GBAR and Antimatter Gravity measurement

Center for Underground Physics Institute of Basic Science

BongHo Kim

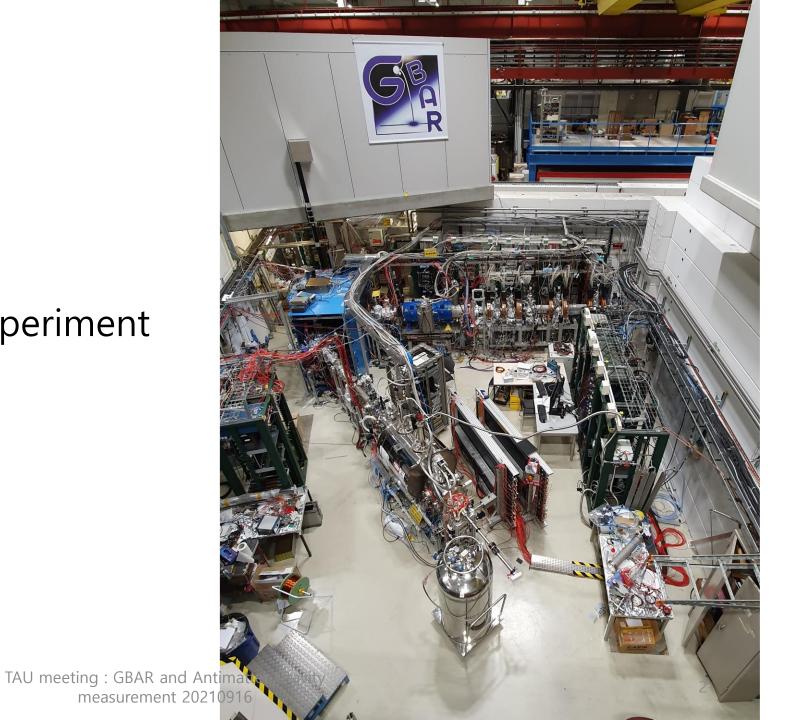




content

motivation

- antimatter gravity experiment
- GBAR experiment
- Conclusion





Antimatter



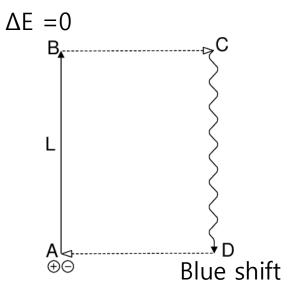
- Just after the discovery of electron by J. J. Thomson, there was questions why there's no negative gold although there's negative electricity (1898)..
- After prediction by Dirac equation in 1927, antiparticles as a pair of particle had detected (positron 1932, antiproton 1955...)
- The question of anti-gravity raised at 1950s and have continuing interests about antimatter's gravity...
- For a period of 100 years, the tested properties and interactions of antimatter are agreed with the standard model which based on CPT symmetry.
- But the universe near neighborhood is totally matter dominant without any trace or remnant of antimatter and shows no proof of symmetry..
- : We are not worried about antimatter comet for example..
- With hardness of production and cooling of antimatter, there can be unknown properties of antiparticles.
- This mysterious status attracts much attention from the public. (with huge applications and fictions)





Antimatter gravity

- Arguments about antigravity were started about 1960
- P. Morrison (1958): violation of energy conservation
- ← not only tensor but with scalar and vector (component of repulsive gravity) may required..
- other arguments by L.I. Schiff(1959), M.L. Good(1961) (M. M. Nieto and T. Goldman, Physics Reports, Volume 205, Issue 5, p. 221-281)
- Villata "General relativity and CPT are compatible only with repulsive gravity" м. villata, EPL 94, 20001 (2011).



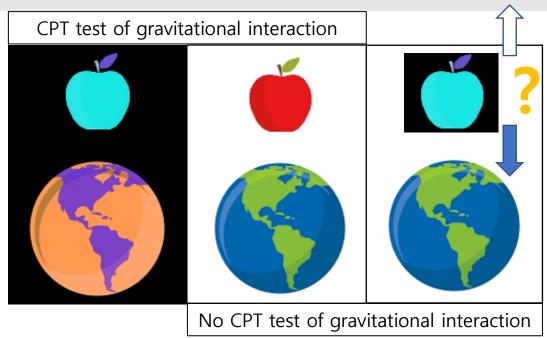
• Above arguments are mainly about antigravity...

Motivation: Cosmological mysteries

- ★ Our understanding is very incomplete and limited...
- Matter and antimatter asymmetry
- Matter domain in observable Hubble volume : $n_B\gg n_{ar{B}}$
- The standard model predicts equal amount of matter and antimatter in the universe after Bigbang..
- Unknown nature for most energy contents of universe
- Dark energy: 69.4%, Dark matter: 26.1%, Ordinary matter: 4.5%...
- Hard to link dark energy to particle physics, no measurement of dark matter..
- Some attempts to make a cosmological model to explain the asymmetry and to keep them apart by gravitational repulsion..

A. Benoit-Levy and G.Chardin, A&A 537, A78(2012), M.Villata, Astrophys. Space Sci. 339, 7 (2012), D. S. Hajdukovic, Int. J. Theor. Phys. 49, 1023 (2010)) can give new sight related to Dark energy and Dark matter..

Test of Weak Equivalence Principle



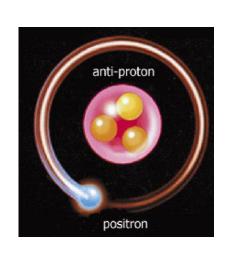
- $m_I = m_G (F = m_I a = -G m_G m'_G / r^2)$,
- $m_I = \overline{m_I} (by CPT)$
- $m_G = m_I = \overline{m_I} = ? \overline{m_G}$

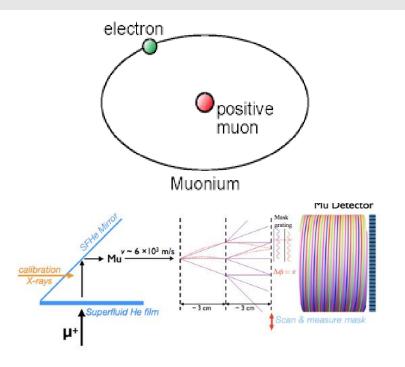
• Weak Equivalence Principle :

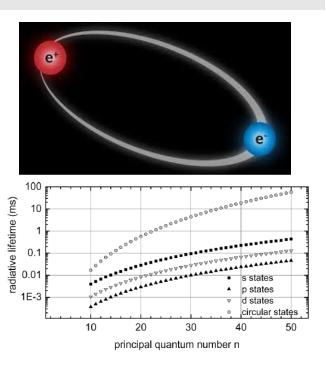
"The trajectory of a <u>point</u> mass in a gravitational field depends only on its initial position and velocity, and is independent of its composition and structure"

- : The gravitational acceleration is not distinguishable with other acceleration locally $<\!m_g\!=\!m_i\!>\!equivalence$ between the inertial mass and the gravitational mass
- Test between matter and earth has upper limit $\Delta(m_g/m_i)/(m_g/m_i)_{Be/Ti} = (0.3\pm1.8)e^{-13}$ (Eötvös)
- Previous attempts (positron, antiproton..) are failed because of relatively small interaction compared with effects by electric charge (patch effect, gas charges..)

