

Mass-varying gauge boson that couples to the dark energy field

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Review on gauged quintessence

- Quintessence gauged under U(1) (quintessence + Dark gauge boson)

$$\mathcal{L} = -|D_\mu \Phi|^2 - V_0(\Phi) - \frac{1}{4} \mathbb{X}_{\mu\nu} \mathbb{X}^{\mu\nu}$$

$$\Phi = \frac{1}{\sqrt{2}} \phi e^{i\eta}$$

$$\mathbb{X}_{\mu\nu} = \partial_\mu \mathbb{X}_\nu - \partial_\nu \mathbb{X}_\mu$$

$$D_\mu = \partial_\mu + ig_X \mathbb{X}_\mu$$

- In Unitary gauge, ($\eta = 0$, $X_\mu = \mathbb{X}_\mu + \frac{1}{g_X} \partial_\mu \eta$)

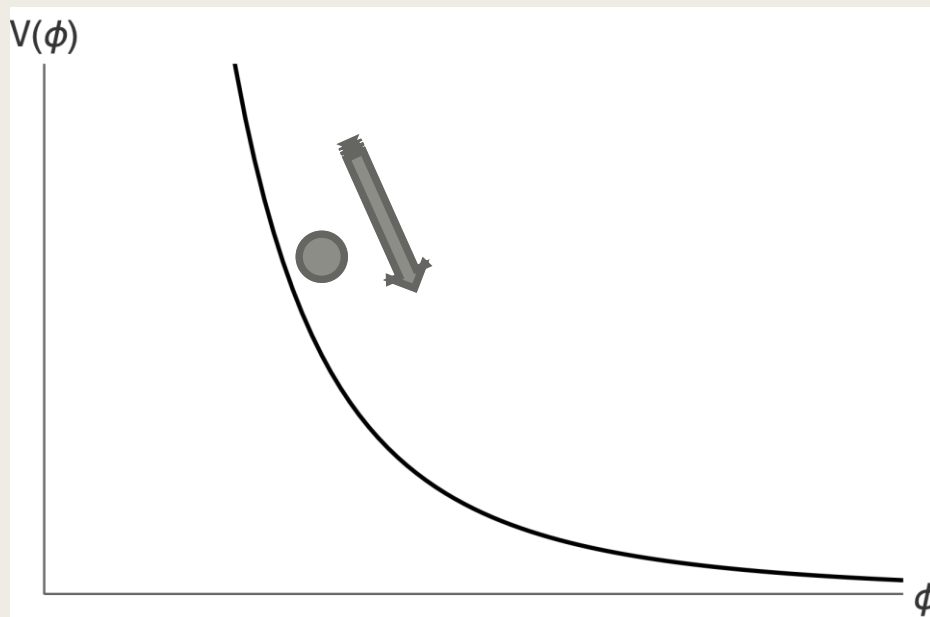
$$\mathcal{L} = \underbrace{-\frac{1}{2}(\partial_\mu \phi)^2 - V_0(\phi)}_{\text{Scalar (quintessence)}} - \underbrace{\frac{1}{4} X_{\mu\nu} X^{\mu\nu} - \frac{1}{2} g_X^2 \phi^2 X_\mu X^\mu}_{\text{Massive vector boson (dark gauge boson)}}$$

Ratra peebles potential (Inverse power)

