

Improved Measurements of (α, p) Reactions with Digital ANASEN

Keilah Davis

PhD Student

Department of Physics & Astronomy

Louisiana State University, USA

(α, p) reactions in X-ray Bursts

- X-ray bursts are the most commonly observed stellar explosions in our Galaxy
- Difficult to compare simulations to observations due to uncertainties in nuclear physics inputs
- Sensitivity studies demonstrate importance of many (α, p) reactions, but most nuclear data comes from indirect studies
- ANASEN designed for direct (α, p) measurements — e.g. $^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$

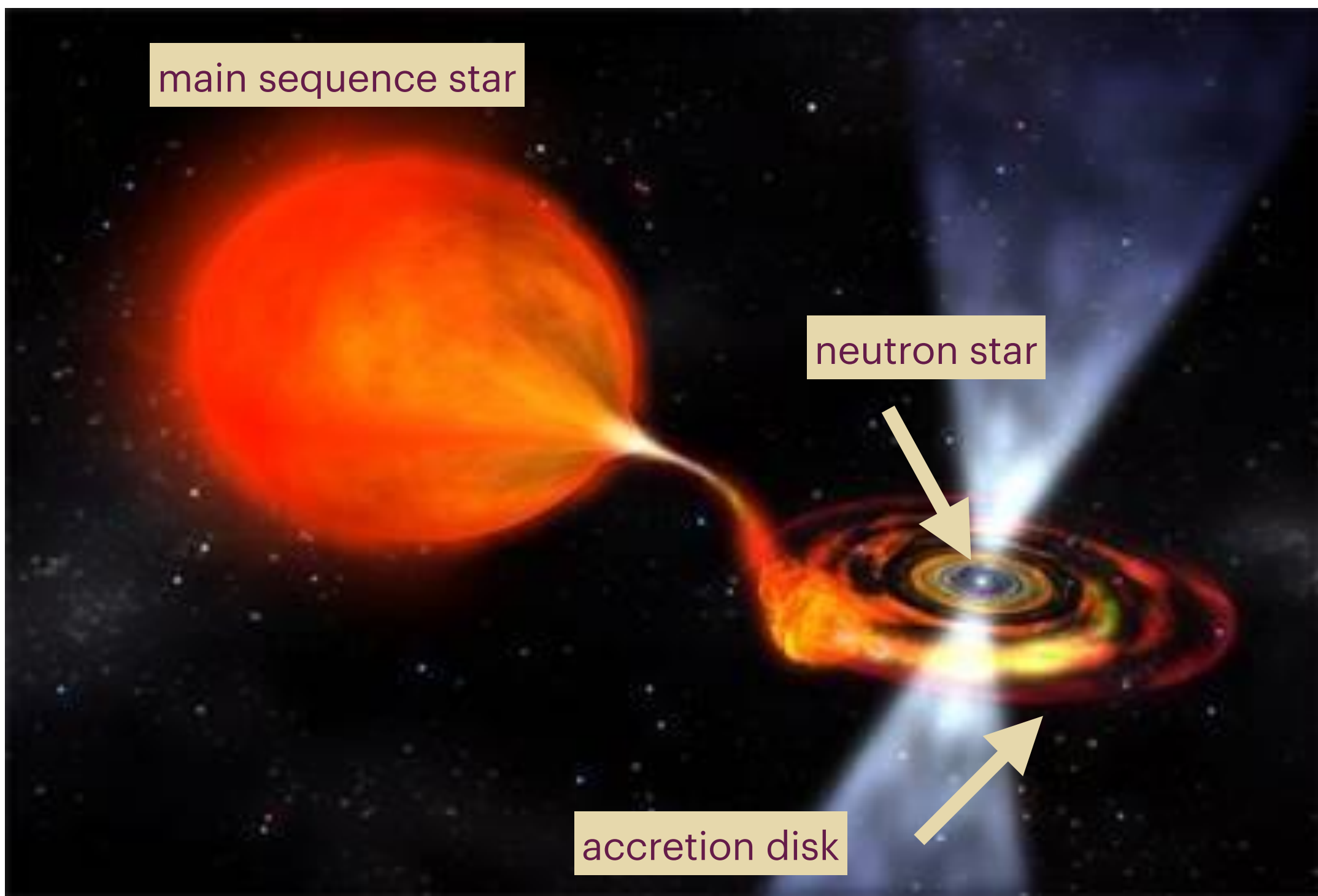


Table 2
Reactions that Impact the Burst Light Curve
in the Multi-zone X-ray Burst Model

Rank	Reaction	Type ^a	Sensitivity ^b	Category
1	$^{15}\text{O}(\alpha, \gamma)^{19}\text{Ne}$	D	16	1
2	$^{56}\text{Ni}(\alpha, p)^{59}\text{Cu}$	U	6.4	1
3	$^{59}\text{Cu}(p, \gamma)^{60}\text{Zn}$	D	5.1	1
4	$^{61}\text{Ga}(p, \gamma)^{62}\text{Ge}$	D	3.7	1
5	$^{22}\text{Mg}(\alpha, p)^{25}\text{Al}$	D	2.3	1
6	$^{14}\text{O}(\alpha, p)^{17}\text{F}$	D	5.8	1
7	$^{23}\text{Al}(p, \gamma)^{24}\text{Si}$	D	4.6	1
8	$^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$	U	1.8	1
9	$^{63}\text{Ga}(p, \gamma)^{64}\text{Ge}$	D	1.4	2
10	$^{19}\text{F}(p, \alpha)^{16}\text{O}$	U	1.3	2
11	$^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$	U	2.1	2