Test of WEP by neutral exotic atom

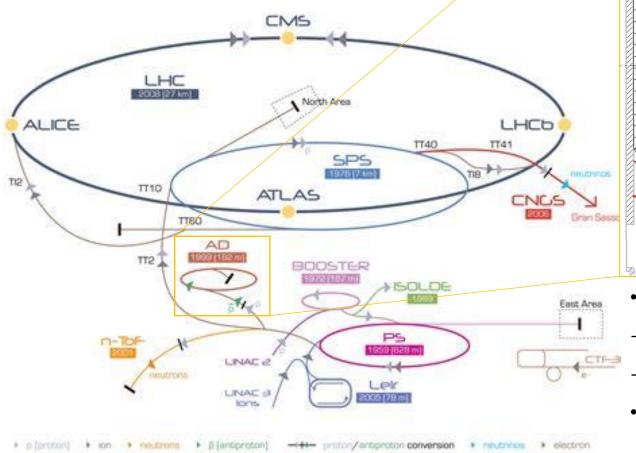




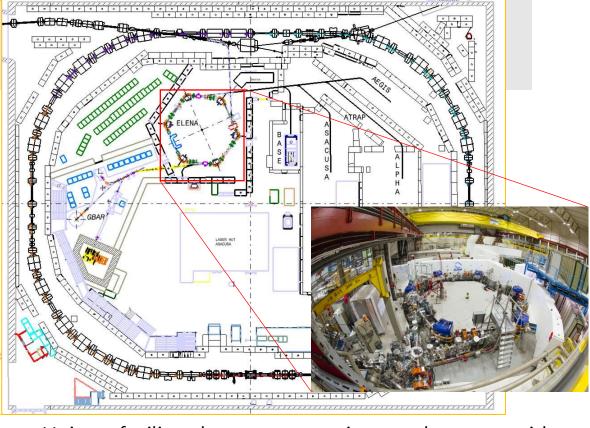


- Neutral exotic atom-based experiment has been prepared : Positronium (e⁺e⁻), Muonium(μ ⁺e⁻) and Antihydrogen(\overline{H}).
- Three constructing experiments using \overline{H} will be compared.
- Conceptual design of gravity measurement by muonium atom at Mage collaboration(muonium beam by superfluid He film to get monochromatic direction (1% precision goal) T. J. Philips, EPJ Web of Conferences 181, 01017 (2018)
- Positronium atom with long-lived Rydberg states (D. B. Cassidy and S. D. Hogan, Int. J. Mod. Phys.: Conf. Ser. 30, 1460259 (2014))

AD & ELENA



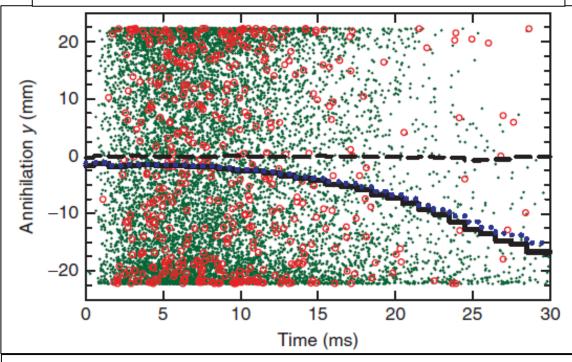
LHC Large Hadron Collider SPS Super Proton Synchrotron PS Proton Synchrotron AD Antiproton Decelerator CTF-3 Dic Test Feedey CNCS Cem Neutrinos to Gran Sasso 15-CLDG histogie Separator Drilling DEvice



- Unique facility: low energy antiproton beam provider
- 26 GeV/c proton + Iridium target → 3.5GeV/c antiproton.
- 5.3MeV (100MeV/c) antiproton beam 2x10⁷#/120s
 - ELENA: 100keV antiproton beam (σ <100ns, 0.5x10⁷#/120s: ½ bunches) (commissioning during last beam time with GBAR experiment)

Previous measurement of antimatter gravity

Nature communications, 4, 1785 (2013)



Red circle: real data (434 annihilation points)

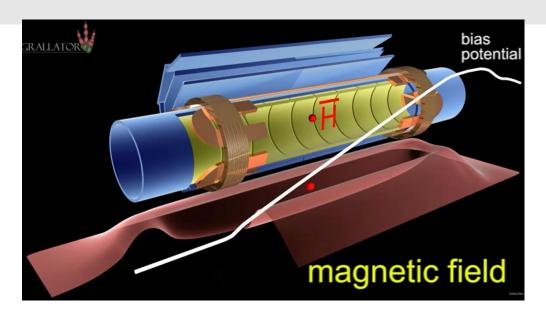
Green point : simulation

Black solid line: averaged simulation (F=100)

Blue dotted line: averaged simulation with detector s

mearing (F=100)

Black dashed line: averaged simulation (F=1)



 Only one anti-hydrogen experiment was done -ALPHA experiment

: $-65 < F (= m_0/m_i) < 110 (95\% significance level)$

• The apparatus is not for gravity measurement : no control for vertical direction.

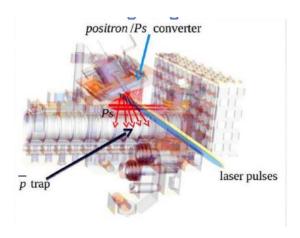
: Too high temperature (0.5K at 0ms), vertical trapping is not optimized

• Need to cool down & make better magnetic field

Gravity test approaches

AEGIS

- Antihydrogen production : cold antiproton + excited o-Ps $\bar{p} + Ps^* \rightarrow \bar{H} + e^-$
- The antihydrogen production cross section is proportional to n_{ps}^{4} (=quantum number).
- The Rydberg Ps has longer lifetime than o-Ps (τ =142ns)



2021-09-16

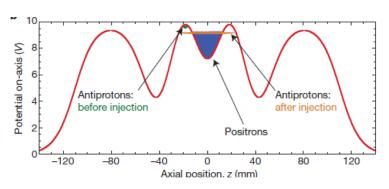
ALPHA-g

 Antihydrogen production : cold antiproton + positron

$$\bar{p} + e^+ \rightarrow \bar{H} + \gamma$$

 $\bar{p} + e^+ + e^+ \rightarrow \bar{H} + e^+$

- Nested-well potential for mixing.
- Antihydrogen production &trapping($\Delta E = -\mu \cdot B$) by octupole field in same trap



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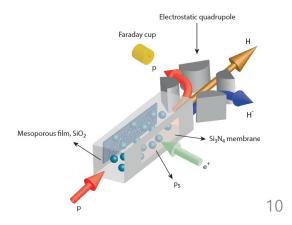
GBAR

 Antihydrogen ion production : antiproton + positronium x 2

$$\bar{p} + \mathrm{Ps}^{(*)} \to \bar{H} + e^-$$

 $\bar{H} + \mathrm{Ps}^{(*)} \to \bar{H}^+ + e^-$

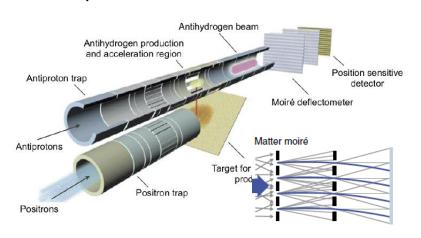
- The antihydrogen ion production cross section is proportional to ρ_{ps}^{-2} .
- Produce dense o-Ps cloud by cavity shaped positronium target.