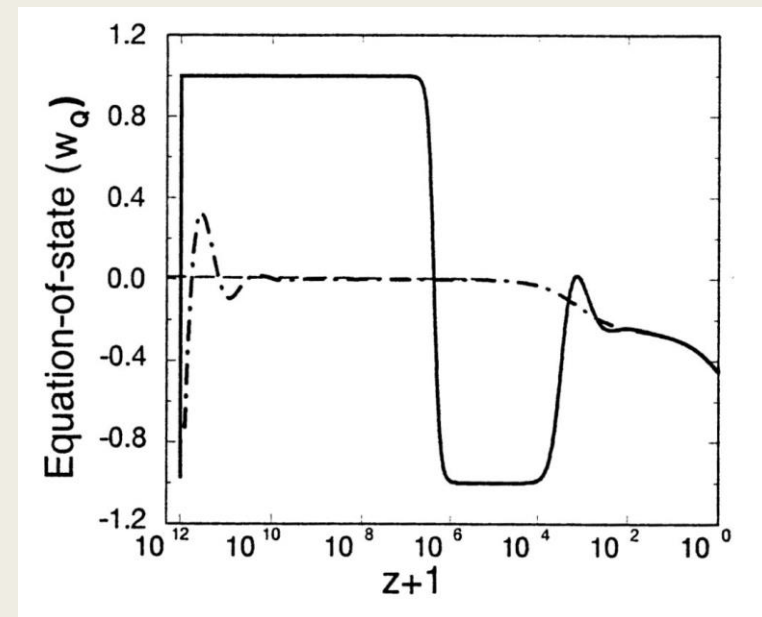
- Ratra-Peebles potential [Ratra and Peebles Phys. Rev. D 37\(1988\)3406](#)

$$V_0(\phi) = \frac{M^{\alpha+4}}{\phi^\alpha}$$

- The ϕ field runs toward to the larger value



- Tracking solution



Gauge Potential

$$\mathcal{L} = -\frac{1}{2}(\partial_\mu\phi)^2 - V_0(\phi) - \frac{1}{4}X_{\mu\nu}X^{\mu\nu} - \boxed{\frac{1}{2}g_X^2\phi^2 X_\mu X^\mu}$$

$\equiv V_{\text{gauge}}$ 1. Give mass to X
2. Additional potential for ϕ

- Gauge potential affects the dynamics of ϕ and X .

$$\ddot{\phi} + 3H\dot{\phi} + \frac{\partial V_0}{\partial\phi} + g_X^2 X_\mu X^\mu \phi = 0,$$

$$\partial_\mu X^{\mu\nu} + 3HX^{0\nu} - g_X^2\phi^2 X^\nu = 0$$

Originated from V_{gauge}

Production of coherent vector boson



Homogeneous in the observable universe

How the quintessence affects the relic vector boson density?

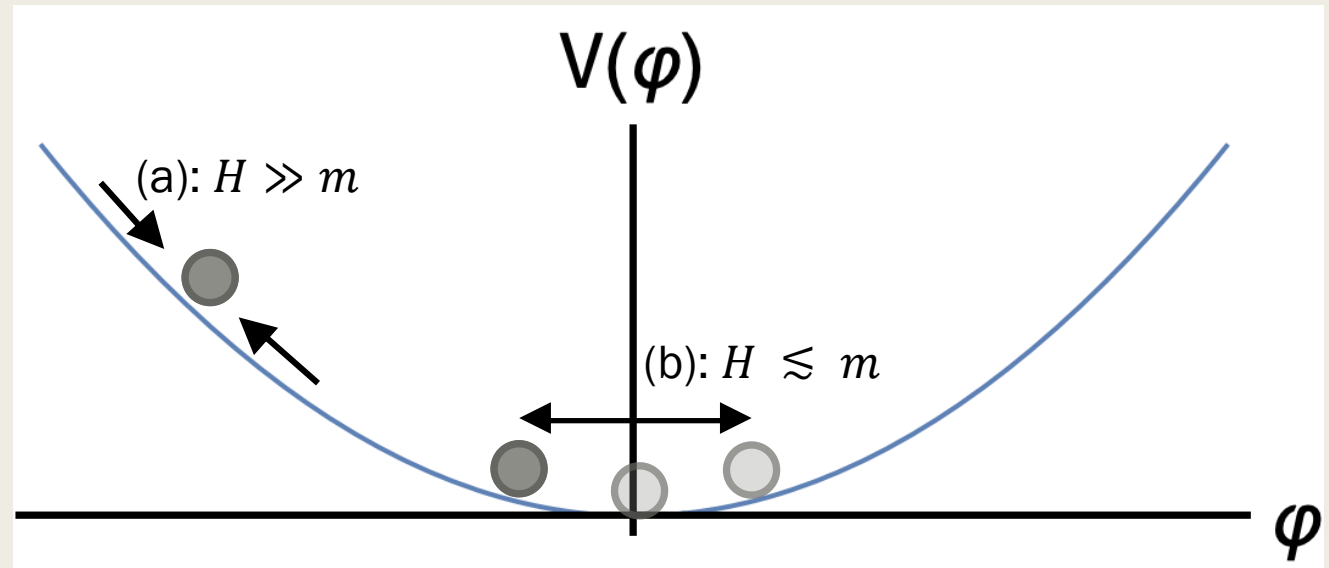
Misalignment Mechanism (Scalar)

