

The Impact of 1D and 3D NLTE Effects on neutron-capture elements

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Maria Bergemann's "Stellar spectroscopy and stellar populations"

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What is NLTE?

- Non-local thermal equilibrium (Non-LTE)

What is NLTE?

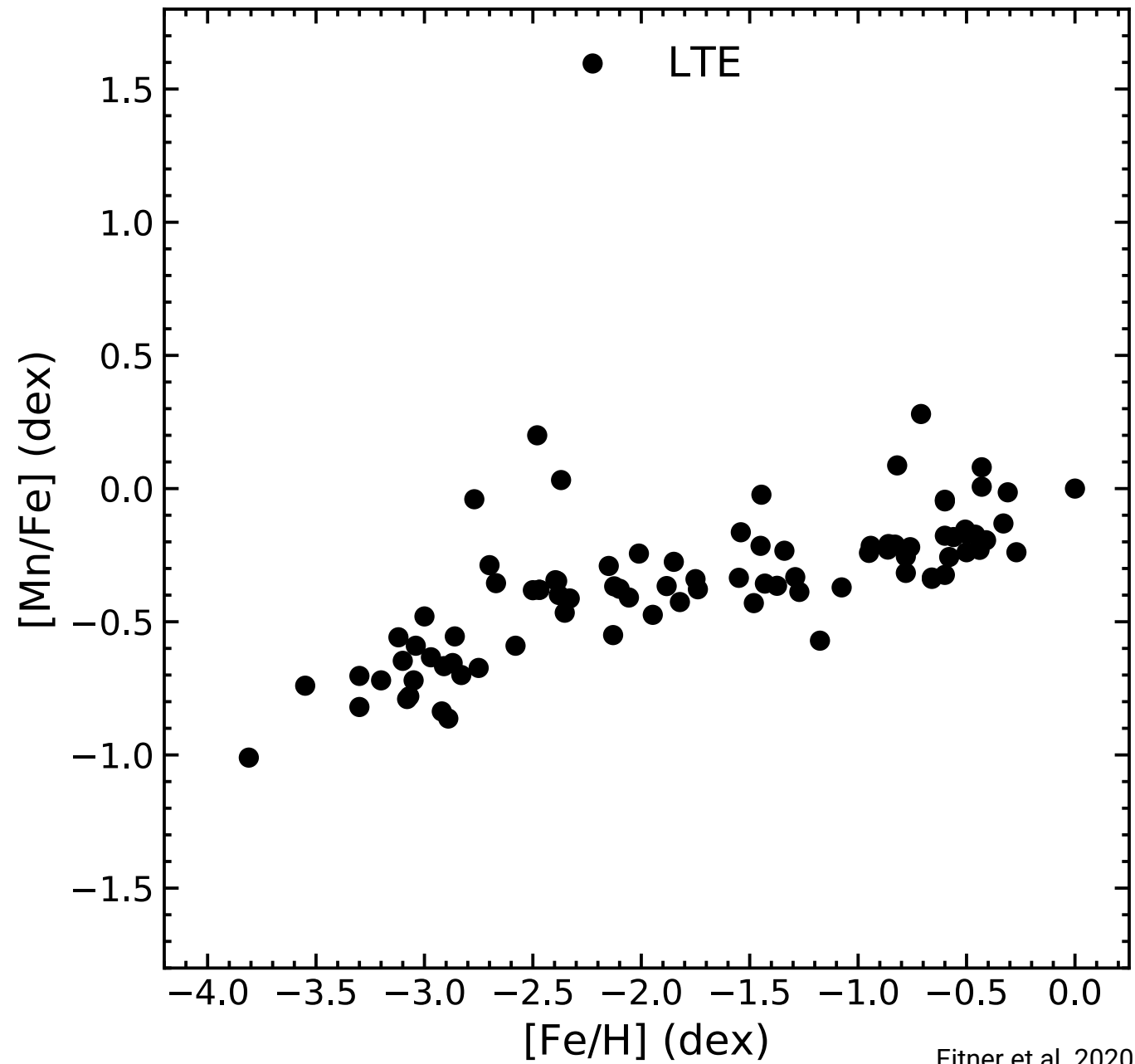
- Non-local thermal equilibrium (Non-LTE)
- No assumption of LTE in the radiative transfer
 - Energy level populations are not assumed to follow Boltzmann distribution
 - More accurate abundances?

What is NLTE?

- Non-local thermal equilibrium (Non-LTE)
- No assumption of LTE in the radiative transfer
 - Energy level populations are not assumed to follow Boltzmann distribution
 - More accurate abundances?
- Usually difficult
 - Takes longer time than LTE fitting
 - Need accurate atomic data, e.g.:
 - Energy levels
 - Bound-bound transitions, $\log gf$
 - Quantum mechanical data
 - E.g. radiative rates for collisions w/ e^- , H

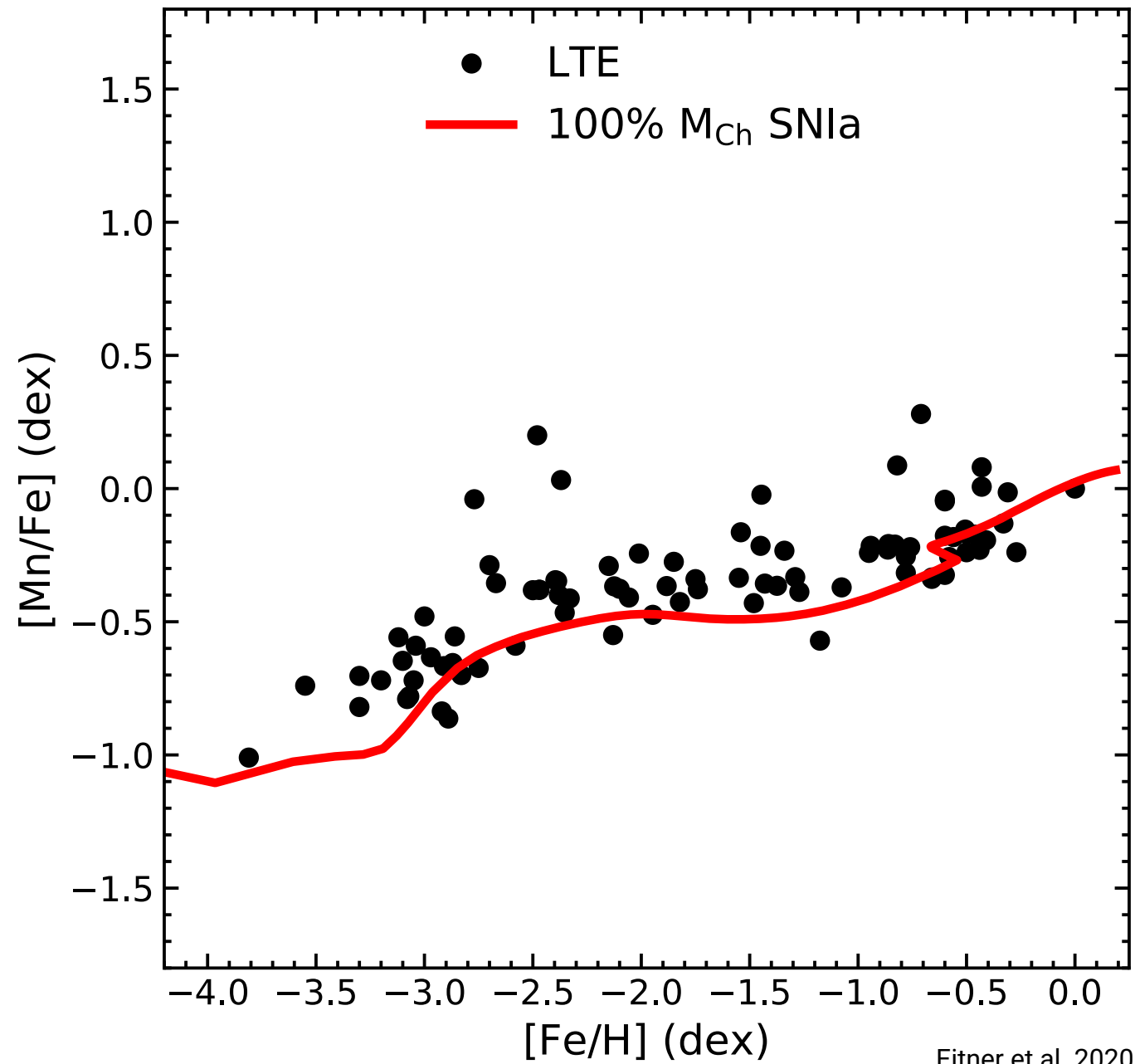
Why do NLTE?

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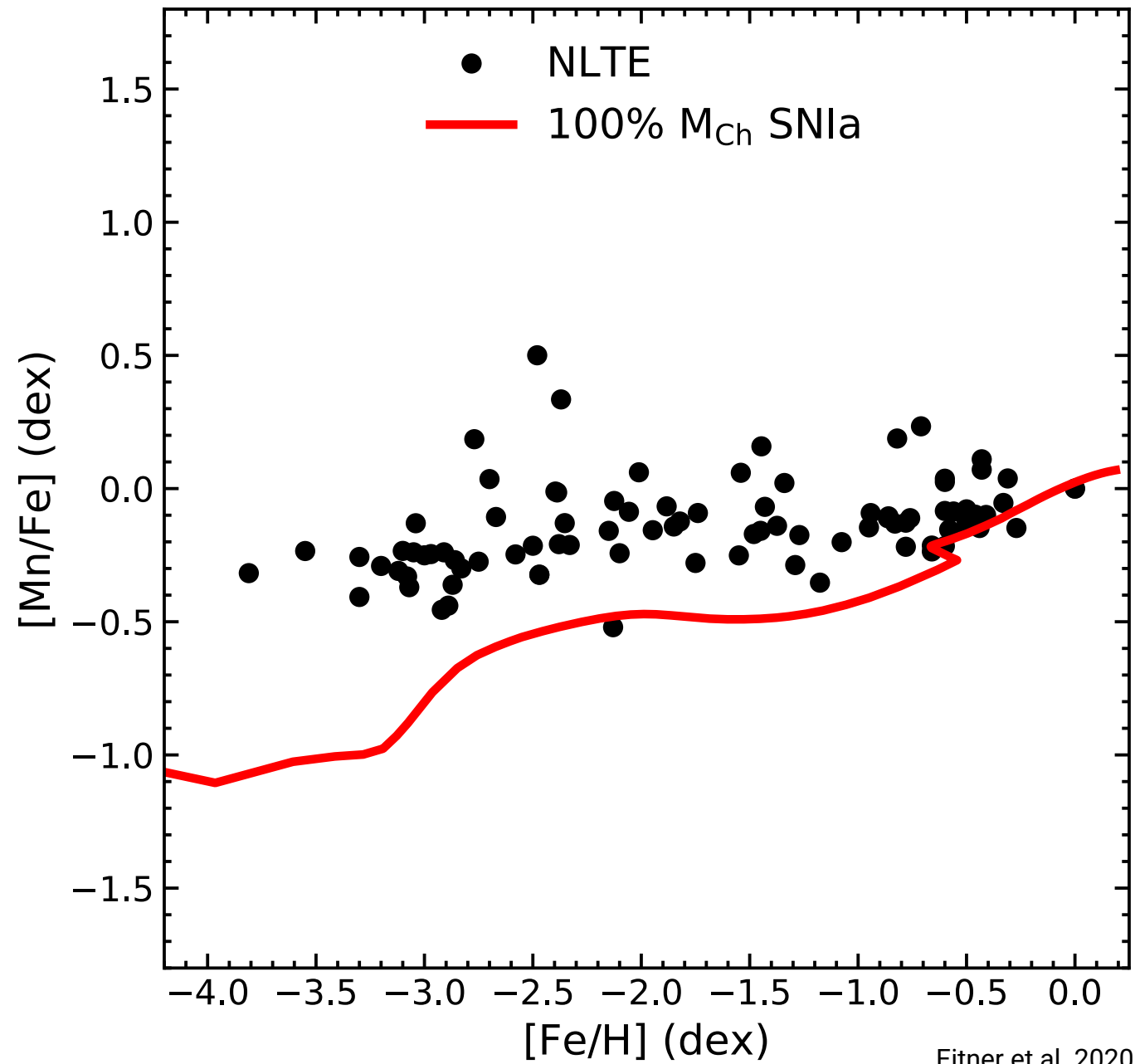
Eitner et al. 2020

Why do NLTE?



Eitner et al. 2020

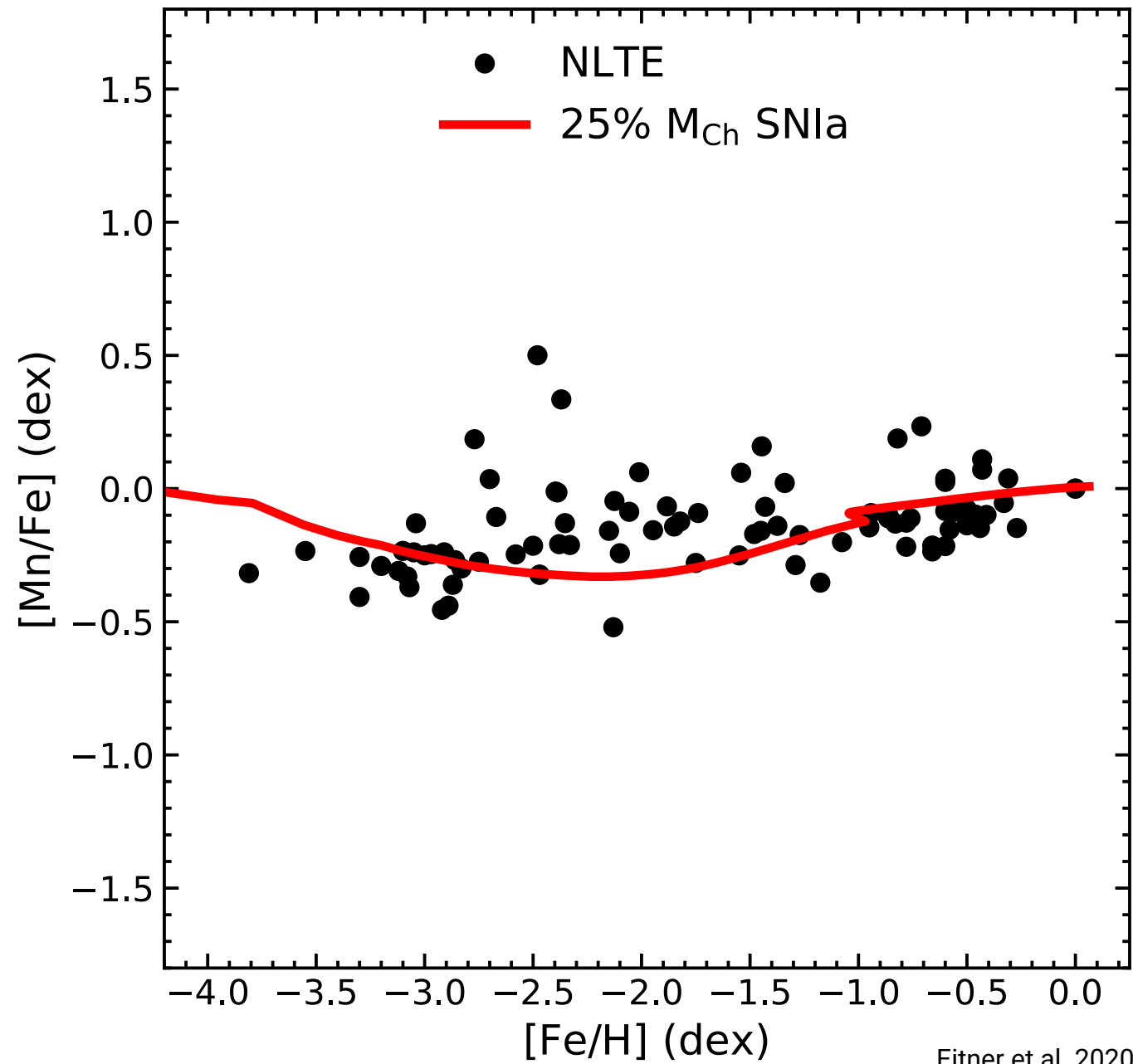
Why do NLTE?