Gravity test approaches

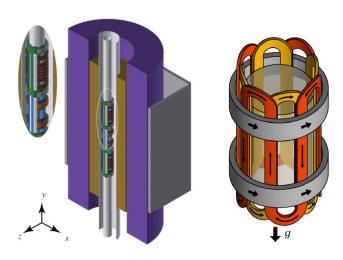
AEGIS

- Pulsed Antihydrogen beam (2021)
- Moire deflectometer tested by \bar{p} . (nature communications 5, 4538 (2014)
- → Pattern will be compared with one from light
- Aim : ~100mK (v~40m/s)
- 1% precision with 1000# \overline{H} .



ALPHA-g

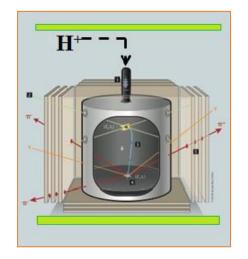
- Antihydrogen trapping (0.5K)
- + Vertical trap (280mm long)
- Aim (1%): sub-50mK
 (v~28m/s) temperature by
 extra cooling by laser &
 precise measurement of
 magnetic field



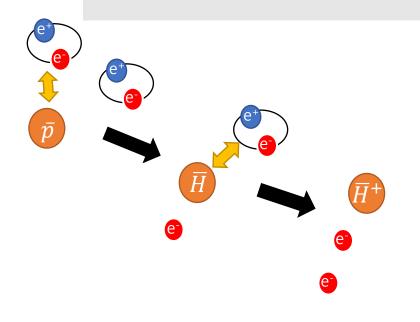
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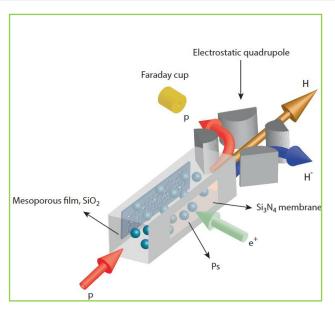
GBAR

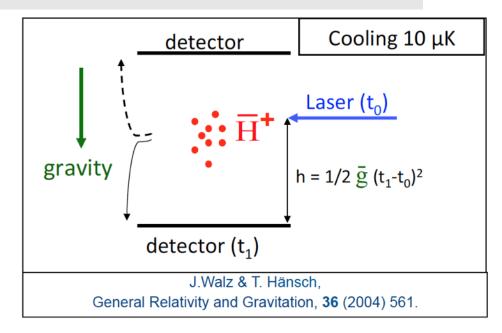
- Antihydrogen ion production
- Trapping and cooling antihydrogen ion.
- Classical Freefall test (z=0.25m)
- Aim: 10uK (v~0.4m/s)
- 1% precision with 1500# \overline{H} .



GBAR experiment approach







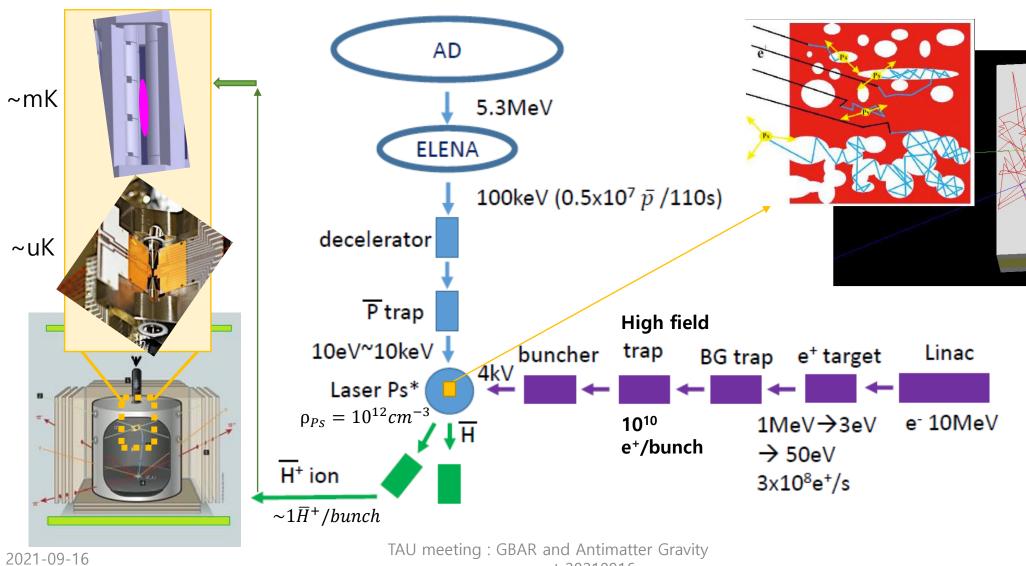
- GBAR (Gravitational Behaviour of Antihydrogen at Rest) experiment aims a direct measurement of the gravitational acceleration of antihydrogen at terrestrial gravitational field by a classical freefall test.
- Antihydrogen ion production is required to get ultra-cold antihydrogen.
- Double charge exchange process between antiproton beam and dense positronium cloud

$$\bar{p} + Ps \rightarrow \bar{H} + e^-$$

 $\bar{H} + Ps \rightarrow \bar{H}^+ + e^-$

- Cooling antihydrogen ion down to 10uK range (ultra-cold) with Be+ to get extremely slow velocity.
- After dropping (by photo-detachment laser) one positron, let the ultra-cold antihydrogen freefall.