Equation of motion

$$\ddot{\varphi} + 3H\dot{\varphi} + m^2\varphi = 0$$

Energy density

$$\rho_{\varphi} = \frac{1}{2}\dot{\varphi}^2 + \frac{1}{2}m^2\varphi^2$$



- (a): $H \gg m$ The field is initially frozen due to Hubble friction.

$$\varphi \sim \text{constant} \quad \rho_{\varphi} \sim \text{constant}$$

- (b): $H \lesssim m$ A coherent oscillation begins as the Hubble parameter drops below the mass.

$$\varphi \propto a^{-3/2} \quad \rho_{\varphi} \propto a^{-3} \quad (\text{Same as the non-relativistic matter})$$

Misalignment Mechanism (Vector)

Equation of motion

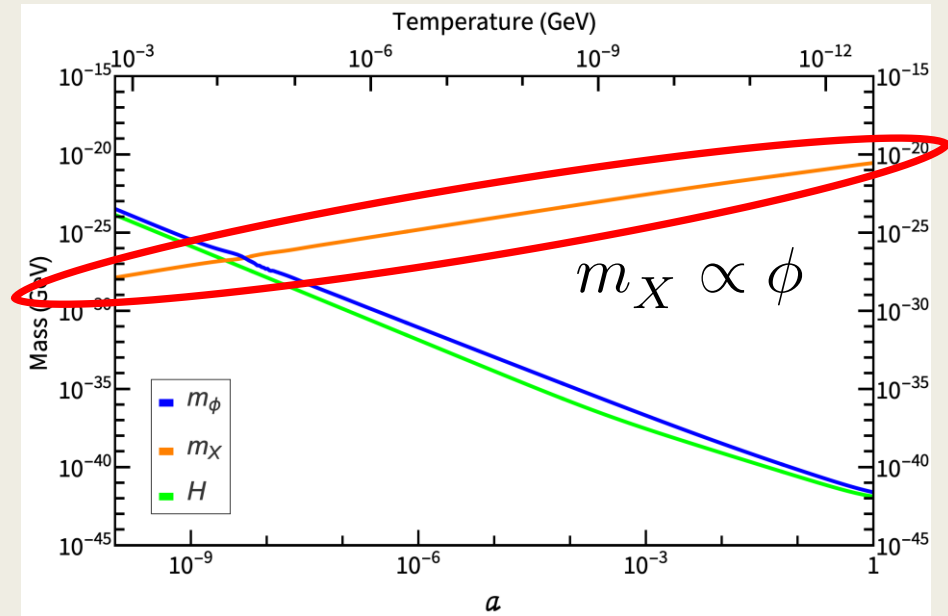
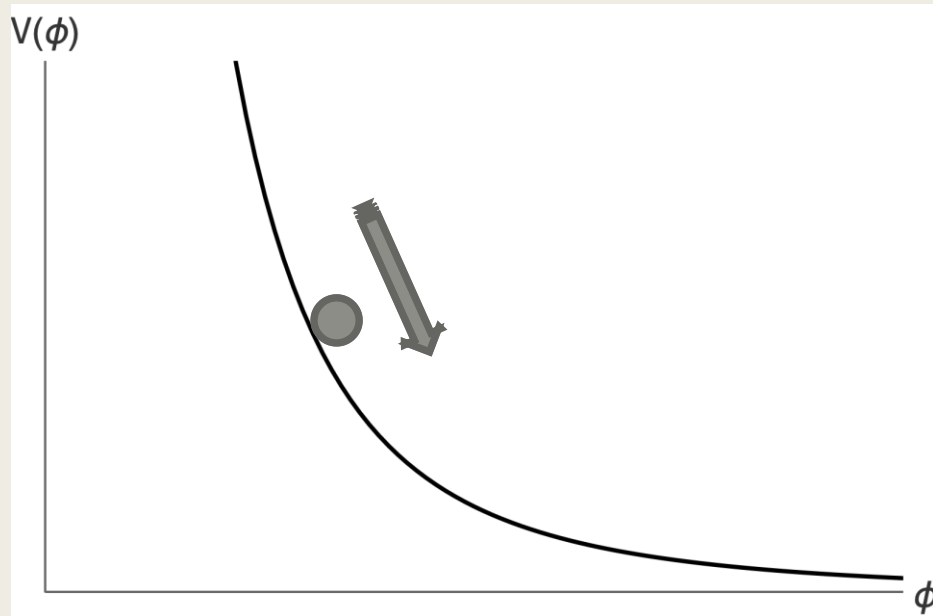
$$\ddot{X} + H\dot{X} + m_X^2 X = 0$$