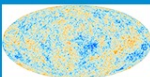
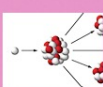
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Why do NLTE?

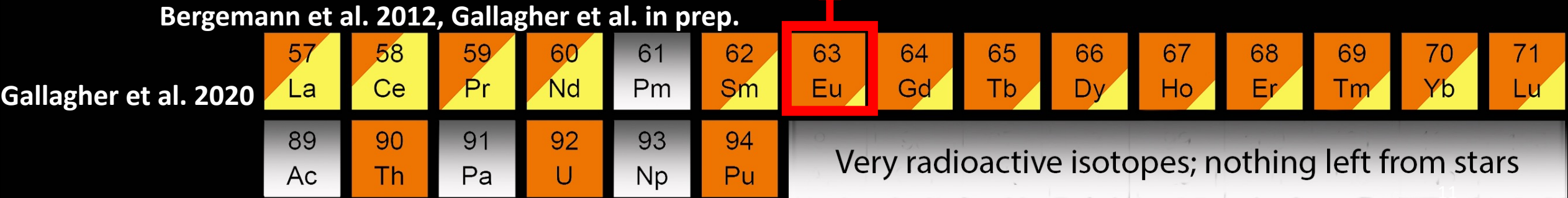
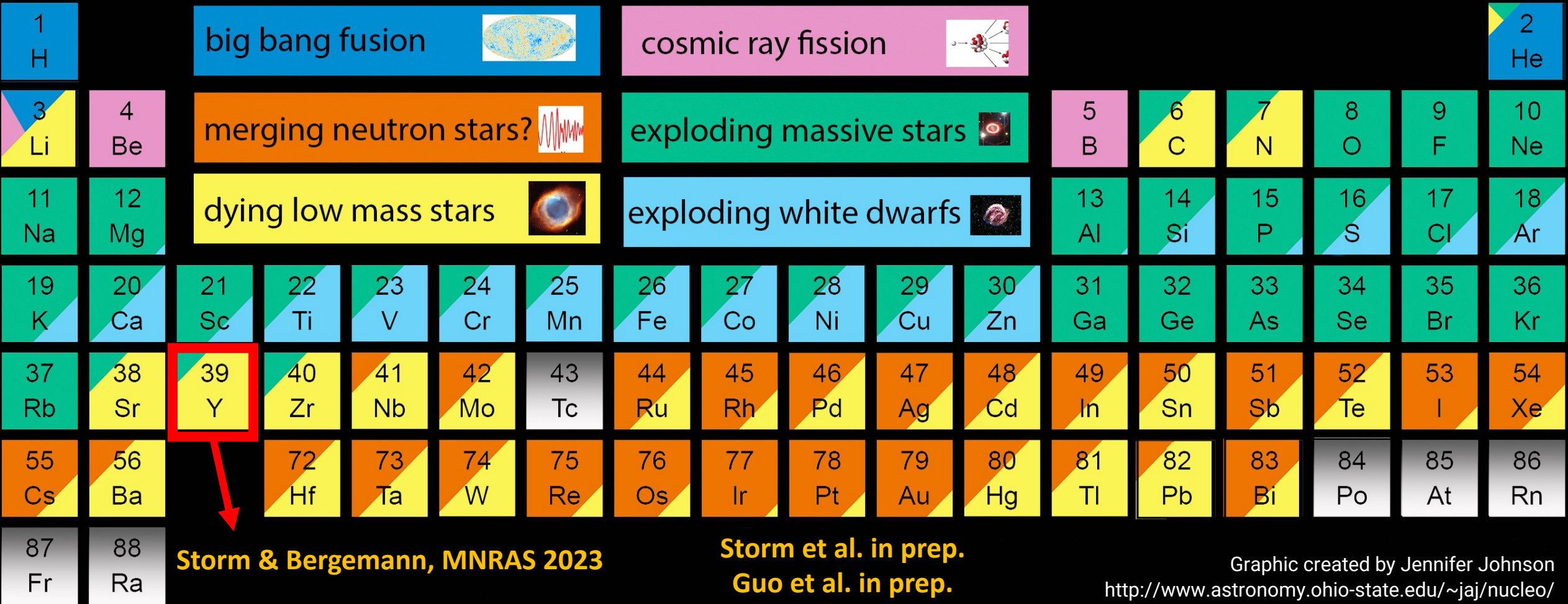
- NLTE (and 3D) may change our understanding of the GCE models



Eitner et al. 2020

1 H	big bang fusion 						cosmic ray fission 						2 He						
3 Li	4 Be	merging neutron stars? 						exploding massive stars 						5 B	6 C	7 N	8 O	9 F	10 Ne
11 Na	12 Mg	dying low mass stars 						exploding white dwarfs 						13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr		
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe		
55 Cs	56 Ba		72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn		
87 Fr	88 Ra																		
Graphic created by Jennifer Johnson http://www.astronomy.ohio-state.edu/~jaj/nucleo/																			
Bergemann et al. 2012, Gallagher et al. in prep.																			
Gallagher et al. 2020		57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu			
		89 Ac	90 Th	91 Pa	92 U	93 Np	94 Pu	Very radioactive isotopes; nothing left from stars											

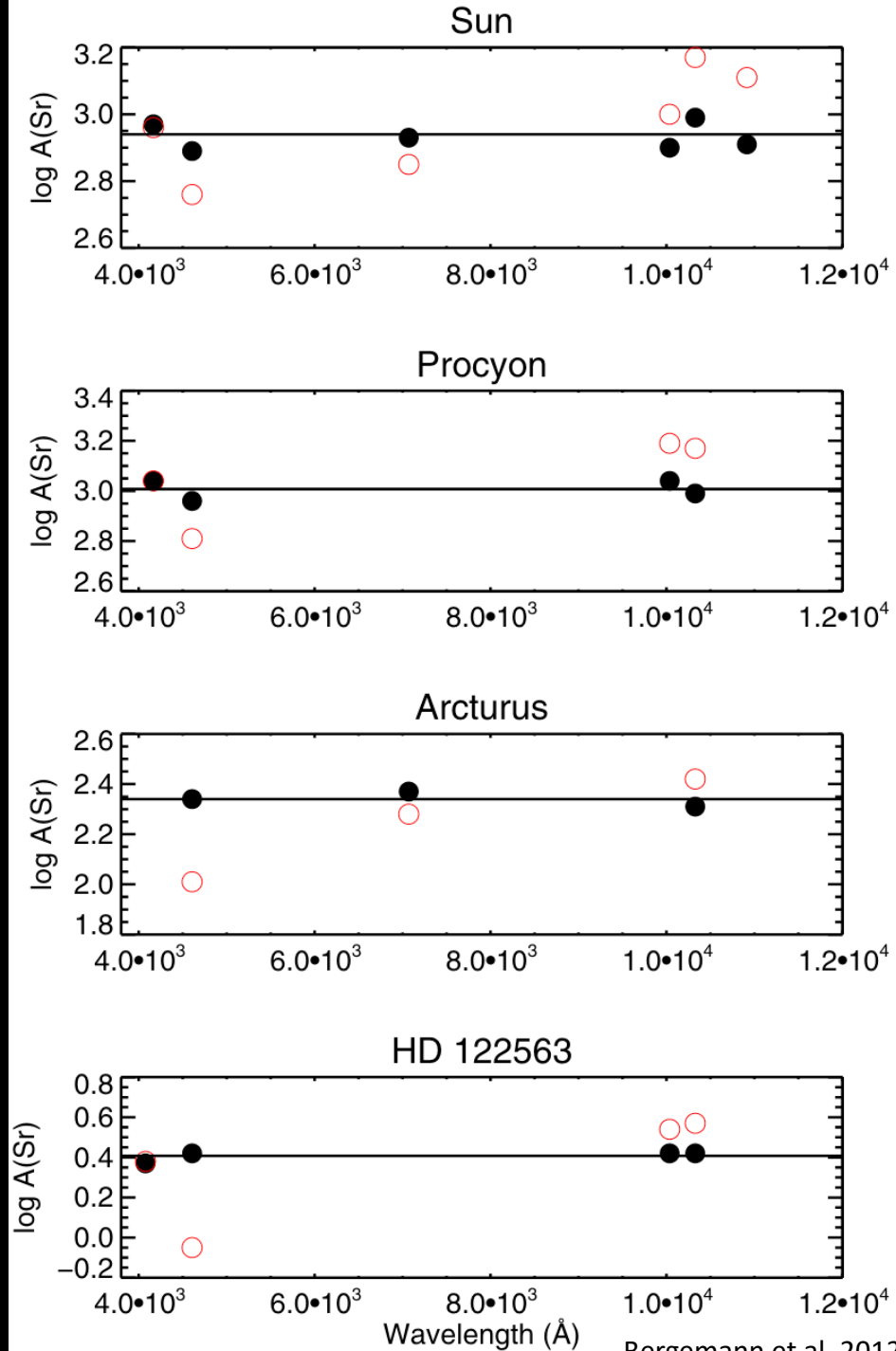
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Strontium, Barium

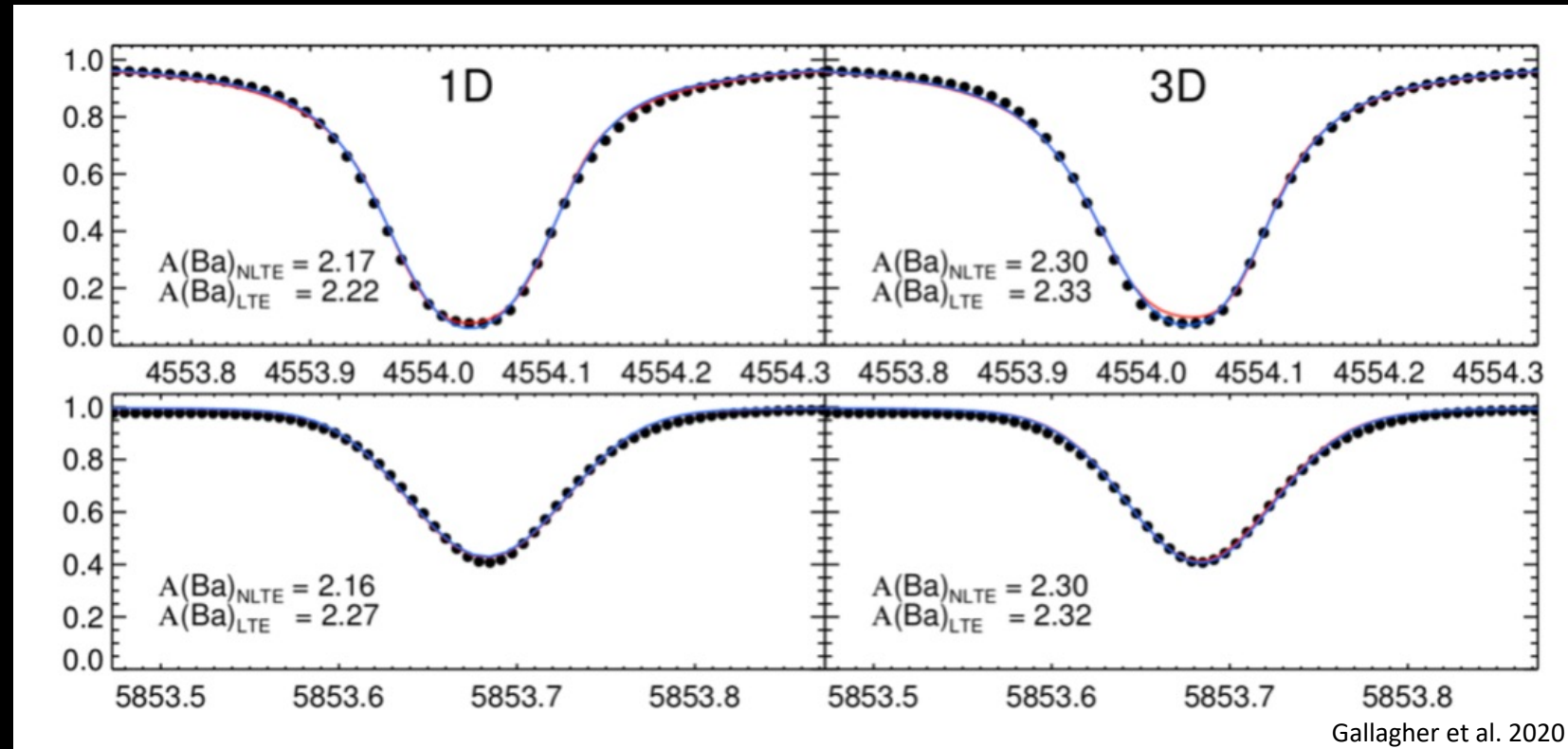
Strontium

- 1st s-process peak element
- Both Sr I and Sr II
- LTE: red circle, NLTE (1D): black
- LTE: up to 0.5 dex discrepancy