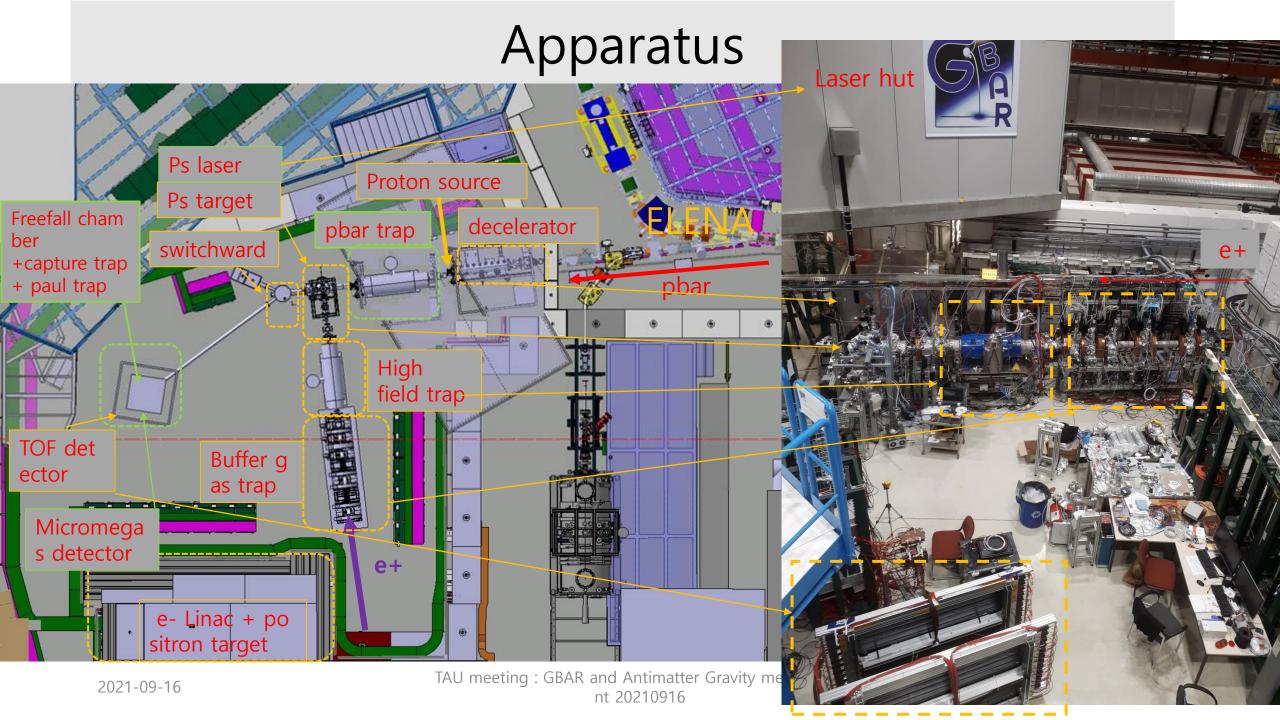
Experiment Scheme



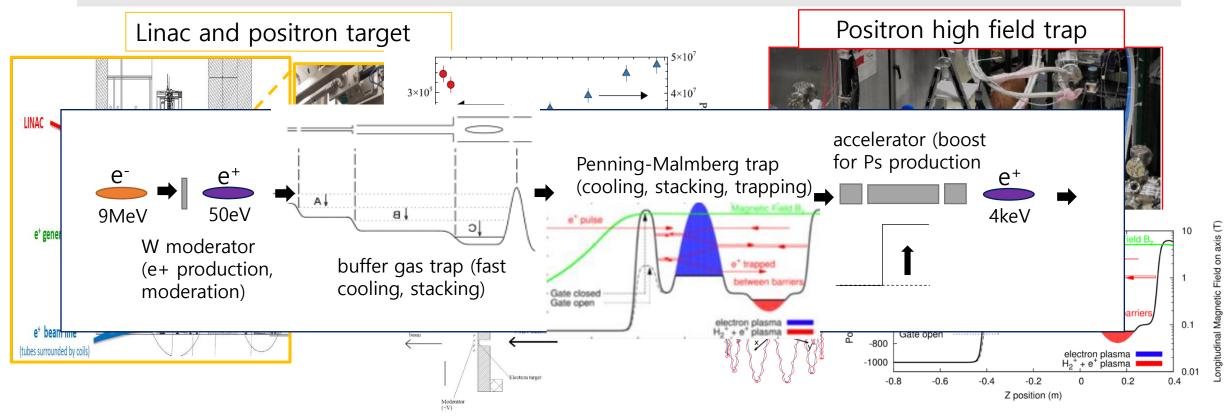
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Green: e+

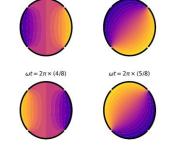
White: Target

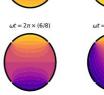


Positron beam for Ps production



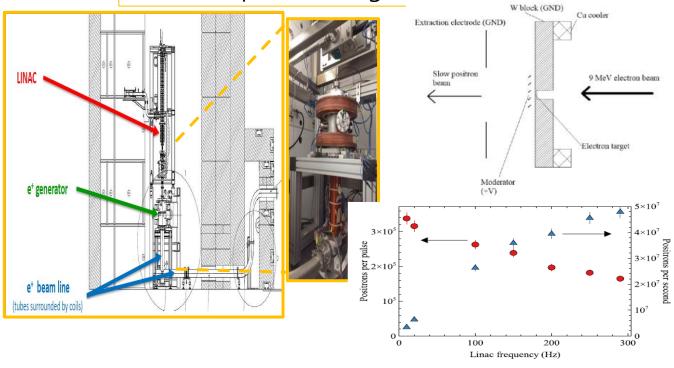
- Positron beam : (Near monoenergetic low energy by W moderator) positron beam generated from bremsstrahlung-induce pair production by 300Hz 9MeV e- linac with a goal to 3 x 10⁸e⁺/s ←To achieve above radioactive sources (<10⁷ e⁺/s)
- Buffer-gas trap for positron accumulation with small energy spread
- Positron accumulation by high field trap : goal to 1x10¹⁰ e⁺ (110s) with electron cooling
- Positron acceleration &bunching by electrostatic lenses with resistor chain



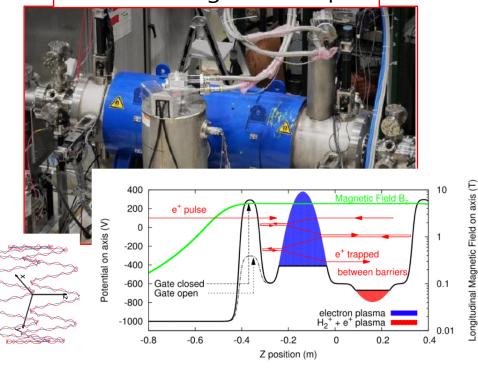


Positron beam for Ps production

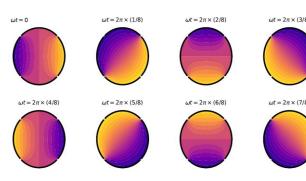




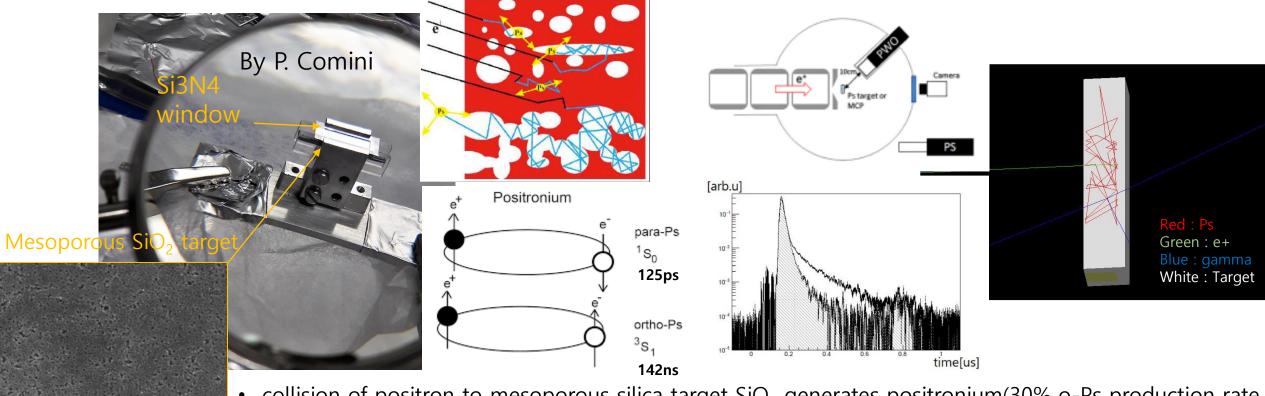
Positron high field trap



- Positron beam: (Near monoenergetic low energy by W moderator) positron beam generated from bremsstrahlung-induce pair production by 300Hz 9MeV e- linac with a goal to 3 x 10⁸e⁺/s ← To achieve above radioactive sources (<10⁷ e⁺/s) M. Charlton et al., Nuclear Inst. and Methods in Physics Research, A 985 (2021) 164657 A (2020)
- Buffer-gas trap for positron accumulation with small energy spread & fast cooling
- Positron accumulation by high field trap: goal to 1x10¹⁰ e⁺ (110s) with electron cooling & compression by rotating wall (detail is in thesis S. Niang, Paris Saclay-University, France (2020))
- Positron acceleration &bunching by electrostatic lenses with resistor chain TAU meeting: GBAR and Antimatter Gravity measurement 20210916