Energy density

$$\rho_X = \frac{1}{2a^2} \left(\dot{X}^2 + m_X^2 X^2 \right)$$

- The field is initially frozen due to Hubble friction. But energy decreases as $\rho_X \propto a^{-2}$
- ❖ This gives $\sim e^{-120}$ suppression during inflation.
- Relic vector density from the misalignment mechanism is negligibly small.

Vector misalignment in the gauged quintessence



- Quintessence (ϕ) field rolls down the potential.
- The mass of the vector boson increases

$$\rho_X = \frac{1}{2a^2} \left(\dot{X}^2 + m_X^2 X^2 \right)$$

$$m_X = g_X \phi$$

Energy density can be amplified!

Coupled dynamics Revisited

$$\ddot{\phi} + 3H\dot{\phi} + \frac{\partial V_0}{\partial \phi} + g_X^2 X_\mu X^\mu \phi = 0,$$

$$\partial_\mu X^{\mu\nu} + 3H X^{0\nu} - g_X^2 \phi^2 X^\nu = 0$$

- Equation of motion of ϕ

$$\ddot{\phi} + 3H\dot{\phi} + \frac{\partial V_{\text{eff}}}{\partial \phi} = 0$$

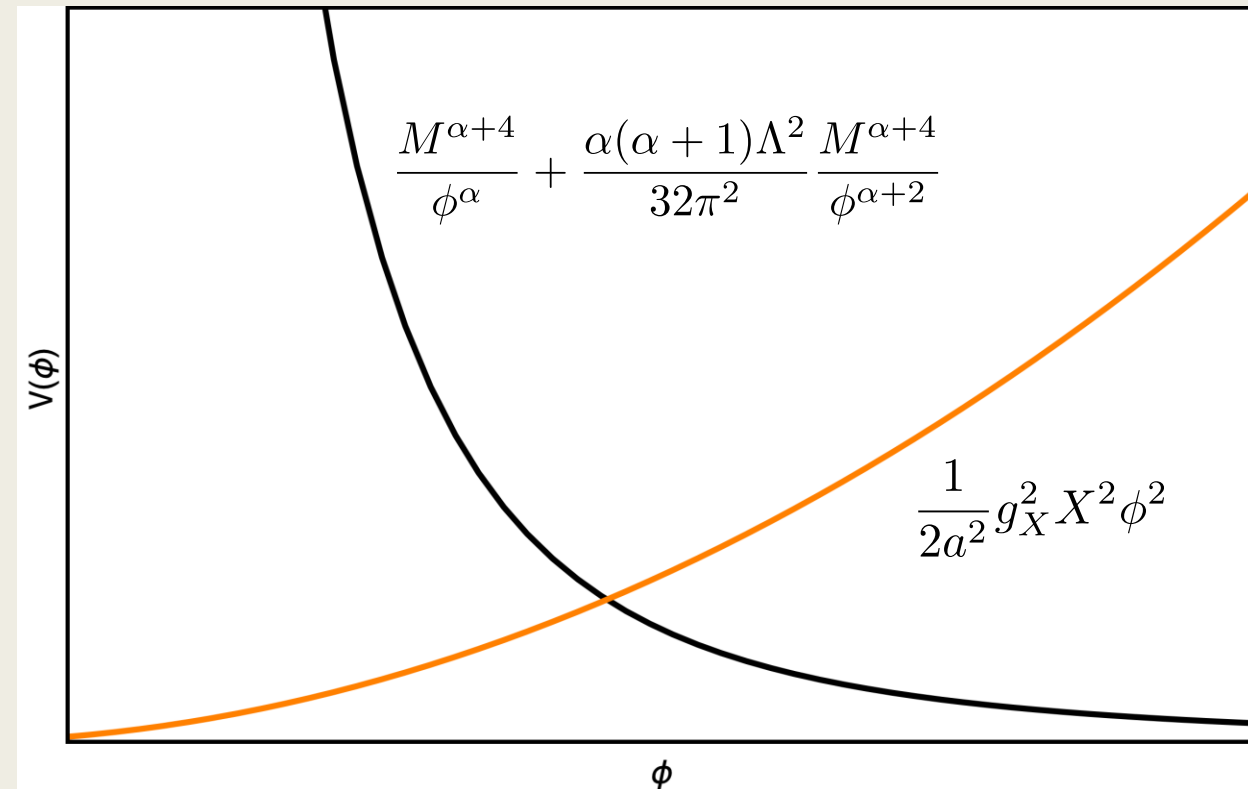