Barium

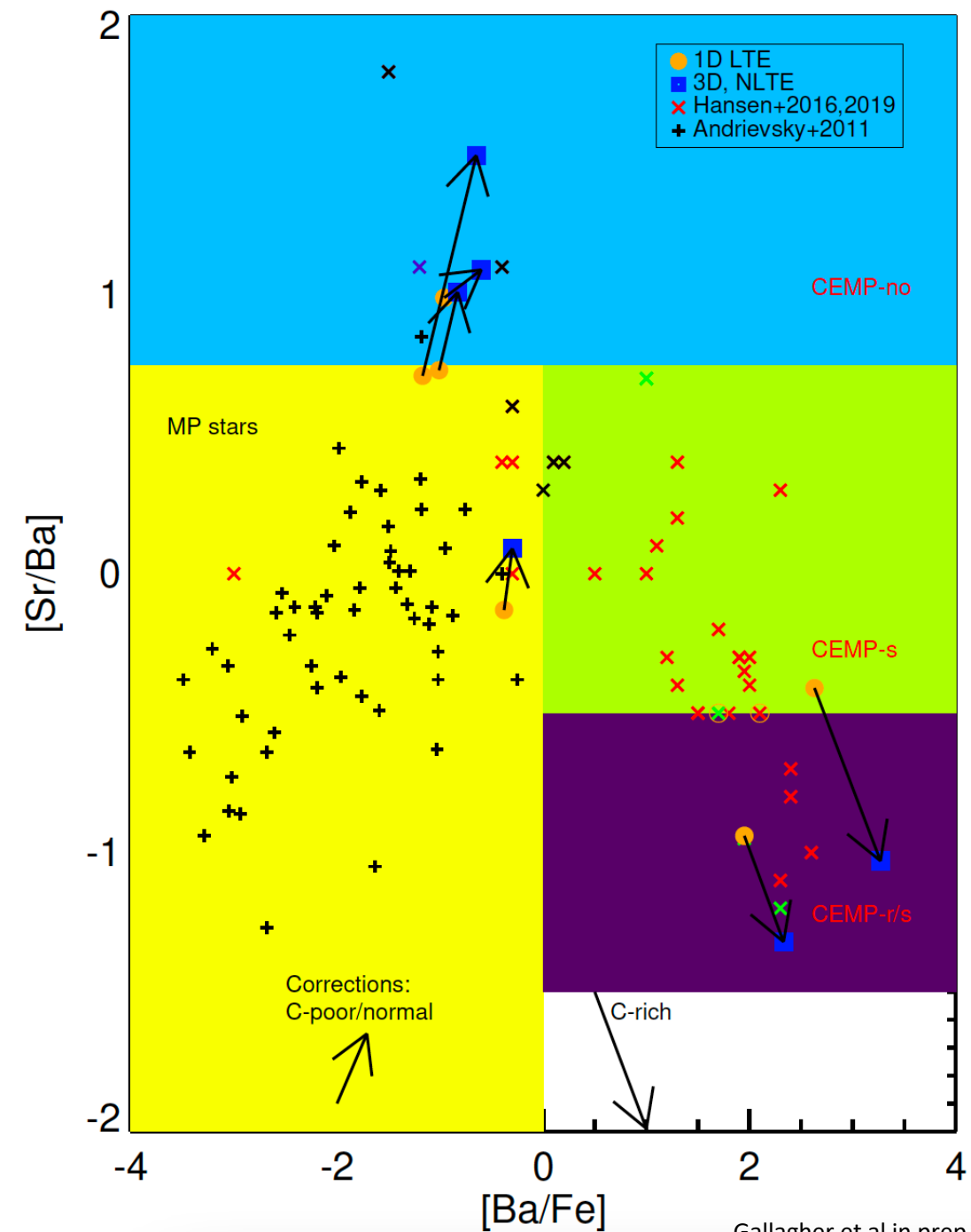
- 2nd s-process peak element
- Very sensitive to physical model type
- Sun:
 - 1D NLTE $A(\text{Ba}) = 2.11$
 - 3D NLTE $A(\text{Ba}) = 2.27$



Gallagher et al. 2020

Barium

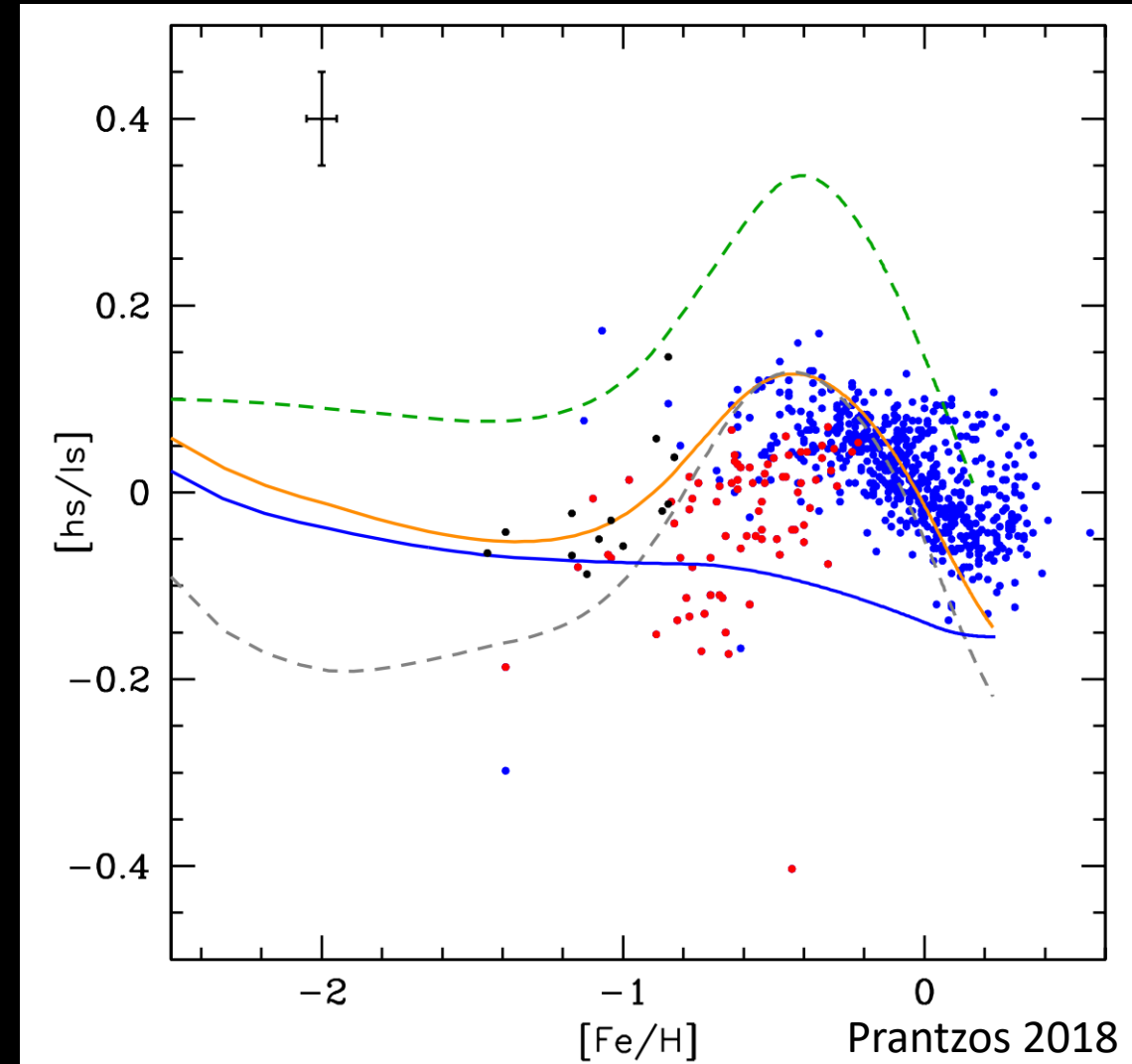
- Classification of the CEMP stars: may change with NLTE effects



Yttrium (1D)

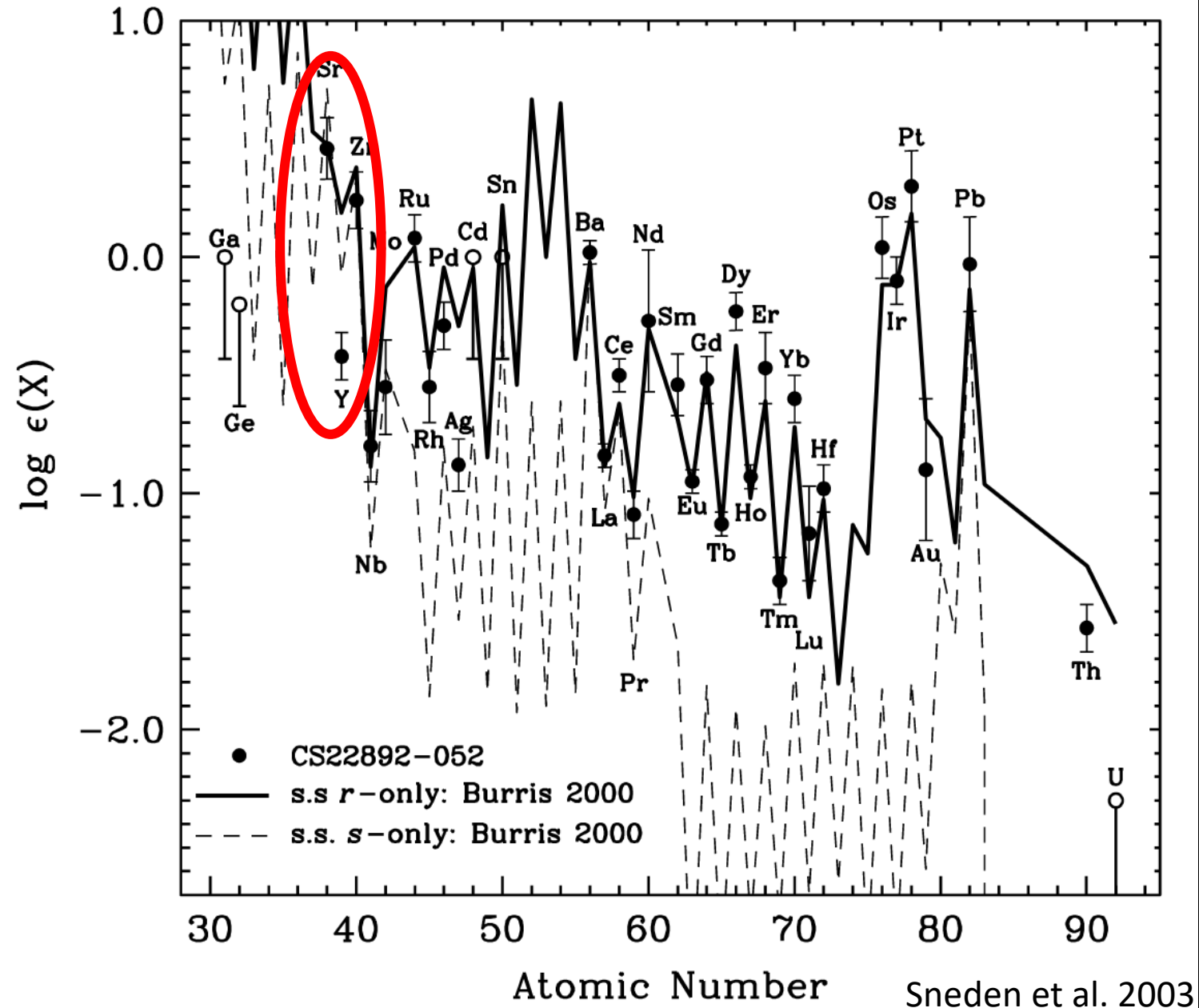
Yttrium

- 1st s-process peak element
- Used in:
 - [Ba/Y]: s-process efficiency (Travaglio 2004, Prantzos 2018)
 - [Y/Mg]: chemical clock for solar like stars (e.g. da Silva et al. 2012, Nissen 2015)



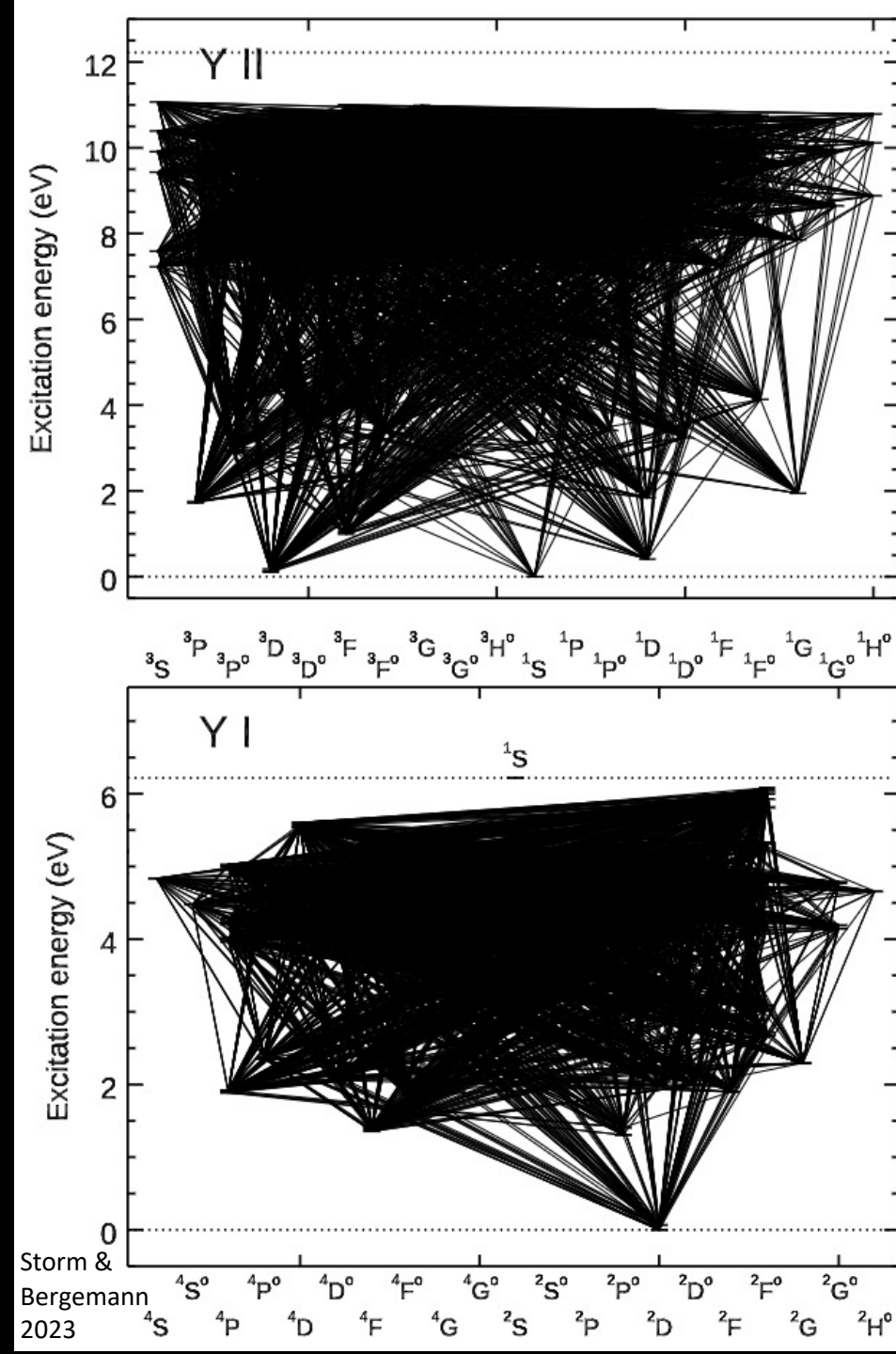
Previous studies

- Y underabundant compared to Sr and Zr?
 - e.g. Sneden et al. 2003
 - CS 22892-052: 0.15-0.30 dex lower than Sr and Zr