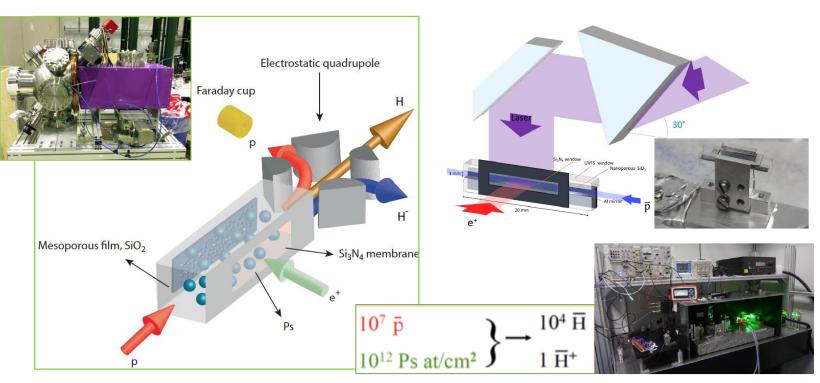


o-Ps production for antihydrogen production

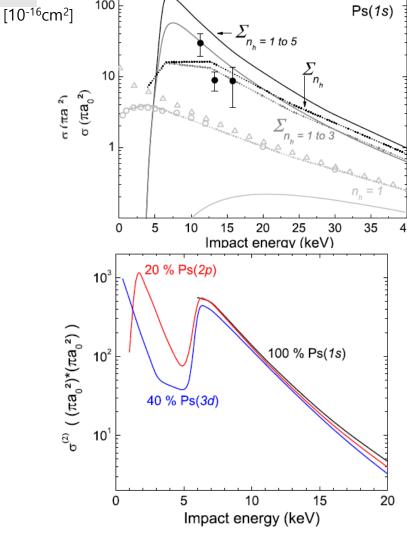


- collision of positron to mesoporous silica target SiO₂ generates positronium(30% o-Ps production rate, 48±5meV energy (PRA81, 052703))
- Ortho-positronium with 142 ns lifetime can be used for interaction with p-bar beam
- Positronium target with cavity shape (1x1x10mm³) is prepared to make high density (10¹²cm⁻³ density required) positronium cloud.
- SSPALS based on PbWO₄ crystal detector for pulse beam detection measurement 20210916

Antihydrogen production

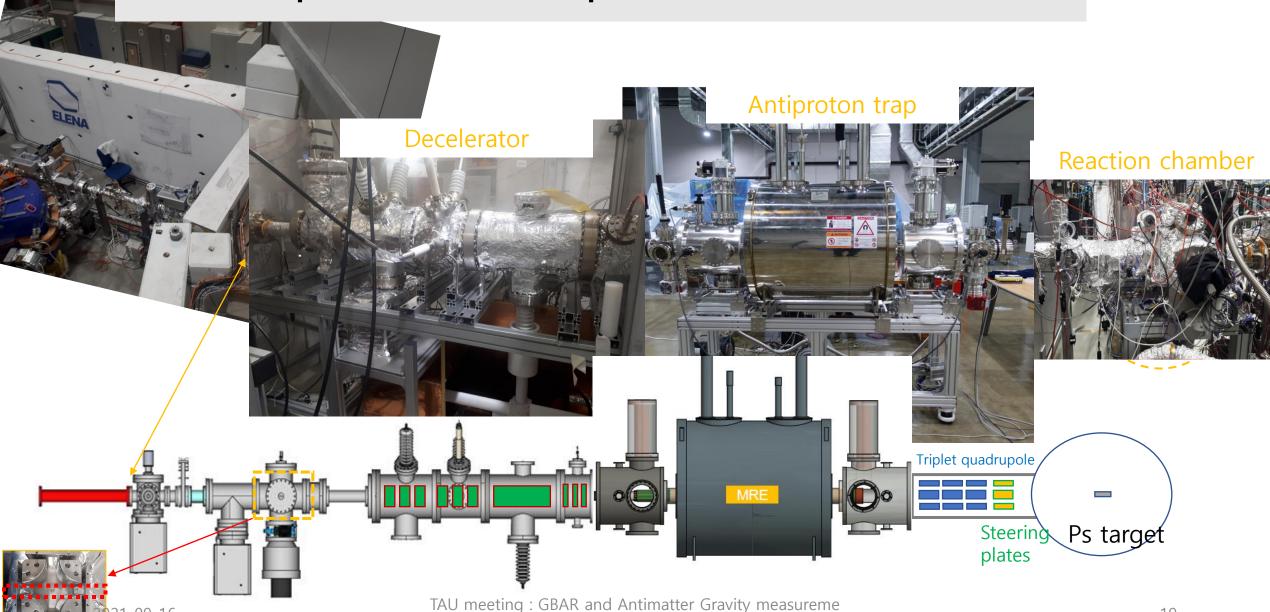


- \bar{H}^+ production can be maximized with Ps(2p, 3d or 1s)
- 10¹²Ps/cm³ inside of target cavity
- About 1# \overline{H} + expected for every 2 minute (1 cycle of AD)
- Study to optimize the production is ongoing for double charge exchange.
- The expected \overline{H} number will be changed by new paper of cross-section : New J. Phys. 23 (2021) 029501
- A main goal of GBAR in this year is the production of \overline{H} and H.