$$V_{\text{eff}} = V_0 + V_{1\text{-loop}} + V_{\text{gauge}}$$

- Equation of motion of X

$$\ddot{X} + H\dot{X} + g_X^2 \phi^2 X = 0$$



Similar to the damped oscillator with time varying frequency ($g_X \phi = m_X$)



X dynamics

* $H \gg m_X$

w_b : background equation of state

$$\bar{X} \equiv X/a, \quad \bar{X} + 3H\bar{X} + \frac{1 - 3w_b}{2} H^2 \bar{X} = 0$$

■ Kinetic energy dominated: $\frac{1}{2a^2} \dot{X}^2 \gg \frac{1}{2a^2} m_X^2 X^2$

$$X \propto a^{-1}, \quad \rho_X \propto H^2 a^{3w_b - 1}$$

■ Potential energy dominated: $\frac{1}{2a^2} \dot{X}^2 \ll \frac{1}{2a^2} m_X^2 X^2$

$$X \sim \text{constant}, \quad \rho_X \propto \phi^2 a^{-2}$$

* $H \ll m_X$

- Coherent Oscillation
(WKB approximation)

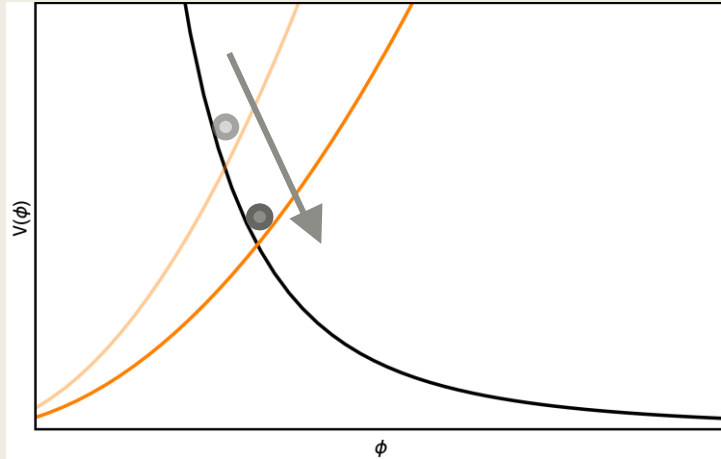
$$X \propto \frac{1}{\sqrt{\phi a^3}}$$

$$\rho_X \propto \phi a^{-3} (\propto m_X a^{-3})$$

Non-relativistic. But mass-varying.