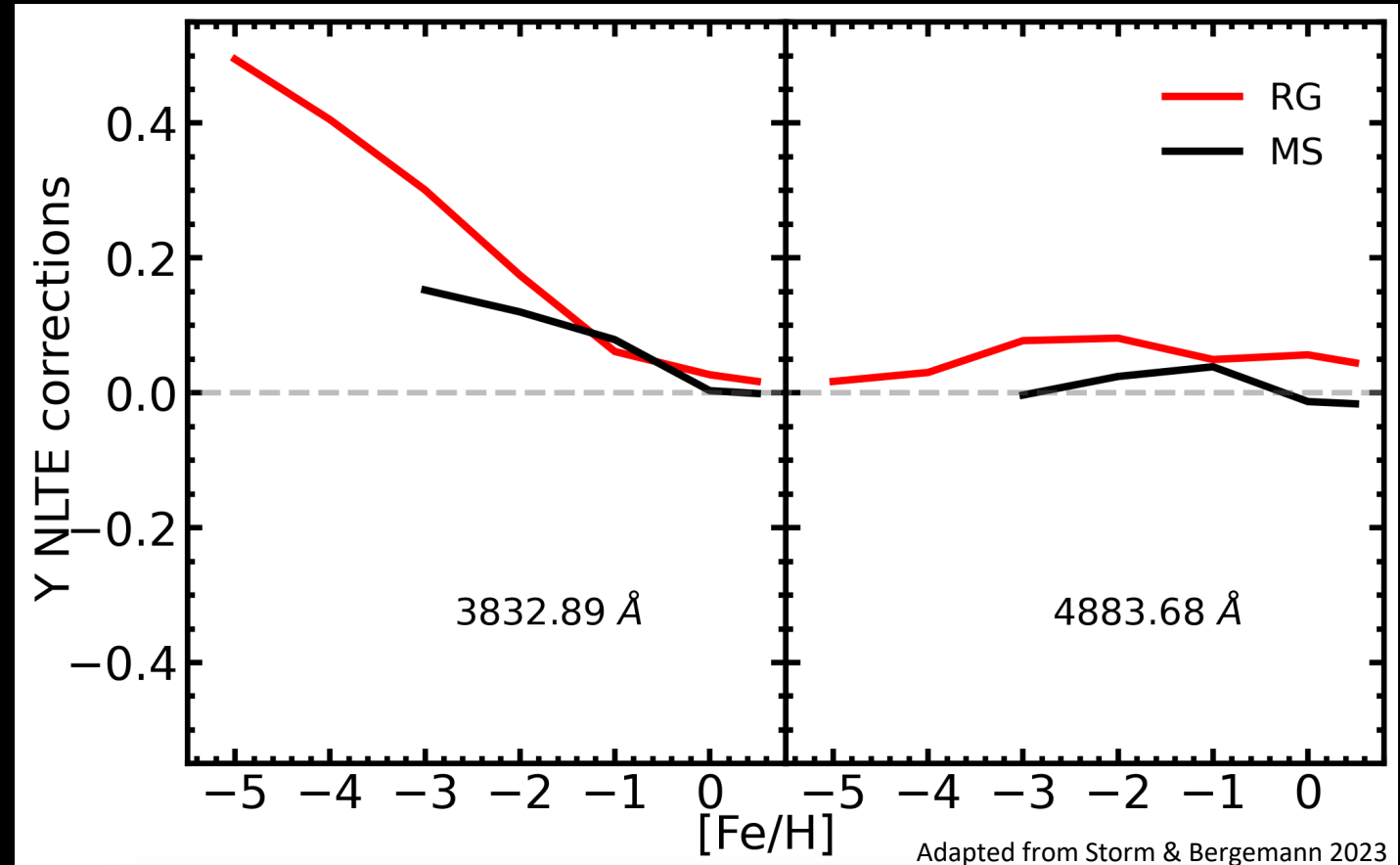
Y model atom

- 423 Y I and Y II levels
- Collisional data with e- and H: classical recipes
- NLTE calculations:
 - MULTI1D
 - Departure coefficients for TurboSpectrum (Plez et al. 2012, Gerber et al. 2023) -> can fit with TSFitPy + TurboSpectrum
 - <https://github.com/TSFitPy-developers/TSFitPy>
 - 15 elements in NLTE for 1D or "average 3D"
 - MULTI3D@DISPATCH (Eitner et al. in prep)



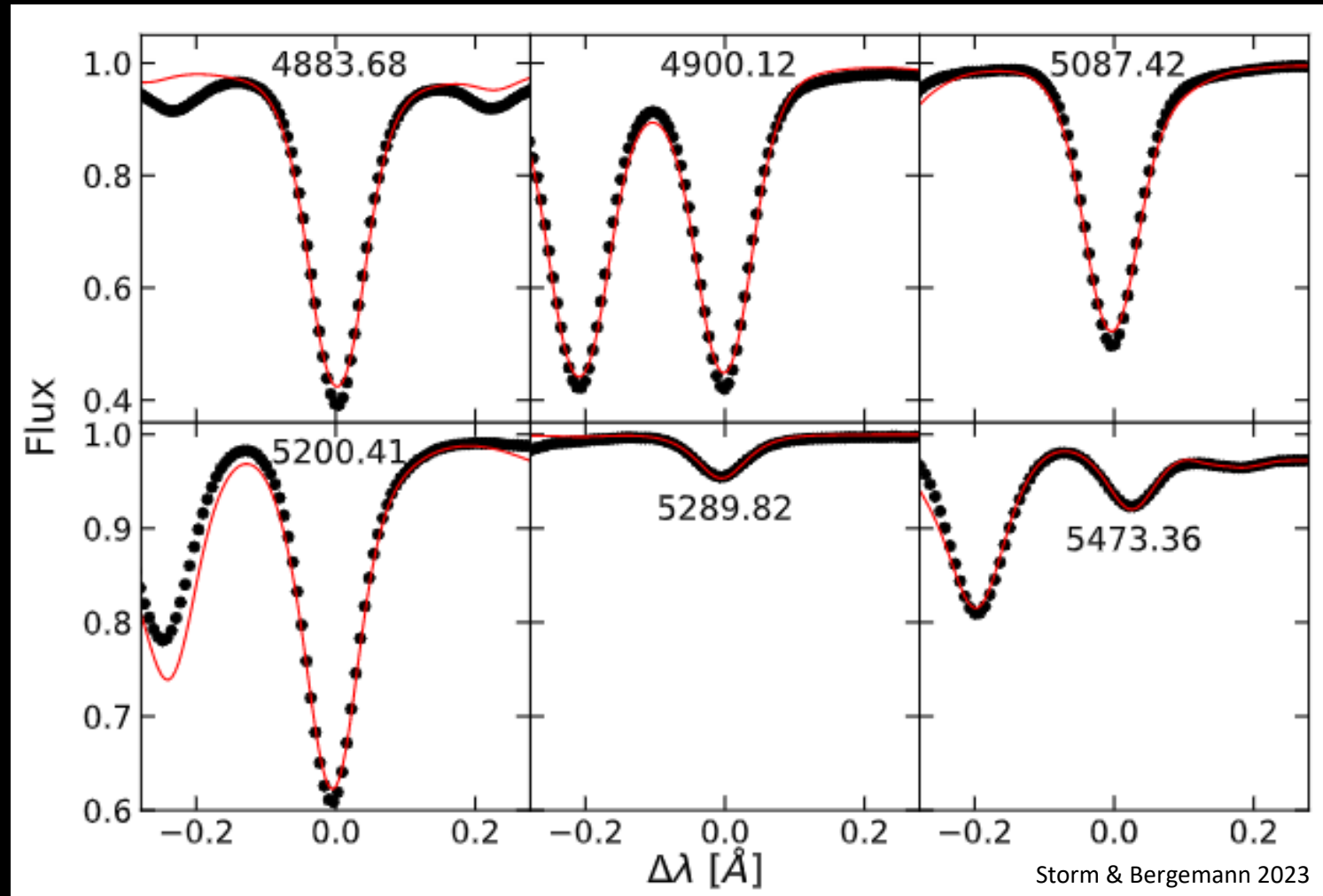
Yttrium NLTE corrections

- Higher for red giants



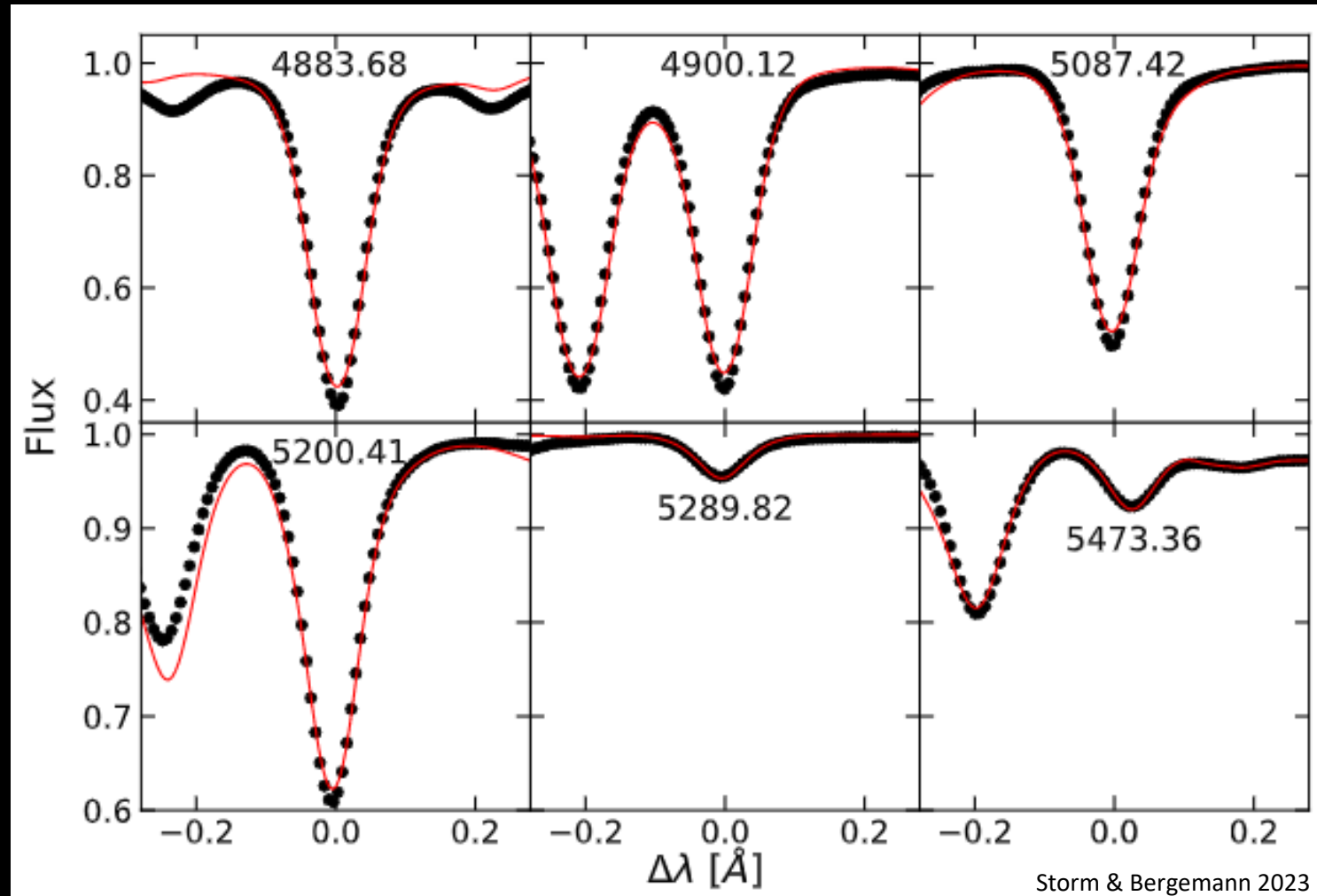
The Sun

- Grevesse et al. 2015:
 - Older loggf values
 - 1D LTE: 2.14 ± 0.06
 - 3D LTE: 2.21 ± 0.05



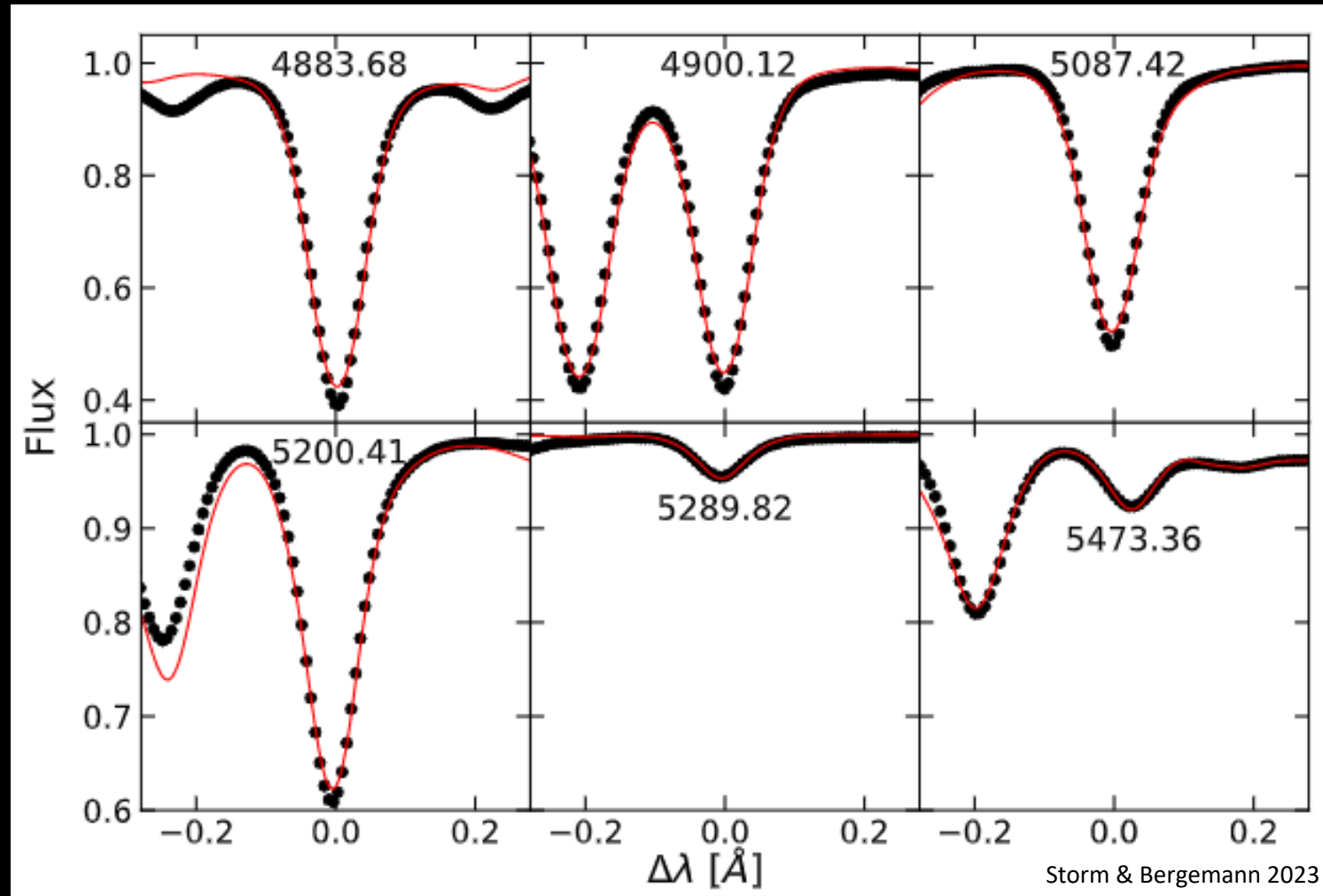
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- Meteoric value (Lodders et al. 2009):
 - 2.17 ± 0.04
- Our A(Y)
 - 1D LTE: 2.08 ± 0.04
 - 1D NLTE: 2.12 ± 0.04
 - 3D NLTE: Next slides

