

Functional Renormalization Group method for Nuclear Matter and Neutron Star

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in collaboration with

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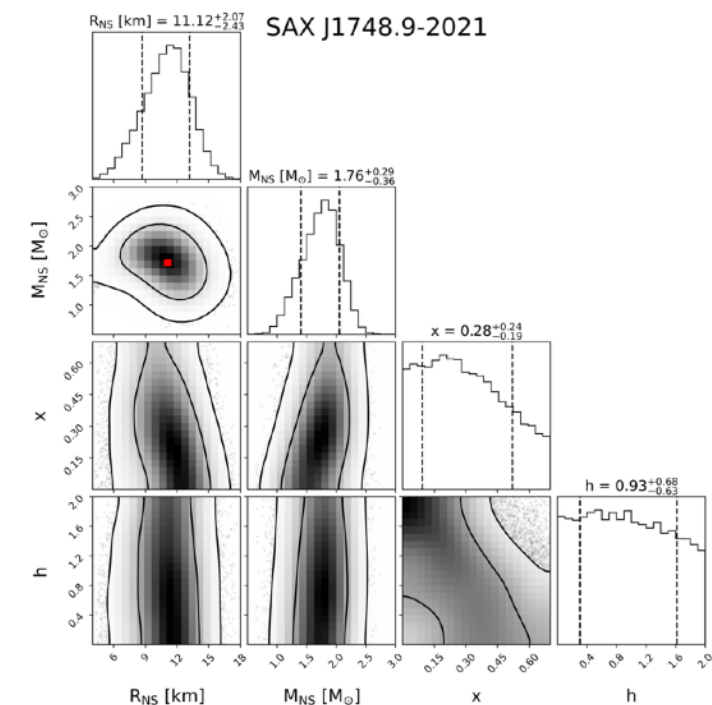
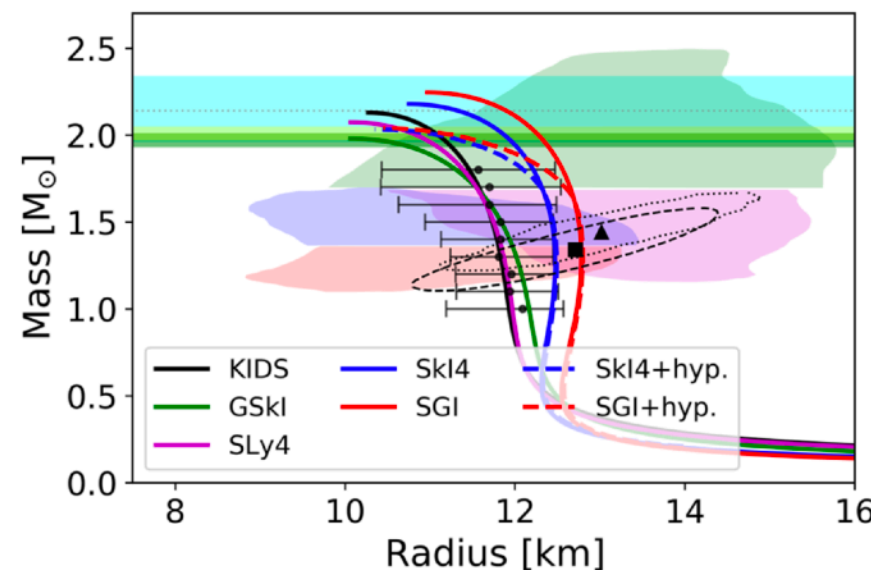
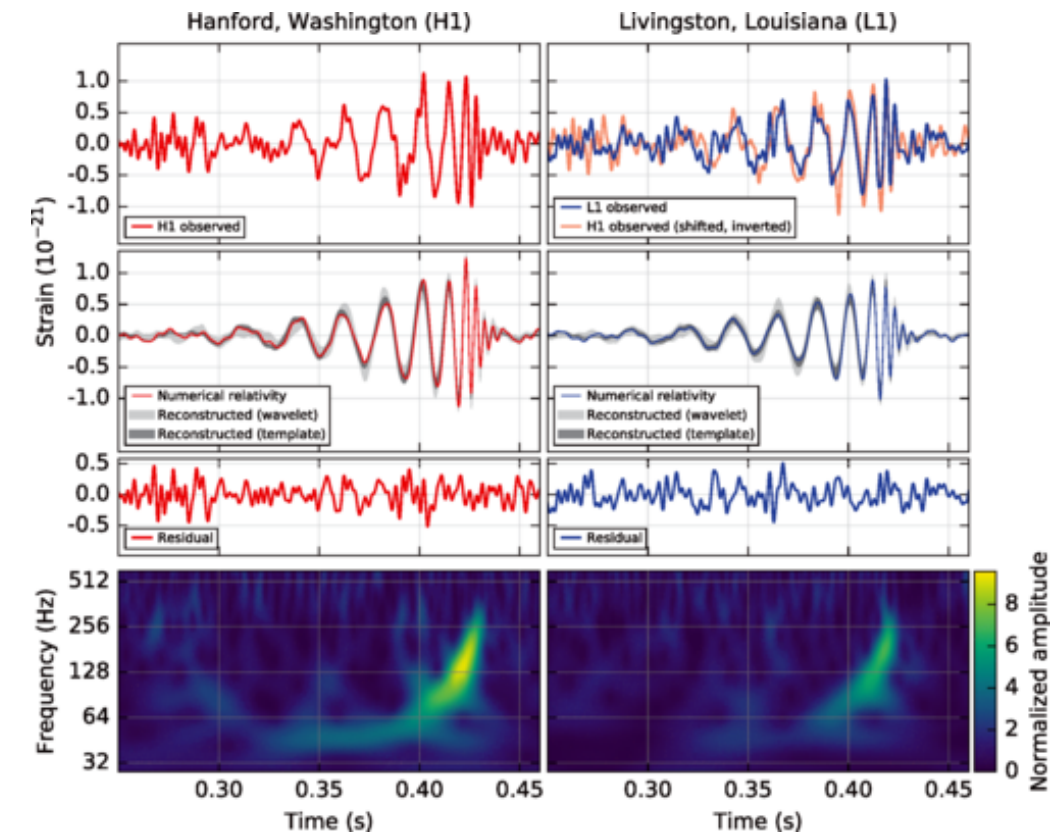
My Research is ...

Gravitational Wave Data Analysis

- Detector Characterization + Machine Learning
- Bayesian Inference on Neutron Star properties : NS mass, radius, tidal deformability, etc.

Dense Nuclear Matter & Neutron Star

- Neutron Star Equation of State
- DJBUU + RAON



Polytropic equation of state

$$p = K \rho^\Gamma = K \rho^{1+1/n}$$

$$\begin{aligned} \epsilon &= (1+a) \rho c_0^2 + \frac{p}{\Gamma-1} \\ &= (1+a) \rho c_0^2 + \frac{K}{\Gamma-1} \rho^\Gamma, \end{aligned}$$

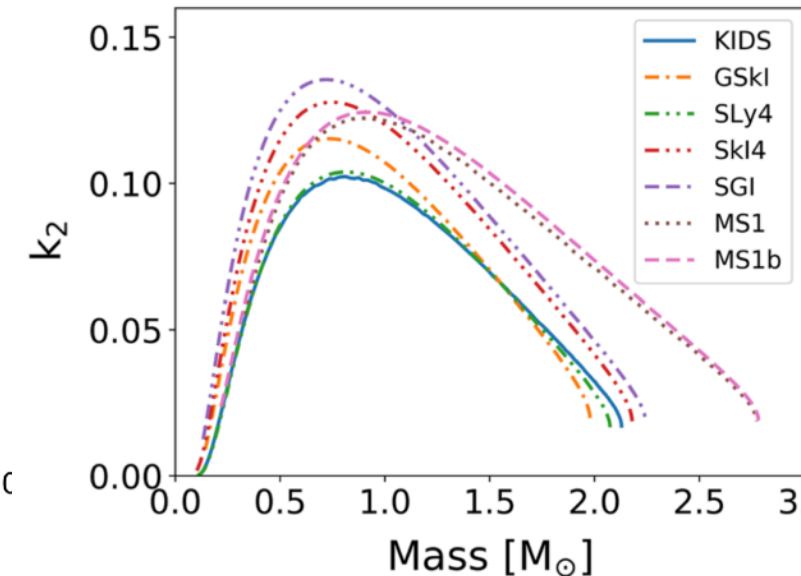
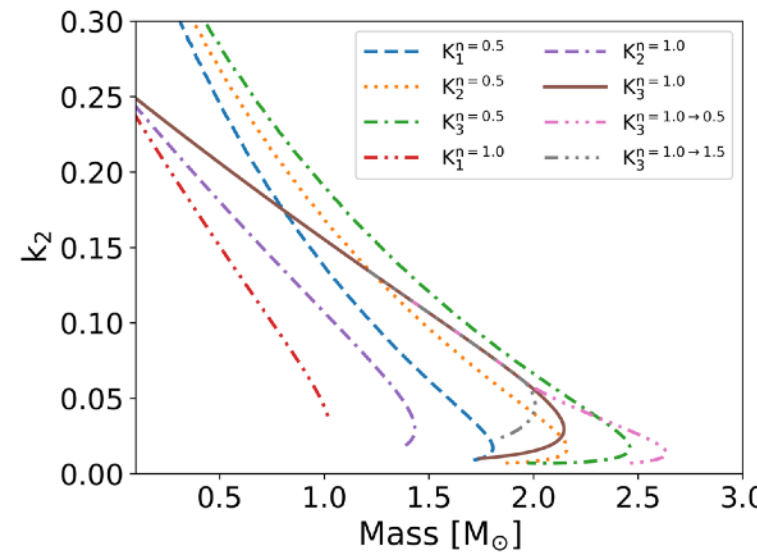
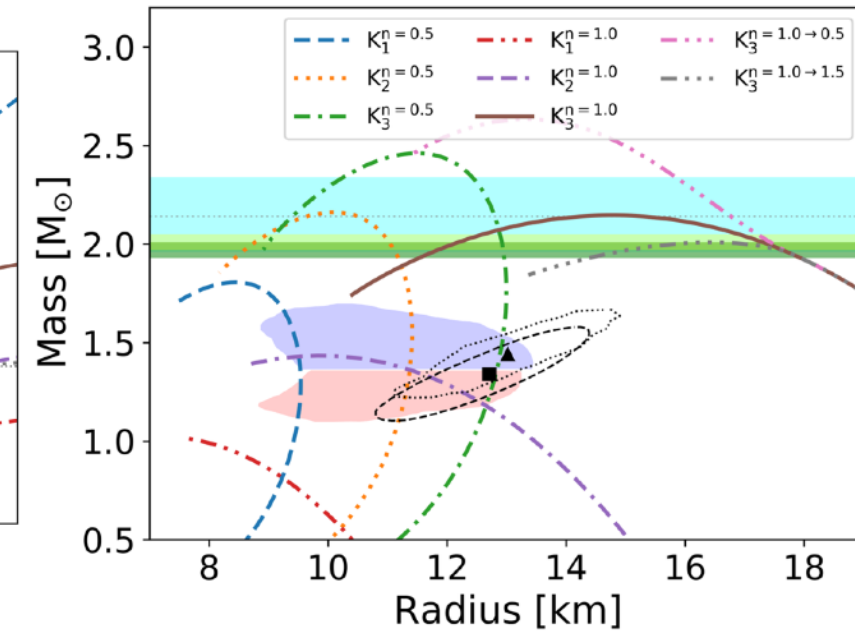
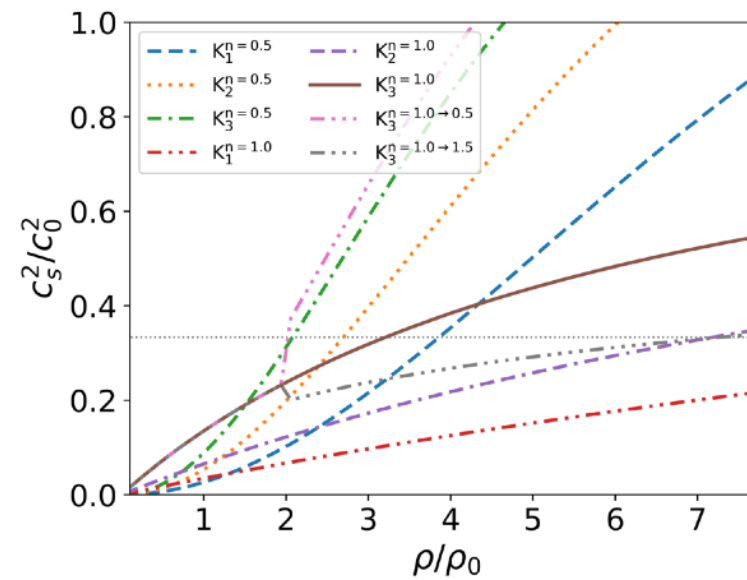
$$\epsilon_2 = (1+a_2) \rho c_0^2 + \frac{K_2}{\Gamma_2-1} \rho^{\Gamma_2},$$

