

AMoRE-I Analysis Result

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On Behalf of the AMoRE Collaboration

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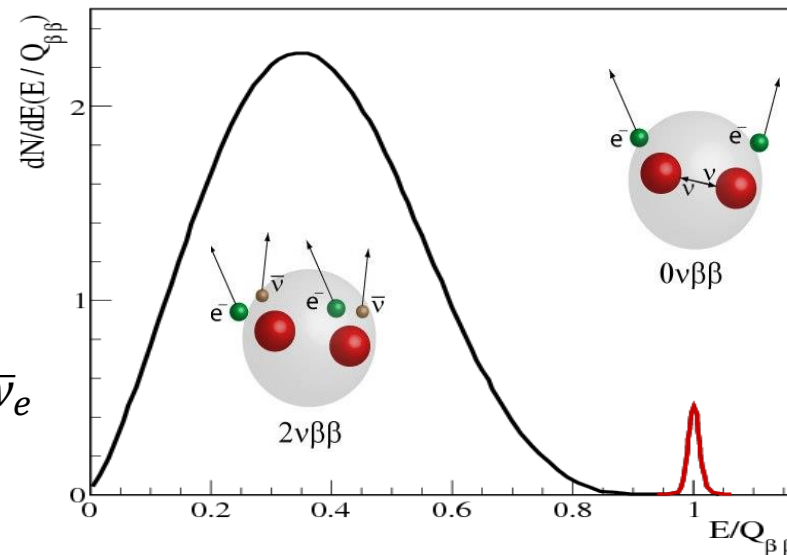
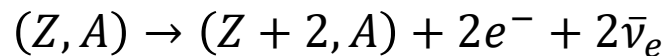
$0\nu\beta\beta$ search using ^{100}Mo

AMoRE:

A search for neutrinoless double beta ($0\nu\beta\beta$) decay of ^{100}Mo using Mo-based scintillating crystals and low-temperature sensors.

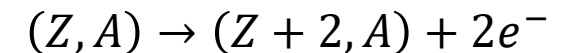
$2\nu\beta\beta$ decay

- 2nd order beta decay
- Rare nuclear decay
- ($>10^{18}$ years of half life)



$0\nu\beta\beta$ decay

- Massive neutrino
- Majorana particle
- Lepton number violation
- Beyond the SM model
- $>10^{25}$ years of half-life

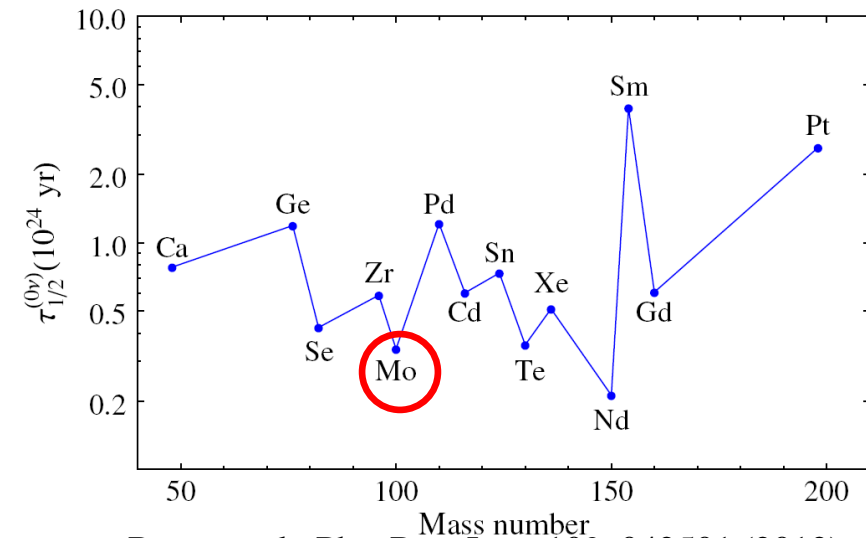


- Neutrinoless double beta decay:
 - Direct measure of Majorana nature of neutrino.
 - Lepton number violation process.
 - Effective neutrino mass.

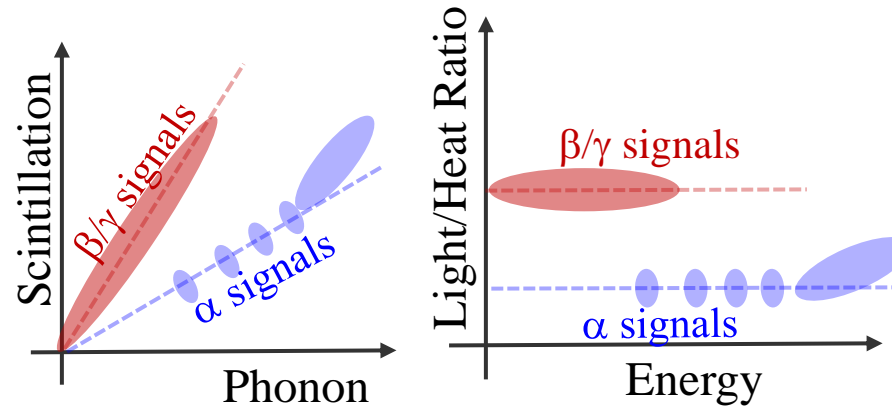
$0\nu\beta\beta$ search using ^{100}Mo

^{100}Mo :

- High $Q_{\beta\beta} = 3034$ keV
- High natural abundance: 9.7 %
- Scintillation crystals with ^{100}Mo enrichment > 95% — XMo_aO_b (XMO):
 - $\text{X}=\text{Ca}, \text{Li}_2, \text{Na}_2, \text{Zn}, \text{Sr}, \text{Pb}, \dots$
 - Detection of light/heat signal \rightarrow rejection of surface- α background.
- Relatively short half life ($0\nu\beta\beta$) in theoretical expectation

