Collective excitations in spherical Mo isotopes

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We journey to a new era of discoveries – far beyond the valley of stability



F. Nunes, Physics Today 74, 5 (2021)



Proton-neutron asymmetry in nuclei leads to the formation of a neutron skin



F. Nunes, Physics Today **74**, 5 (2021)

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Neutron skin is linked to exotic dipole excitations, pygmy dipole resonances





Dipole mode evolution reveals a link to neutron skin thickness and PDRs



PDR strength below 15 MeV varies from close to 0 % for the ¹⁶O to about 16 % for ²²O



PDR strength below 10 MeV smoothly increases from ¹⁰⁰Sn to ¹³²Sn

We need more systematic studies of dipole mode evolution across various nuclei





Where is the exact location of the PDR excitation energy? What gives us clues to predict its position?

Is the PDR a collective excitation or resonance structure/oscillation? And how does its degree of collectivity evolve as neutron numbers increases?



Nuclear pygmy modes serve as doorways to nucleosynthesis

Neutron capture rates of nuclei involved in the astrophysical *r*-process



Neutron capture cross sections for $^{205}Pb(n,\gamma)^{206}Pb$



Mo isotopes are ideal for studying single-particle and collective excitations



U We investigate Mo isotopes, which exhibit diverse behaviors with excess protons and neutrons

- With 42 protons, Mo isotopes are close to the semi-magic number 40, making them ideal for studying both single-particle or collective excitations
- □ Mo isotopes participate in multipole astrophysical processes, including *s*-, *r*-, and *p*-processes



Collectivity is analyzed in terms of two quasiparticle configurations within HFB+QRPA

Nucleon-nucleon interaction: Gogny D1M force



Mean field + pairing correlations





Transition densities describe dynamics of each excitation



To treat collective excitations like GDRs and PDRs, we compute transition matrix elements from the ground state to excited states $\langle \psi_{e.s.} | \hat{O}_{E1} | \psi_{g.s.} \rangle$ where $\hat{O}_{E1} = \frac{eZ}{A} \sum_{n=1}^{N} r_n Y_{1m}(\hat{r}_n) - \frac{eN}{A} \sum_{p=1}^{Z} r_p Y_{1m}(\hat{r}_p)$, Then, main results:

- 1) Dipole response functions
- 2) Transition densities of each excited state





Skin evolution: neutron skins for ⁹⁰⁻⁹⁸Mo and proton skins for ⁸²⁻⁹⁰Mo



- □ Skin thickness is quantified as $\delta r = \sqrt{\langle R_n^2 \rangle} \sqrt{\langle R_p^2 \rangle}$
- \square ⁹⁰Mo serve as a reference due to its nearly zero skin thickness
- $\hfill \hfill \hfill$
- □ The figures illustrate:
 - (a) A gradual increase in neutron skins from ⁹⁰Mo to ⁹⁸Mo
 - (b) A gradual increase in proton skins from ⁹⁰Mo to ⁸²Mo from 5 to 8 fm relative to ⁹⁰Mo



Small enhancements of dipole strengths appear in the lower tail of GDR





Transition densities reveal exotic pygmy character of states



The dynamics of the PDR completely differ from those of the GDR !

□ GDR: protons and neutrons out of phase

DPDR: neutrons and protons in phase inside and at the surface only neutrons survive



We identify pygmy candidates linked to different types of skins



We observe proton pygmy candidates in isotopes with a proton skin, and neutron pygmy candidates in isotopes with a neutron skin



Microscopic structure of the pygmy states differs from the GDR



- The percentages indicate contributions of major 2qp configurations
- The structure of the pygmy states is dominated by specific 2qp transitions, whereas the giant states correspond to excitations from many different transitions
- For example, the structure of a giant dipole state in ⁹⁴Mo shows a mixing of many distributed 2qp configurations: 25% of (1f7/2, 2d5/2), 19% of (1f5/2, 1g7/2), 11% of (1f9/2, 1h11/2), 10% of (1f7/2, 1g9/2), etc
 In contrast, the structure of a pygmy dipole state in ⁹⁴Mo
- In contrast, the structure of a pygmy dipole state in ⁹⁴Mo shows a few major 2qp configurations:

66 % of (2d5/2,2f7/2), 11% of (2d5/2,3p3/2), etc

The states near 16 MeV in ^{82,84}Mo show single particle behavior



We plan another exotic dipole excitations, toroidal dipole resonances





Summary and Outlook

Exploring neutron- or proton-rich nuclei far from stability presents many interesting phenomena, e.g. neutron skins

- We have investigated the electric dipole response of Mo isotopes within Gogny HFB+QRPA framework to explore the existence and characters of pygmy dipole states
 - Identified candidates of the PDR states and their microscopic structure through transition densities
 - Studied the relationship between pygmy dipole states and neutron and proton skins
 - Examined collective behaviors associated with proton and neutron skins

U We plan to:

- Investigate PDR states in deformed nuclei
- Study additional excitation mode, toroidal dipole resonances
- Studying collective excitations in astrophysically relevant nuclei, such as Mo isotopes, will deepen our understanding of the origin of heavy elements in the cosmos.



Thank you

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