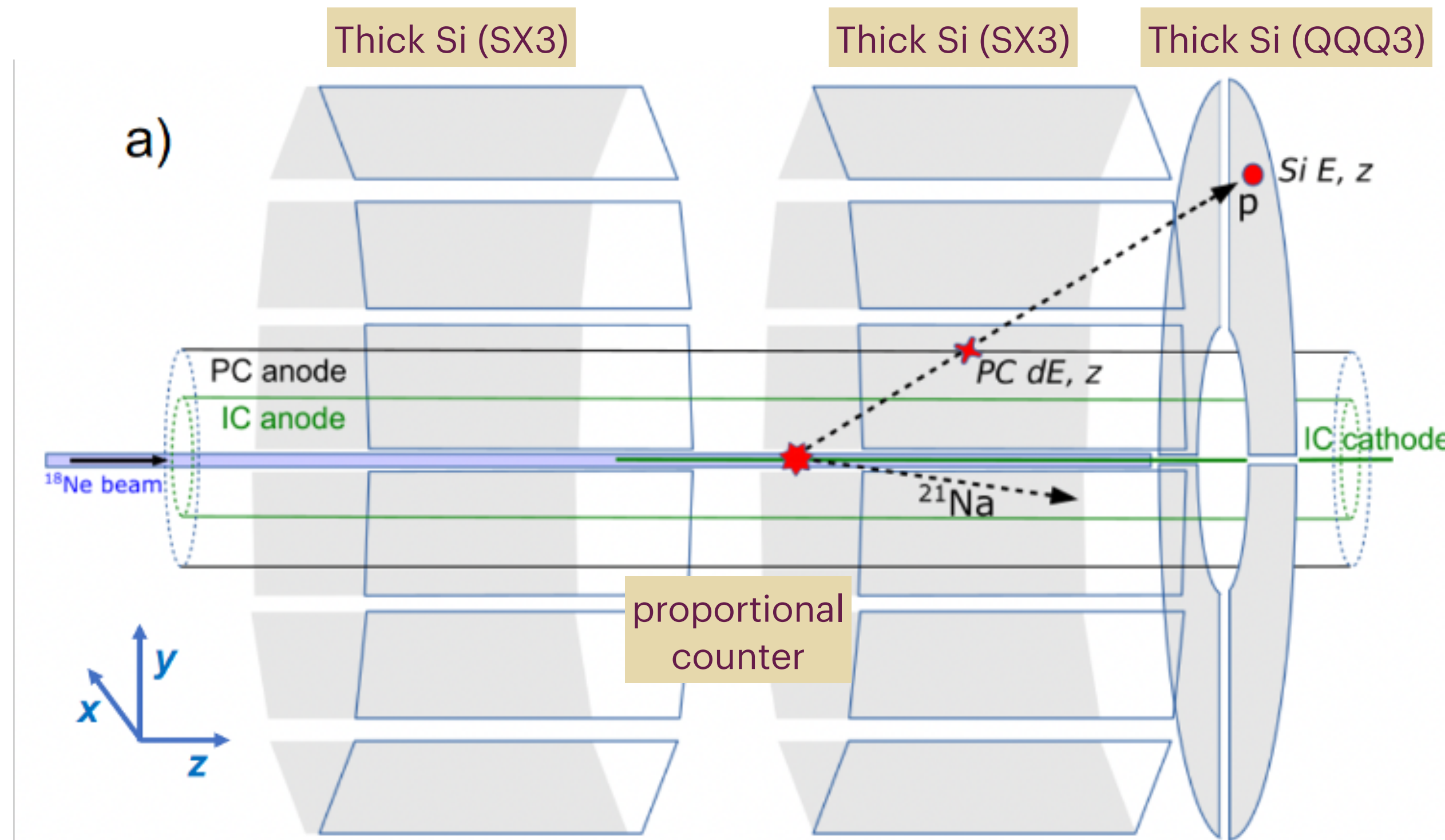
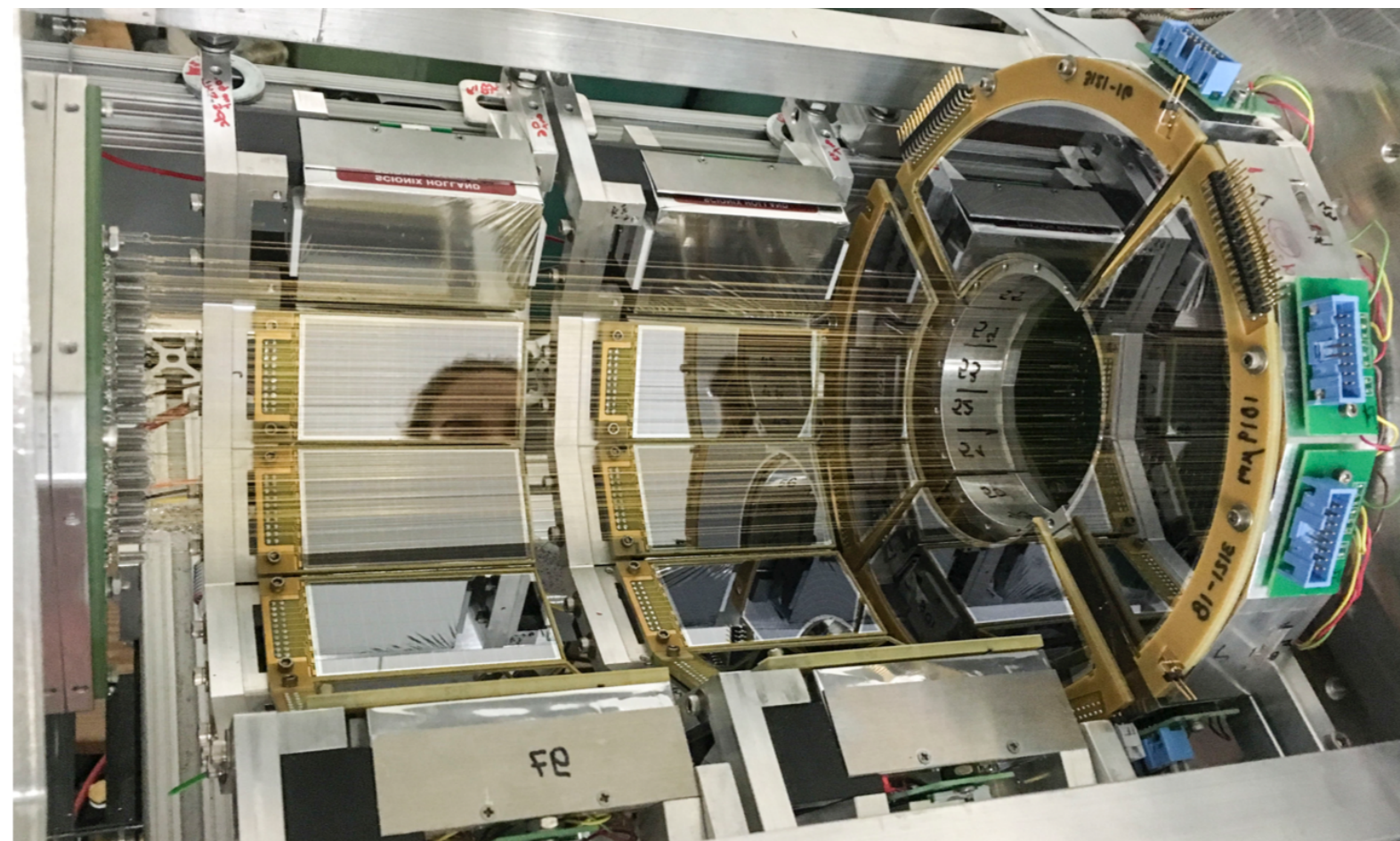
Cyburt et al., ApJ 830, 55 (2016)

Parikh et al., ApJS 178, 110 (2008)

$^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$ Cross Sections using ANASEN

Array for Nuclear Astrophysics and Structure with Exotic Nuclei

- Recently published direct measurement at Florida State University
- Active He-CO₂ gas target/detector
- Cylindrical proportional counter for ΔE
- Thick Si-strip detectors for residual E
- ^{18}Ne beam intensity 10^3 pps, scans $E_{cm} = 2.5\text{--}4.0$ MeV simultaneously

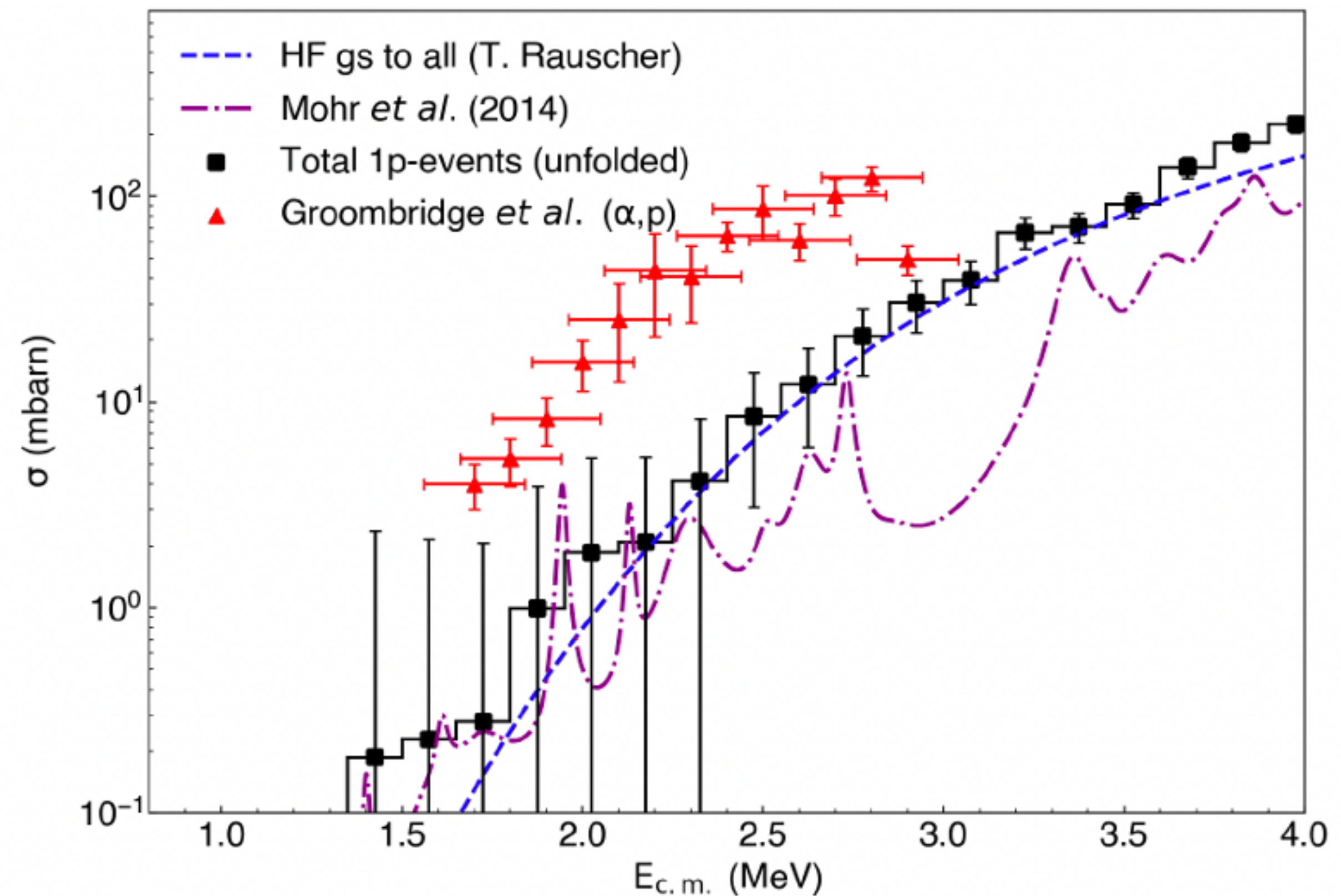


Anastasiou et al., Phys Rev C 105, 055806 (2022)

$^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$ Cross Sections using ANASEN