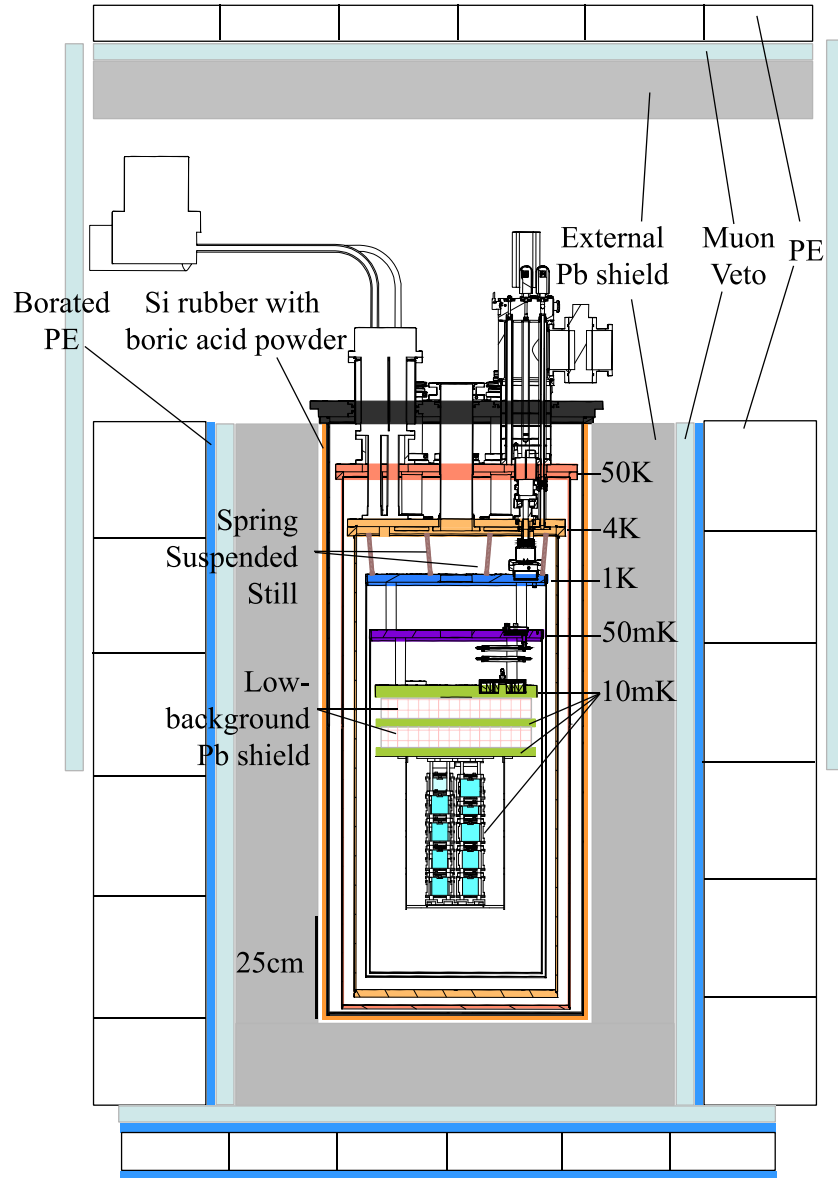


Barea et al., Phy. Rev. Lett. 109, 042501 (2012)

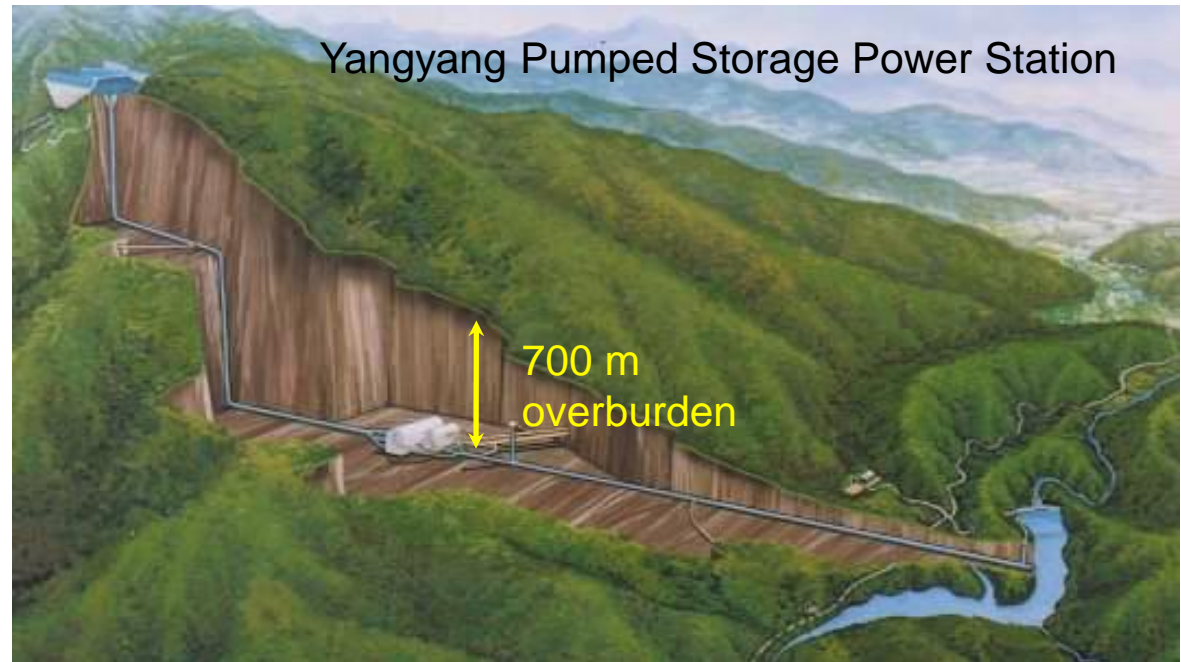


$\beta\beta$ -decay nuclei with $Q > 2$ MeV	Q (MeV)	Abund. (%)
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	4.271	0.187
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	2.040	7.8
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	2.995	9.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Ru}$	3.350	2.8
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	3.034	9.7
$^{110}\text{Pd} \rightarrow ^{110}\text{Cd}$	2.013	11.8
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	2.802	7.5
$^{124}\text{Sn} \rightarrow ^{124}\text{Ge}$	2.228	5.8
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	2.528	34.2
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	2.479	8.9
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	3.367	5.6

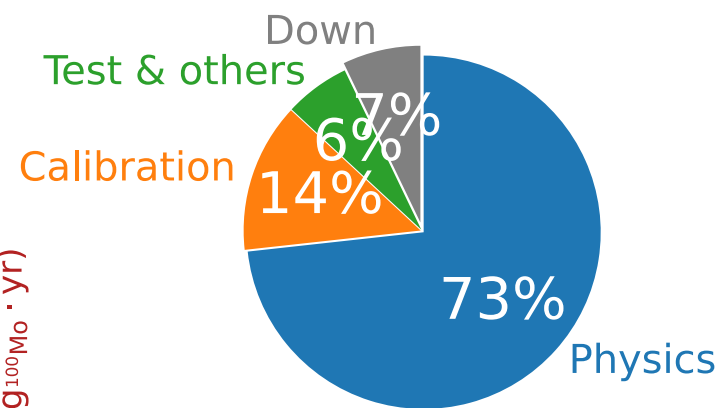
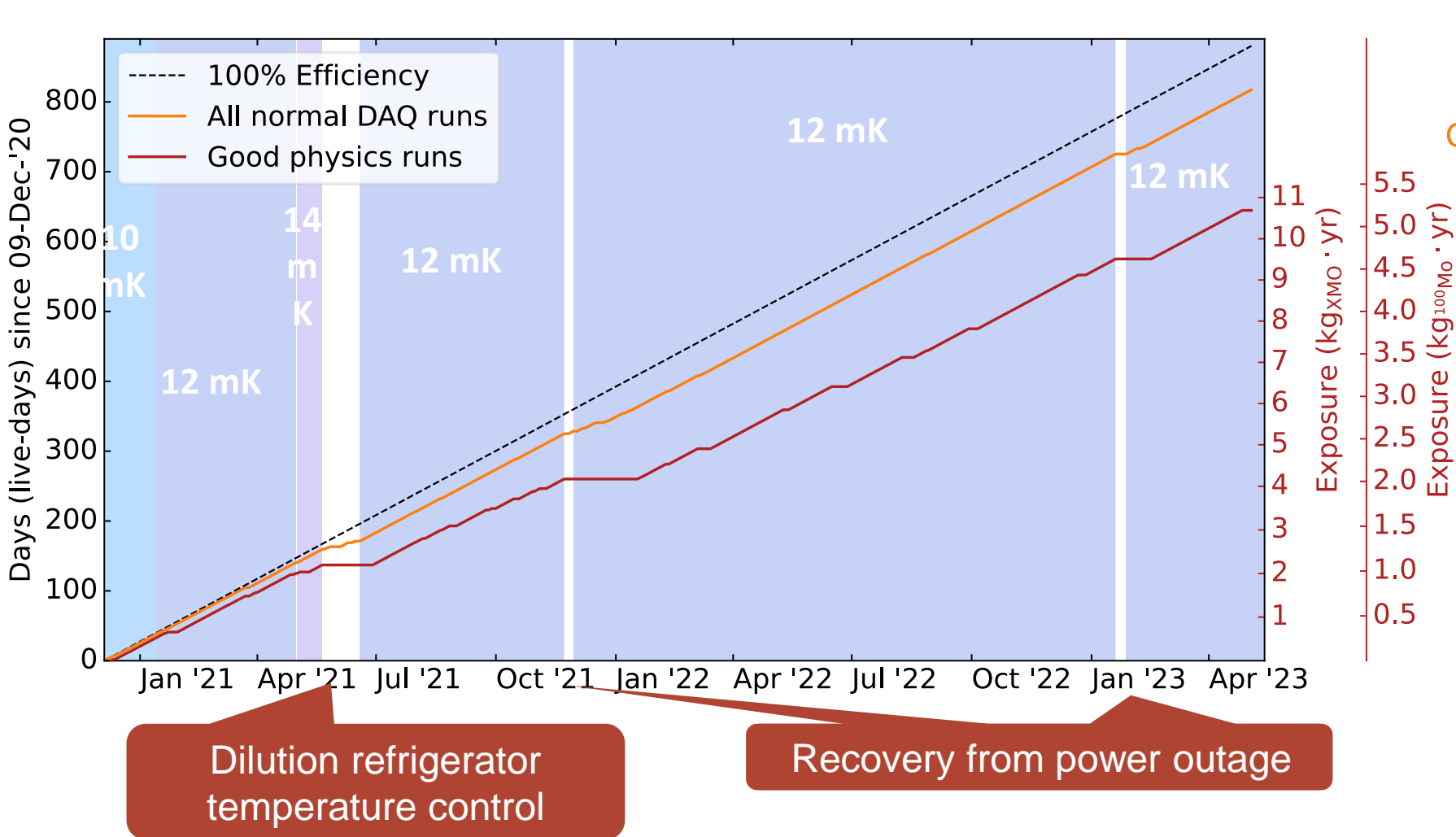
Cryostat, Shielding & Underground