$$a_2 = \left(\frac{1}{\Gamma_1-1} - \frac{1}{\Gamma_2-1} \right) K_2 \rho_t^{\Gamma_2-1}$$

$$\frac{c_s^2}{c_0^2} = \frac{dp}{d\epsilon} = \Gamma \frac{p}{\epsilon + p}$$

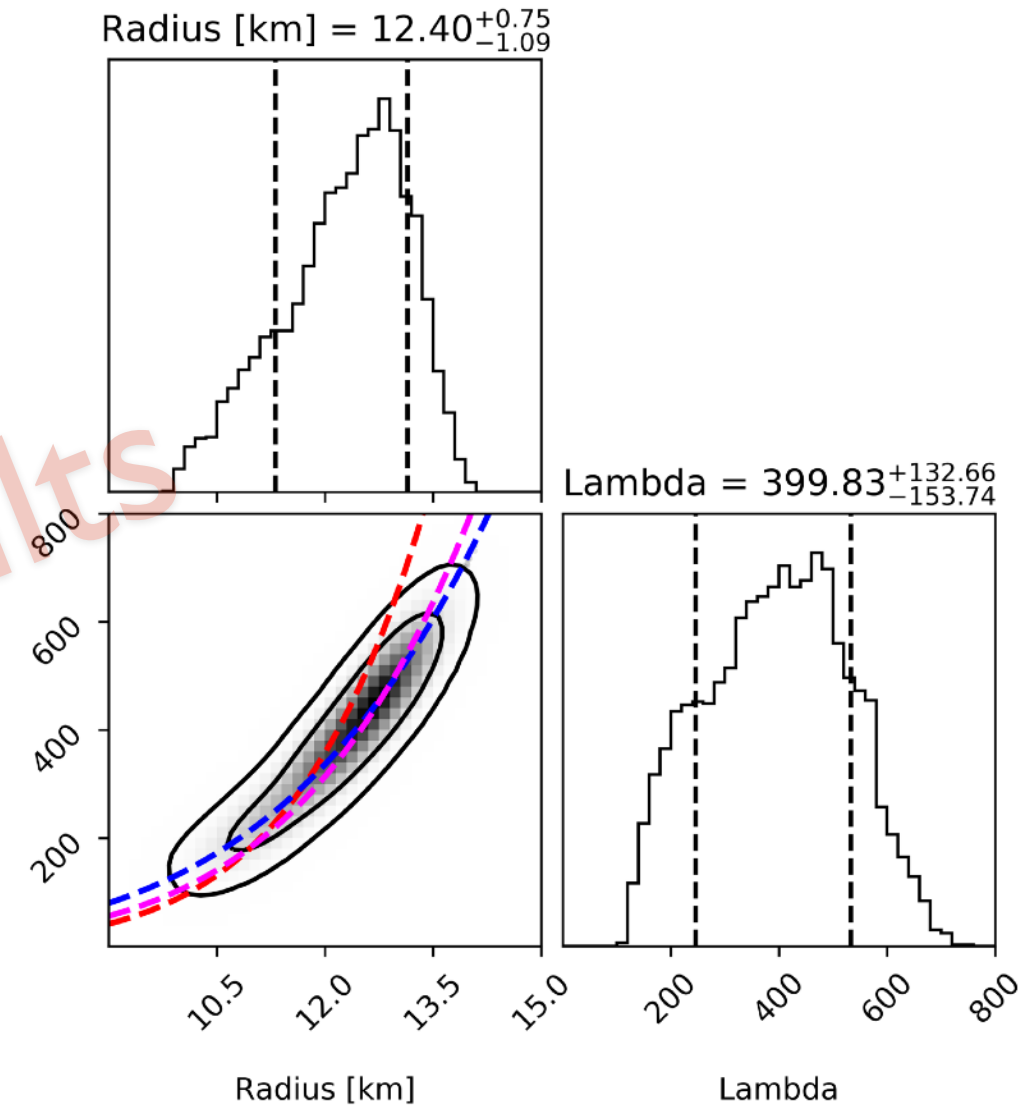
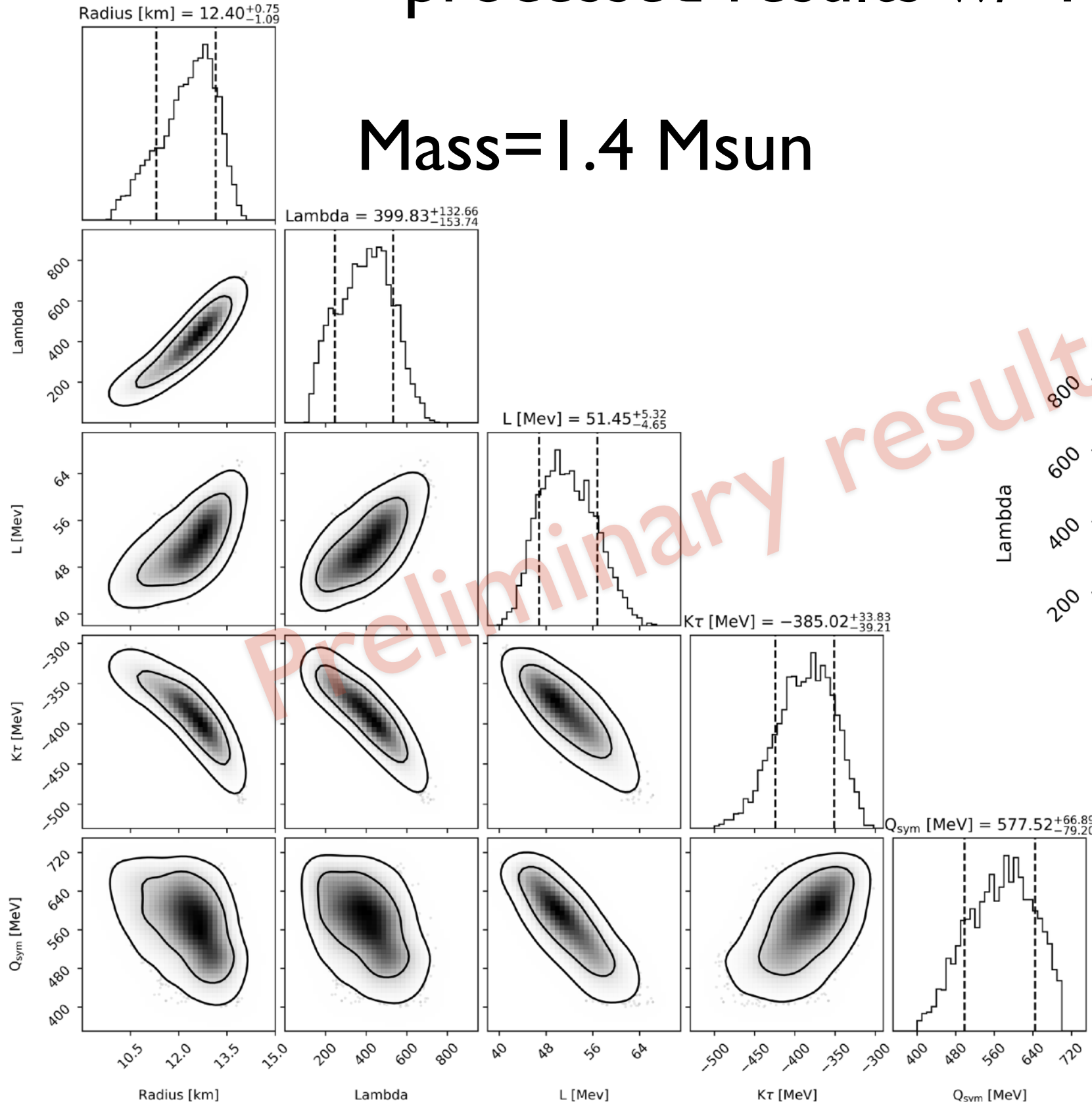
$K_1 < K_2 < K_3$

M. Kim et al. (JKPS 2021)



Posterior samples and Post-processed results w/ KIDS

Mass = 1.4 Msun



Red line: $\Lambda(1.4M_{\odot}) = 2.88 \times 10^{-6} (R/\text{km})^{7.5}$

PhysRevLett.120.172703.pdf

Blue line: $\Lambda(1.4M_{\odot}) = 1.35 \times 10^{-3} (R/\text{km})^{5.0}$

Implication of PREX 2 - PhysRevLett.126.172503 (2021)
 $\Rightarrow \Lambda \sim R^{4.8}$

Magenta line: $\Lambda(1.4M_{\odot}) = 1.05 \times 10^{-4} (R/\text{km})^{6.0}$

More reliable theoretical approach is ...