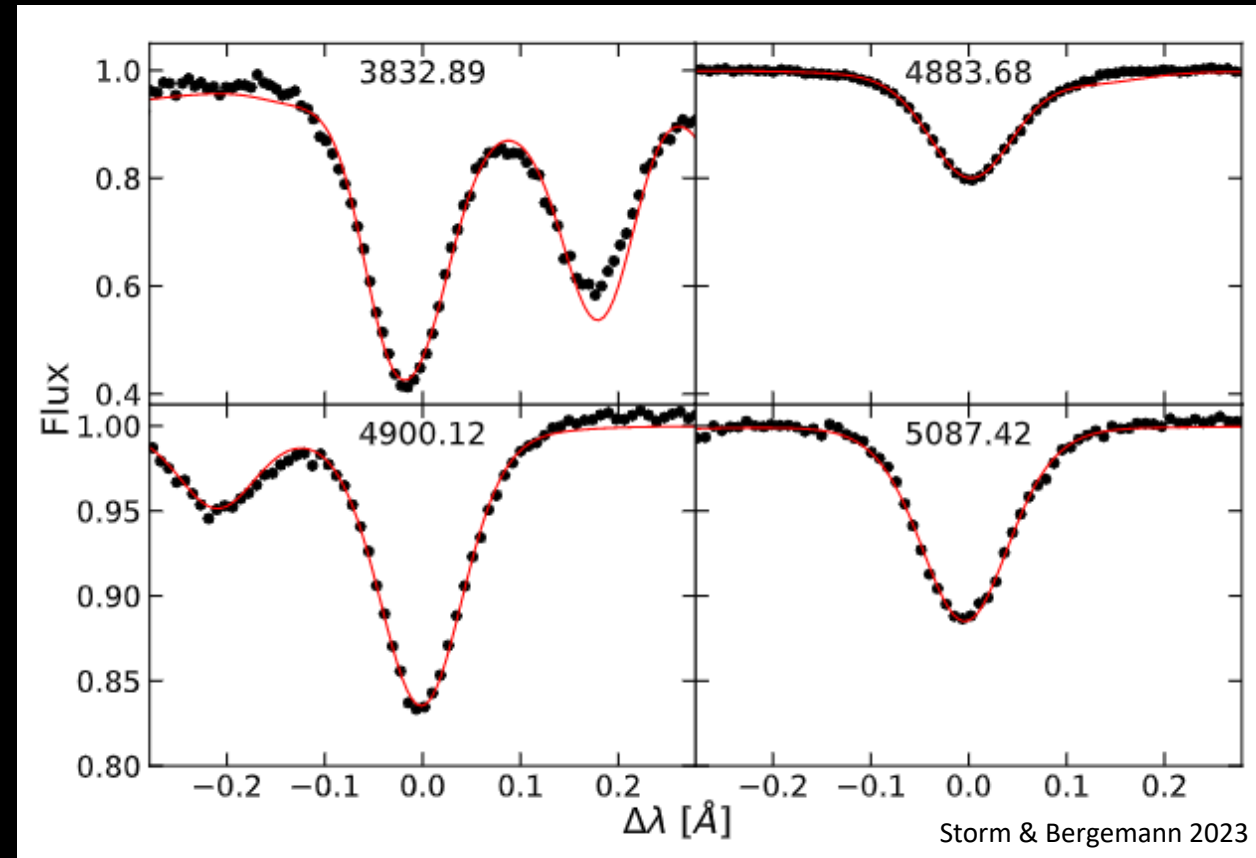


Fitting benchmark HD122563 (1D)

Yttrium

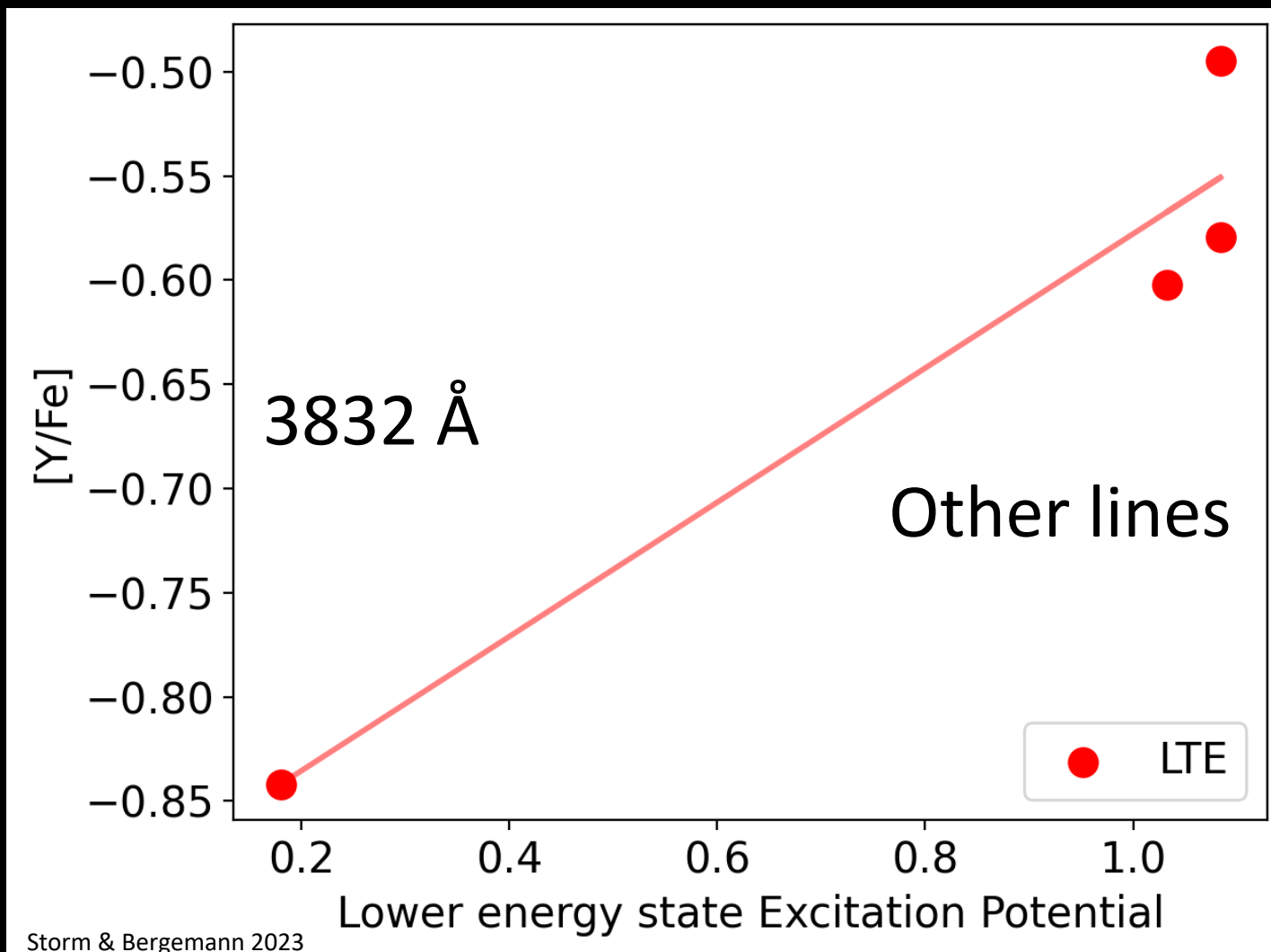
HD 122563

- Well known Gaia/4MOST benchmark star
- $T_{\text{eff}} = 4600$, $\log g = 1.4$, $[\text{Fe}/\text{H}] = -2.6$
- Previous studies:
 - Aoki 2005: $A(Y) = -0.84 \pm 0.24$
 - Honda 2006: $A(Y) = -0.93 \pm 0.09$



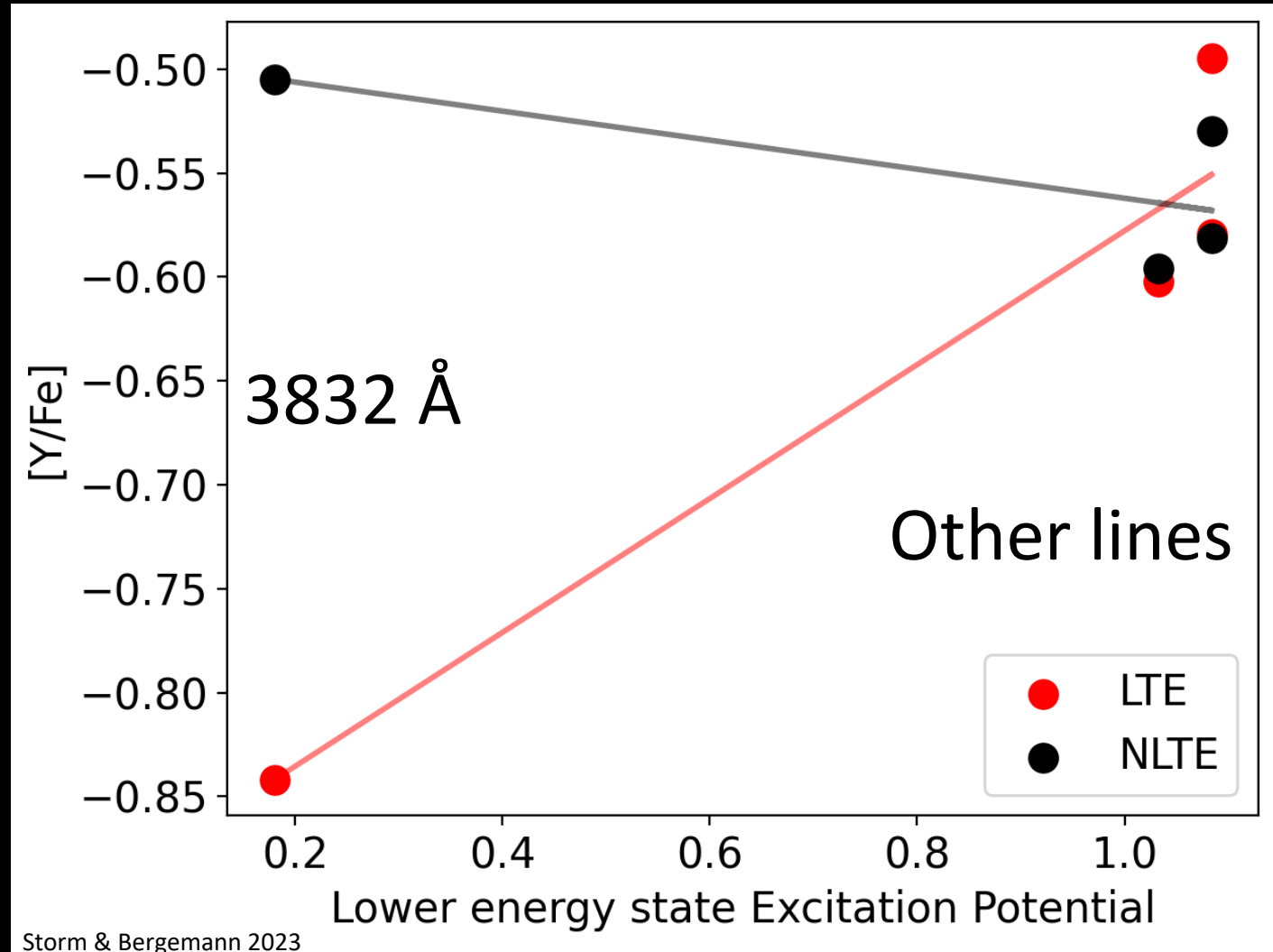
HD 122563

- Our results A(Y):
 - 1D LTE: -1.02 ± 0.13



HD 122563

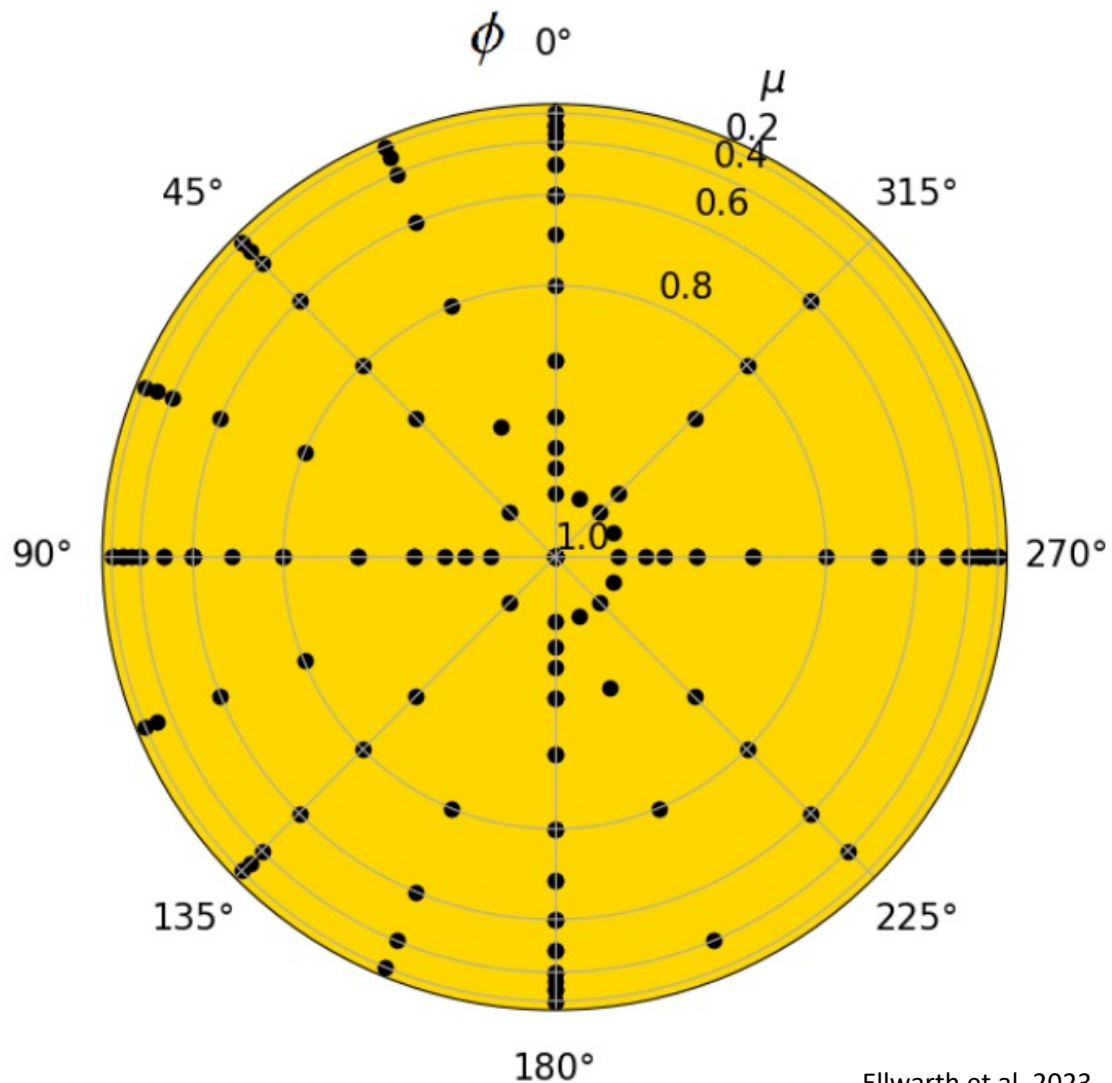
- Our results A(Y):
 - 1D LTE: -1.02 ± 0.13
 - 1D NLTE: -0.89 ± 0.04
 - Full agreement with previous studies, but lower spread