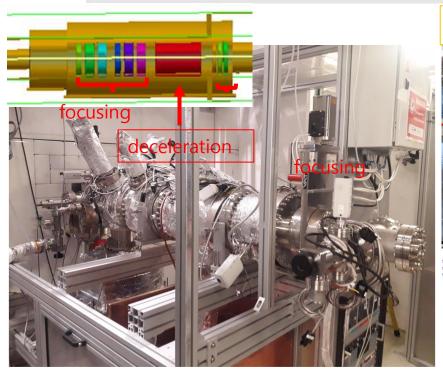
P.Comini et. al, Hyperfine Interactions, 228, 159-165 (2014)

Antiproton and proton beam line

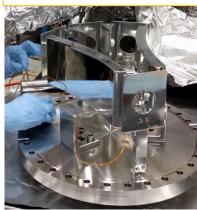


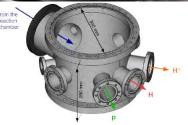
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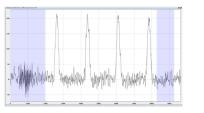
Antiproton and proton beam line

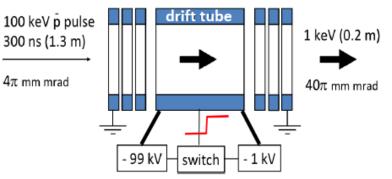


Switchward

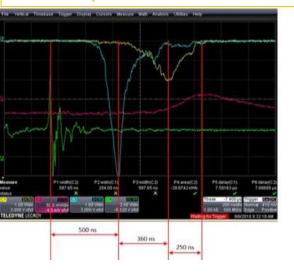


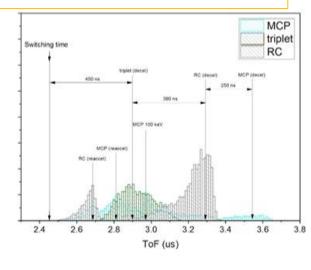






Antiproton beam deceleration test& simulation





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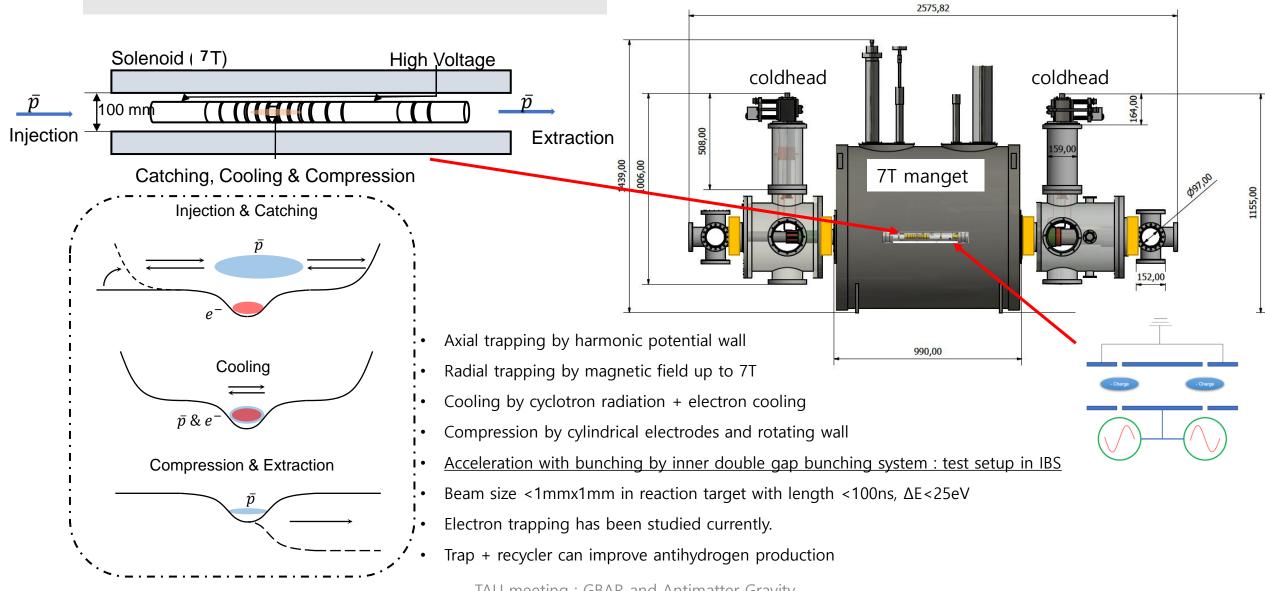
- First beam from ELENA was tested with decelerator in the GBAR line before LS2. A. Husson et al., <u>Nuclear Inst. and Methods in Physics Research</u>, A 1002 (2021) 165245 **A** (2021)
- Deceleration by electrostatic tube with switching circuit from 100keV (300ns) to 1~10keV with high efficiency: Has been tested now.
- Switchward to separate antihydrogen ion from antihydrogen and antiproton

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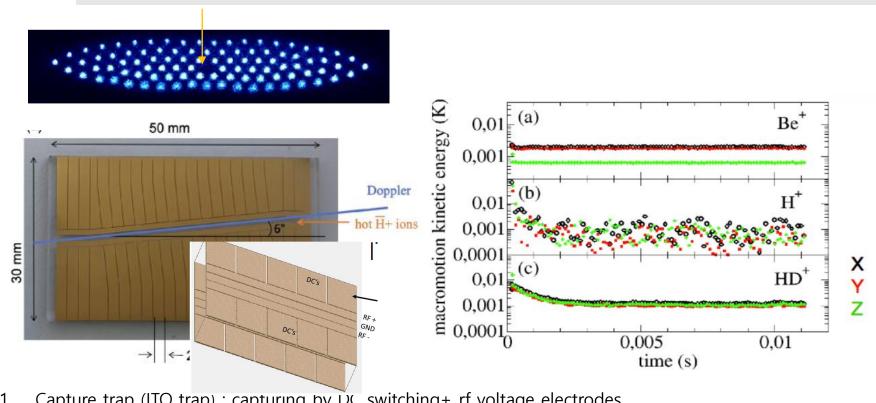
Antiproton trap



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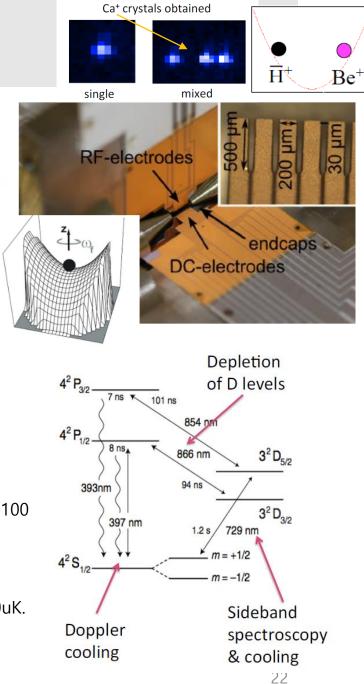
Cooling traps



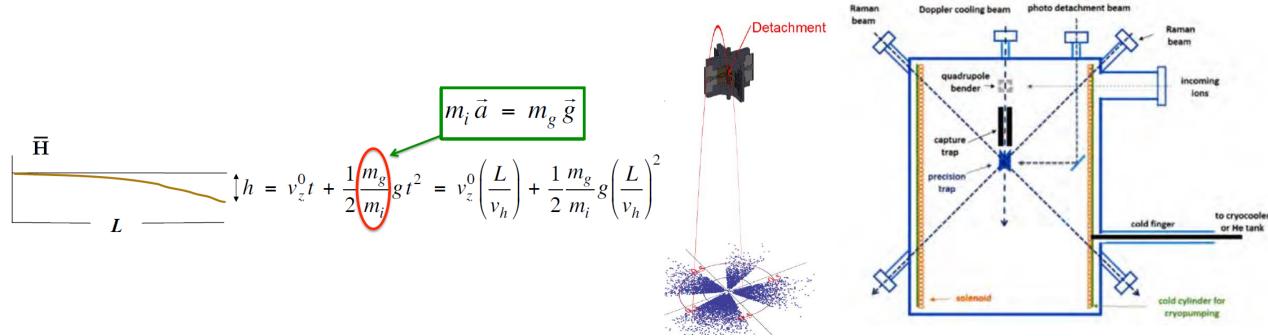
Capture trap (ITO trap): capturing by DC switching+ rf voltage electrodes

Sympathetic Doppler cooling by cooled Be+ ions (>10,000 laser(313nm) cooled Be+/HD+ ions (Wigner crystal), 100 neV, T~mK by rf heating)

- (L. Hilico et. al., Int. J. Mod. Phys. Conf. Ser. 30, 1460269 (2014))
- ← HD+ to reduce mass ratio for better coupling in simulation (similar charge to mass ratio required)
- 2. Precision trap : ion as a quantum harmonic oscillator, Raman sideband cooling for Be $^+/\overline{H}$ ion pair to T \sim 10uK. (W. Schnitzler et. al, Physical Review Letters 102, 070501 (2009).)