Quintessence dynamics

$$\ddot{\phi} + 3H\dot{\phi} + \frac{\partial V_{\text{eff}}}{\partial \phi} = 0$$

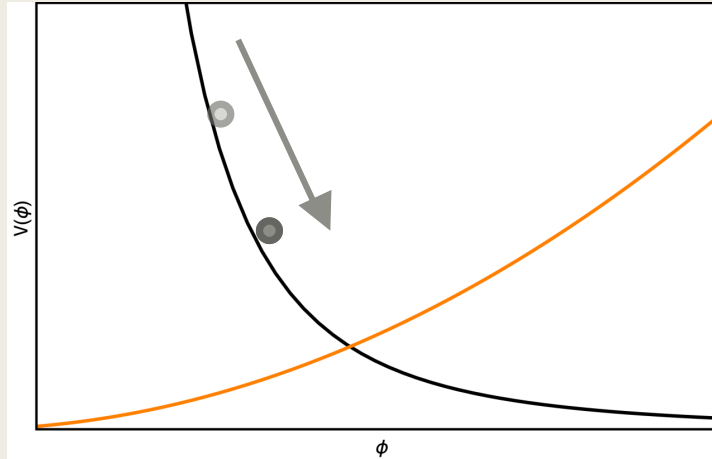


$$* \quad m_{\phi} \gtrsim H, \quad \frac{\partial^2 V_{\text{gauge}}}{\partial \phi^2} \gtrsim H^2$$

ϕ follows the minimum

$$\phi \propto a^{\frac{2}{\alpha+4}} \quad (H \gg m_X)$$

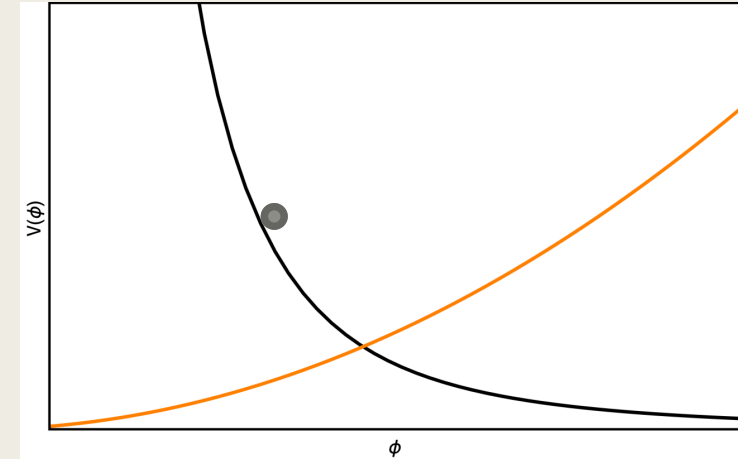
$$\phi \propto a^{\frac{2}{\alpha+3}} \quad (H \ll m_X)$$



$$* \quad m_{\phi} \sim H, \quad \frac{\partial^2 V_{\text{gauge}}}{\partial \phi^2} \ll H^2$$