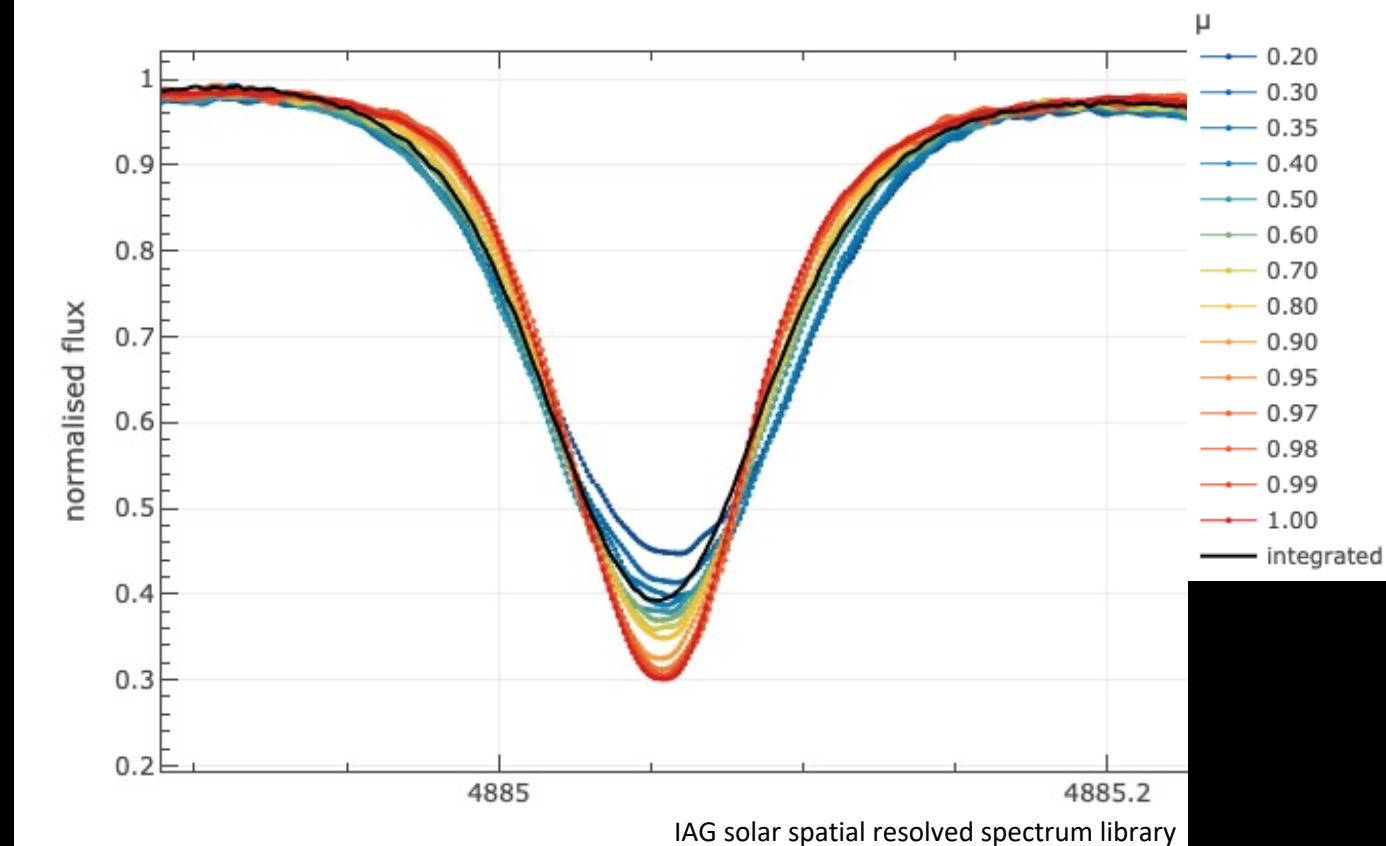
Centre-to-limb variation (3D)

Yttrium

Centre-to-limb variation (CLV)



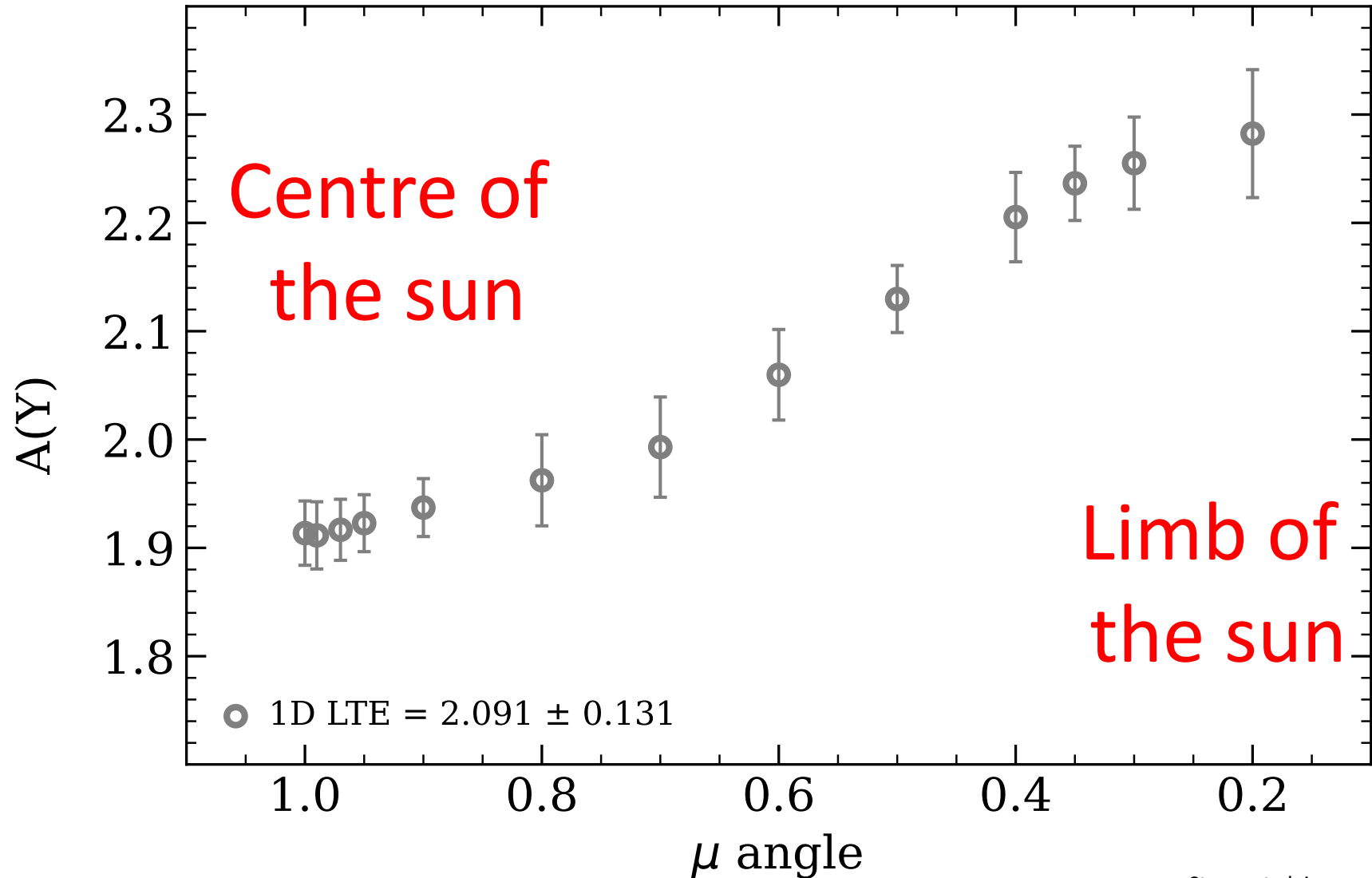
Ellwarth et al. 2023



storm@mpia.de

CLV

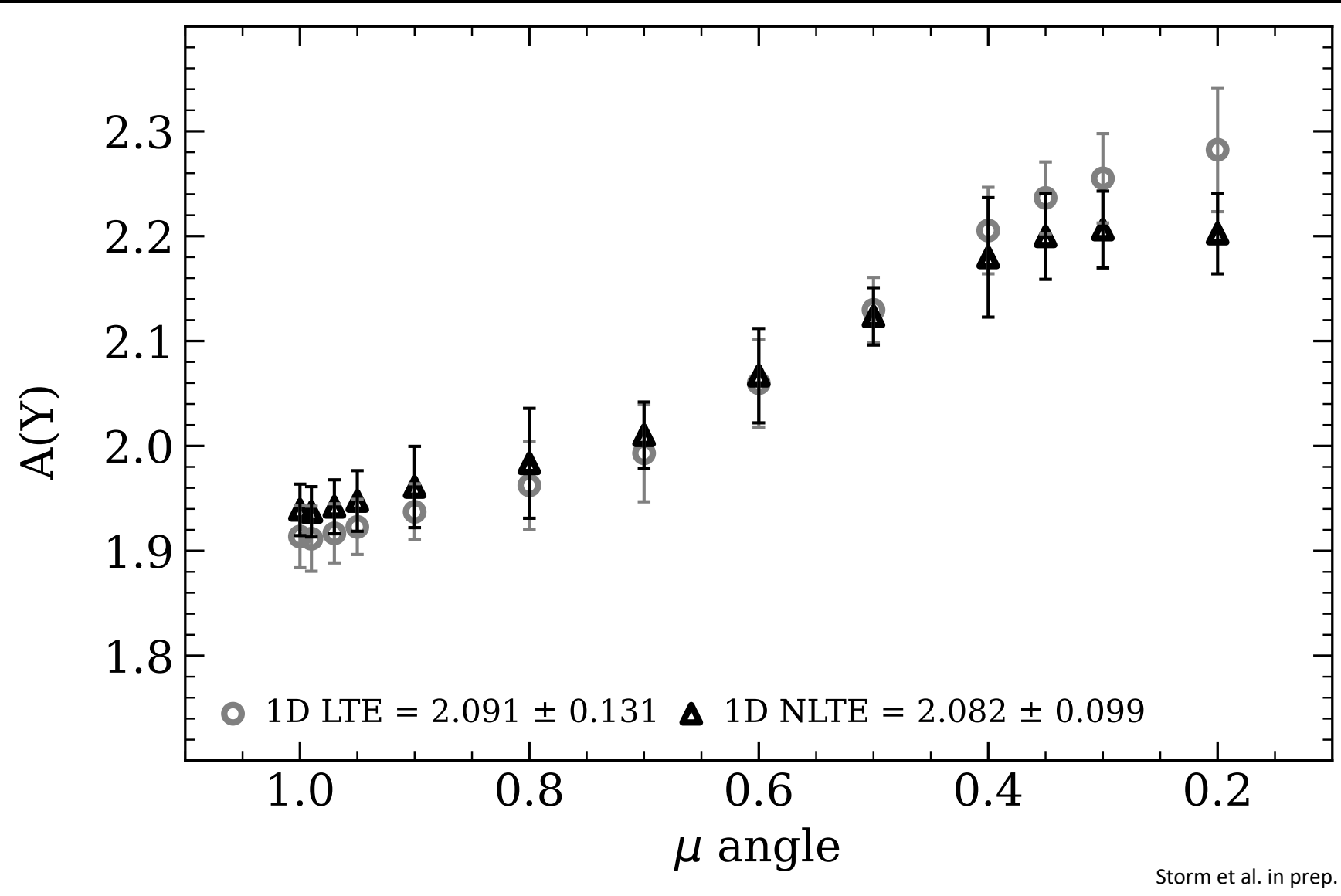
- Solar centre-to-limb variation
- Updated model atom
 - With Y+H quantum-mechanical rates (from Wang et al. 2023, subm.)



Storm et al. in prep.