- Cryogen-free dilution refrigerator for AMoRE-pilot (2016~2018) and AMoRE-I.
- $\sim 1 \mu\text{W}$ cooling power.
- Pb (γ), boron, and polyethylene (n).
- Plastic scintillator muon counters.
- Yangyang Underground Laboratory (Y2L) at 700 m depth.

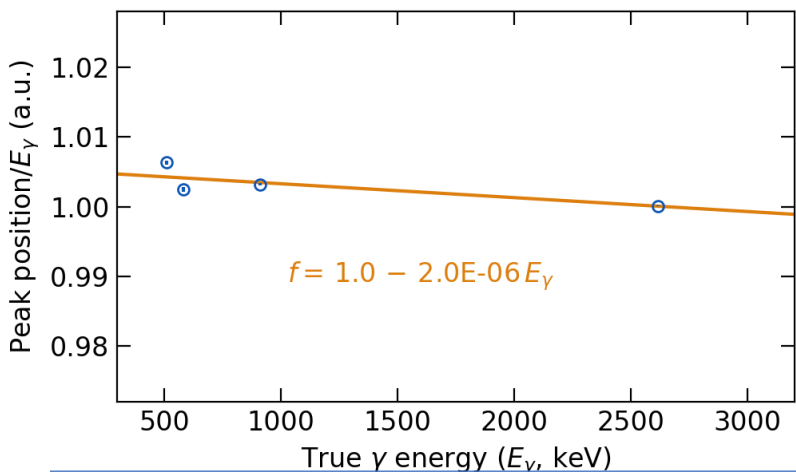
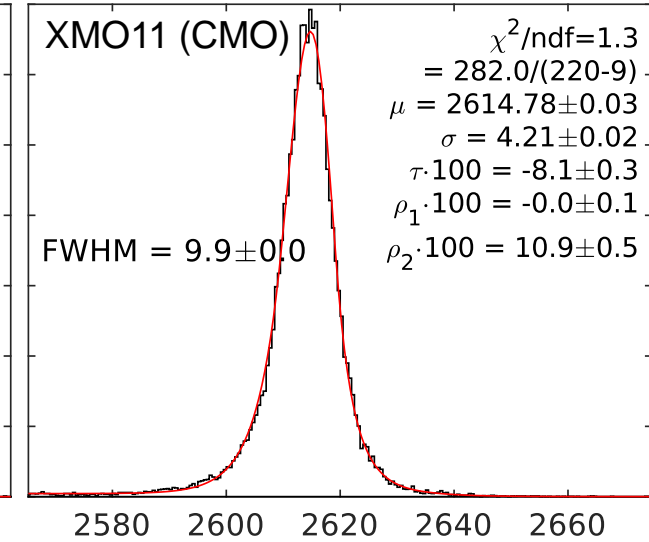
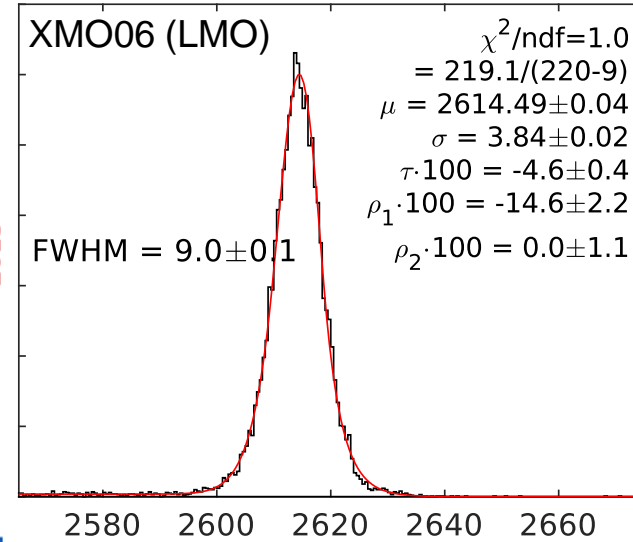
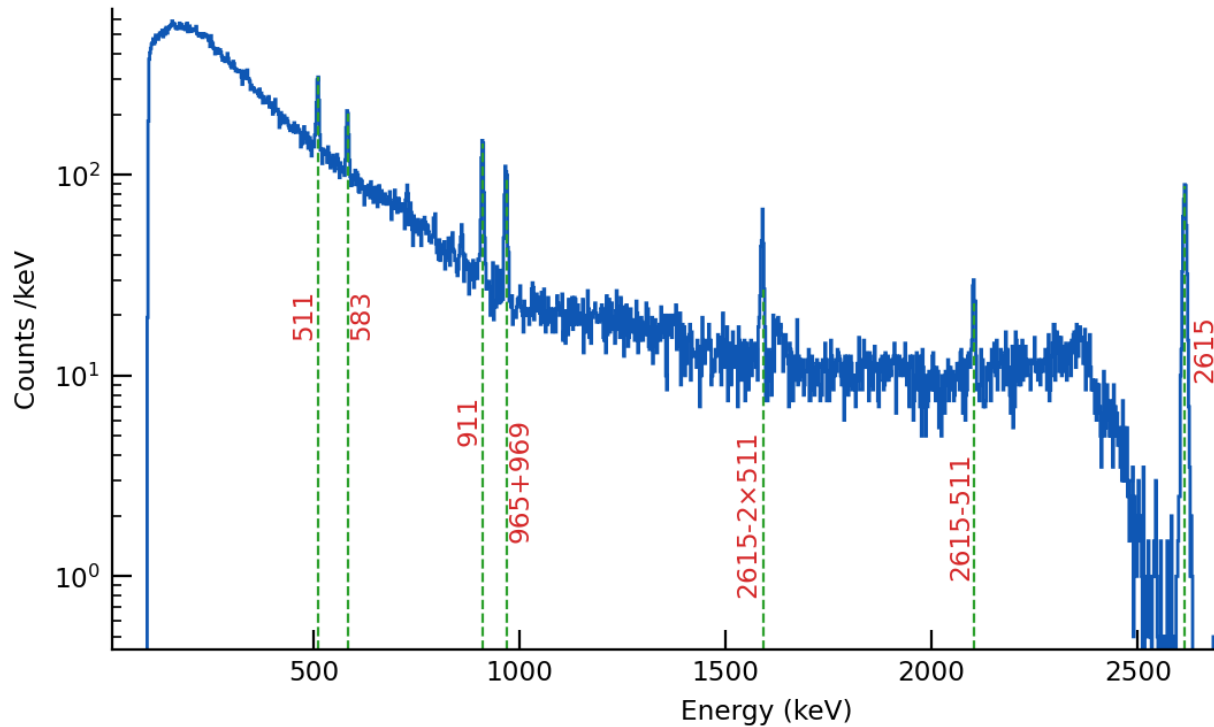