Functional Renormalization Group (FRG) method

$$\Gamma_k = \int d^4x \mathcal{L}|_{U \rightarrow U_k},$$

Zhang et al., PHYSICAL REVIEW D 96, 114029 (2017)

Example: two flavor quark-meson
model w/ omega meson

$$\partial_k \Gamma_k = \frac{1}{2} \text{Tr} \left[\frac{\partial_k R_k}{\Gamma_k^{(2)} + R_k} \right]$$

$$\begin{aligned} \mathcal{L} = & \bar{\psi} [i\gamma_\mu \partial^\mu - g_s(\sigma + i\gamma_5 \boldsymbol{\tau} \cdot \boldsymbol{\pi}) - g_v \gamma_\mu \omega^\mu + \mu \gamma_0] \psi \\ & + \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma + \frac{1}{2} \partial_\mu \boldsymbol{\pi} \cdot \partial^\mu \boldsymbol{\pi} - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ & - U(\sigma, \boldsymbol{\pi}, \omega), \end{aligned}$$

$$\begin{aligned} \partial_k U_k^\phi(T, \mu) = & \frac{k^4}{12\pi^2} \left\{ \frac{3[1 + 2n_B(E_\pi)]}{E_\pi} + \frac{1 + 2n_B(E_\sigma)}{E_\sigma} \right. \\ & \left. - \frac{2\nu_q[1 - n_F(E_q, \mu_{\text{eff}}^k) - n_F(E_q, -\mu_{\text{eff}}^k)]}{E_q} \right\} \end{aligned}$$

2 flavor Quark-Meson model w/ ω meson (I)

Zhang et al., PHYSICAL REVIEW D 96, 114029 (2017)

Flow eq.

$$\partial_k U_k^\phi(T, \mu) = \frac{k^4}{12\pi^2} \left\{ \frac{3[1 + 2n_B(E_\pi)]}{E_\pi} + \frac{1 + 2n_B(E_\sigma)}{E_\sigma} - \frac{2\nu_q[1 - n_F(E_q, \mu_{\text{eff}}^k) - n_F(E_q, -\mu_{\text{eff}}^k)]}{E_q} \right\} \quad \phi^2 \equiv \sigma^2 + \pi^2.$$

where single-particle energies are

$$E_\pi = \sqrt{k^2 + 2U'_k},$$

$$E_\sigma = \sqrt{k^2 + 2U'_k + 4\phi^2 U''_k},$$

$$E_q = \sqrt{k^2 + g_s^2 \phi^2},$$

$$U'_k \equiv \frac{\partial U_k}{\partial \phi^2}$$

$$U_k = U_k^\phi + U_k^\omega$$

$$U_k^\omega = -\frac{1}{2}m_v^2 \omega_{0,k}^2$$

$$n_B(E) = \frac{1}{e^{\beta E} - 1}$$

$$n_F(E, \mu) = \frac{1}{e^{\beta(E-\mu)} + 1}$$

2 flavor Quark-Meson model w/ ω meson (2)

$$\partial_k \omega_{0,k} = -\frac{2g_v k^4}{\phi^2 m_v^2 E_q} \frac{\partial}{\partial \mu} (n_F(E_q, \mu_{\text{eff}}^k) + n_F(E_q, -\mu_{\text{eff}}^k))$$

Initial condition w/ $\Lambda=500$ MeV,

$$U_{\Lambda}^{\phi} = \frac{\lambda}{4} \phi^4$$

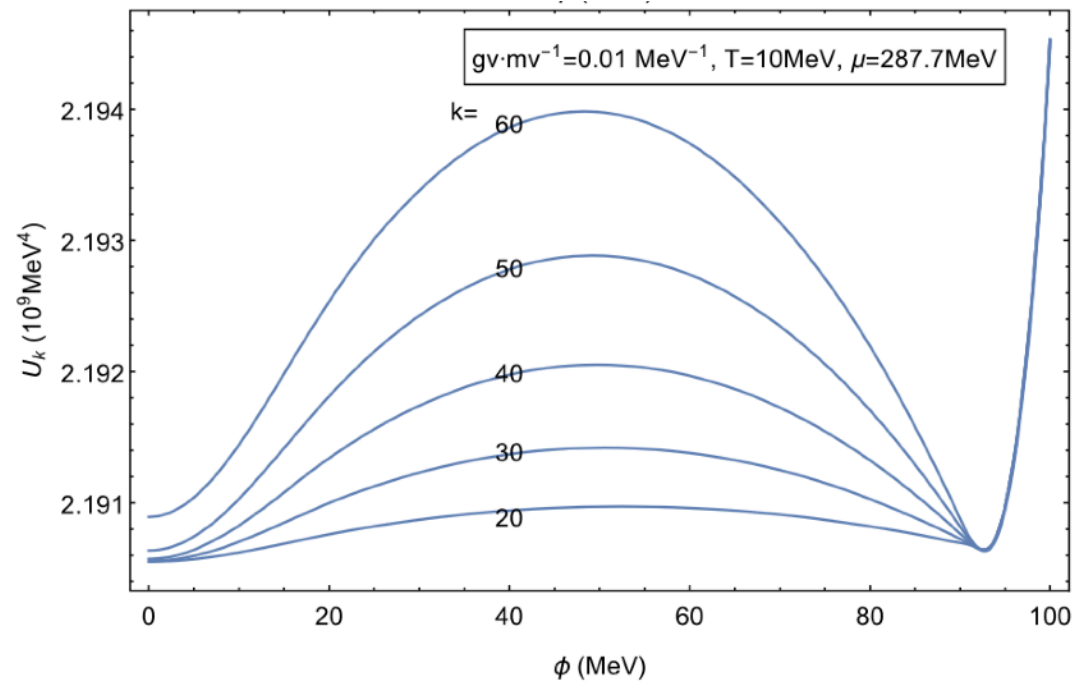
$$g_s = 3.2, \lambda = 8$$

$$\omega_{0,k=\Lambda}(\phi) = 0$$

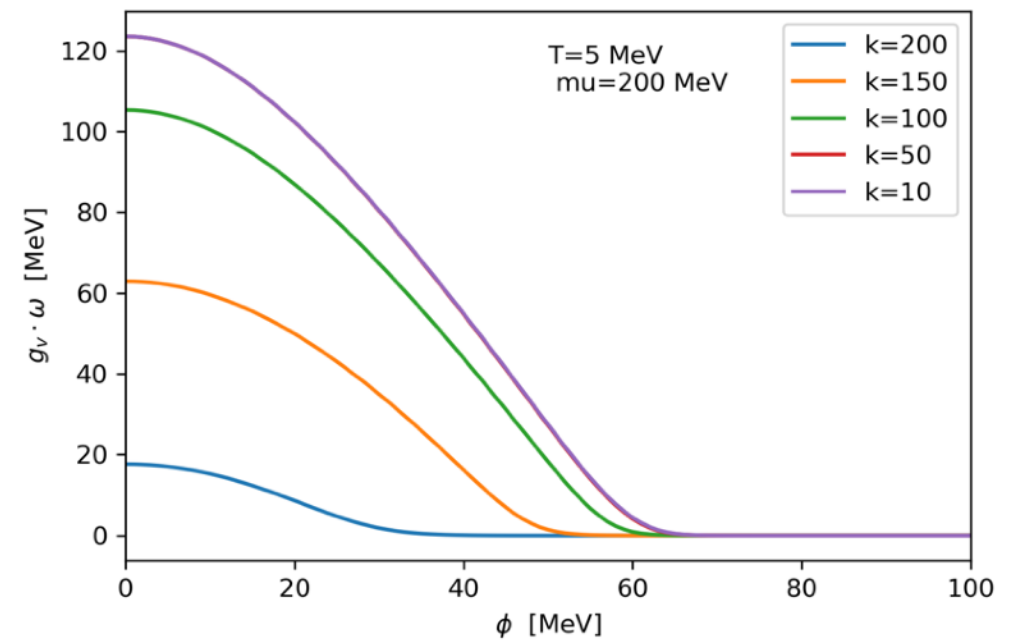
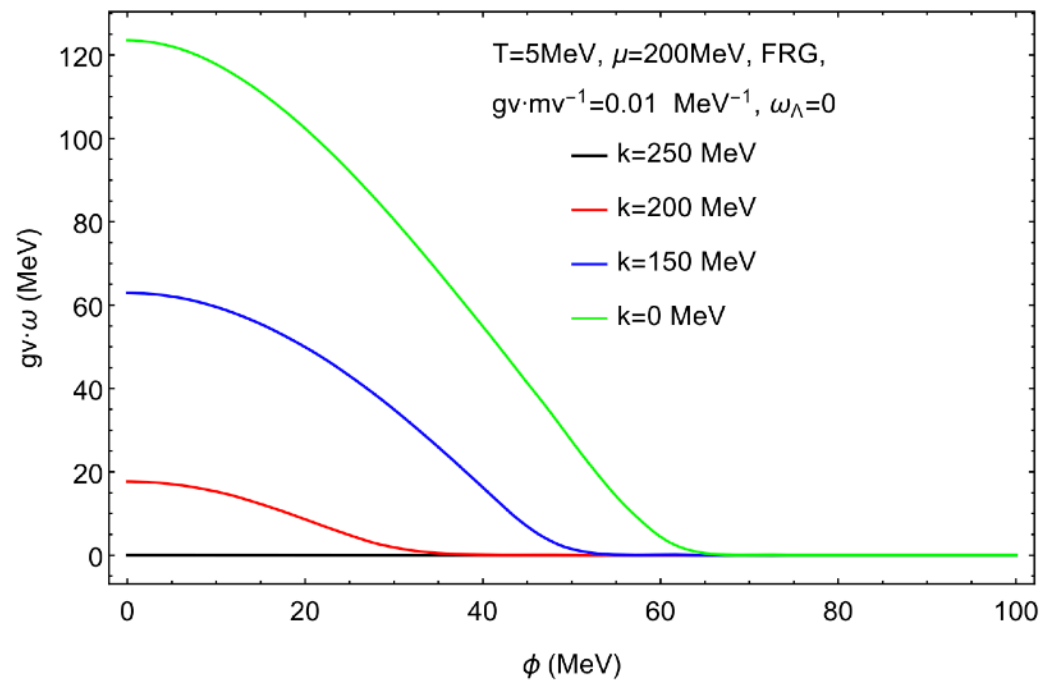
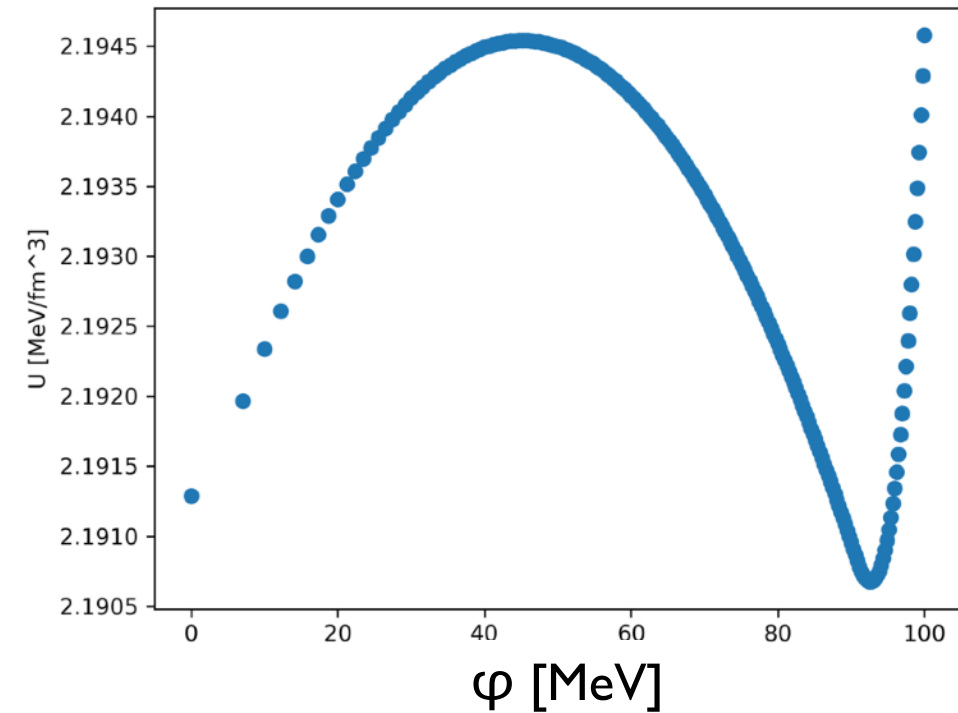
$$\phi = [0, 100] \text{ Mev}$$