Array for Nuclear Astrophysics and Structure with Exotic Nuclei

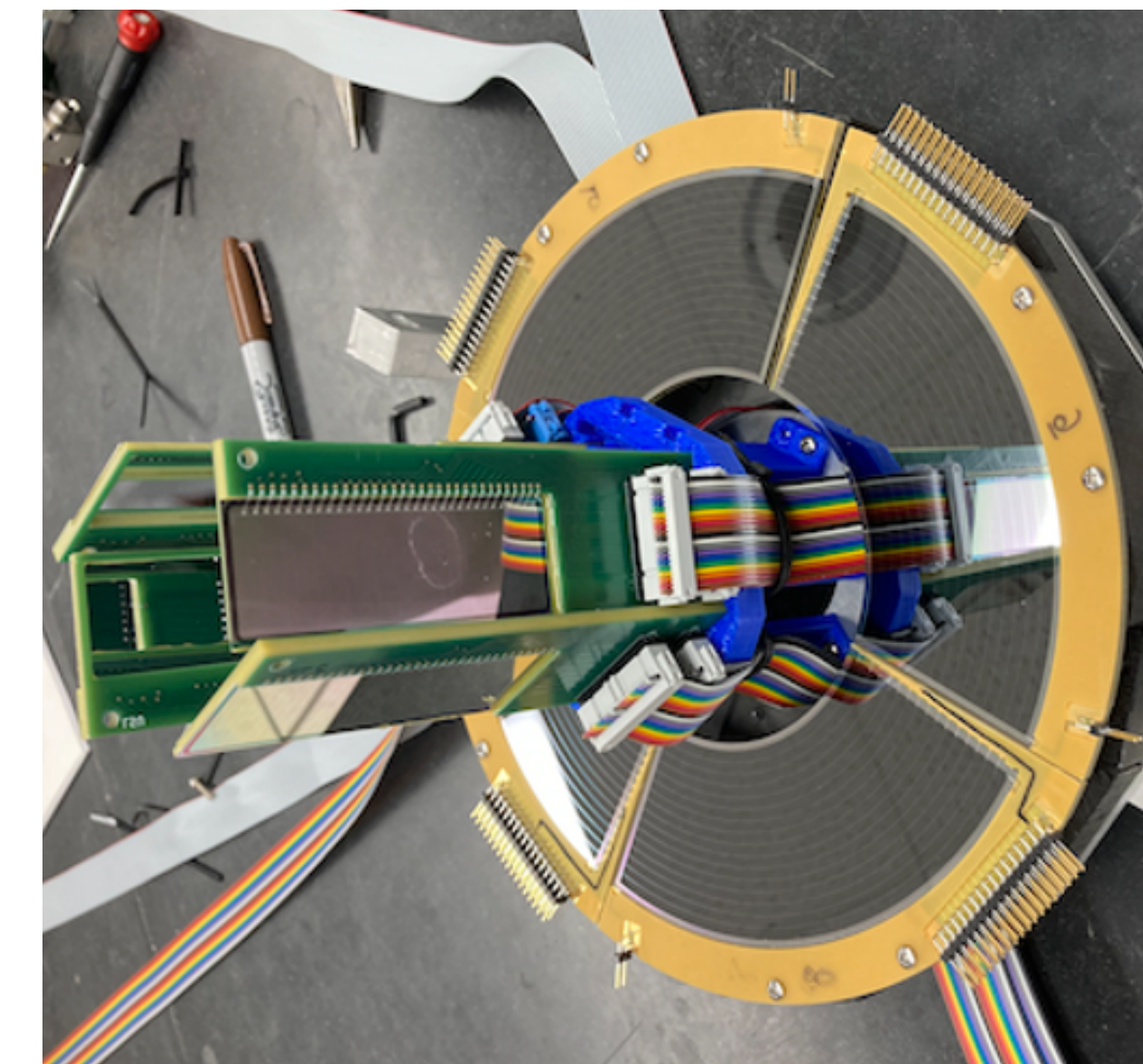
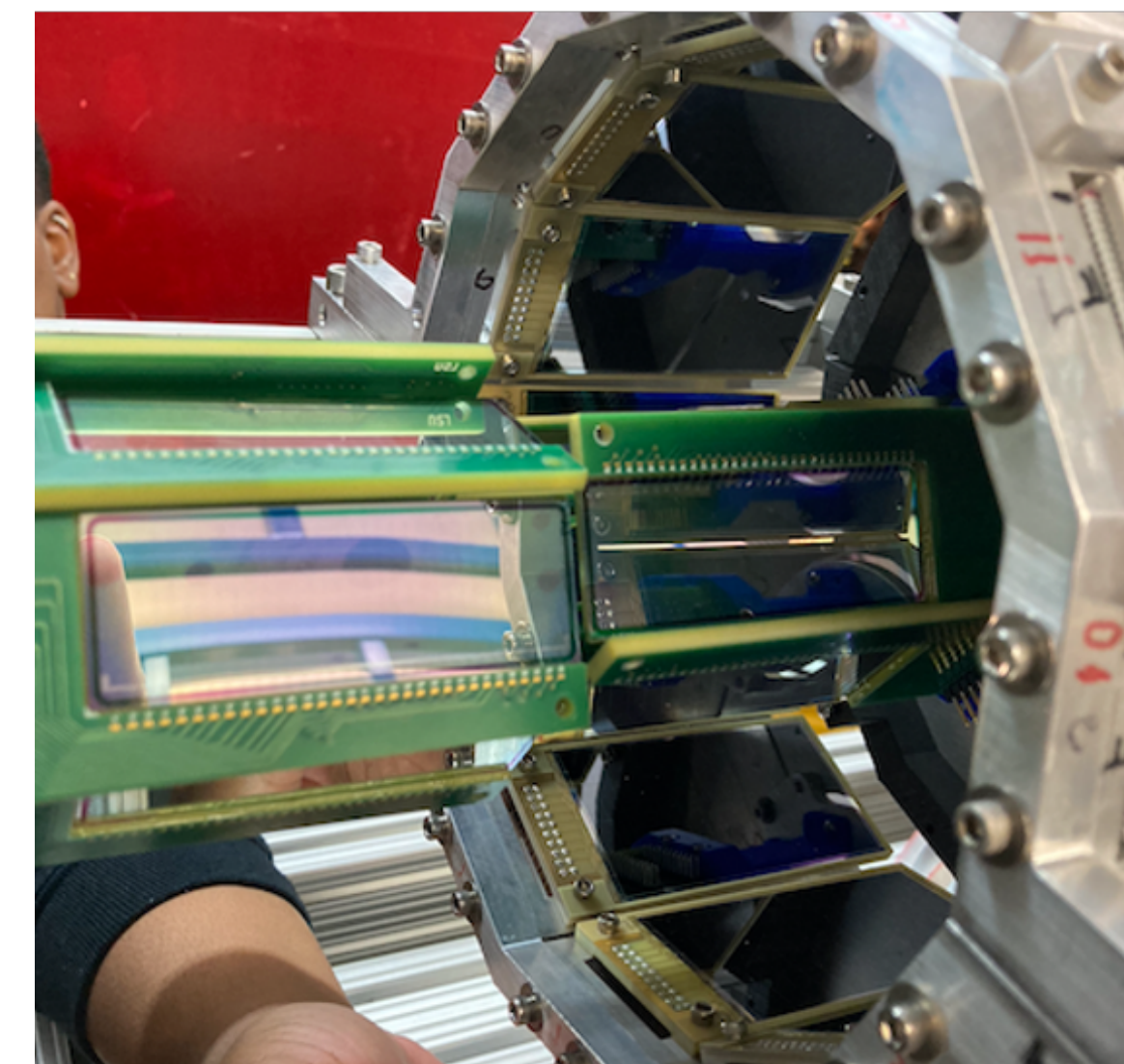
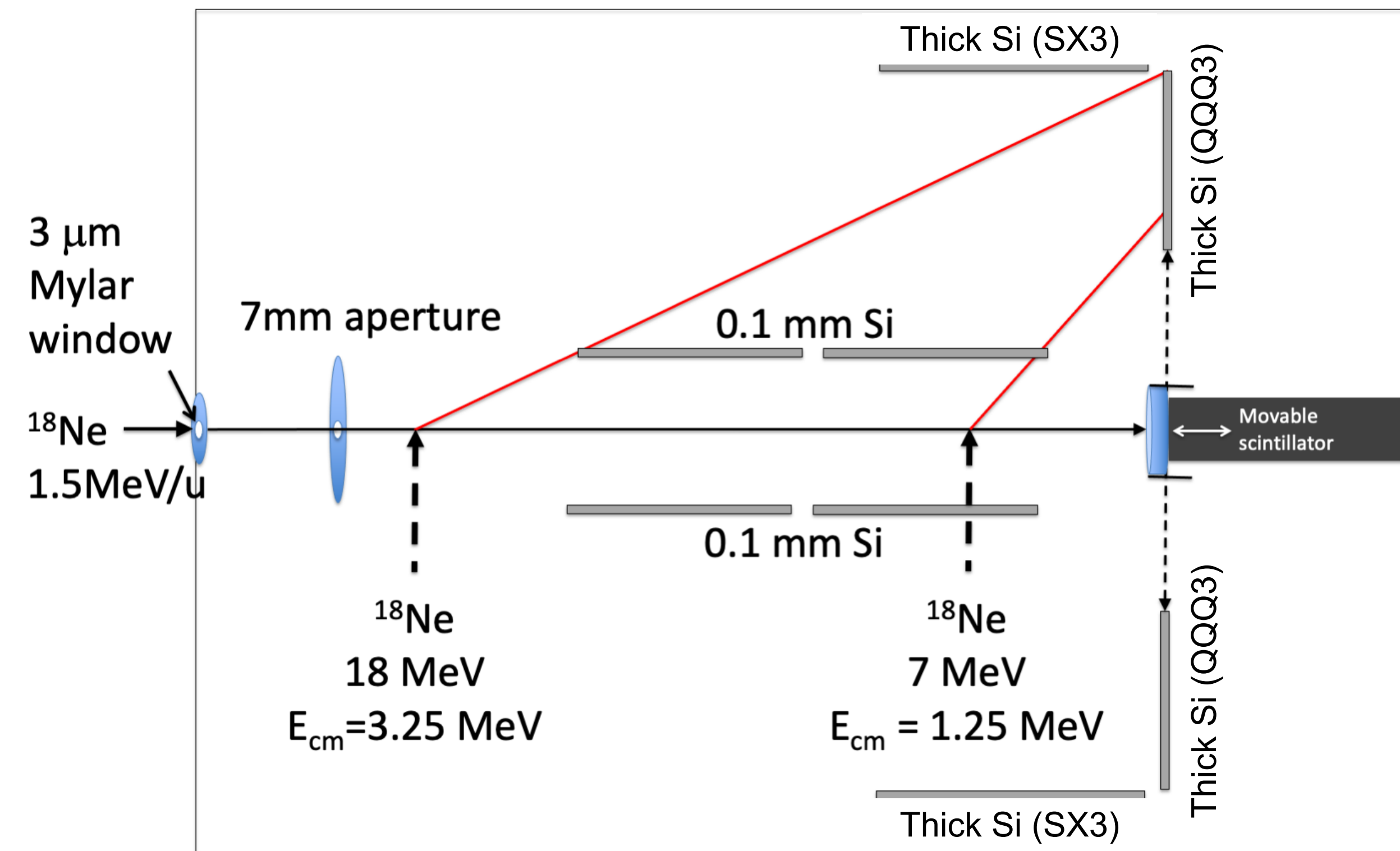
- Recently published direct measurement at Florida State University
- Active He-CO₂ gas target/detector
- Cylindrical proportional counter for ΔE
- Thick Si-strip detectors for residual E
- ^{18}Ne beam intensity 10^3 pps, scans $E_{cm} = 2.5\text{--}4.0$ MeV simultaneously
- Disentangled contributions from 2p-events
- Energy resolution $\delta E_{cm} \approx 1.5$ MeV limited by position resolution
- Only upper limits for $E_{cm} < 2.4$ MeV