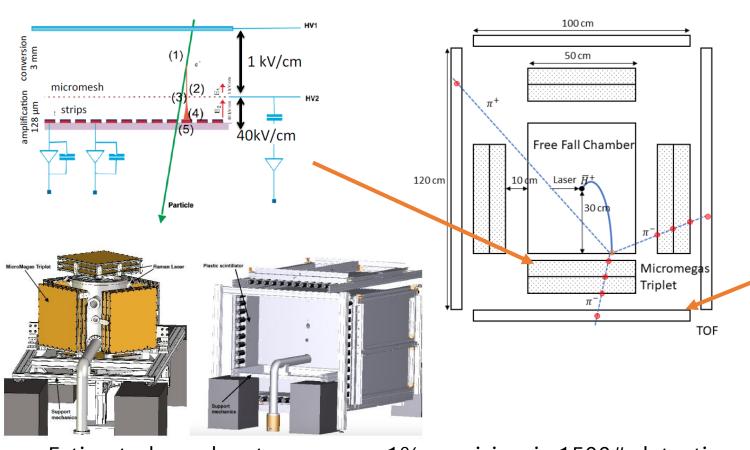
Freefall measurement

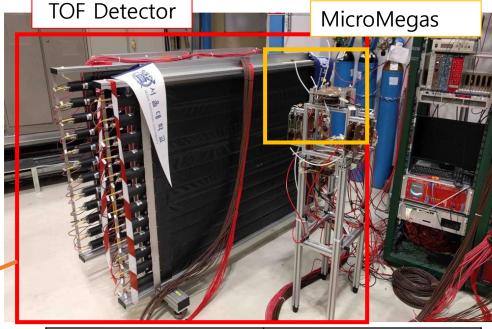


- Initial direction: temperature (0.44m/s for 1MHz) + photon recoil(0.24m/s) + positron emission (\sim 0.3m/s for E_c = 1ueV)
- Length: 10cm, height: 30cm (cf Aegis: L = 1m, h = 20um)
- Small magnetic field gradient, UHV, etc for systematic uncertainties have been studied
- Freefall time : about 500ms for $|\bar{g}| = g$
- uncertainty : $(\Delta z_0, \Delta v_{z0} t \ll z), \frac{\Delta \overline{g}}{\overline{g}} \cong \sqrt{(\frac{\Delta z}{z})^2 + (\frac{2\Delta t}{t})^2} \sim 0.4$ (for single \overline{H}), Δz (T(10uK), ΔE (1ueV):positron emission) better description in G. Dufour et. al., Eur. Phys. J. C (2014) 74:2731

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Freefall measurement

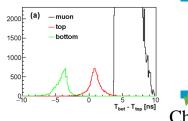


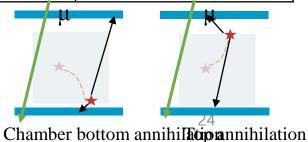


Temperature	Sign decision
10uK	~30 events
1mK	~6000 events

- Estimated number to measure 1% precision is 1500# detection of freefall antihydrogen (10uK)
- Simulation with deep learning has been studied
- Almost all cosmic ray background can be removed.

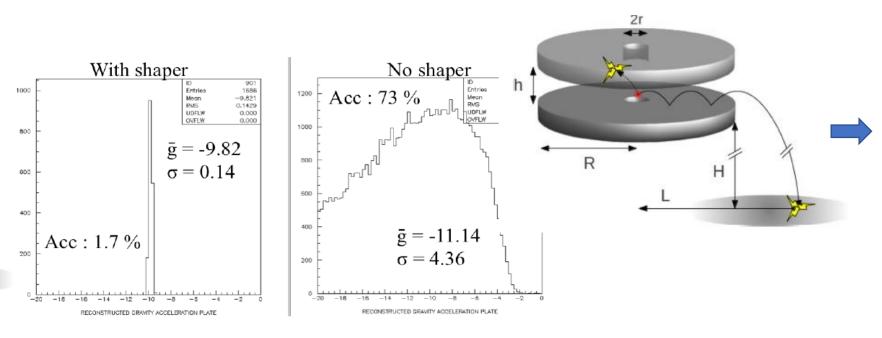
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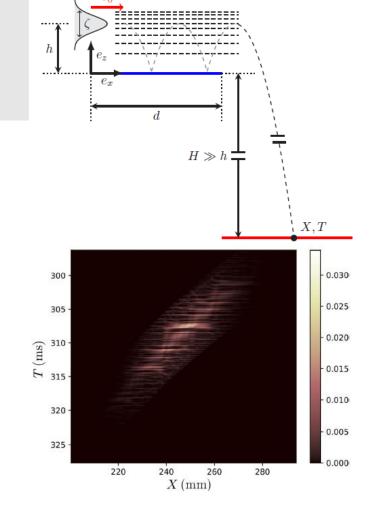




measurement 20210916

Further improvement





- Can be improved with proper shaper
- Quantum bouncing of antihydrogen (by Casimir-polder fore) : demonstrated with ultra cold neutron (<u>V. V. Nesvizhevsky et. al., Nature</u> volume 415, pages297–299 (2002)) :
- + shaping the distribution of velocity: precision be below 10⁻³ (G. Dufour et. al., Eur. Phys. J. C (2014) 74:2731)
- Quantum interference to improve the accuracy (10⁻⁵): detail is in P.-P. Crepin, Phys. Rev. A 99,042119

GBAR collaboration













70 members18 institutes8 countries







Korean group 4 Institutes 6 Ph. D 7 GS









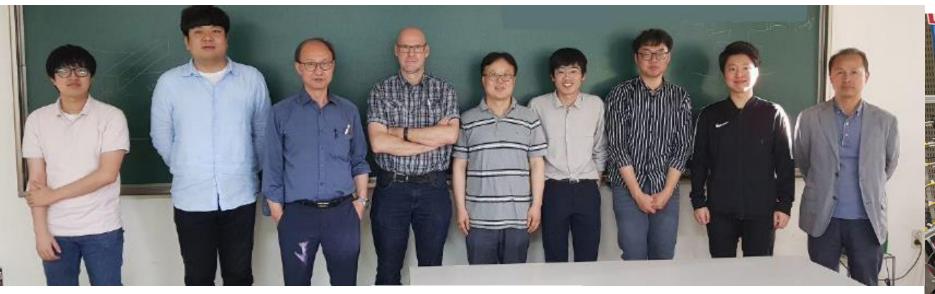








GBAR Korean collaboration













conclusion

- After the discovery of antimatter, there was arguments about antigravity.
- Recently, unprecedent experimental efforts are arose and the oretical rebellion started.
- Currently, experiments to measure the antihydrogen gravitation nal acceleration with 1% precision have competed.
- GBAR experiment will measure gravitational acceleration by cl assical freefall test using ultra-cold antihydrogen.