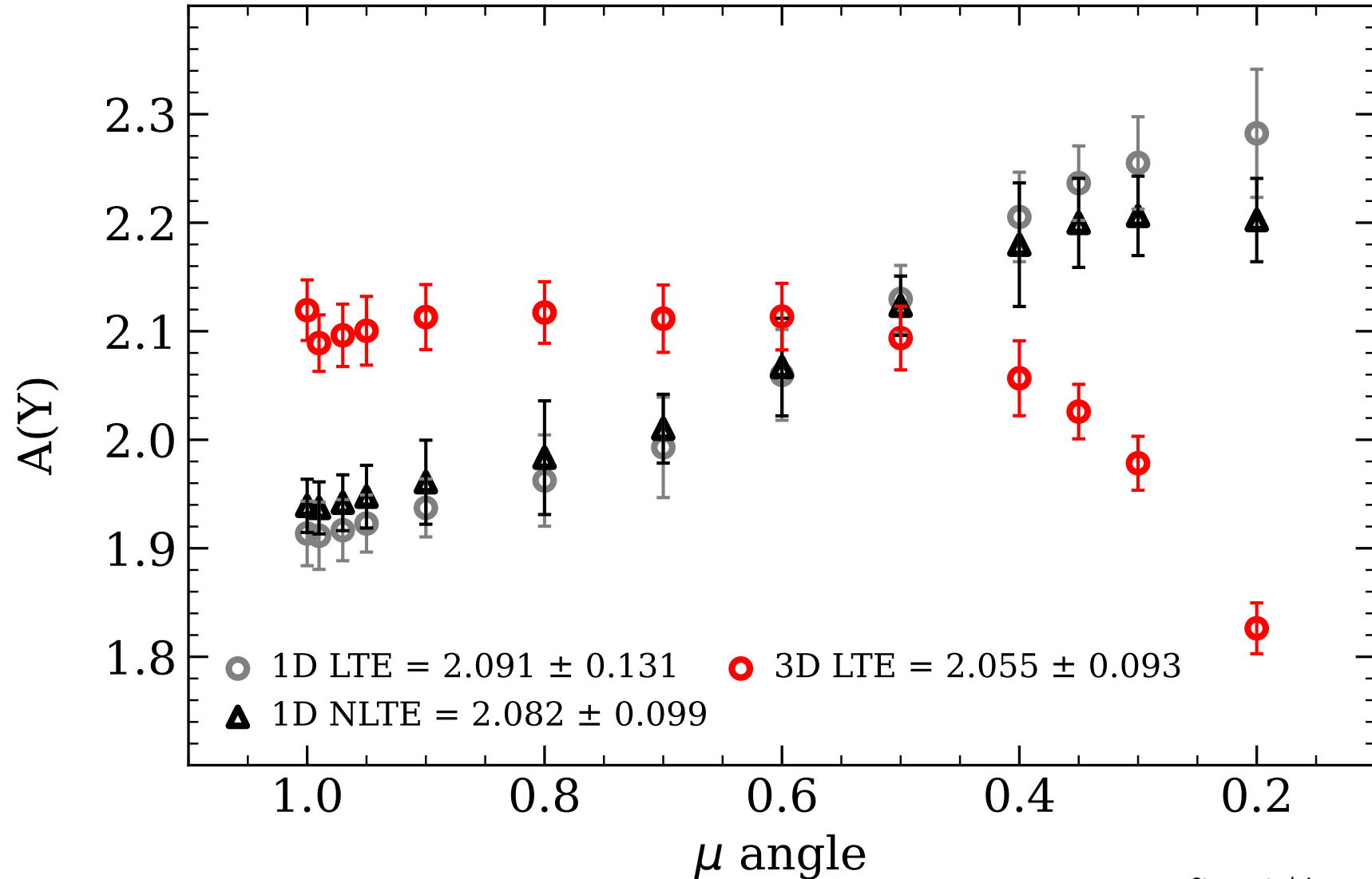
CLV

- Solar centre-to-limb variation



CLV

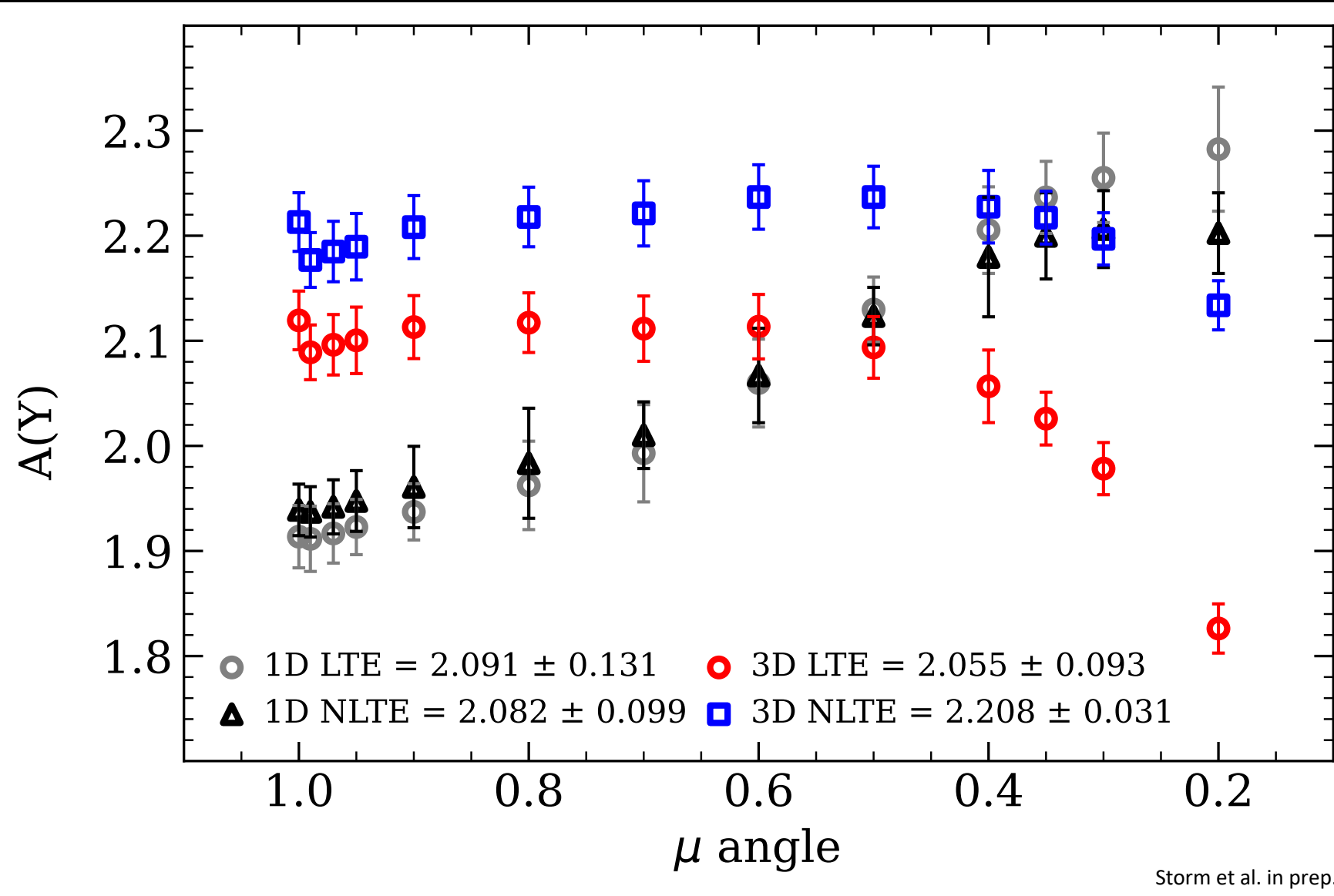
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Storm et al. in prep.

CLV

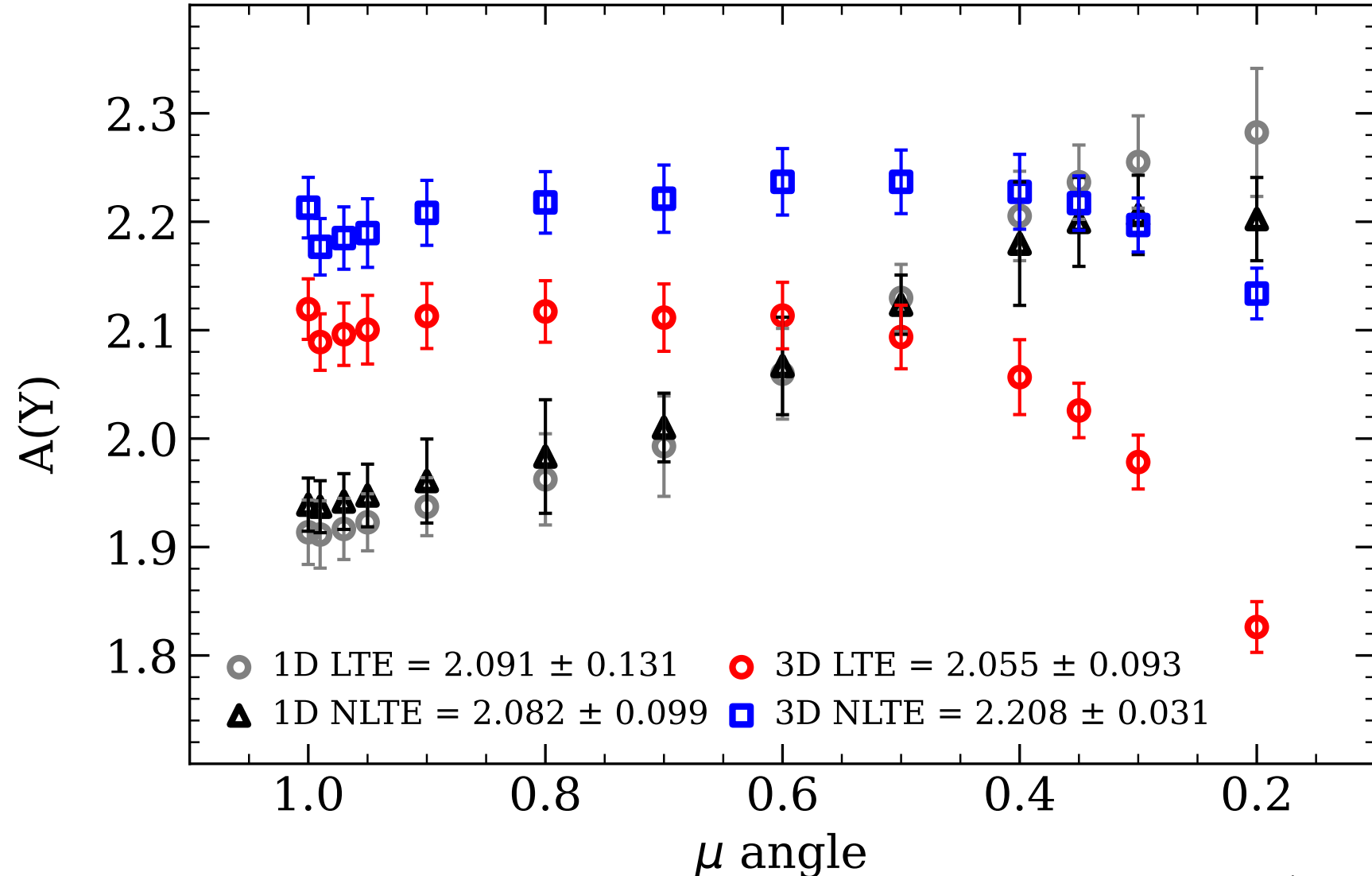
- Solar centre-to-limb variation



Storm et al. in prep.

CLV

- Repeat for other lines:
 - 4883
 - 4900
 - 5087
- Our rederived solar **3D NLTE**:
 - $A(Y) = 2.20 \pm 0.04$
 - ± 0.07 systematic
 - But! Depends on loggf choice
 - Need very accurate loggf values!!

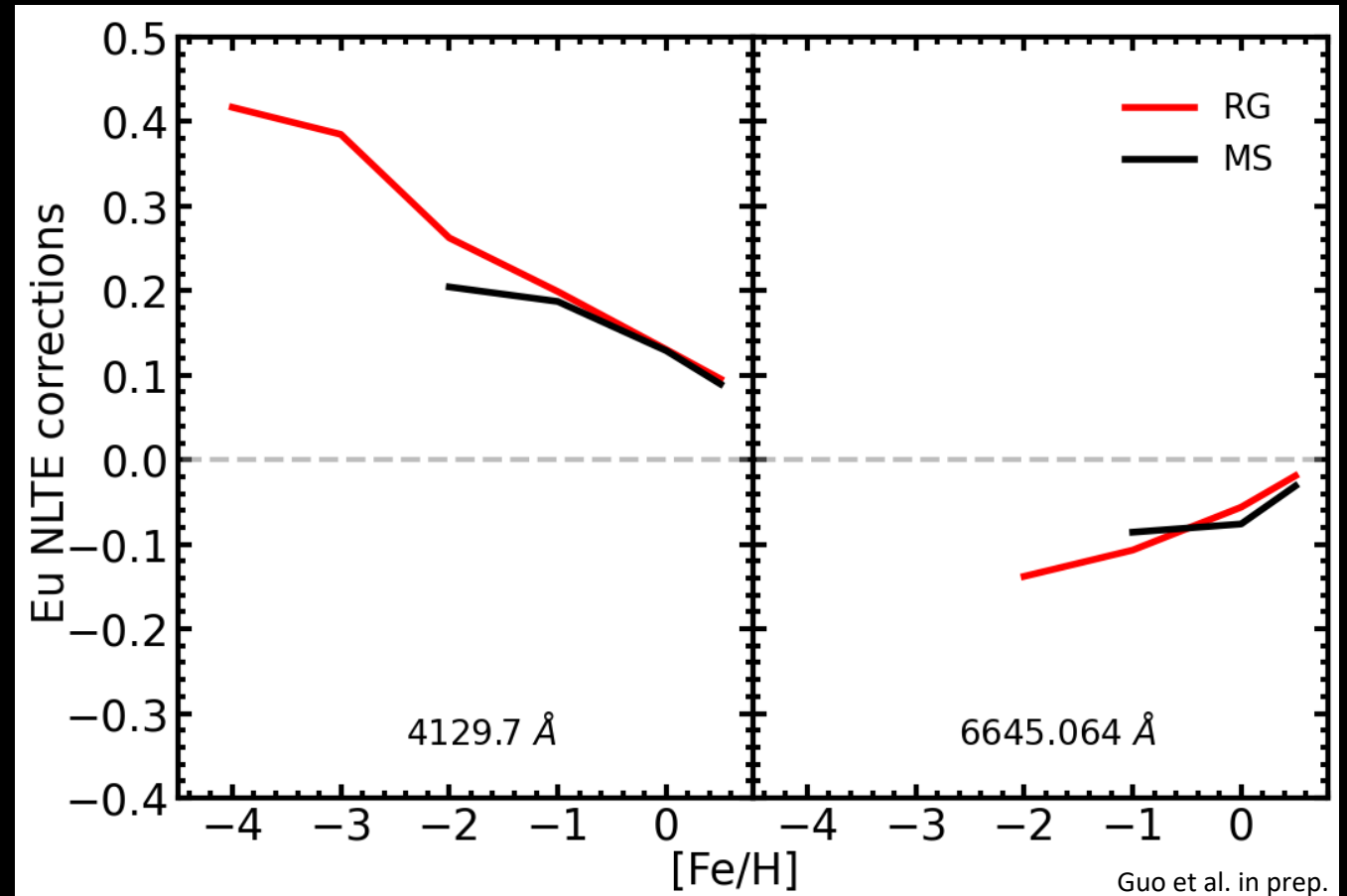


Storm et al. in prep.