AMoRE-I data taking



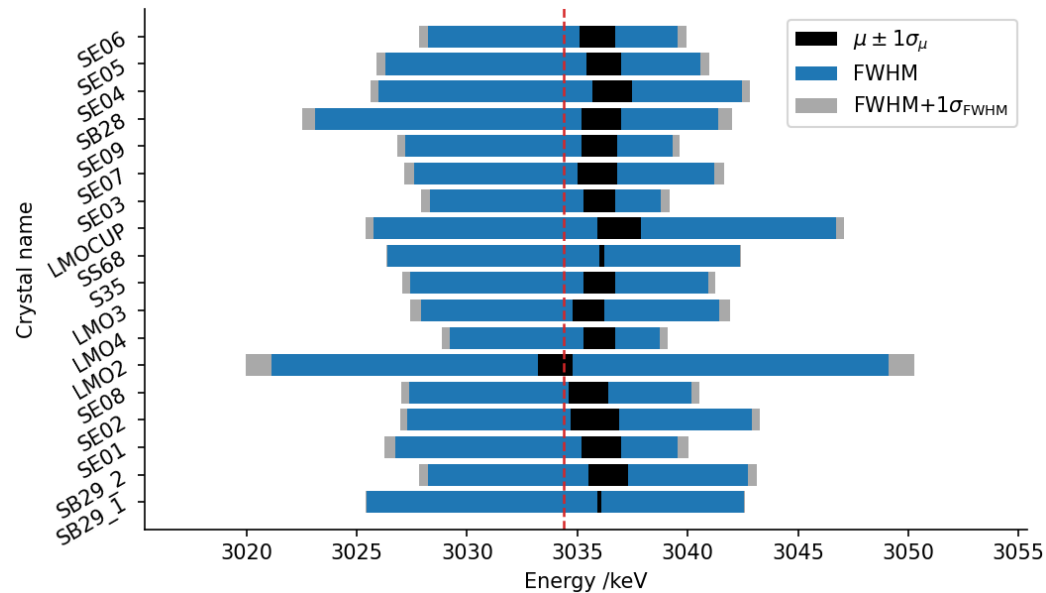
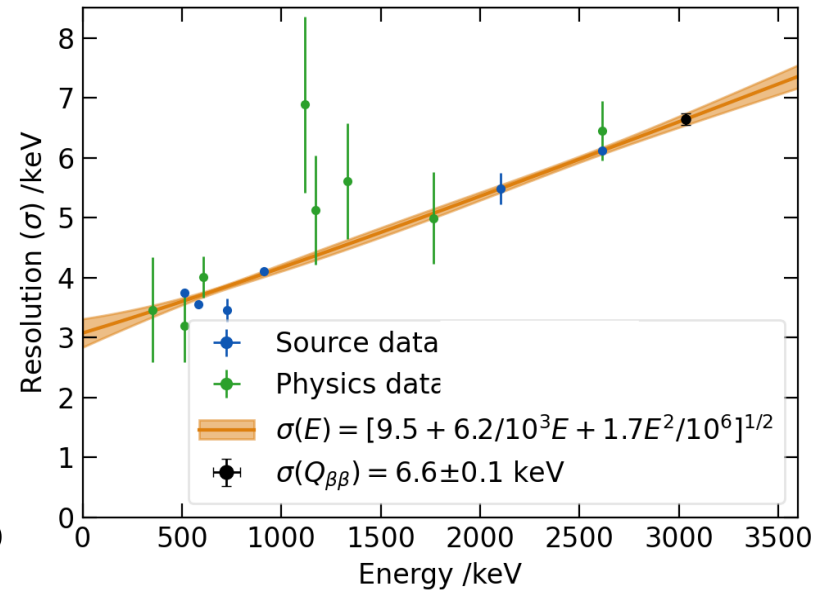
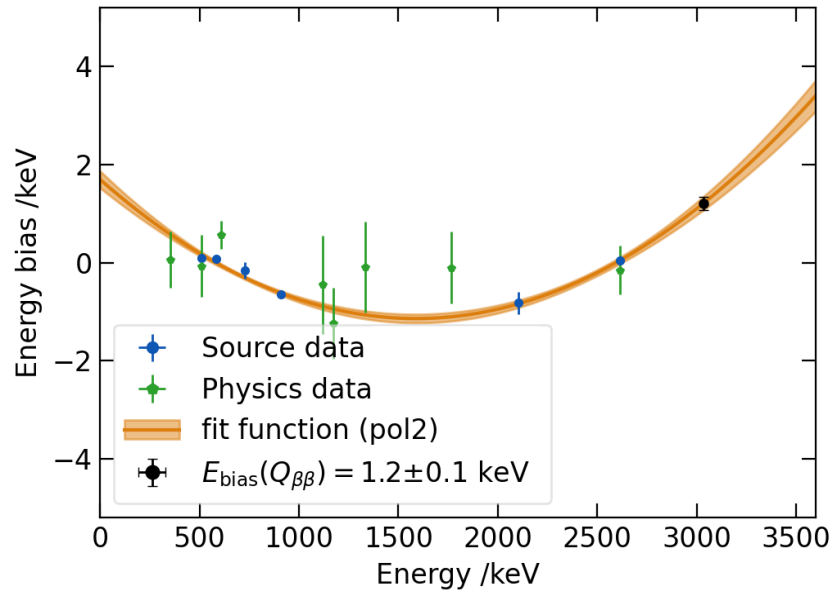
12 mK physics data	Exposure ($\text{kg}_{\text{XMO}} \cdot \text{yr}$)
Total	10.16
- Bad channels/runs	8.76
- deadtime & anti-coincidence (= live exposure)	8.02
in ^{100}Mo	3.88 ($\text{kg}_{\text{ISO}} \cdot \text{yr}$)

Energy Calibration



- Calibration source: ^{232}Th -rich welding rods
- Slight non-linearity between signal amplitude and energy.
- Bukin function instead of gaussian/exponentially modified gaussian – better fit to right tails
<https://root.cern.ch/doc/master/classRooBukinPdf.html>