Test Results - Quark-Meson model

Zhang et al., PHYSICAL REVIEW D 96, 114029 (2017)



Our results



FRG in an Extended Walecka model

$$\begin{aligned}\mathcal{L} = & \bar{\psi}[i \not{\partial} - M_N + g_\sigma \phi - g_\omega \not{\omega} - g_\rho \not{\vec{b}}]\psi, \\ & + \frac{1}{2}(\partial_\mu \phi \partial^\mu \phi - m_\sigma^2 \phi^2) - U(\phi) \\ & + \frac{1}{2}m_\omega^2 \omega^2 + \frac{1}{2}m_\rho^2 \vec{b} \cdot \vec{b} - \frac{1}{2}F_{\mu\nu}F^{\mu\nu} - \frac{1}{2}G_{\mu\nu}G^{\mu\nu}\end{aligned}$$

$$k \frac{\partial U_{k,\chi}}{\partial k} = \frac{k^5}{12\phi^2} \left[\frac{1 + 2n_B(E_\phi)}{E_\phi} - \sum_{i=n,p} 4 \frac{1 - \sum_{r=\pm 1} n_F(E_N - r\mu_{i,\text{eff}})}{E_N} \right]$$

$$g_\omega \omega_{0,k} = \frac{f_\omega}{3\pi^2} \int_k^\Lambda dp \frac{p^4}{E_N} \sum_{r=\pm 1} \frac{\partial}{\partial \mu} [n_F(E_N - r\mu_{p,\text{eff}}) + n_F(E_N - r\mu_{n,\text{eff}})]$$

$$g_\rho \rho_{0,k} = \frac{f_\rho}{3\pi^2} \int_k^\Lambda dp \frac{p^4}{E_N} \sum_{r=\pm 1} \frac{\partial}{\partial \mu} [n_F(E_N - r\mu_{p,\text{eff}}) - n_F(E_N - r\mu_{n,\text{eff}})]$$

Parameters for our calculation

Initial condition w/ $\Lambda=1400$ MeV,

$$g_\omega \omega_0(\Lambda, \mu, T) = 0$$

$$g_\rho \rho_0(\Lambda, \mu, T) = 0$$

$$U_\Lambda^\phi = \frac{1}{3} \tilde{a} \phi^3 + \frac{1}{4} \tilde{b} \phi^4$$

For the reference,

Liu et al., PHYSICAL REVIEW C **65** 045201

TABLE I. Parameter sets.

Parameter	Set I	Set II	NL3
f_σ (fm ²)	10.33	same	15.73
f_ω (fm ²)	5.42	same	10.53
f_ρ (fm ²)	0.95	3.15	1.34
f_δ (fm ²)	0.00	2.50	0.00
A (fm ⁻¹)	0.033	same	-0.01
B	-0.0048	same	-0.003

Parameters to be optimized,

$$f_i \equiv \frac{g_i^2}{m_i^2} \quad \tilde{a}, \tilde{b}$$

$$\rho_0 = 0.16/\text{fm}^3$$

$$E/A = -16\text{MeV}$$

$$M^* = 0.75M_N$$

$$K_v = 240\text{MeV}$$

Thermodynamic Properties

$$\Omega(T, \mu_{n,p}) = U(T, \mu_{n,p}, \sigma) \quad \text{with} \quad \sigma = \langle \phi \rangle$$

$$p = -\Omega(T, \mu_{n,p}), \quad n_i = -\frac{\partial \Omega(T, \mu_{n,p})}{\partial \mu_i}, \quad (\text{in our calculation, } \Delta\mu=0.1 \text{ MeV})$$

$$s = -\frac{\partial \Omega(T, \mu_{n,p})}{\partial T}, \quad \epsilon = -p + Ts + \sum_{i=n,p} \mu_i n_i.$$

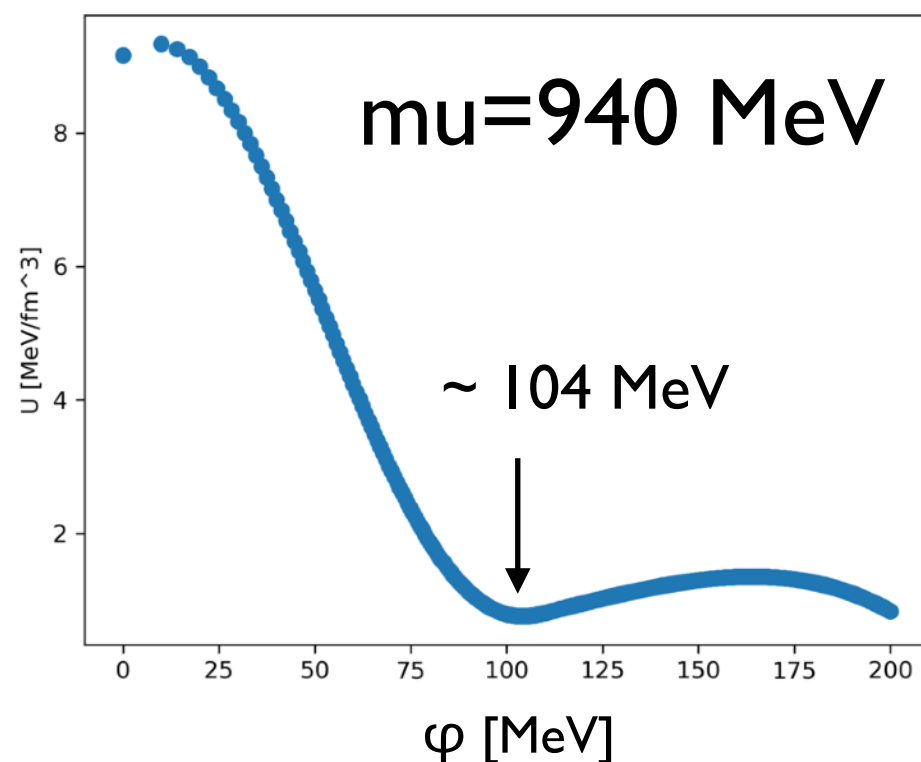
Test Results (I) for SNM ($\mu=\mu_p=\mu_n$)

For the reference,

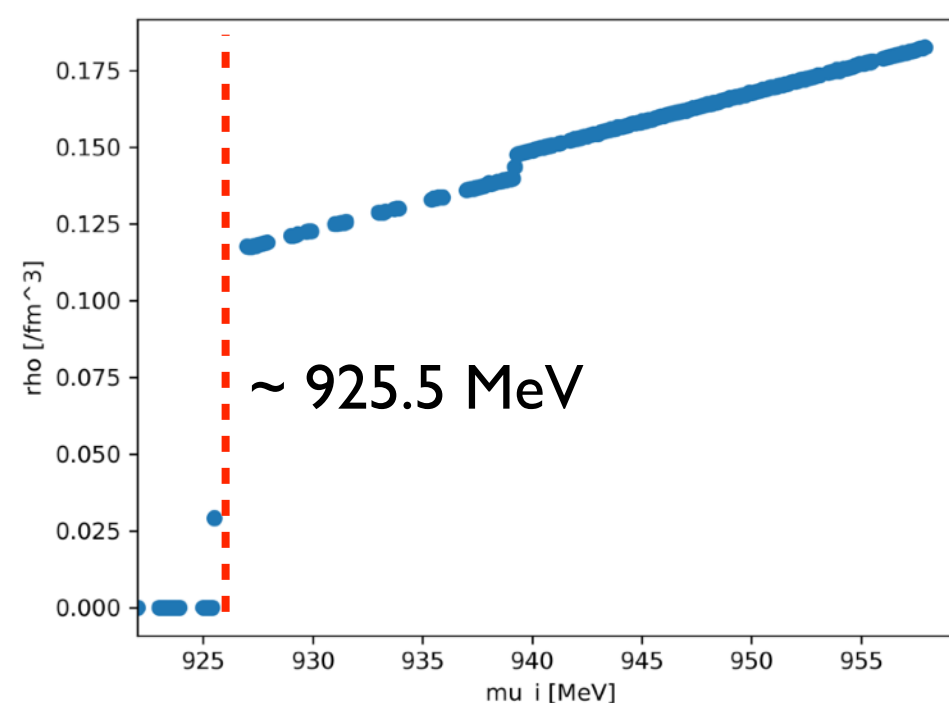
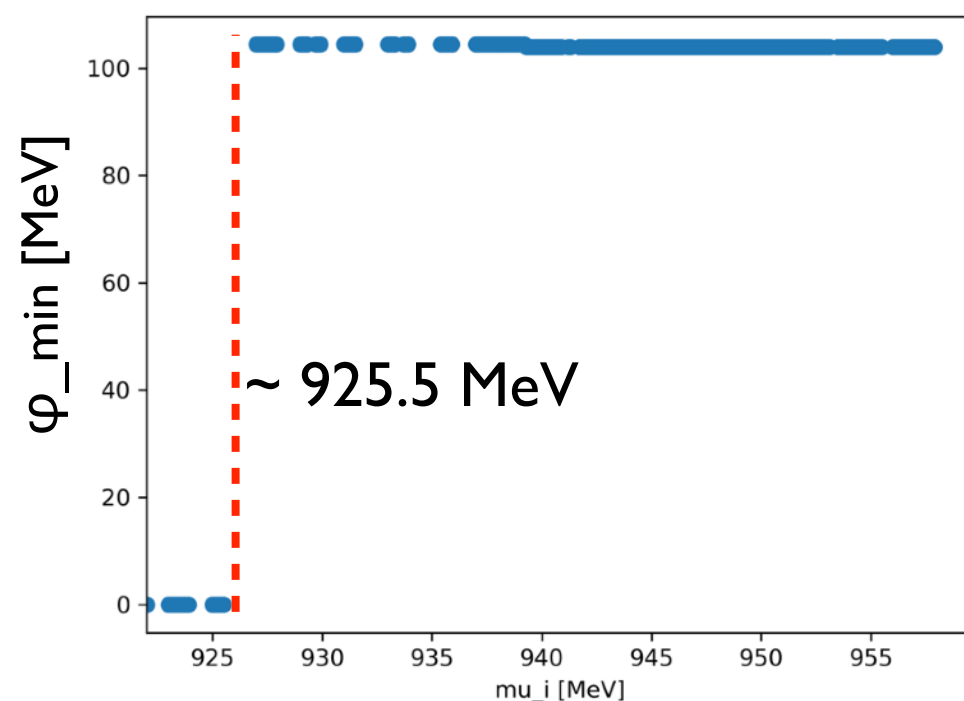
Liu et al., PHYSICAL REVIEW C **65** 045201