ϕ follows the tracking solution of Ratra-Peebles potential

$$\phi \propto a^{\frac{3(1+w_b)}{\alpha+1}}$$

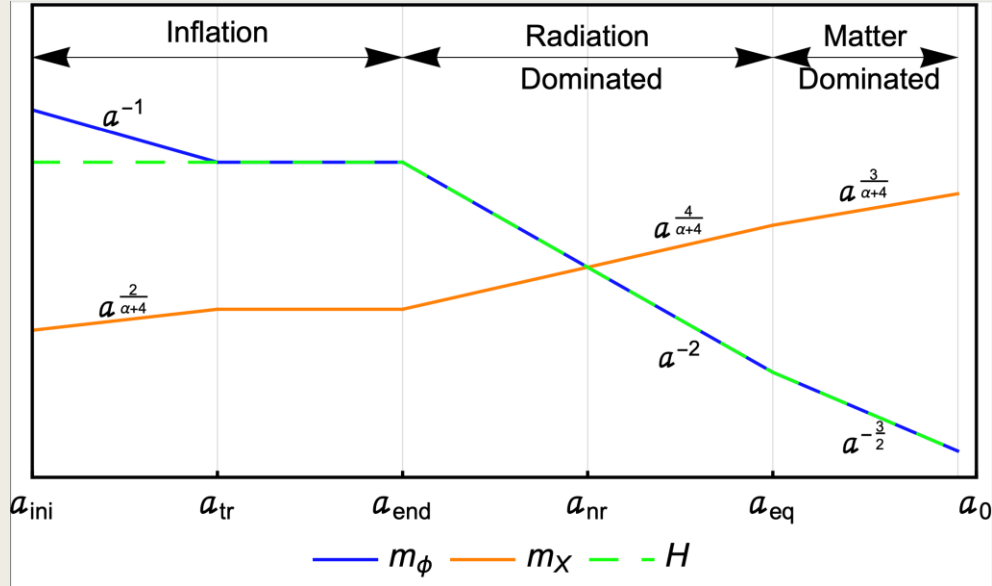


$$* \quad m_{\phi} \ll H$$

$\dot{\phi} \ll H\phi$: ϕ is frozen

$$\phi \sim \text{constant}$$

Viable Scenario



	\bar{X}	$\phi \propto m_X$	ρ_X
$a_{\text{ini}} < a < a_{\text{tr}}$	$\propto a^{-1}$	$\propto a^{\frac{2}{\alpha+4}}$	$\propto a^{-\frac{2\alpha+4}{\alpha+4}}$
$a_{\text{tr}} < a < a_{\text{end}}$	$\propto a^{-1}$	$\approx \text{constant}$	$\propto a^{-2}$
$a_{\text{end}} < a < a_{\text{nr}}$	$\propto a^{-1}$	$\propto a^{\frac{4}{\alpha+4}}$	$\propto a^{-\frac{2\alpha}{\alpha+4}}$
$a_{\text{nr}} < a < a_{\text{eq}}$	$\propto a^{-\frac{3\alpha+16}{2\alpha+8}}$	$\propto a^{\frac{4}{\alpha+4}}$	$\propto a^{-\frac{3\alpha+8}{\alpha+4}}$
$a_{\text{eq}} < a < a_0$	$\propto a^{-\frac{3\alpha+15}{2\alpha+8}}$	$\propto a^{\frac{3}{\alpha+4}}$	$\propto a^{-\frac{3\alpha+9}{\alpha+4}}$