Anastasiou et al., Phys Rev C 105, 055806 (2022)

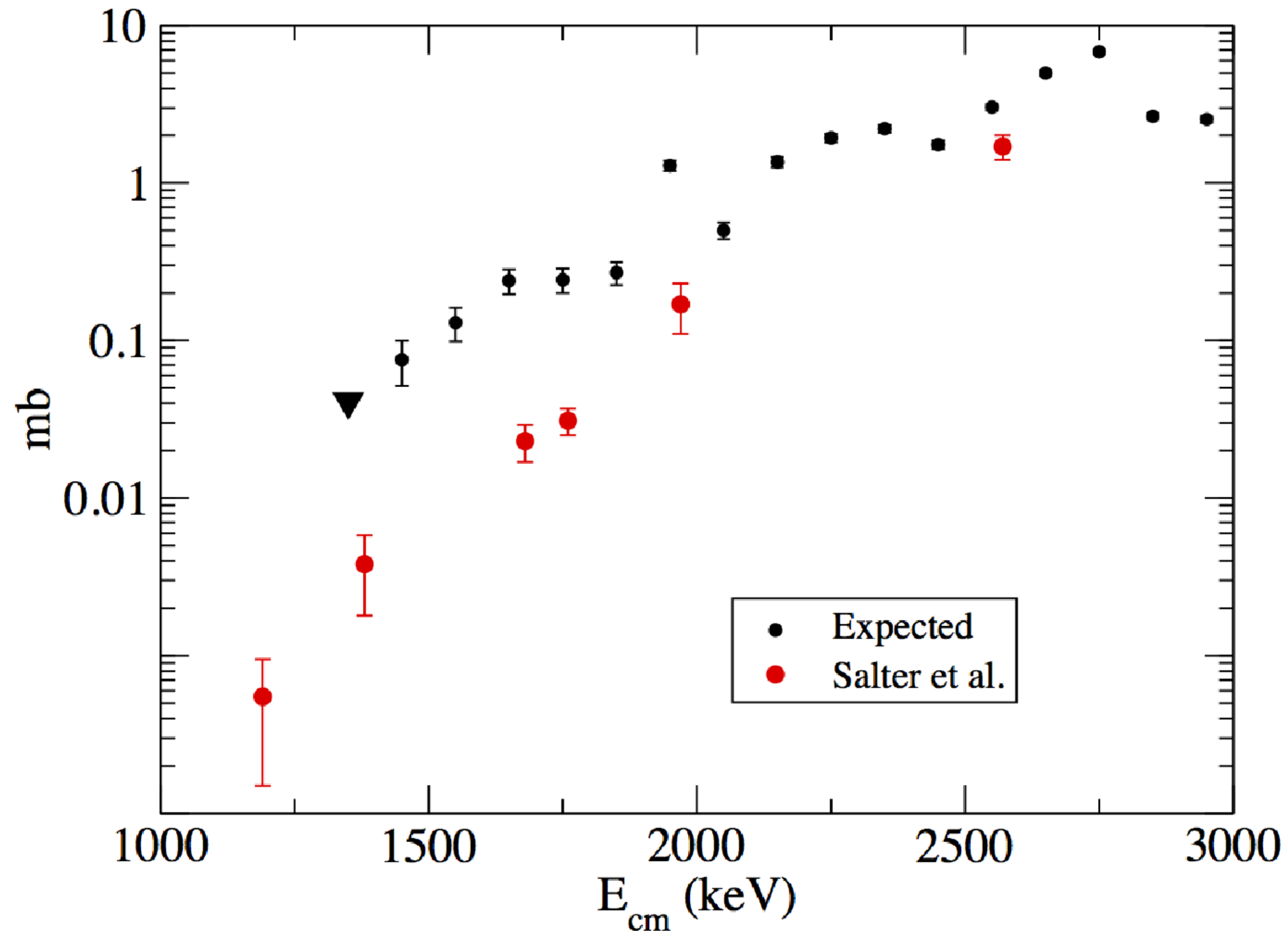
Upgrades for Improved Measurements

- Pure ^4He gas target
- 85- μm Si detectors, 2-mm strips for ΔE (CSIC-Barcelona)
- Thick Si-strip detectors for residual E (Micron)
- Sacrificed solid angle to focus on higher energy protons at forward angles
- Simulated $\delta E_{cm} \approx 150$ keV, although energy- and angle-dependent



