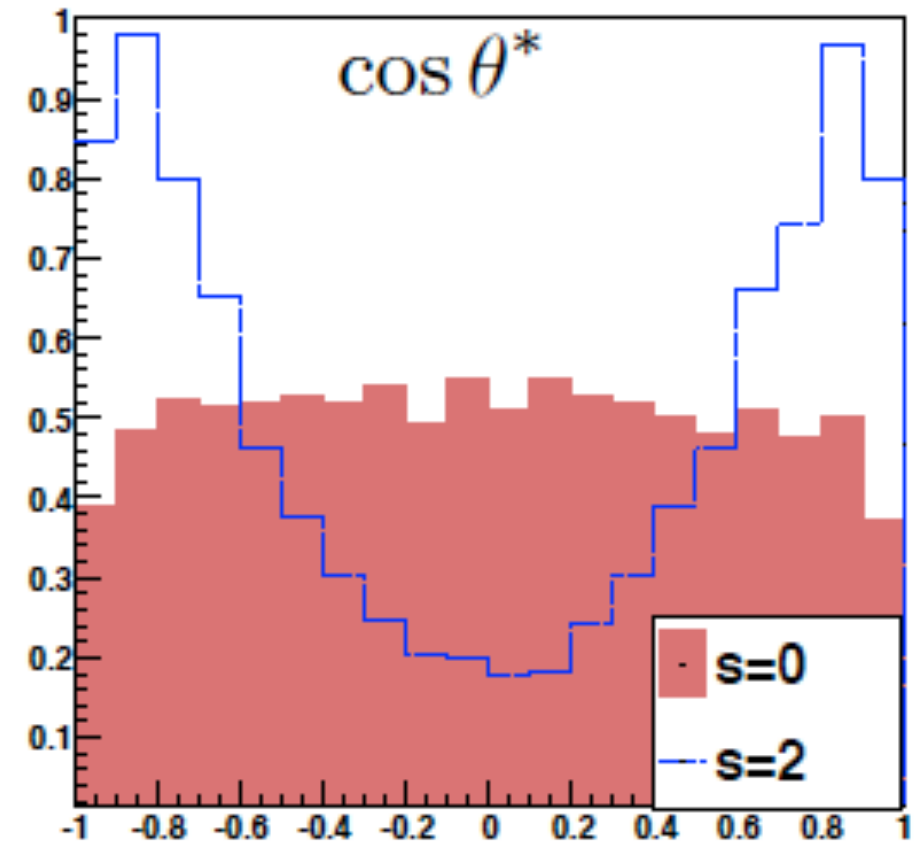
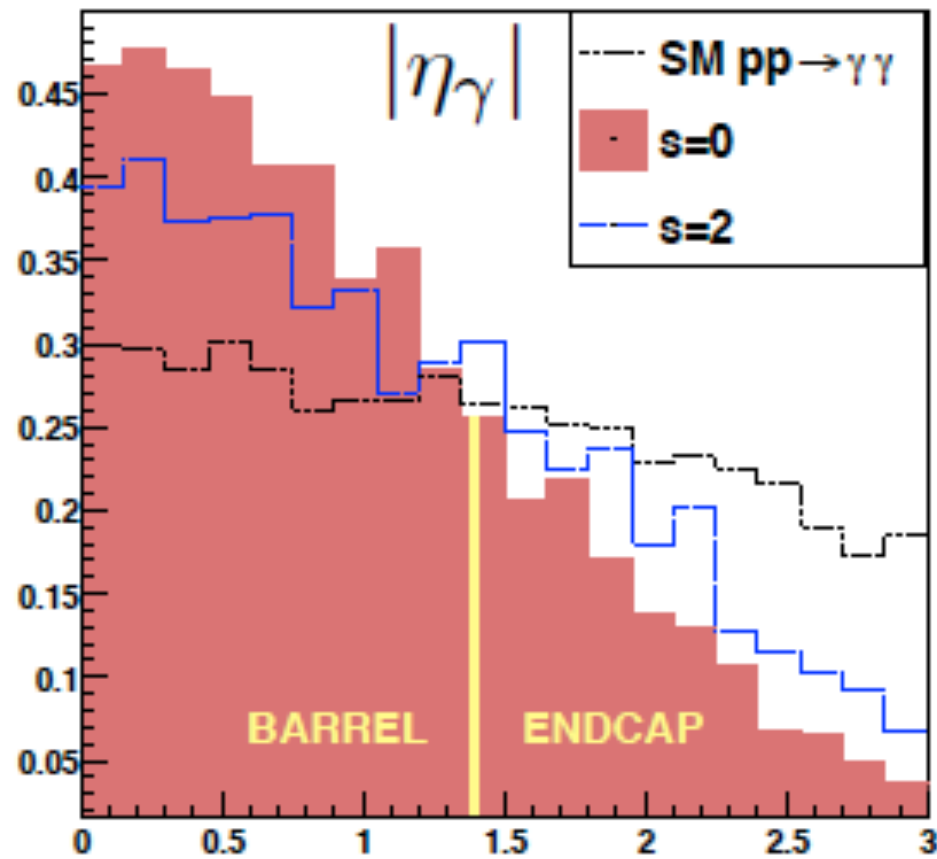
e.g. diboson resonance [D.Kim et al, 1507.06312]

$$\frac{d\sigma}{d\cos\theta^*} \sim \begin{cases} 1, & \text{for } gg \rightarrow 0^+, 0^- \rightarrow W^+W^- \\ 1 + \cos^2\theta^*, & \text{for } q\bar{q} \rightarrow 1^+, 1^- \rightarrow W^+W^- \\ 1 + 6\cos^2\theta^* + \cos^4\theta^*, & \text{for } gg \rightarrow 2^+ \rightarrow W^+W^- \end{cases}$$



θ^* : the angle of decay product wrt the beam direction in C.M. of decaying particle.

Diphoton distribution



“rapidity” $m_{\gamma\gamma} = 750 \pm 50 \text{ GeV}$

- More events of spin-2 in end-cap region than spin-0 case. (CMS 2.6σ excess comes from EBEE.)
- Angular distribution distinguishes spin: spin-0 $\sim$ symmetric, spin-2 $\sim$ boosted decay.

Conclusions

- We considered the interpretation of the diphoton resonance in EFT.
- We presented a concrete example for spin-2 resonance where KK graviton plays a role of dark matter mediator.
- Gluon coupling can be constrained by mono-jet and dijet searches.
- Invisible decay rate is small in the region of correct DM relic density.
- Gluon or photon coupling can be constrained by indirect detection for scalar or vector DM.