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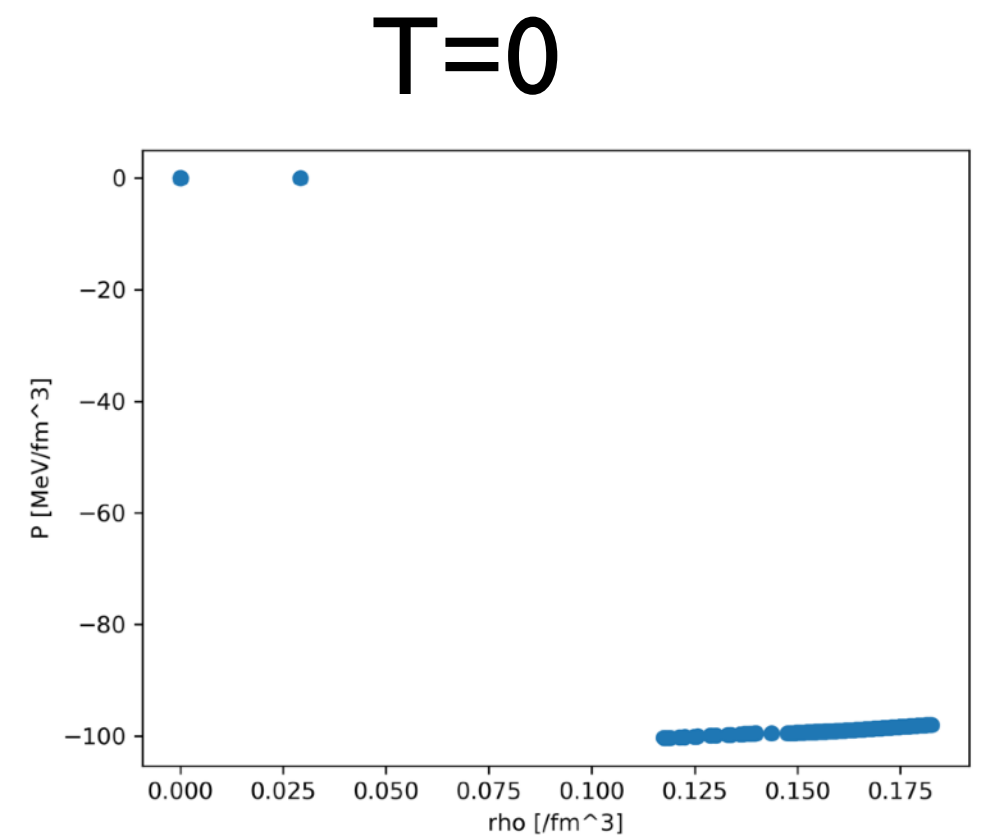
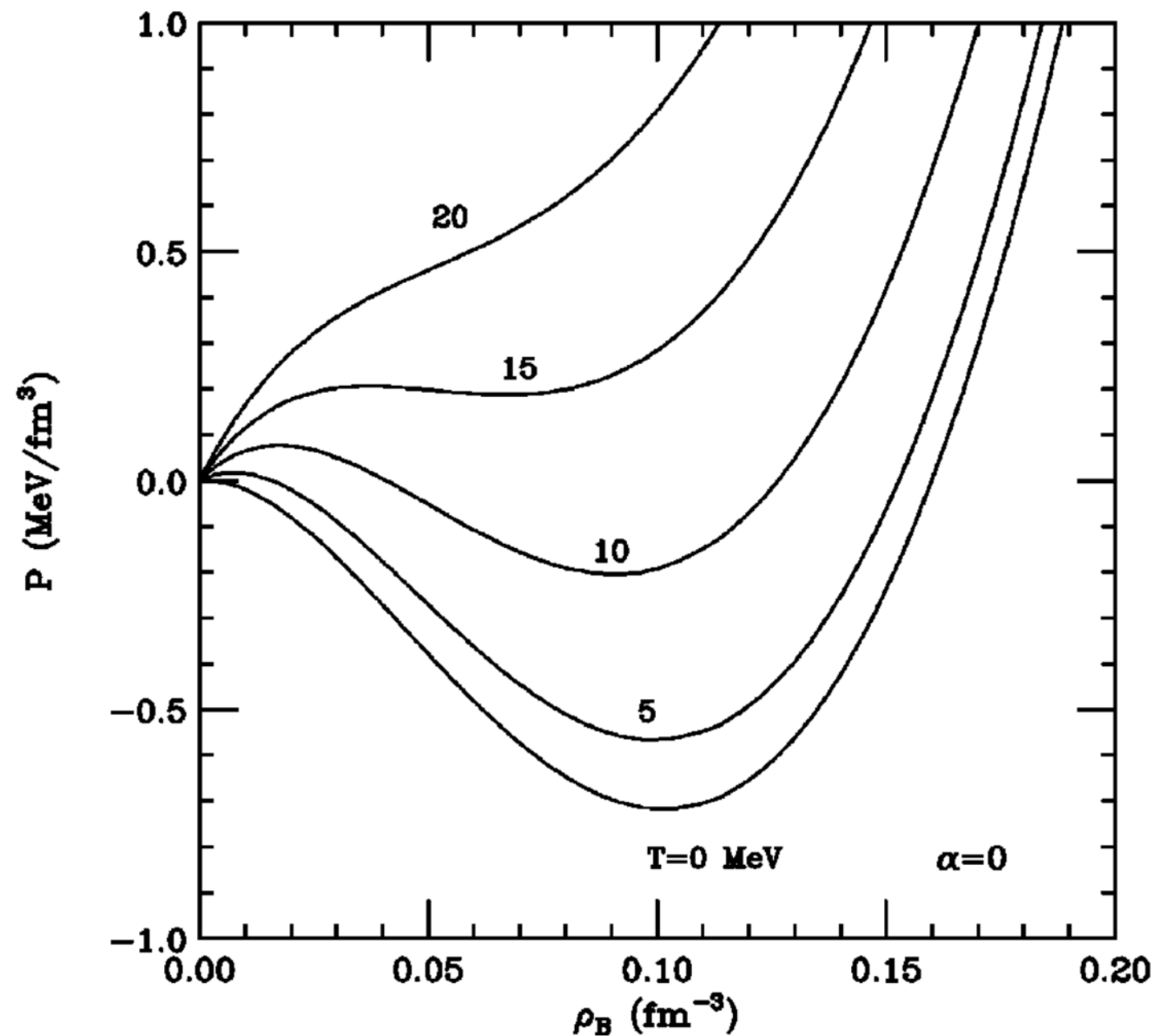


$\Delta\mu=0.1$ MeV



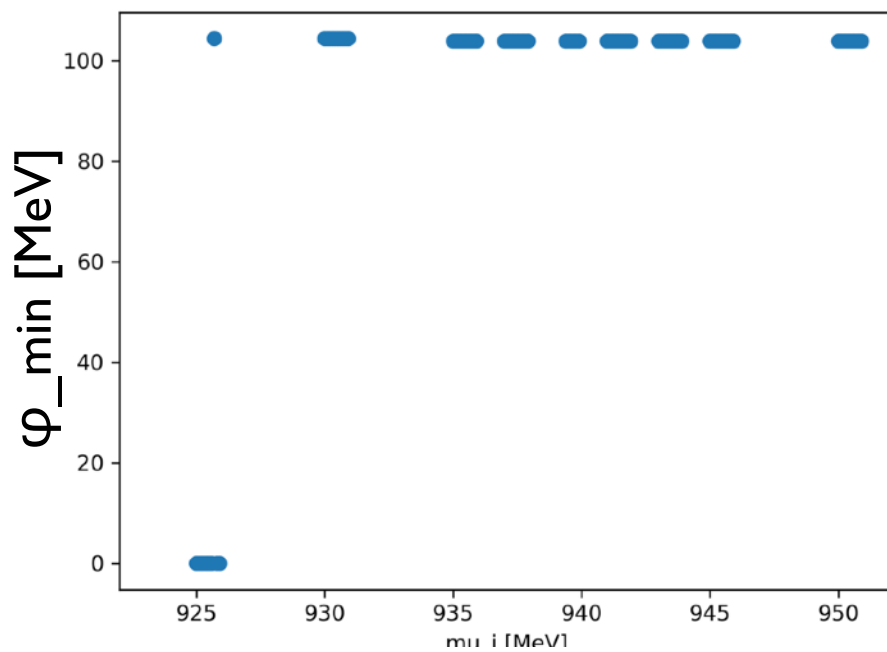
Test Results (2) for SNM

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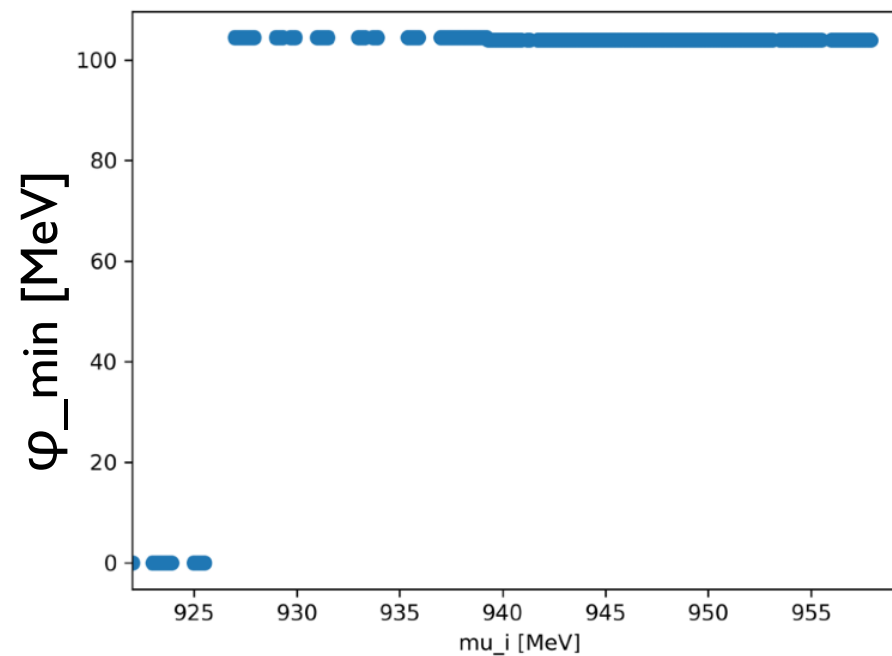