Predicted Excitation Function

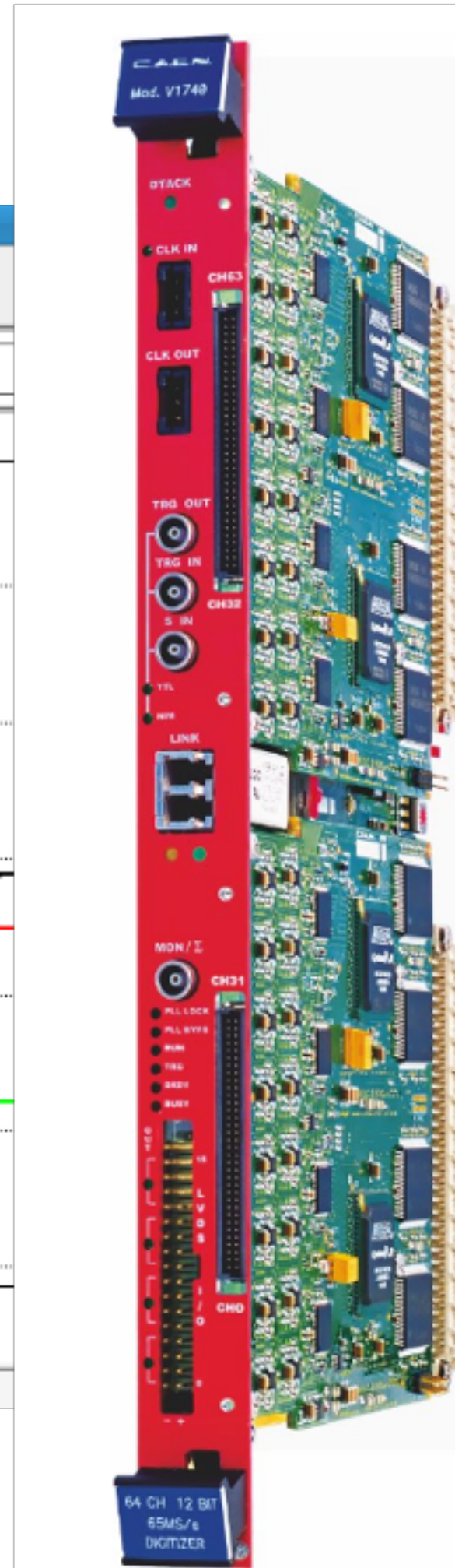
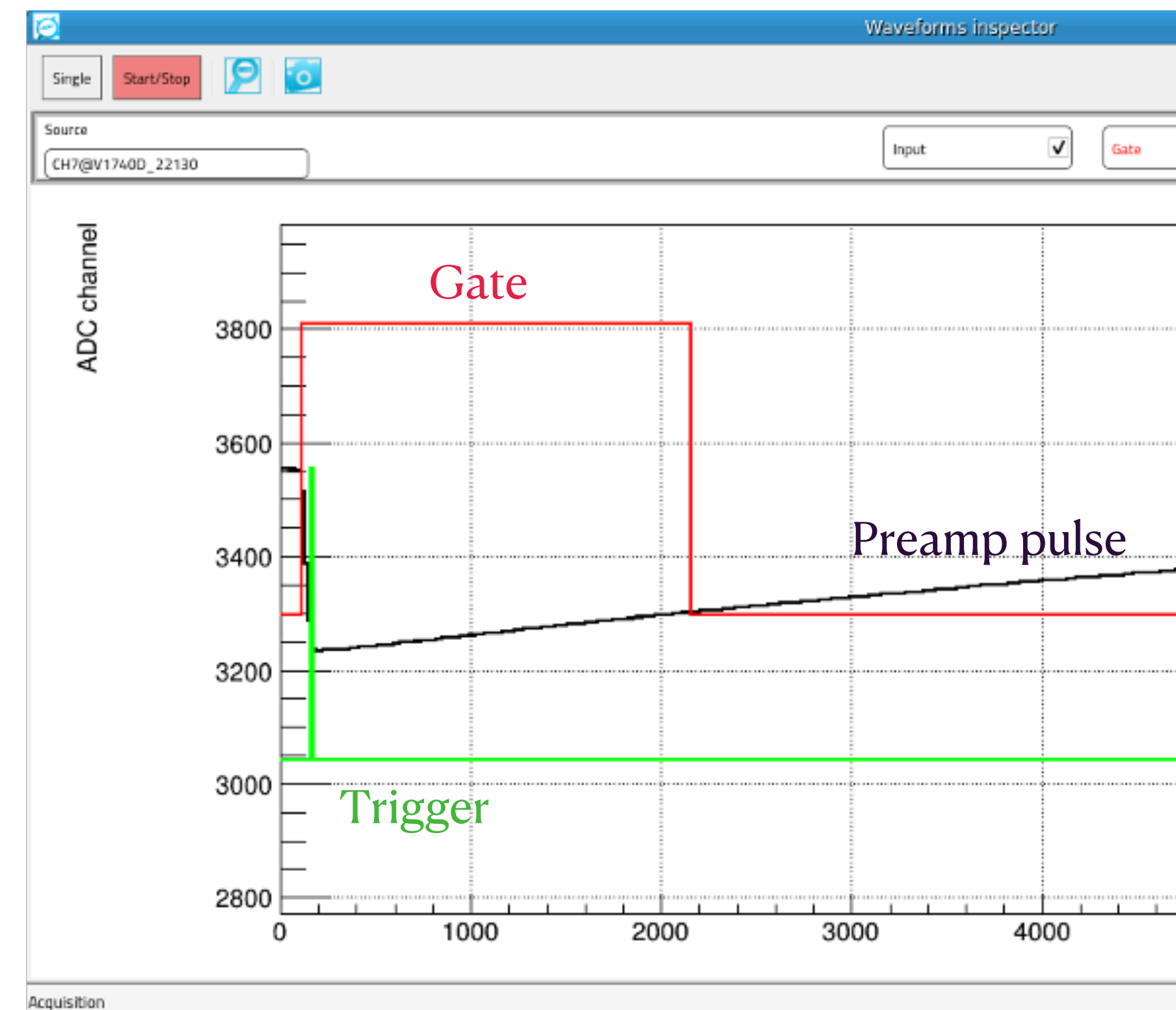
- With >100x intensity at TRIUMF-ISAC I, ^{18}Ne beam scans $E_{cm} = 1.25\text{--}3.25$ MeV



New Digital DAQ for ANASEN



- LSU 72-channel preamp box
- Caen 1740D digitizers (64 channels/module, 12 bit, 62.5 MSamples/sec, ~\$200/channel)
- Caen's CoMPASS software for data acquisition
- QDC firmware integrates preamp pulses
- ~1% energy resolution up to 1 kHz per channel rates
- Record all hits, coincidences analyzed in post-processing