ROI estimation



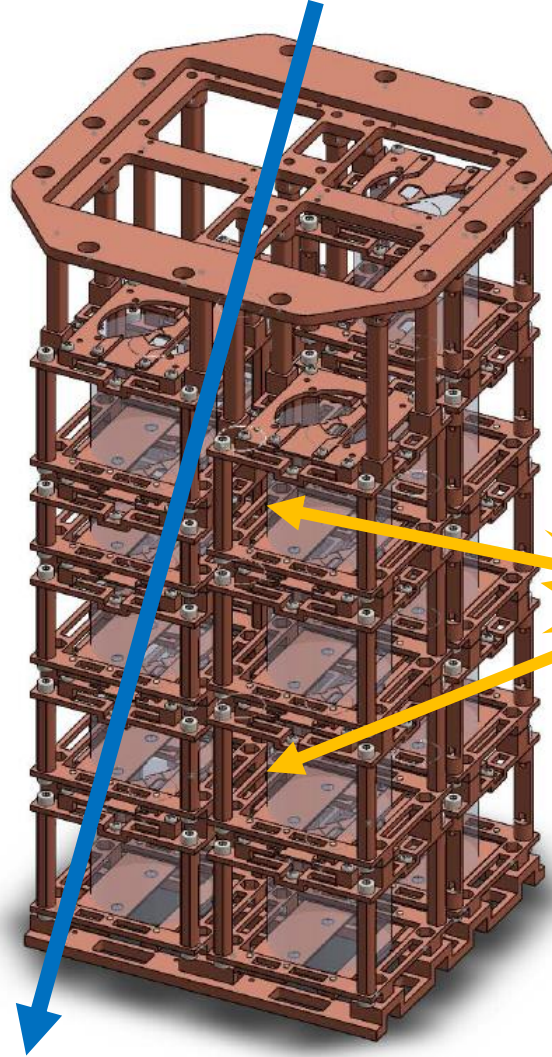
- $E_{\text{bias}} \sim 2^{\text{nd}}$ polynomial of energy
 - $\sigma(E) = \sqrt{b_1 + b_2 E + b_3 E^2}$
- (Fittings done only for source data)



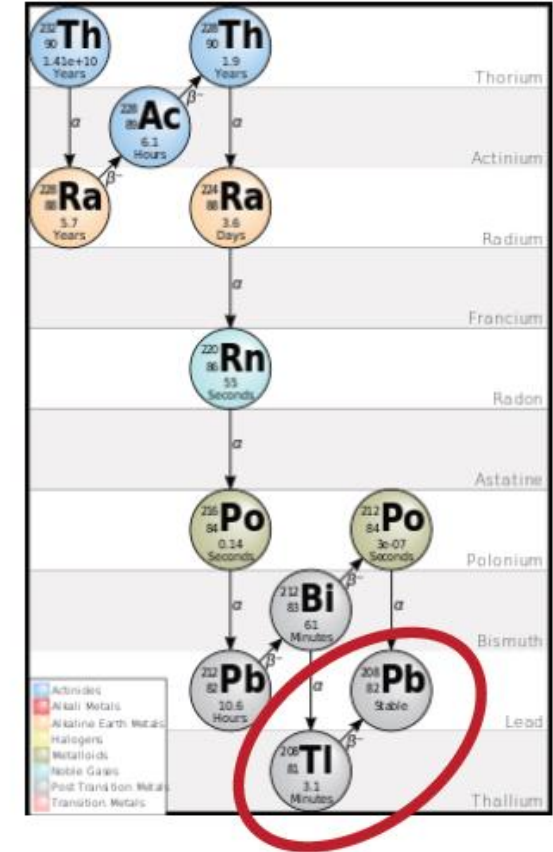
$E_{\text{bias}}(Q_{\beta\beta}): -0.4 \sim 2.5 \text{ keV}$
 $\text{FWHM}(Q_{\beta\beta}): 9.5 \sim 28.0 \text{ keV}$

Particle Identification

Anti-coincidence

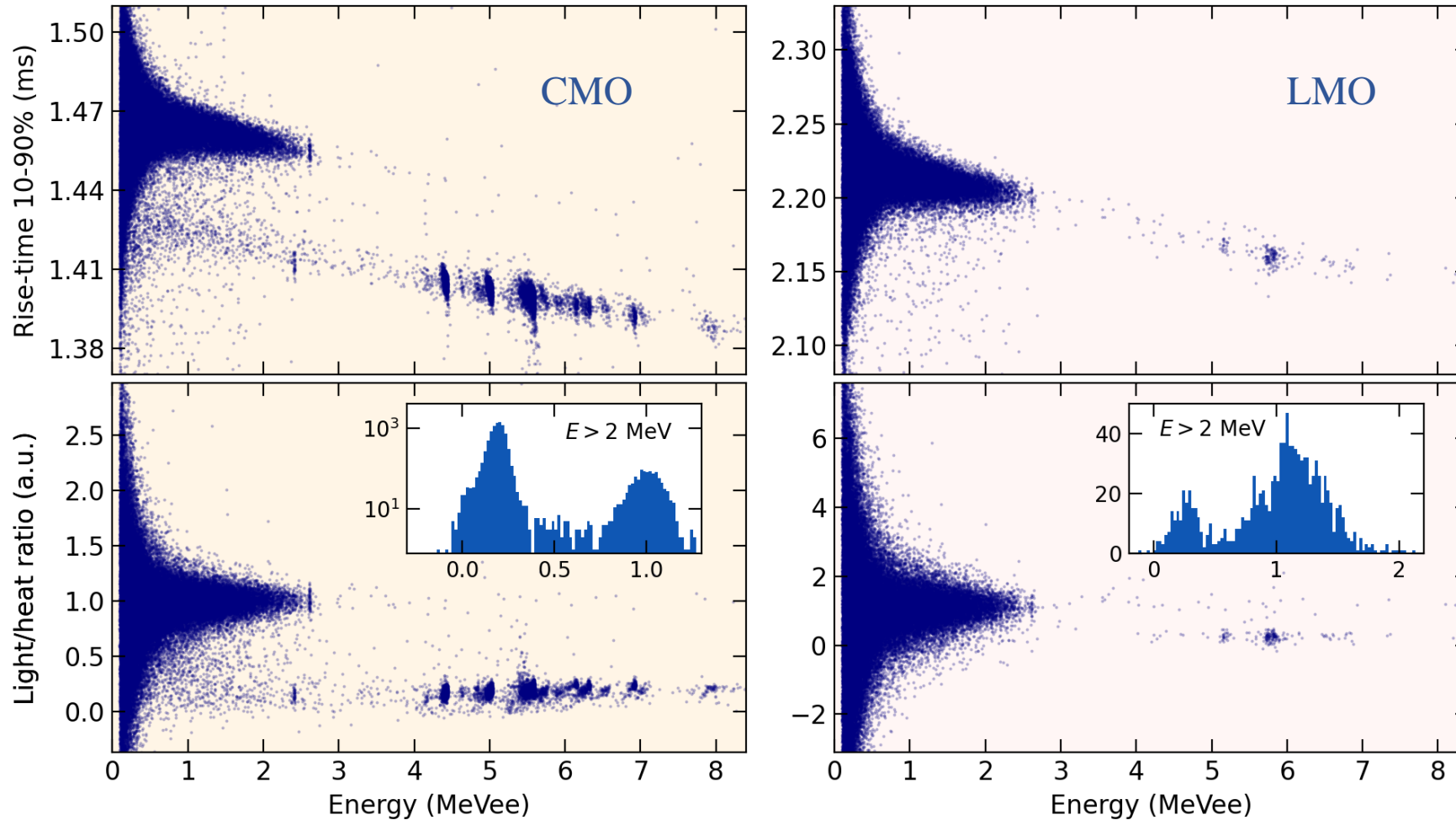


- Multiple hits:
Events within $\pm 2\text{ms}$ time window is thought to be coincidental
- $\epsilon \sim 99.8\%$
- Muon tagging data from Muon Veto System (MVS) installed outside of the cryostat
- $\pm 10\text{ ms}$ windows are rejected
- $\epsilon \sim 99.8\%$