Test Results (3) for SNM

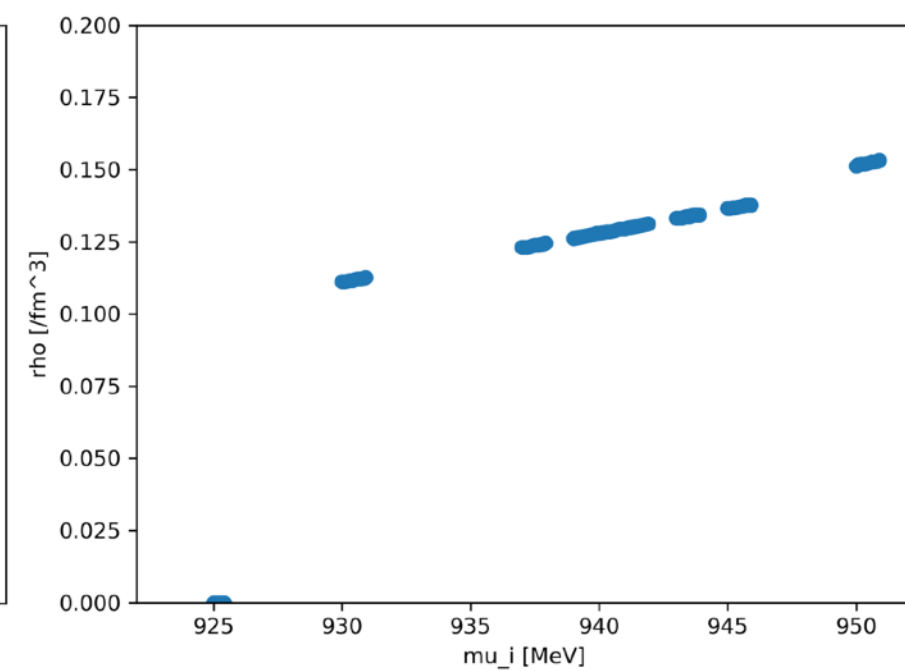
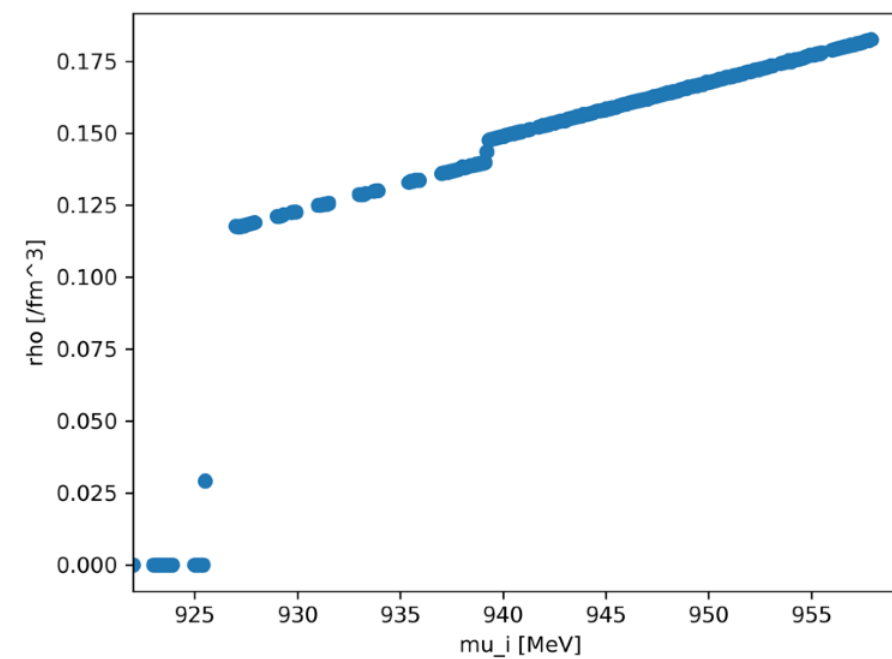
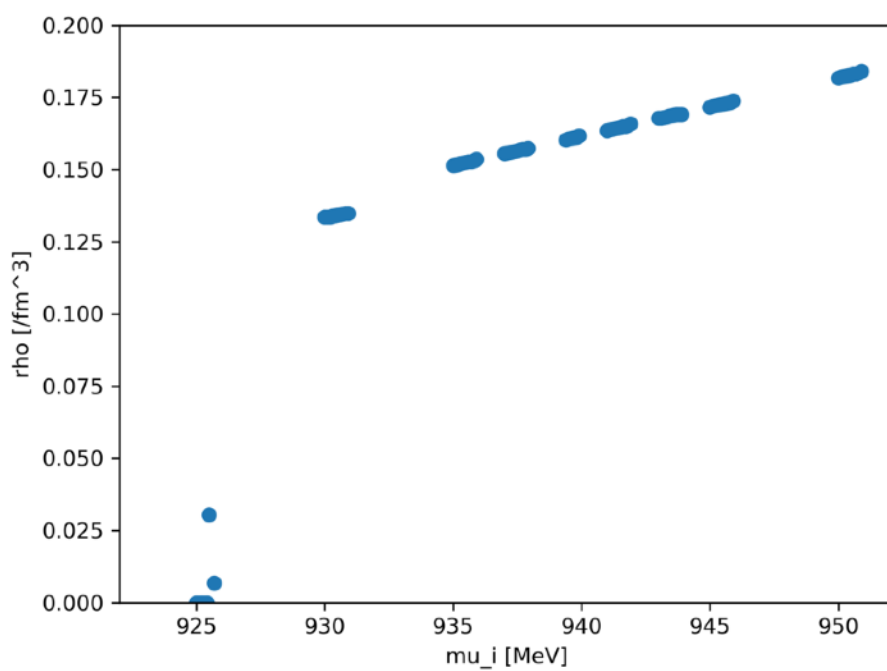
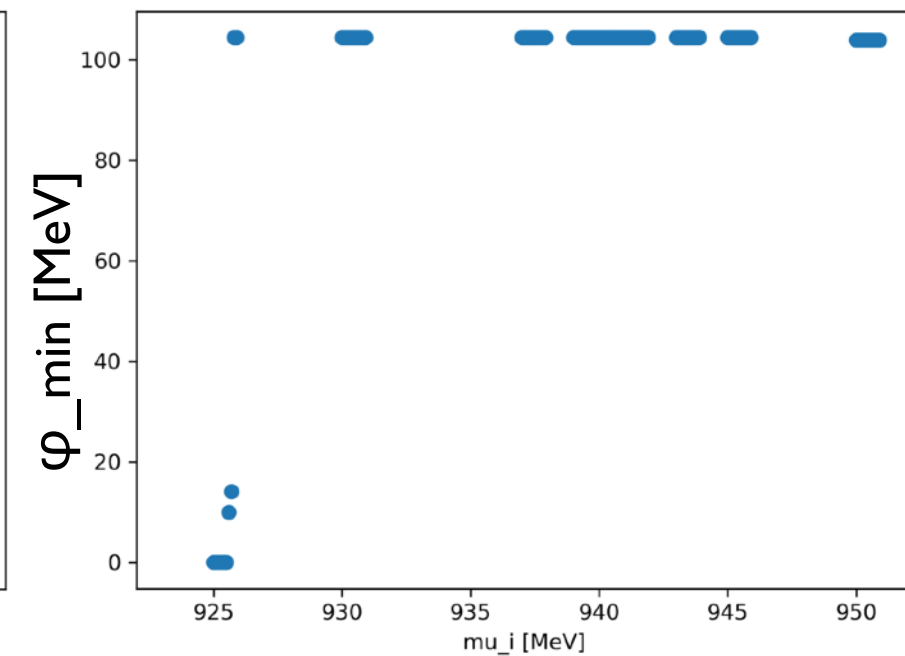
$f_w=5.0 \text{ fm}^2$