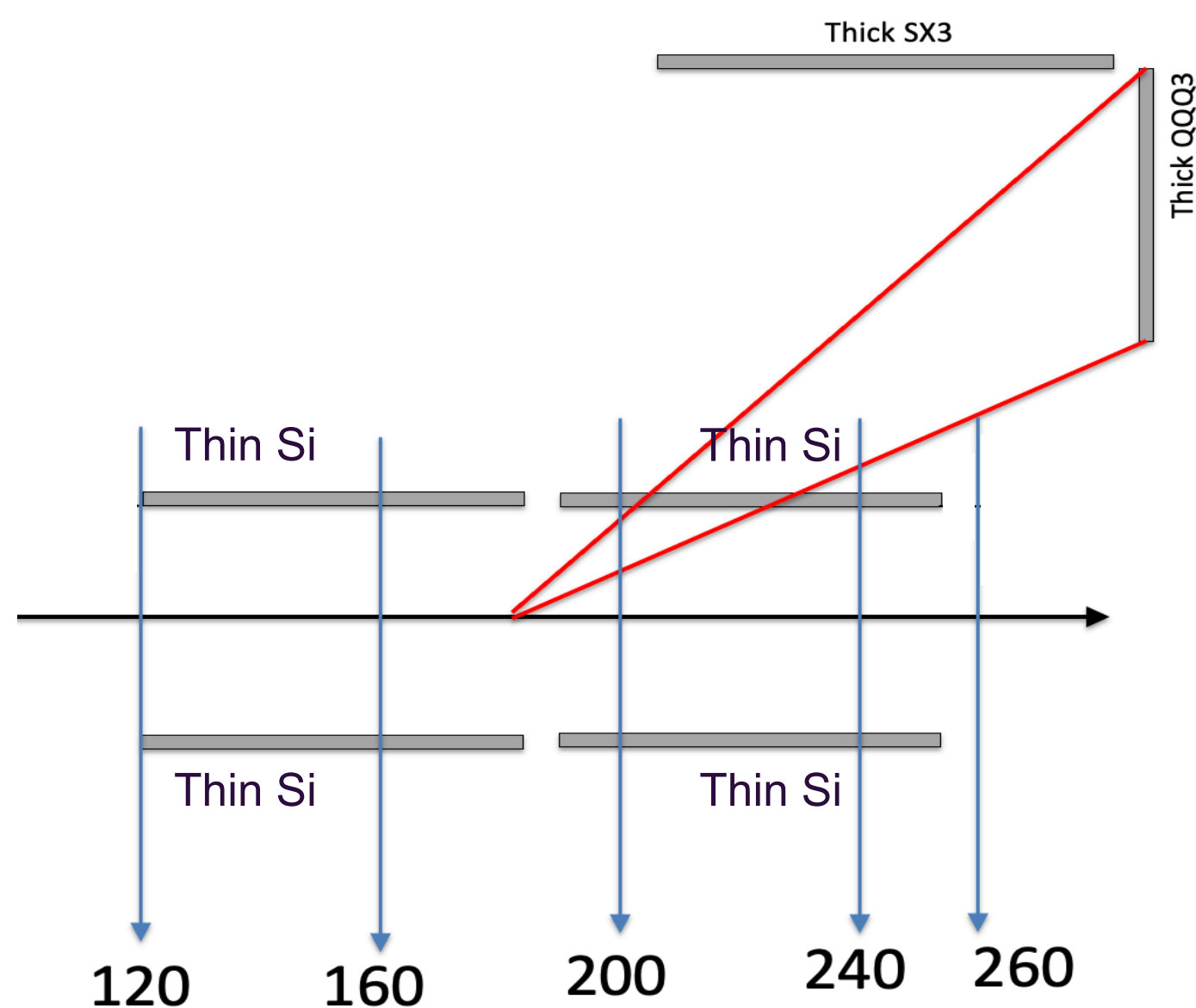




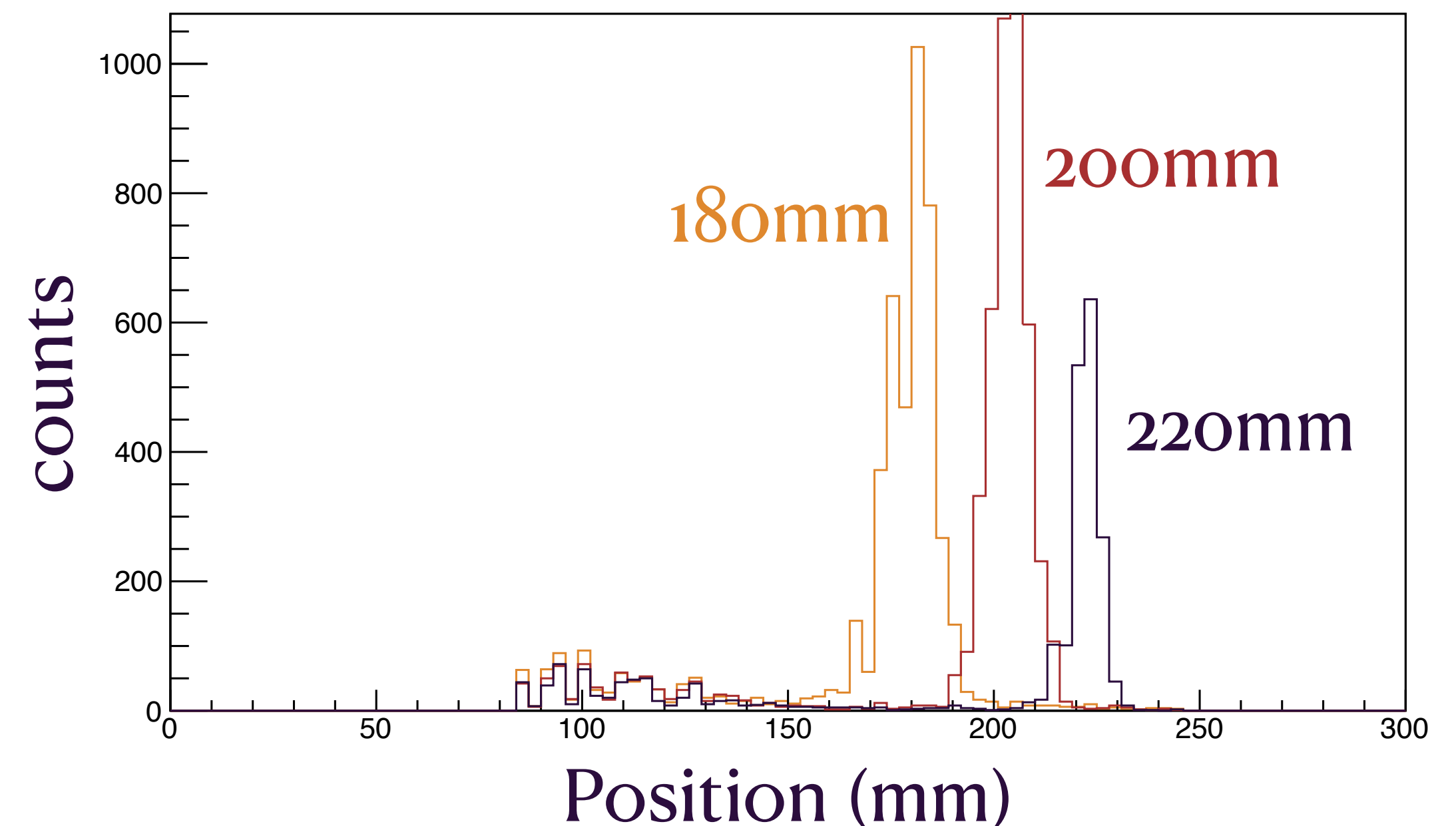
Proton Scattering with Solid ^{181}Ta Target

Position resolution improved to 1 cm FWHM

- 10 MeV protons
- Position resolution improved to 1 cm FWHM, or $\delta E_{cm} \approx 120$ keV for ^{18}Ne
- Caveat: slightly optimistic case with high energy protons at larger angles

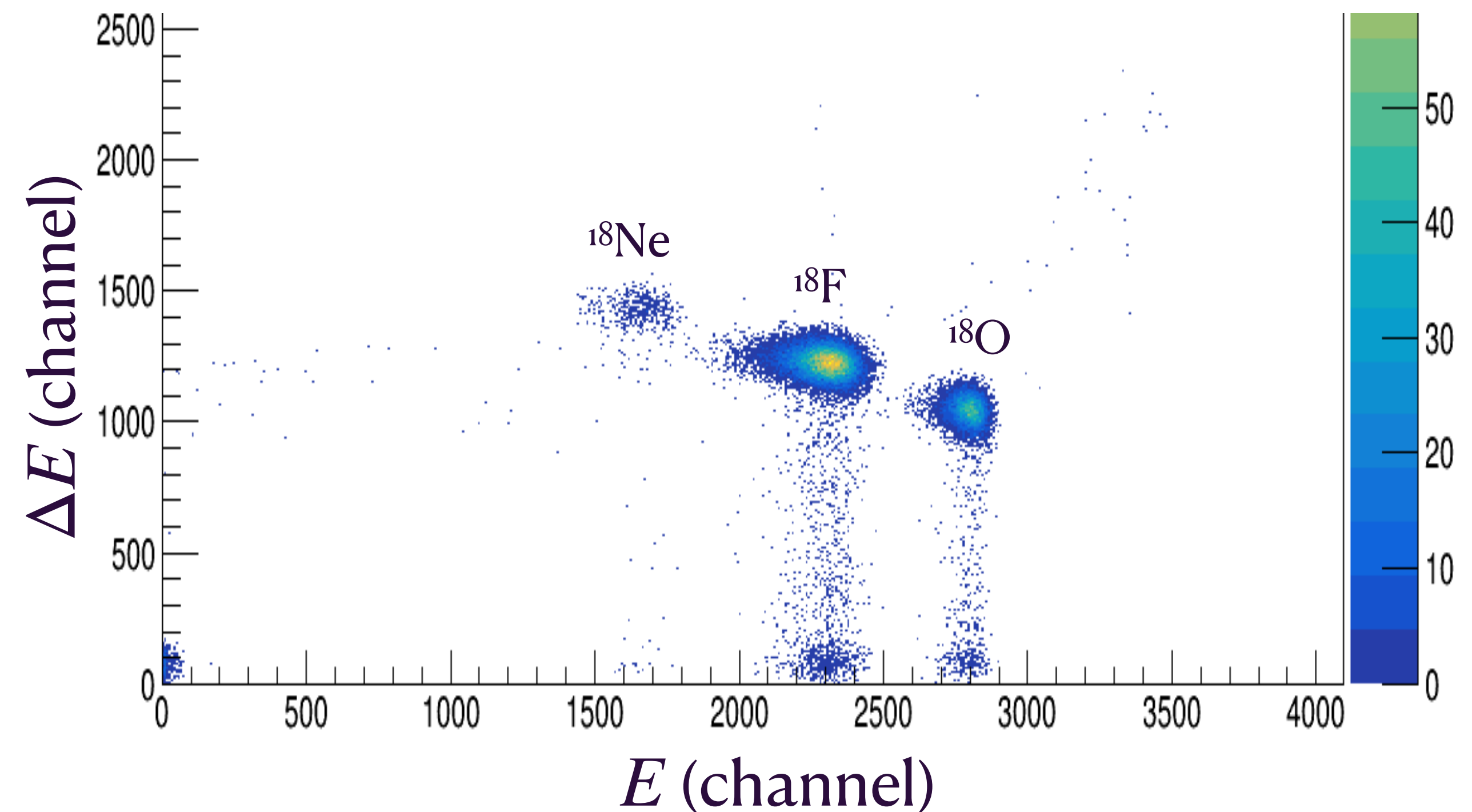


Target distance from window (mm)