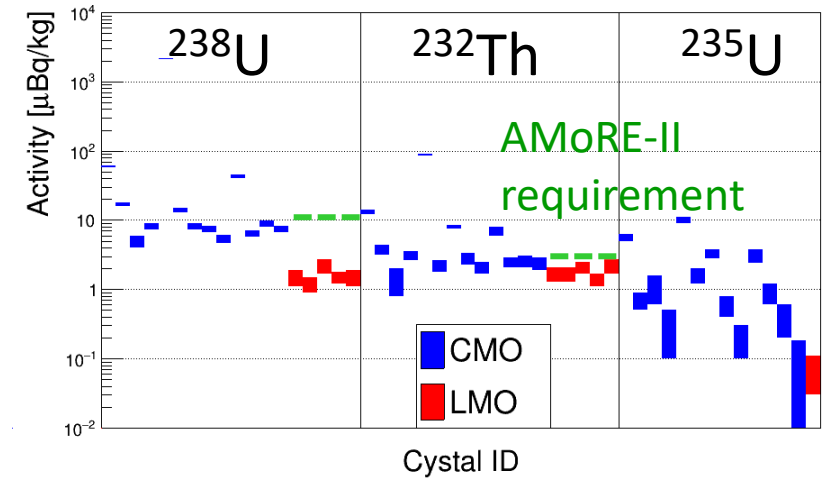
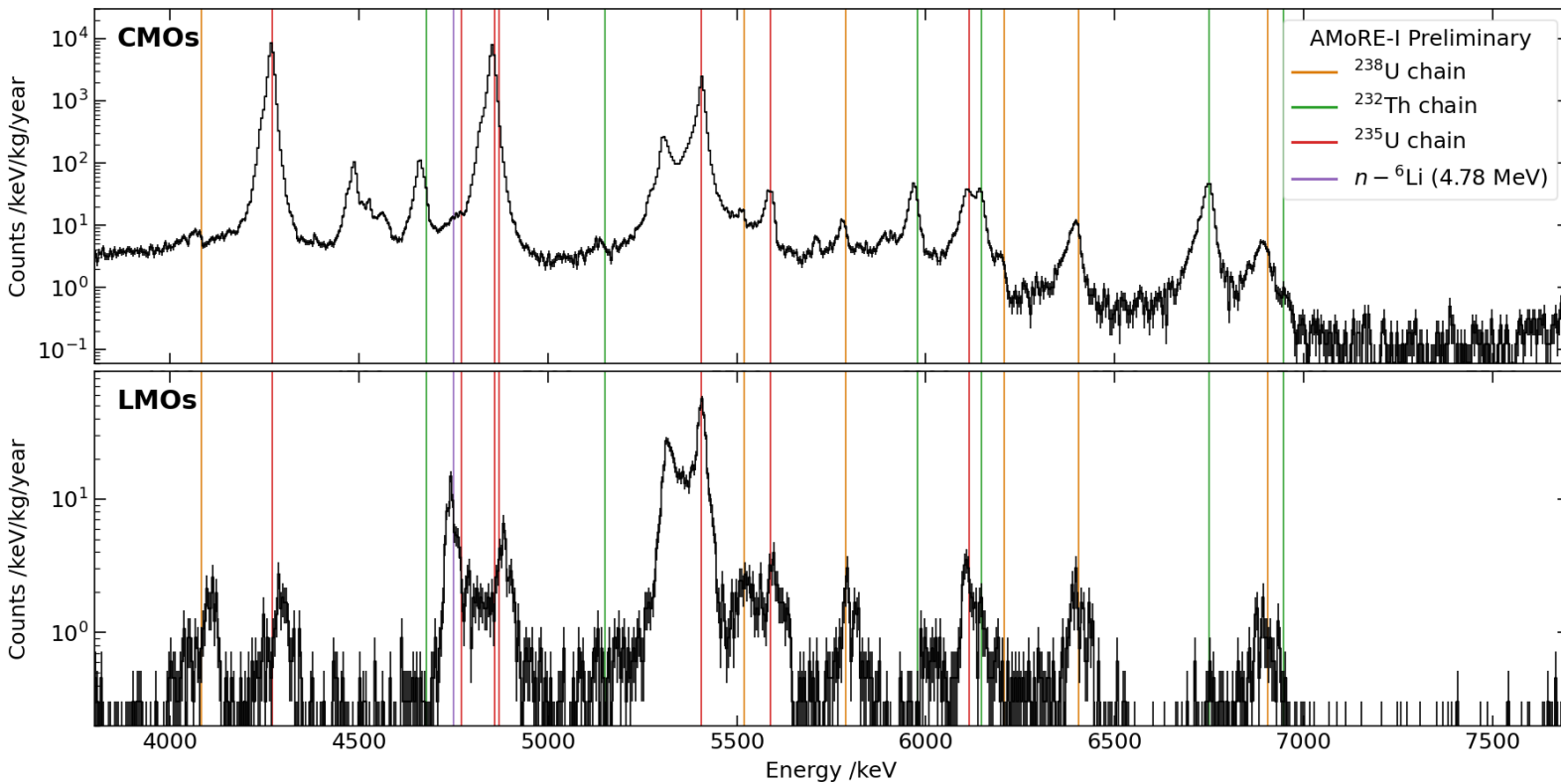
- $\text{Bi-212} \rightarrow \text{Tl-208}$ (α 35.94%): $T_{1/2}=60.55\text{ min}$, 6207.26 keV
- $\text{Tl-208} \rightarrow \text{Pb-208}$ (β): $T_{1/2}=3.053\text{ min}$, $Q \sim 5\text{ MeV}$
- 20 min. window after each Bi-212 candidates are rejected
- $\epsilon \sim 98\%$ ($>99\%$ for many crystals, $\sim 82\%$ for the worst case (SB29))

Particle Identification (PaID)