$f_w=5.42 \text{ fm}^2$

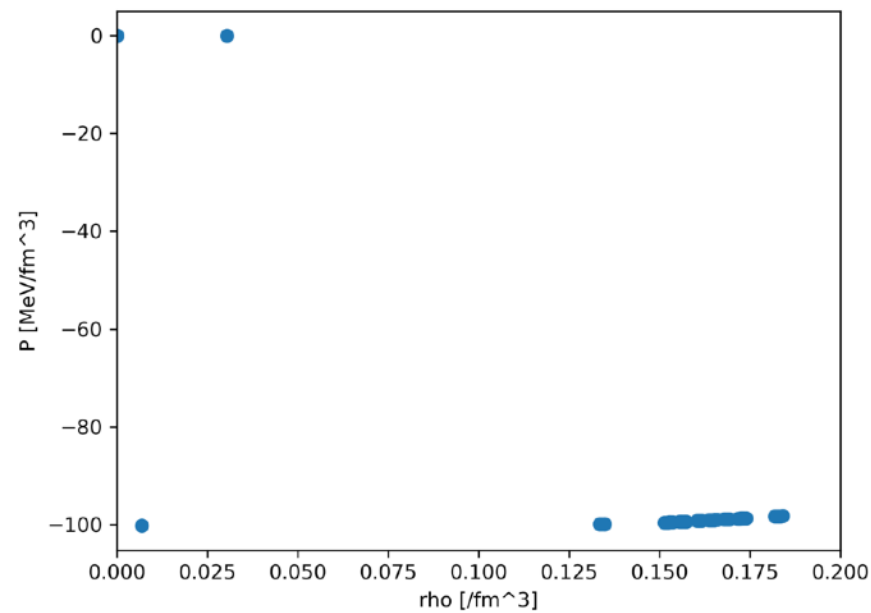


$f_w=6.0 \text{ fm}^2$

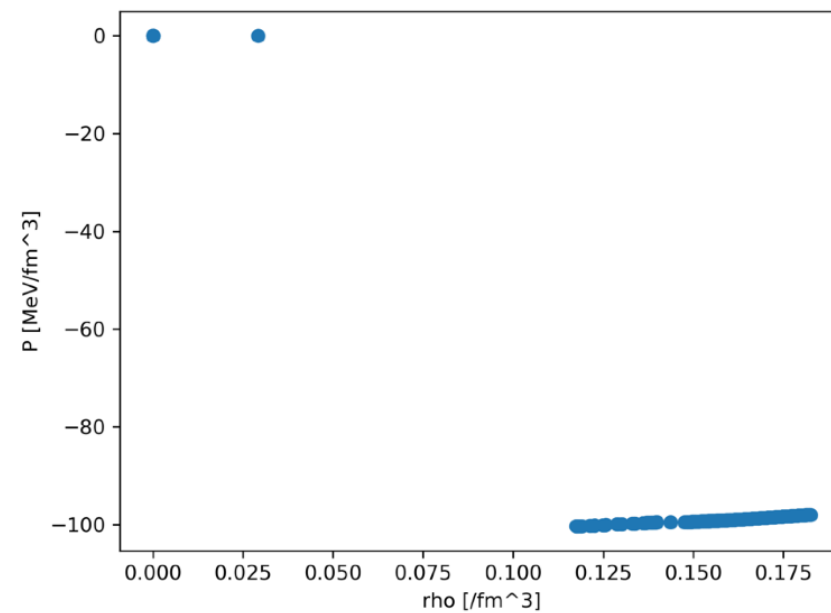


Test Results (4) for SNM

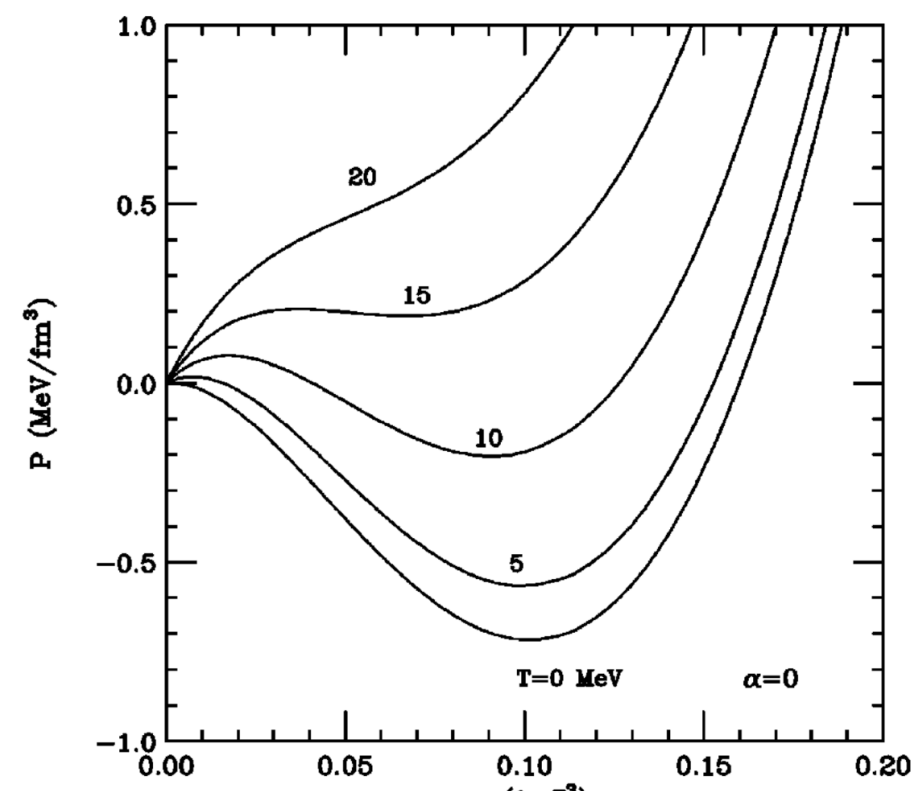
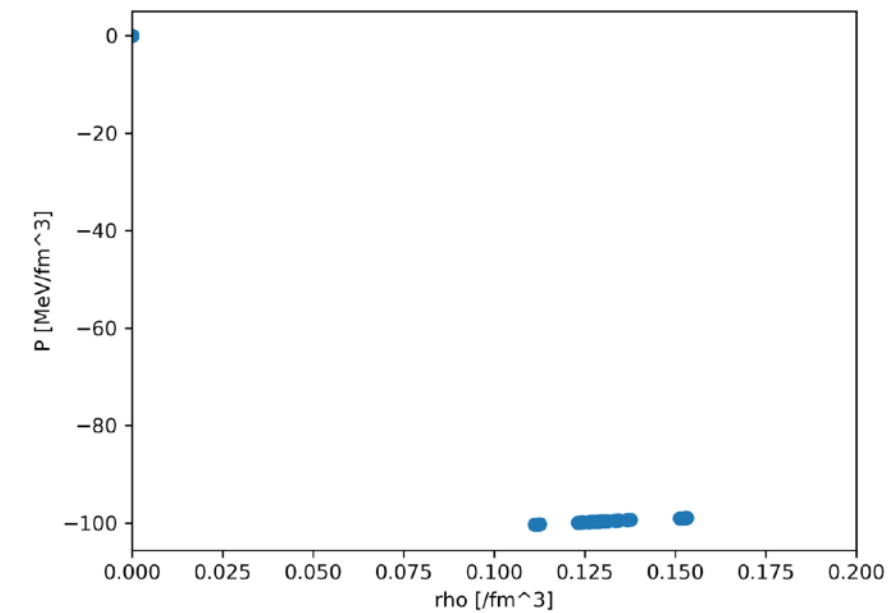
$f_w=5.0 \text{ fm}^2$



$f_w=5.42 \text{ fm}^2$

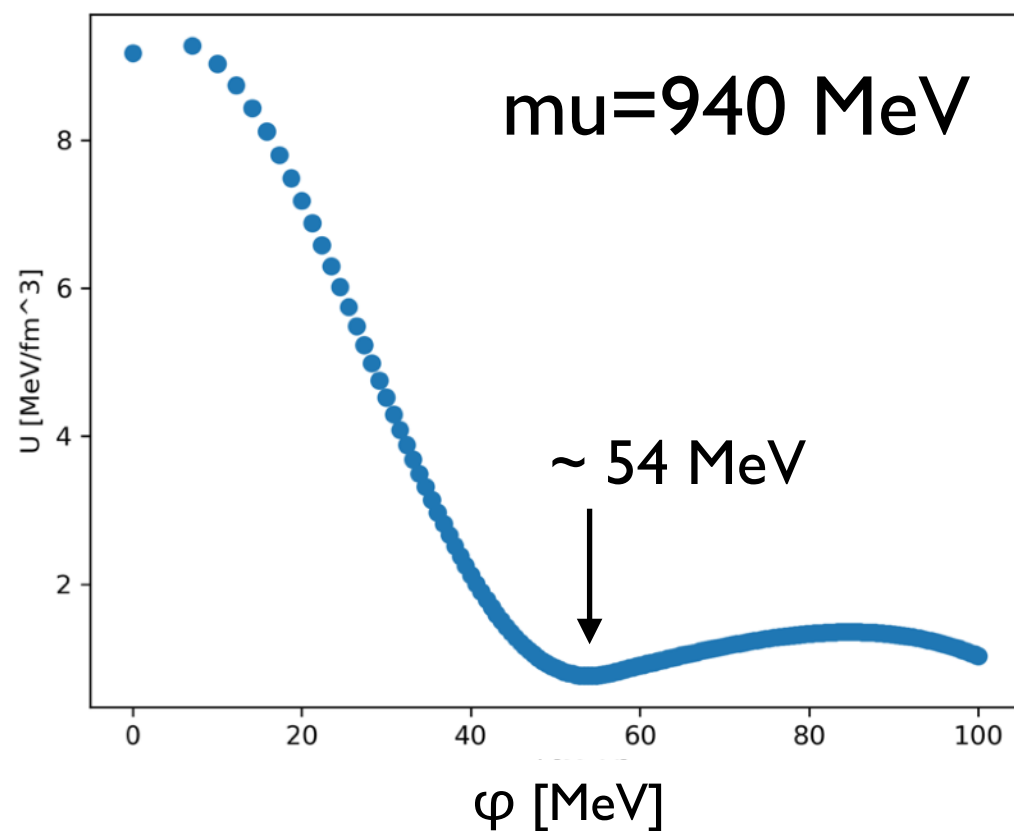


$f_w=6.0 \text{ fm}^2$

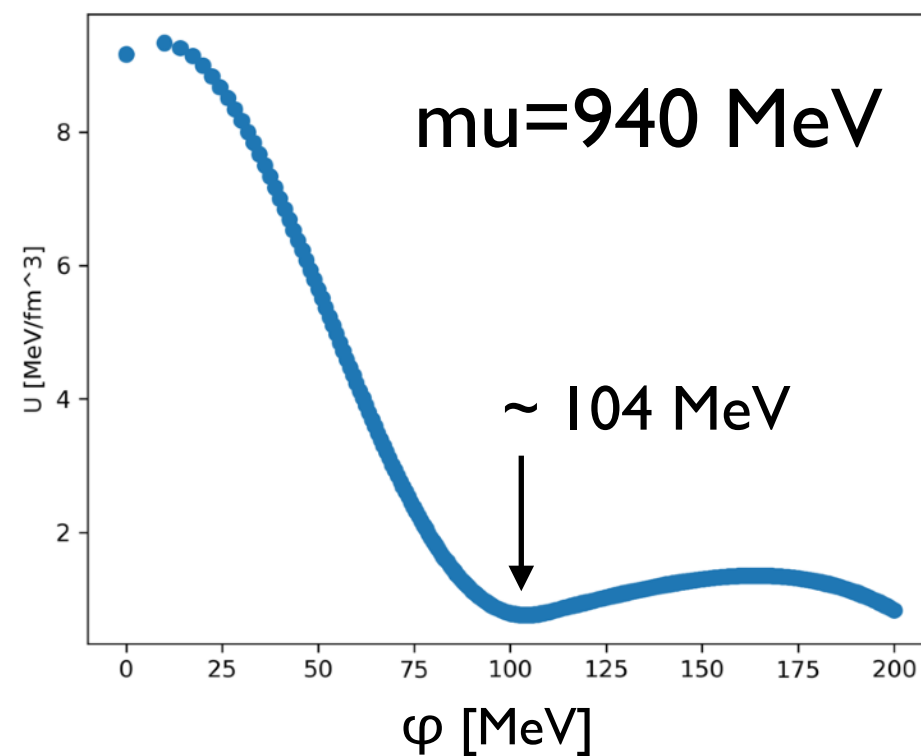


Test Results (5) for SNM

$$g_s = 16$$

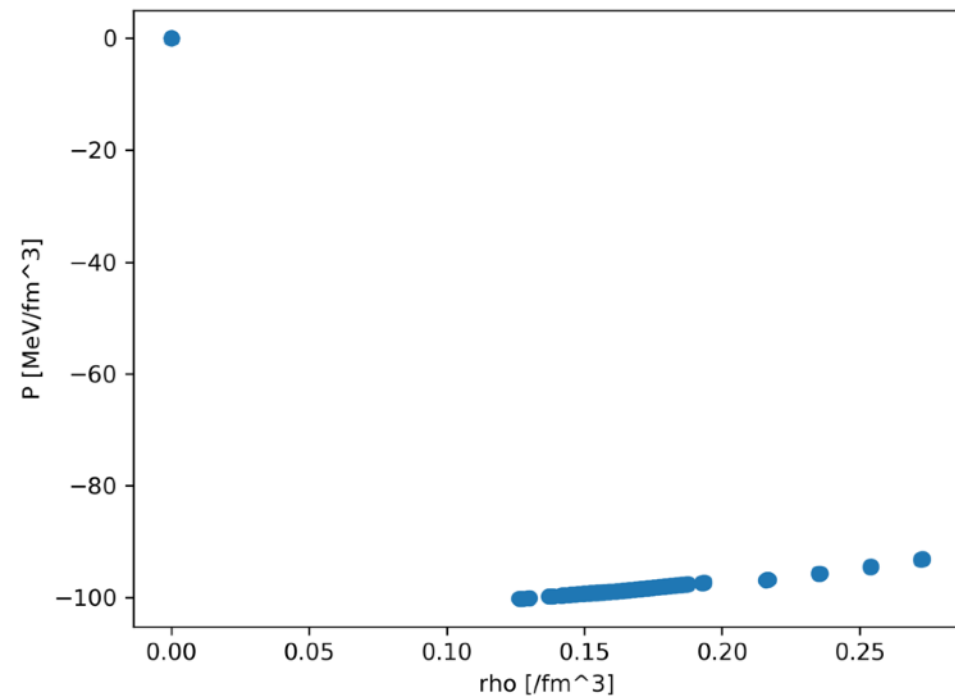
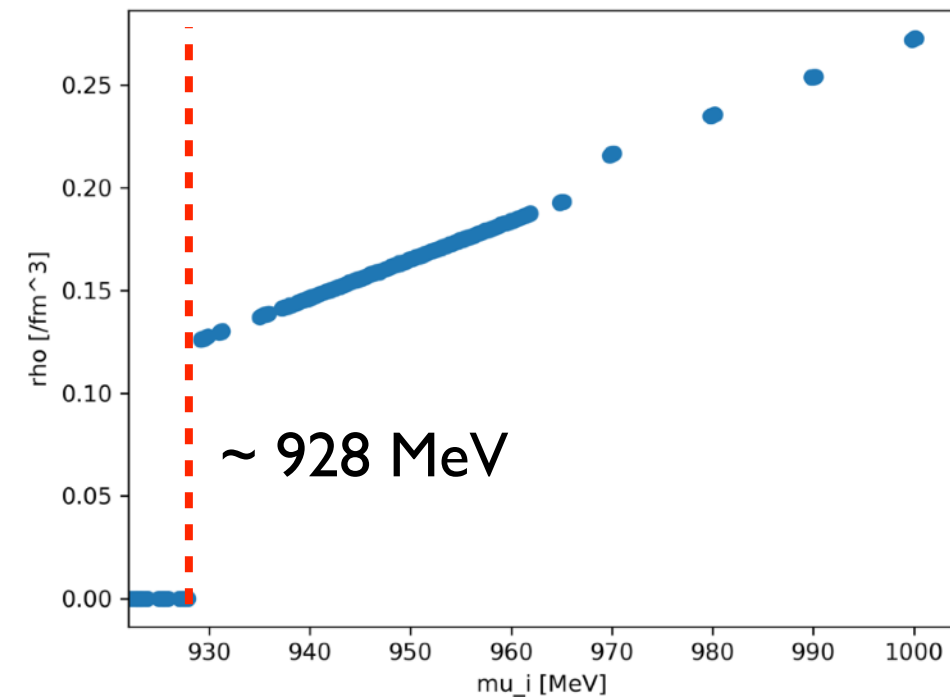