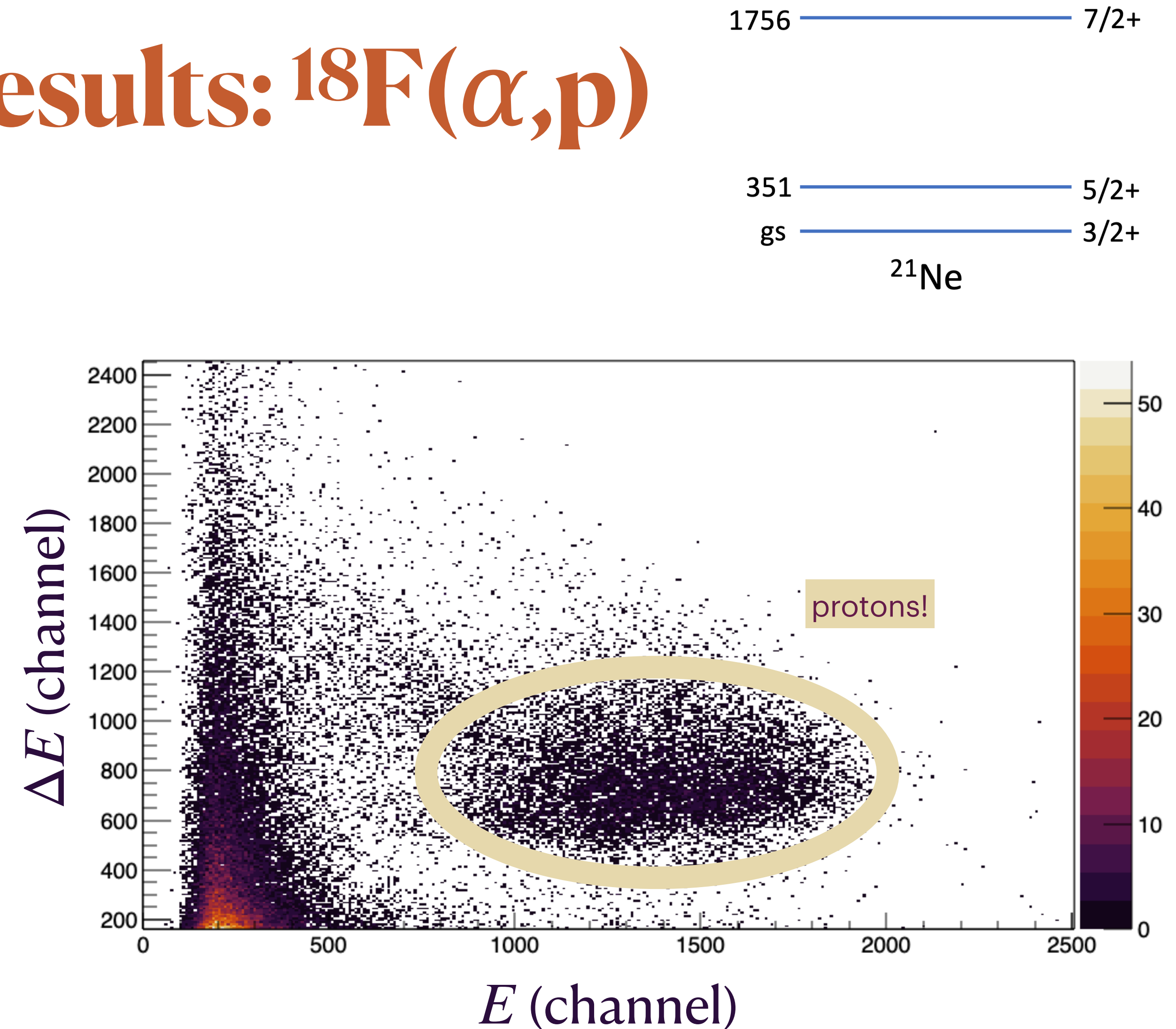
Radioactive Beam: ^{18}F

- Unable to achieve sufficient ^{18}Ne intensity due to technical difficulties
- $^{18}\text{F}(\alpha, p)$ is a potential proton source in AGB nucleosynthesis and accreting white dwarfs if rate closer to upper limit
- Need to measure $^{18}\text{F}(\alpha, p)$ for background in ^{18}Ne experiment
- Optimized ^{18}F beam with ~2:1 ratio $^{18}\text{F}:^{18}\text{O}$ with intensity $\sim 10^5\text{--}10^6$ pps
- Can ignore ^{18}O contamination due to $^{18}\text{O}(\alpha, p)$ Q-value



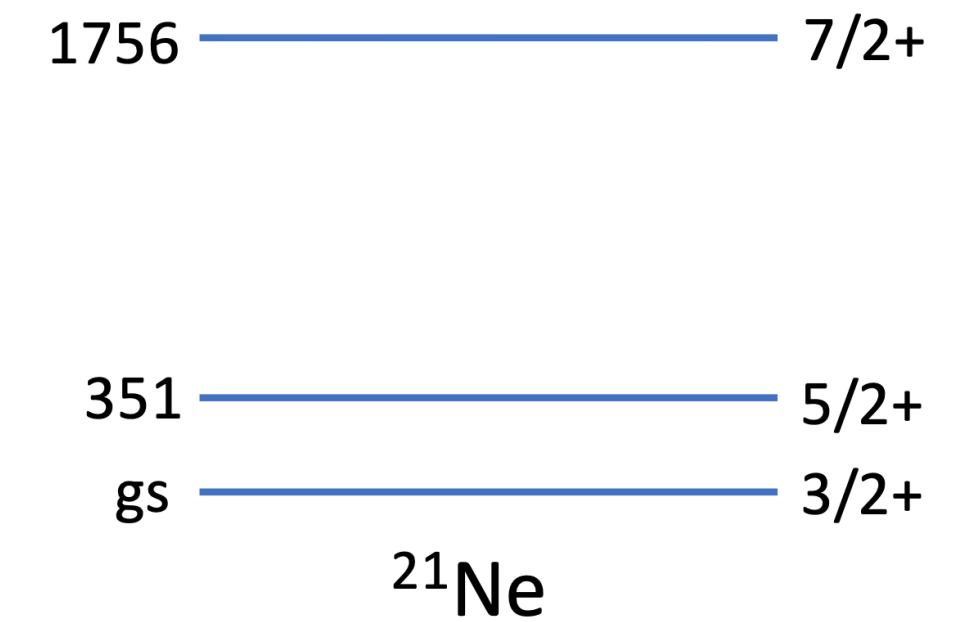
Preliminary Results: $^{18}\text{F}(\alpha, p)$

- Protons clearly identified by ΔE - E
- Expect to populate ground and 1st excited states in ^{21}Ne

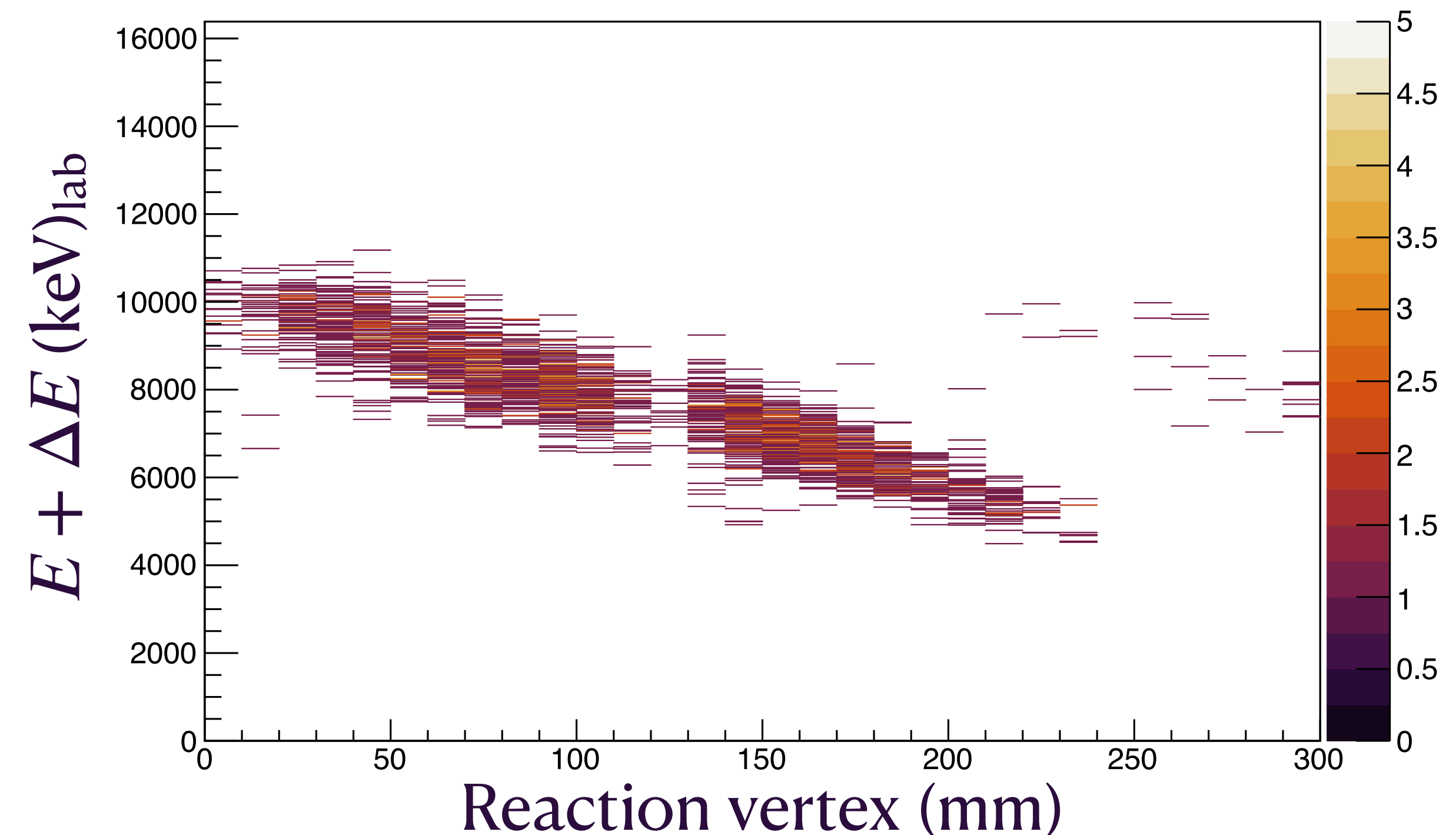


Preliminary Results: $^{18}\text{F}(\alpha, p)$

- Protons clearly identified by ΔE -E
- Expect to populate ground and 1st excited states in ^{21}Ne
- However, it's not clear that we can resolve these states