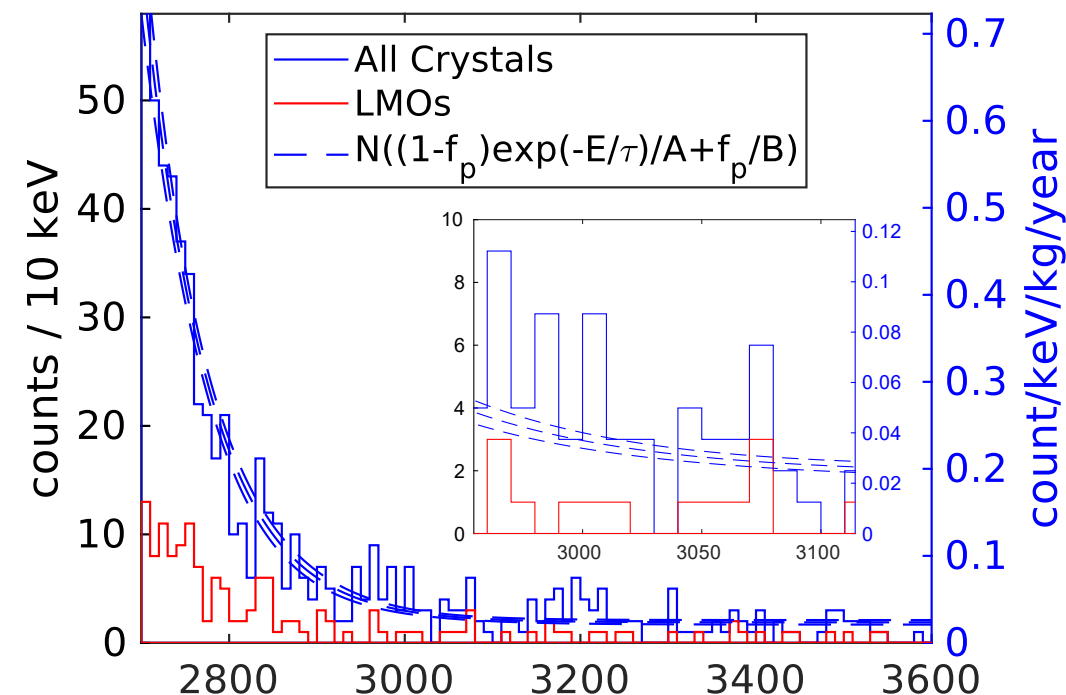
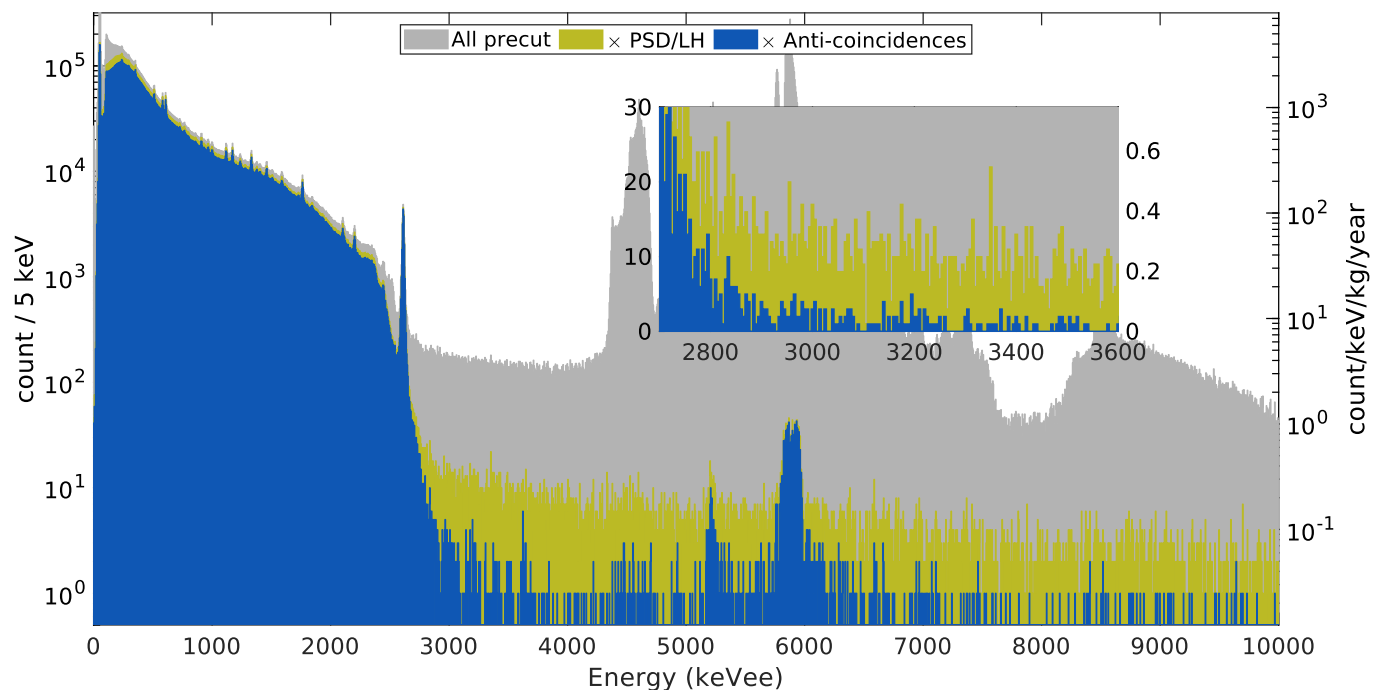


- CMO shows better discrimination power thanks to a higher light yield.
- $\epsilon_{\text{PaID,ROI}} = 92.9\sim 99.2\%$ with ± 3 median absolute deviations (MAD) range of PSD & L/H (91.6 % if normally distribution that ± 3 MAD ($\sim \pm 2\sigma$) gives 95.70 % C.L.)

α background

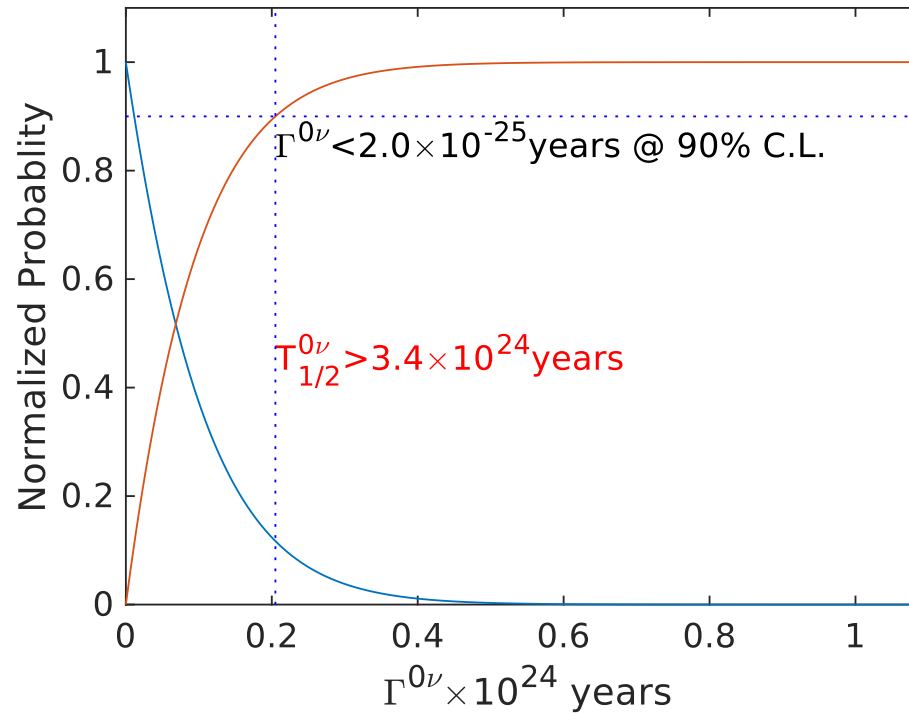


- Overall, the radioactive contamination by U/Th of LMO is measured to be substantially lower than that of CMO.