Thank you

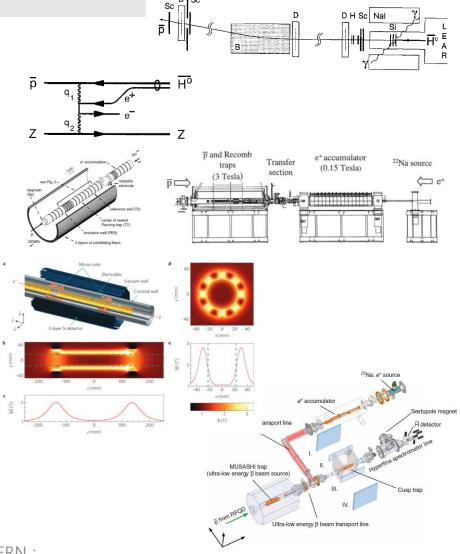
Motivation to study antihydrogen

- One of the best particle for fundamental physics as a direct tool of comparison
- : neutral antimatter composed of only antiparticles
- CPT test between matter and antimatter
- : Can extend Standard model
- Weak equivalence principle (WEP) test by antimatter
- : The gravitational interaction between matter and antimatter (antimatter and antimatter) has never been tested.

 WEP is only proven between matter and matter.

Recent breakthroughs of H

- 1995 : First production of antihydrogen at LEAR (Phys. Lett. B 368, 251) with stochastic cooling and electron cooling
- 2002 : Cold antihydrogen production at ATHENA(Nature 419, 456-459), ATRAP(Phys, Rev. Lett 89, 213401)
- 2010~2011: Trapping antihydrogen (Nature 468, 673) up to 1000s (Nature physics, 7, 558-564) at ALPHA
- 2013 : Antihydrogen gravity test by ALPHA (Nature communications 4, 1785)
- 2014 : antihydrogen beam source by ASACUSA (Nature communications 5, 3089)
- 2021 : Antihydrogen laser cooling by ALPHA(Nature 592, 35-42)



CPT test

Spectroscopy of antihydrogen

1S-2S transition: hydrogen 4.2e-15

- antihydrogen $2x10^{-10}(2017)$ → $2x 10^{-12} (2018)$

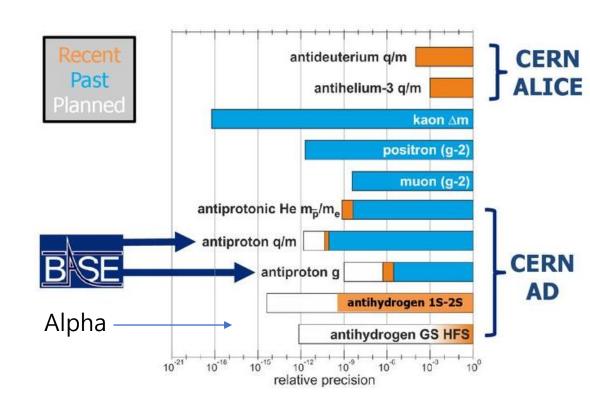
Hyperfine splitting : observed 2% $(2P_{1/2}-2P_{3/2})$

(nature 548, 66-69 (2017), nature 578, 375 (2020)) by ALPHA

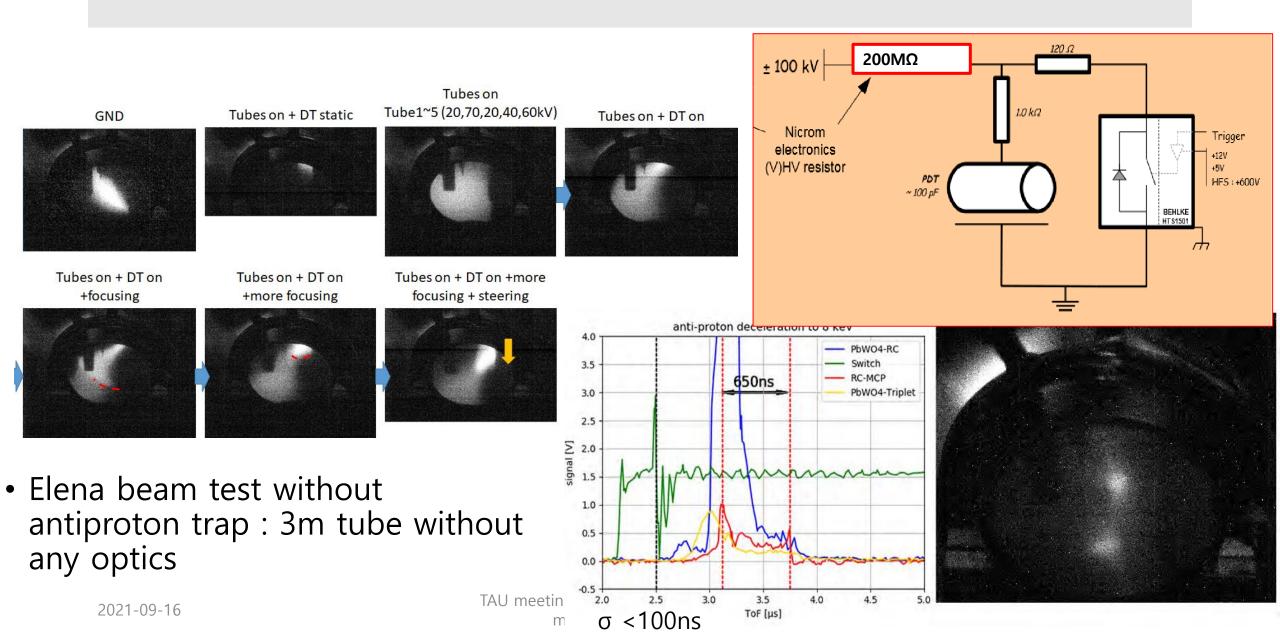
- Lamb shift: agreed a level of 11% (2S_{1/2}-2P_{1/2}) (nature 578, 375 (2020)) by ALPHA
- Proton & antiproton CPT test by BASE

(nature 524, 196-199 (2015)) and updated information is not in the slide.

- Antideuteron, antihelium, muon, positron, kaon..
- (There's also many fantastic results for exotic antiprotonic helium, etc..)



Elena beam test before LS2



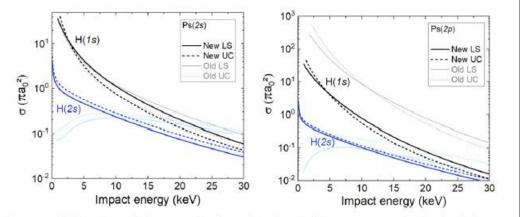


Figure 2. Antihydrogen ion production cross sections from positronium excited in a state $n_p = 2$ as a function of antihydrogen impact energy. See the caption of figure 1 for the detailed legend of the graph.