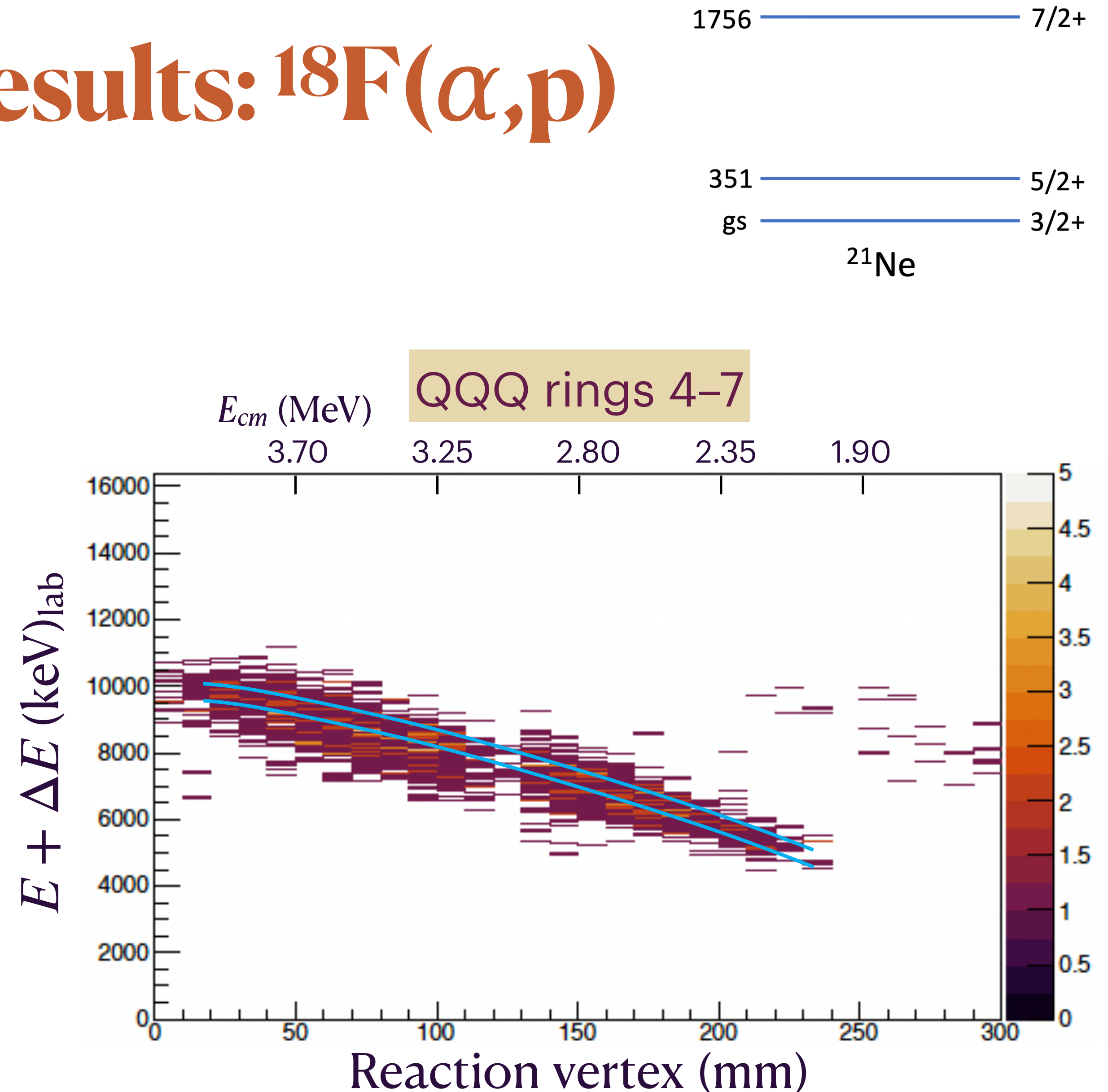


QQQ rings 4-7



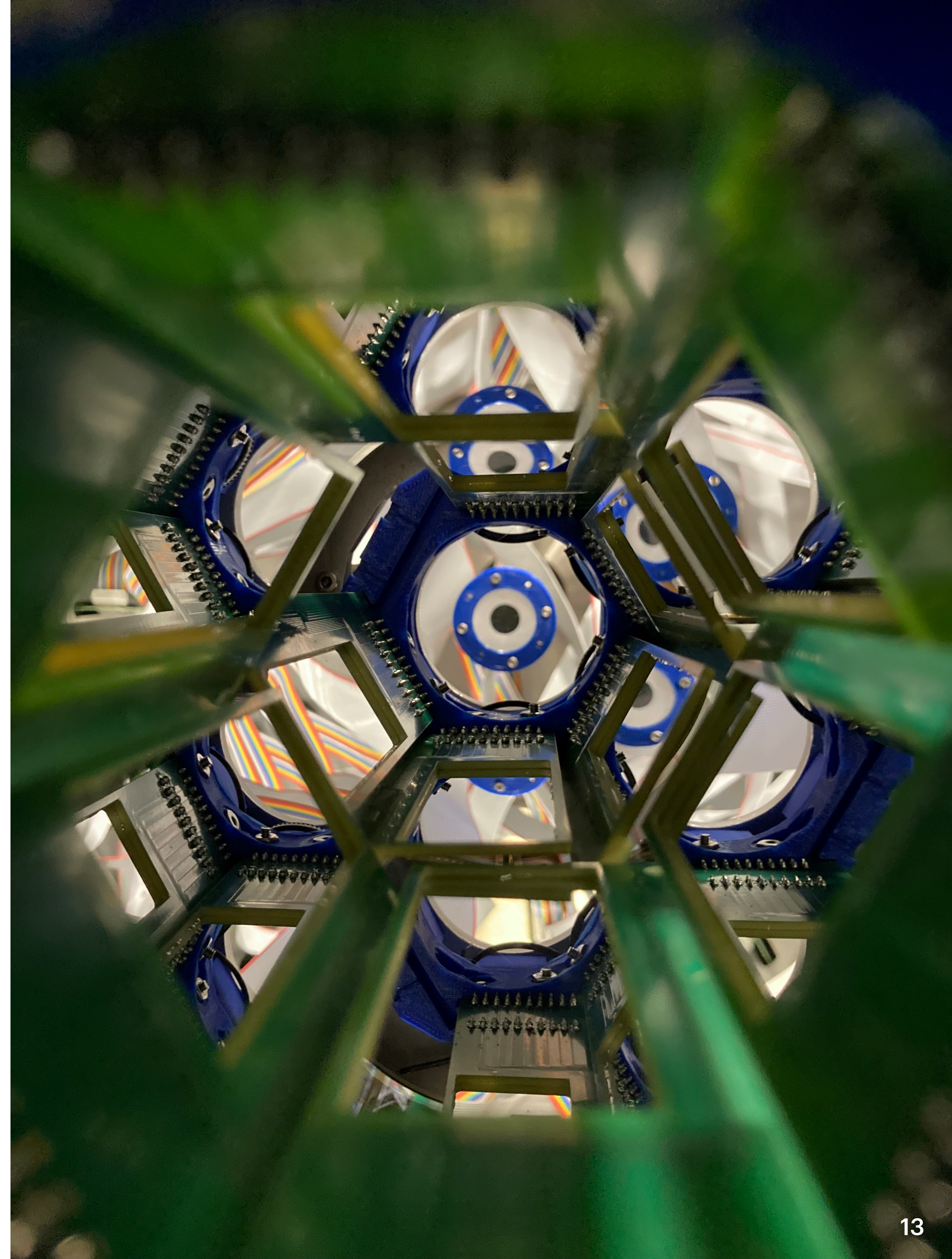
Preliminary Results: $^{18}\text{F}(\alpha, p)$

- Protons clearly identified by ΔE -E
- Expect to populate ground and 1st excited states in ^{21}Ne
- However, it's not clear that we can resolve these states
- Straggling in ΔE pushing our energy resolution closer to 500 keV
- In progress: analyzing SX3 data and extracting excitation function



Conclusions & Future Work

- Digital ANASEN marks an improvement for (α, p) direct measurements
- Straggling may be reduced with updated geometry
- Proposed measurements:
 - $^{14}\text{O}(\alpha, p)$ at Florida State University (Fall 2023)
 - $^{18}\text{Ne}(\alpha, p)$ at TRIUMF
 - $^{30}\text{S}(\alpha, p)$ at FRIB



Acknowledgements

Jeff Blackmon, Catherine Deibel, Gemma Wilson, Sudarsan Balakrishnan
Louisiana State University

Vignesh Sitaraman, Gordan W. McCann, Ingo Wiedenhoefer, Lagy Baby, Chris Esparza
Florida State University

Ivan Tolstukhin
Argonne National Laboratory

Annika Lennarz, Soham Chakraborty
TRIUMF

Jan Glorius
GSI Helmholtzzentrum für Schwerionenforschung

Dan Bardayan, Scott Carmichael
University of Notre Dame

Steve Pain
Oak Ridge National Laboratory

Reuben Russell
University of Surrey