$$f_w = 10.33 \text{ fm}^2 \longrightarrow g_s = 8.28$$

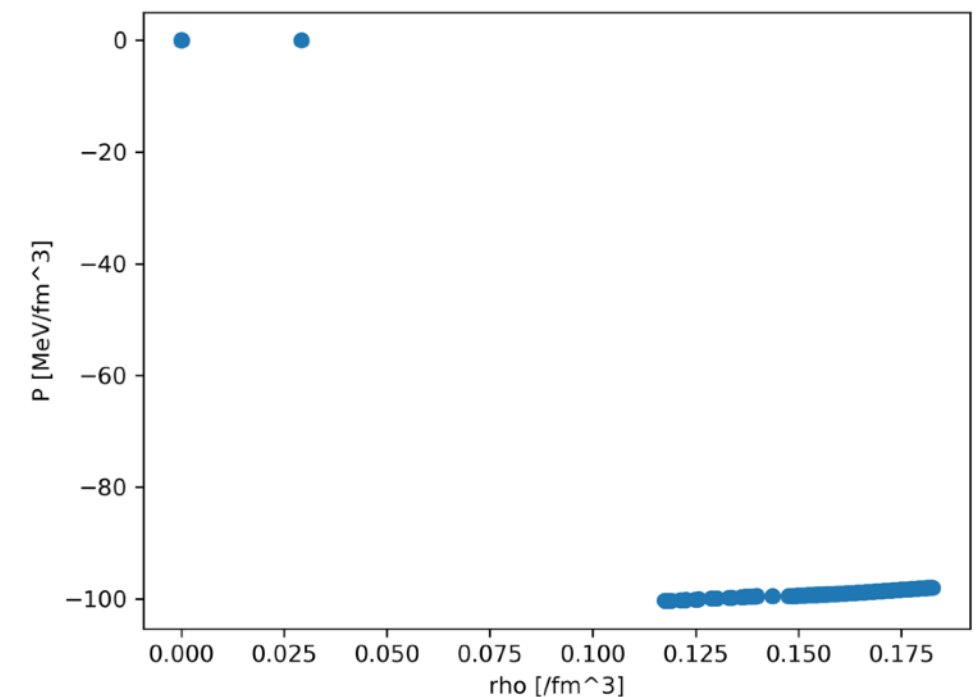
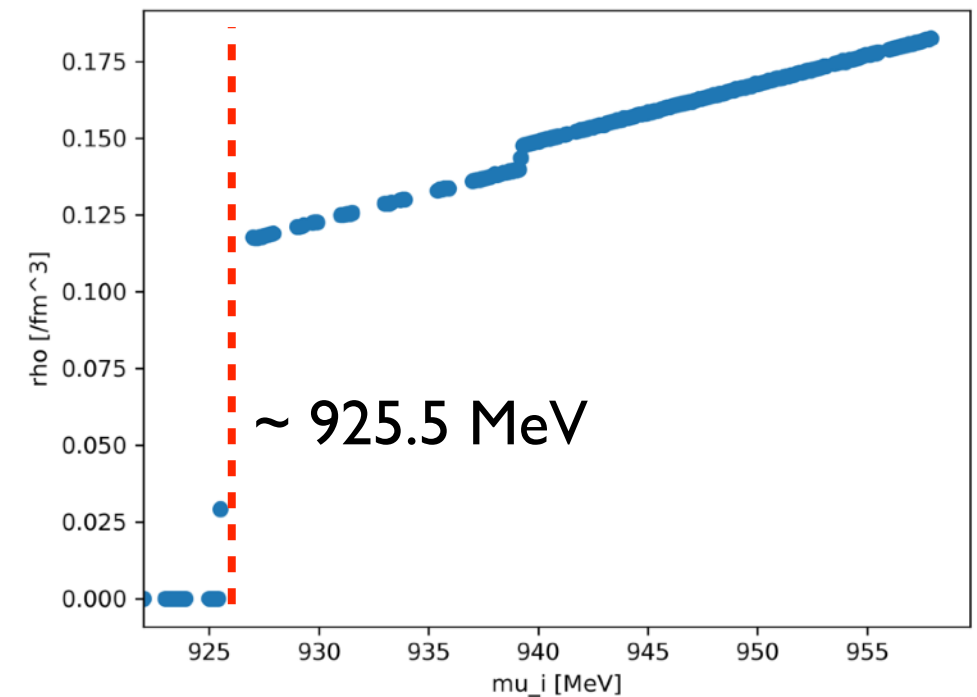


Test Results (6) for SNM

$$g_s = 16$$



$$f_w = 10.33 \text{ fm}^2 \longrightarrow g_s = 8.28$$



Test Results (7) for SNM

Initial condition

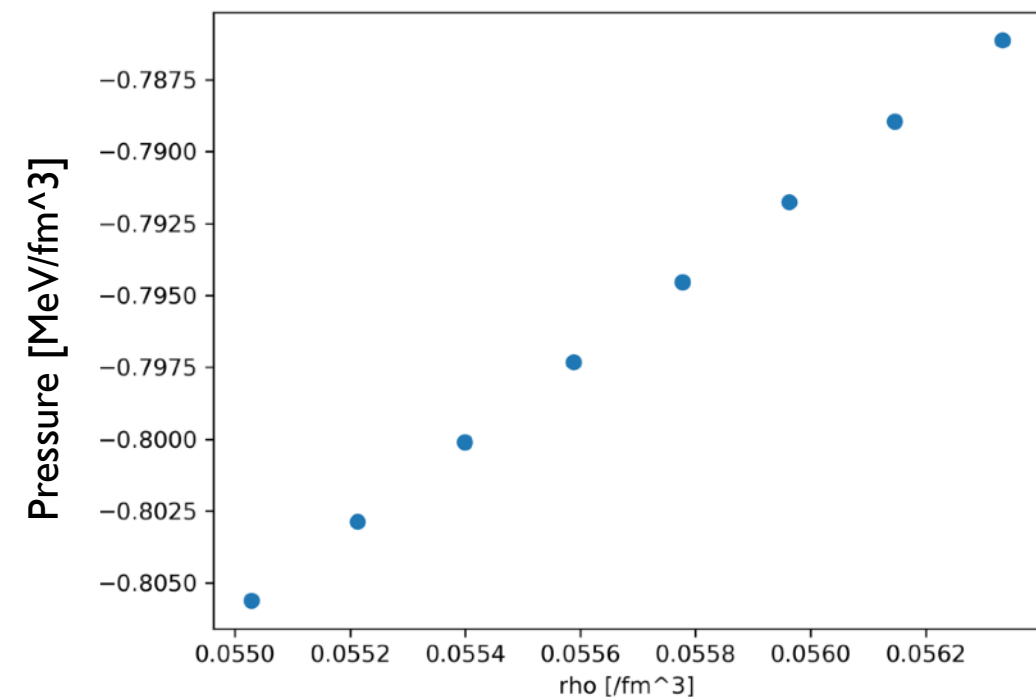
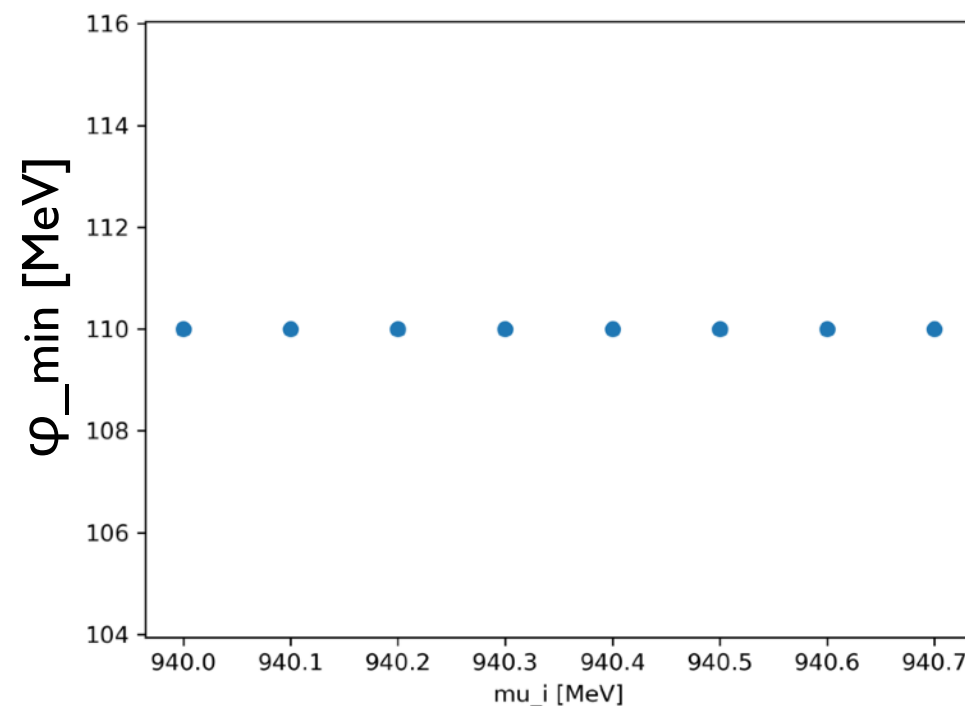
$$U_{\Lambda}^{\phi} = \frac{1}{3}\tilde{a}\phi^3 + \frac{1}{4}\tilde{b}\phi^4$$

$A \rightarrow A/100, B \rightarrow B/100$

Liu et al., PHYSICAL REVIEW C **65** 045201

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Summary and future plans

1. FRG calculation is still on-going.
2. We need to optimize model parameters.
 - Number of calculation : $\sim 10^6$, running time of a single job : 10~60 min.
 - Bayesian optimization and/or deep learning approach
3. Future plans
 - Nuclear matter equation of state $\longrightarrow$ NS mass, radius, tidal deformability
 - Application to Finite Nuclei
 - Application to Heavy Ion Collision w/ DJBUU

Thank you for your attention.