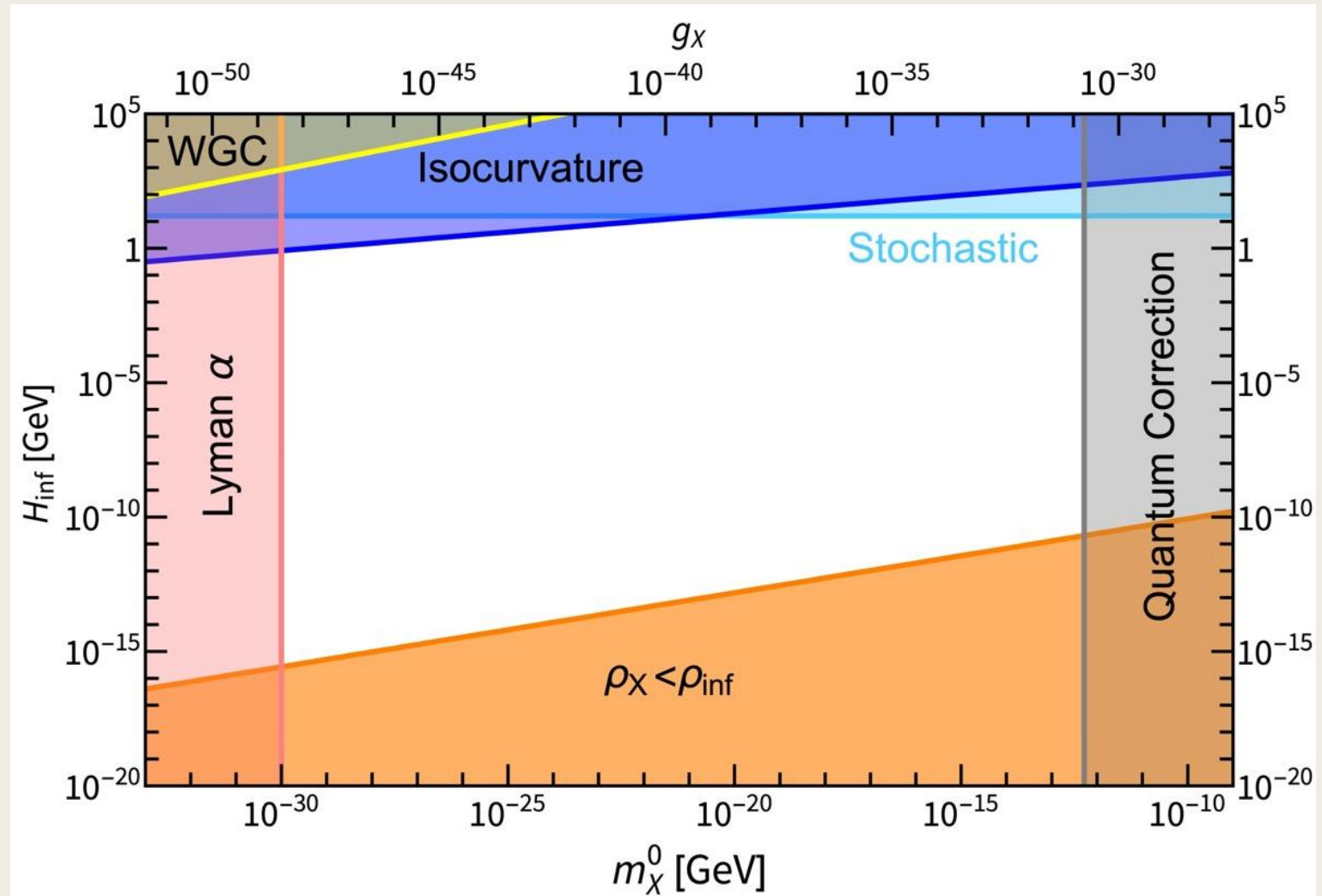
- This scenario requires initial potential energy domination of X $\frac{1}{2a^2}\dot{X}^2 \ll \frac{1}{2a^2}m_X^2 X^2$
- ϕ field joins the tracking solution during the inflation

Misalignment production in the gauged quintessence

$$\alpha = 1, M = 2.2 \times 10^{-6} \text{ GeV}$$

$$V_0 = \frac{M^{\alpha+4}}{\phi^\alpha}$$

$$\Omega_X h^2 \sim 0.12$$



- The mass-increasing effect allows to produce sufficient amount of vector boson density

Summary and remark

- Gauged quintessence: quintessence + $U(1)$ gauge symmetry (dark gauge boson)
 - The mass of dark gauge boson can increase during cosmic history

- Misalignment production of coherent vector boson is extremely suppressed
 - The mass-increasing effect can amplify the energy density of vector boson

So, a large amount of relic density can be produced (comparable to CDM)
- ❖ We expect the mass-increasing effect generically help the misalignment production of vector boson

Thank you!