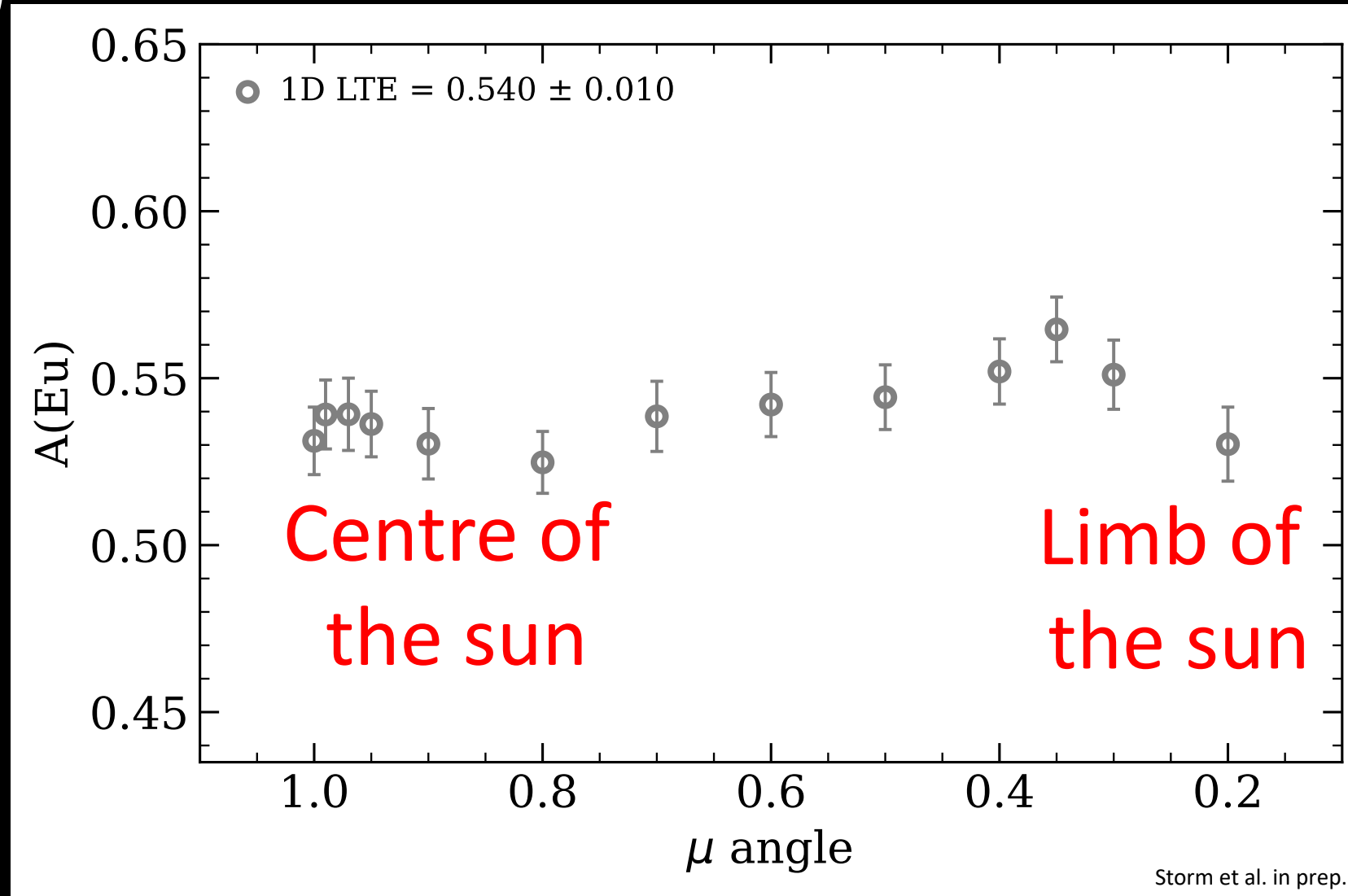
Europium

Europium (1D)

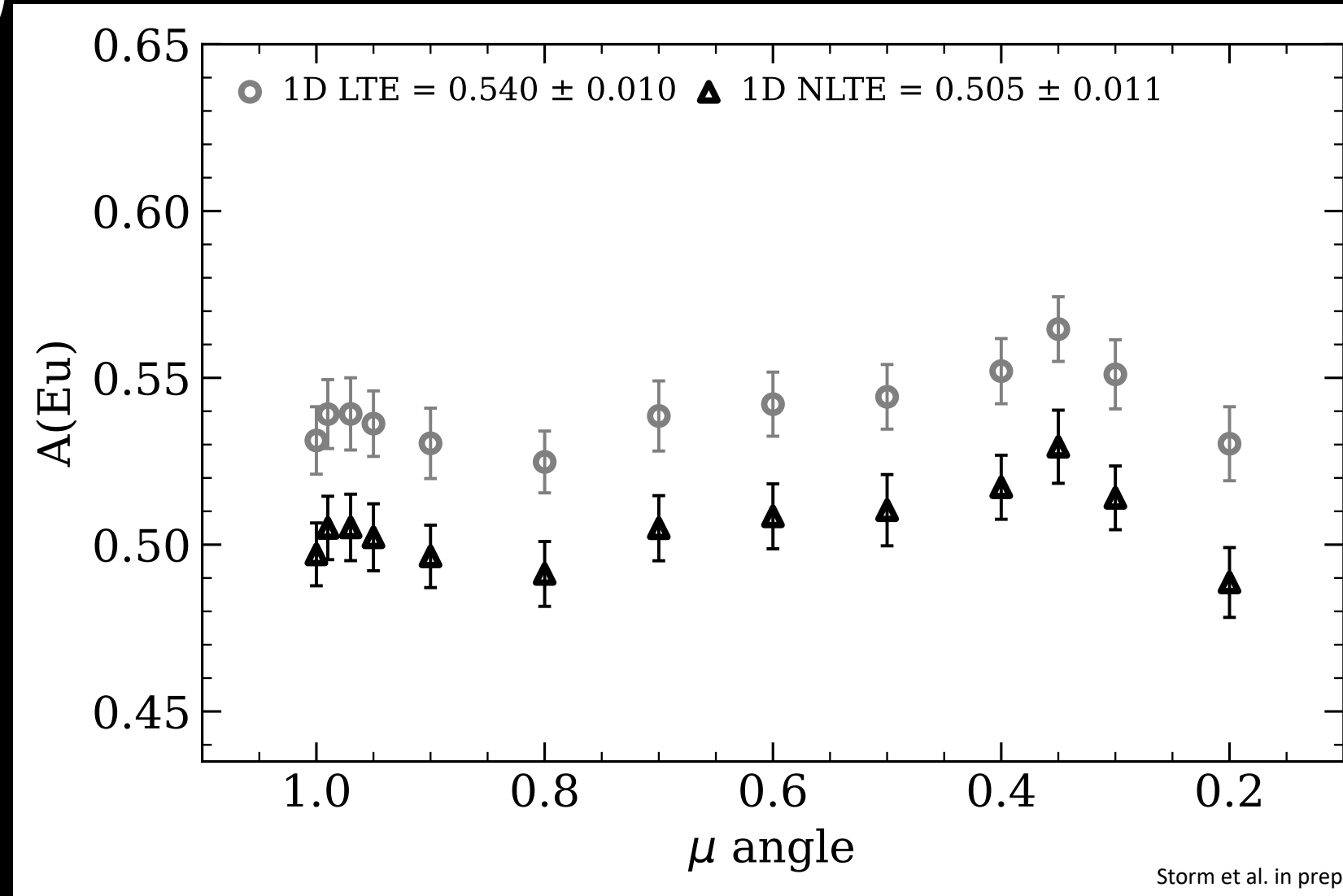
- Mainly r-process element
- New model atom with some quantum-mechanical rates



CLV variation (3D)

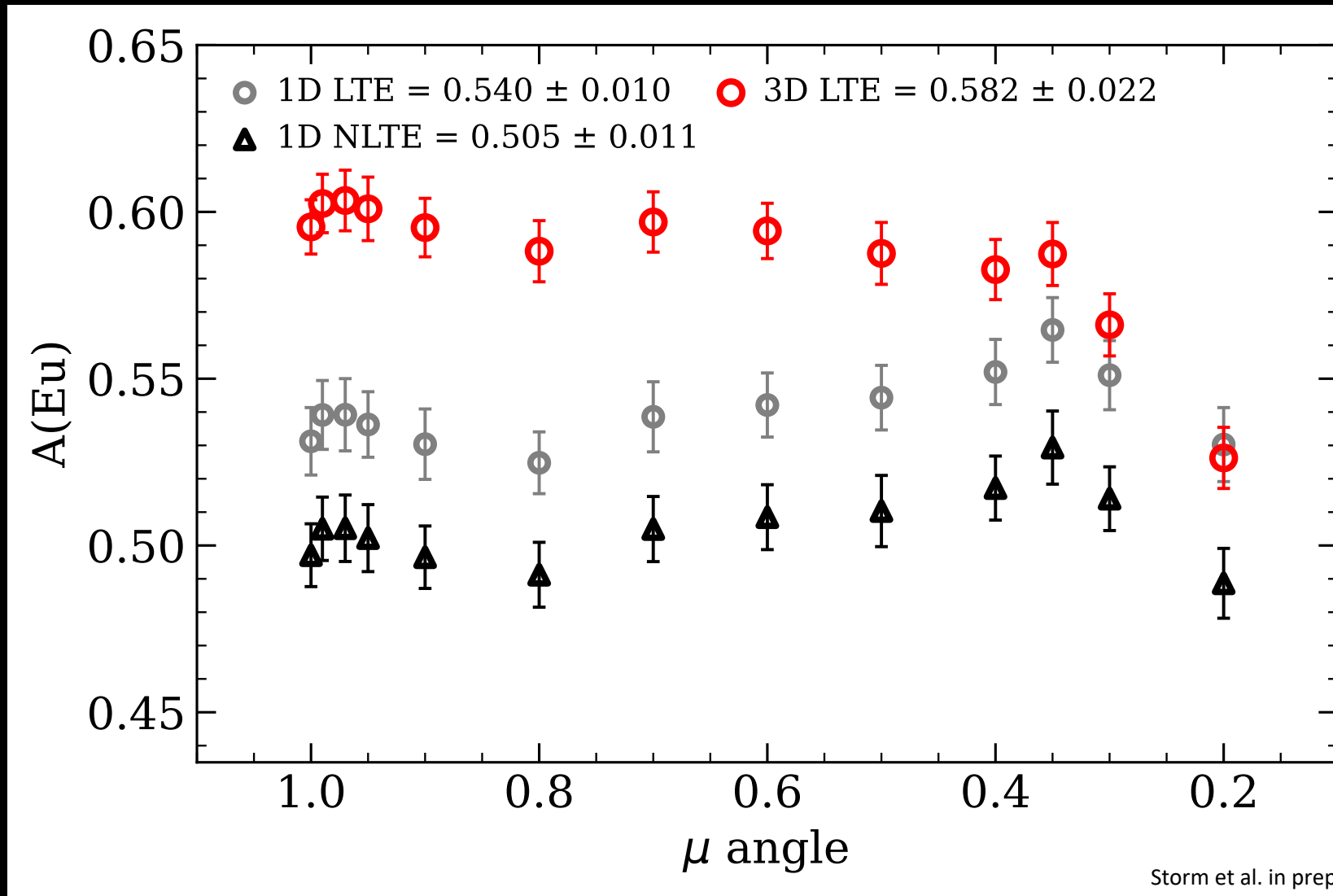


CLV variation (3D)



Storm et al. in prep.

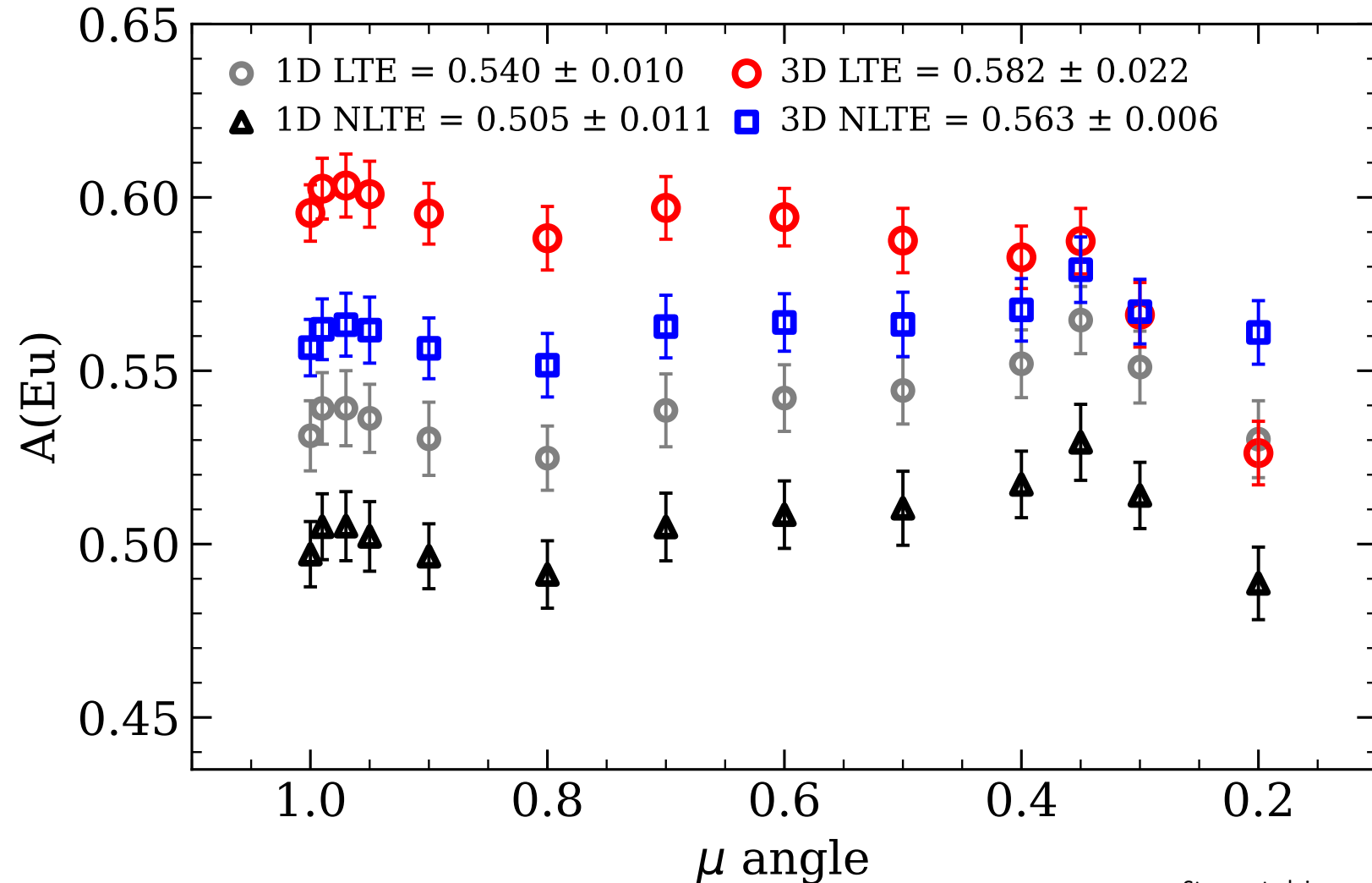
CLV variation (3D)



Storm et al. in prep.

CLV variation (3D)

- 3D NLTE: most consistent abundance
 - Lowest scatter
- Our solar **3D NLTE**:
 - $A(\text{Eu}) = 0.56 \pm 0.01$
 - ± 0.07 systematic



Storm et al. in prep.

Conclusions

Conclusions

- NLTE and 3D modelling improves consistency of the results
- Y and Eu
 - NLTE effects especially important for near-UV lines
 - Higher for RG stars
- 3D NLTE
 - Most consistent results for center-to-limb variation
- Rederived solar Y and Eu abundances (**3D NLTE**):
 - $A(Y) = 2.20 \pm 0.04$
 - $A(Eu) = 0.56 \pm 0.01$
- You can also do fitting in NLTE: TSFitPy (GitHub)