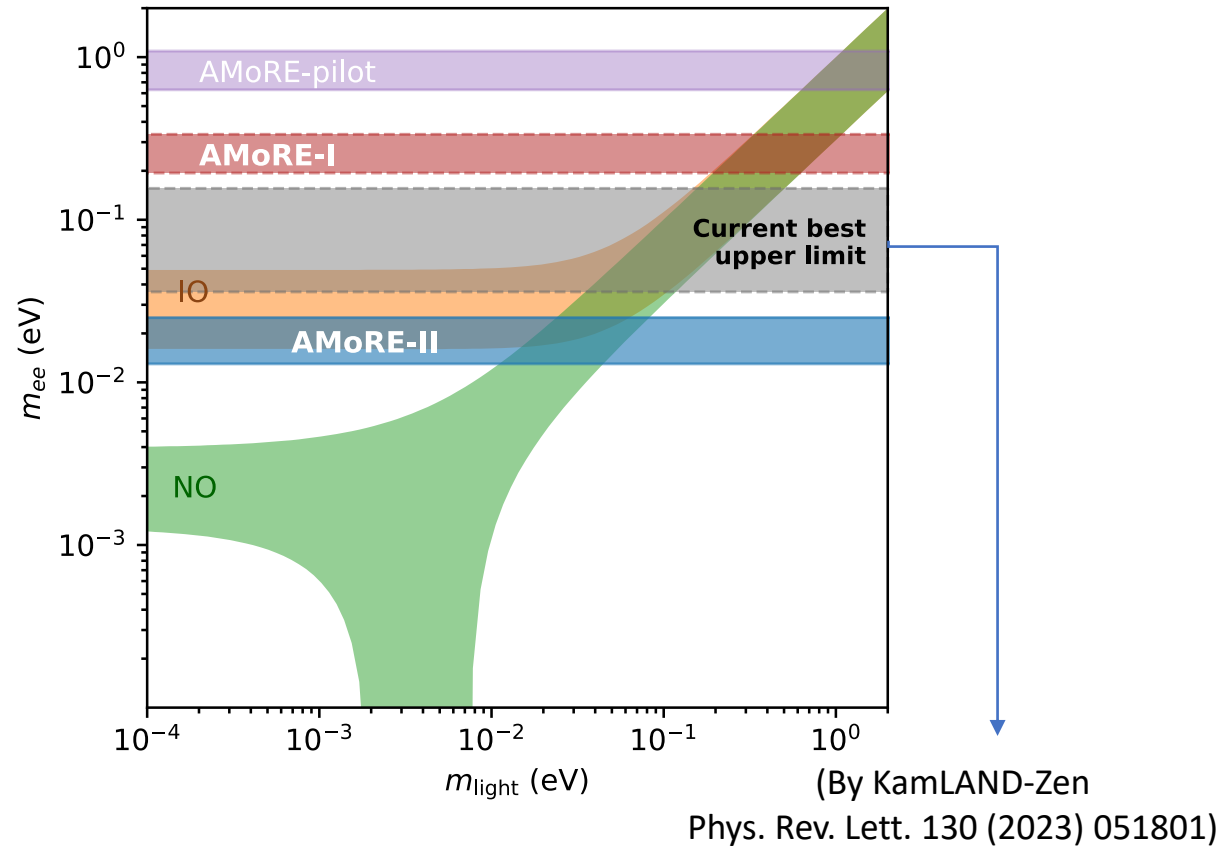
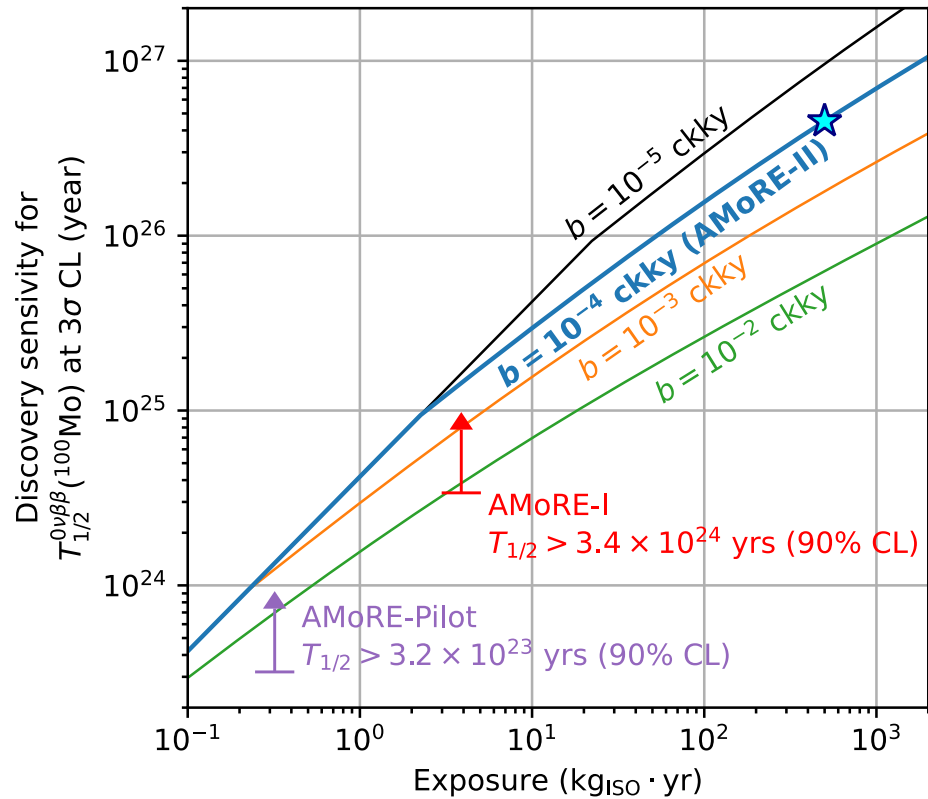
- All crystal excluding one LMO (for very poor β/α discrimination power)
 - 13 CMO + 4 LMO: exposure = $8.02 \text{ kg}_{\text{XMoO}_4} \cdot \text{yr} = 3.88 \text{ kg}^{100\text{Mo}} \cdot \text{yr}$.
- Anti-coincidence cuts reject events:
 - coincident at multiple crystals within 2 ms ($\epsilon \sim 99.8\%$),
 - within 10 ms after a muon counter event ($\epsilon \sim 99.8\%$),
 - within 20 minutes after a ^{212}Bi α -decay event candidate ($\epsilon \sim 98\%$).

Live exposure	Bkg. @ $Q_{\beta\beta}$ / ckky
Total (8.02 $\text{kg}_{\text{XMoO}_4}$ yr)	0.032 ± 0.003
CMO (6.19 $\text{kg}_{\text{XMoO}_4}$ yr)	0.031 ± 0.003
LMO (1.83 $\text{kg}_{\text{XMoO}_4}$ yr)	0.037 ± 0.006



- $\text{ROI} = |E - Q_{\beta\beta}| < 2.5 \Delta E_{\text{FWHM}}$, $\epsilon_{\text{containment}} \sim 81\%$.
- Background = 0.032 ± 0.003 counts/keV/kg/year, from ROI side-band.
- Unbinned likelihood for $\Gamma^{0\nu}$ ($= \ln 2 / T_{1/2}$) for each crystal, with signal shape and background rate constrained from calibration and sideband data, respectively.
- $T_{1/2}^{0\nu} > 3.4 \times 10^{24}$ years at 90% C.L.

Limits & Sensitivities



- AMoRE-II for $T_{1/2}^{0\nu} > 5 \times 10^{26}$ years by 100 kg of $^{100}\text{Mo} \times 5$ years running.
- Reduction of background level down below 10^{-4} ckky.

- AMoRE searches for $0\nu\beta\beta$ using ^{100}Mo based scintillating crystals at the low temperature.
- AMoRE-I took data for 29 months.
- Li_2MoO_4 shows lower discrimination power than CaMoO_4 with either PSD parameter or light/heat ratio while the alpha rate is much lower.
- Background rejections such as muon veto, multiple-hit-tagging and alpha(Bi212)-tagging efficiently suppress the background at ROI.
- Result of AMoRE-I:
 - Mass \times time live exposure: 8.02 (3.88) kg \cdot yr XMoO_4 (^{100}Mo).
 - Background level ~ 0.032 counts/keV/kg/year at $Q_{\beta\beta}$.
 - Resolution: 9.5 – 28.0 keV at $Q_{\beta\beta}$
 - PaID efficiency is better than normally distributed case ($>91.6\%$)
 - Half-life limit: $T_{1/2}^{0\nu} > 3.4 \times 10^{24}$ yr.

Thank you!