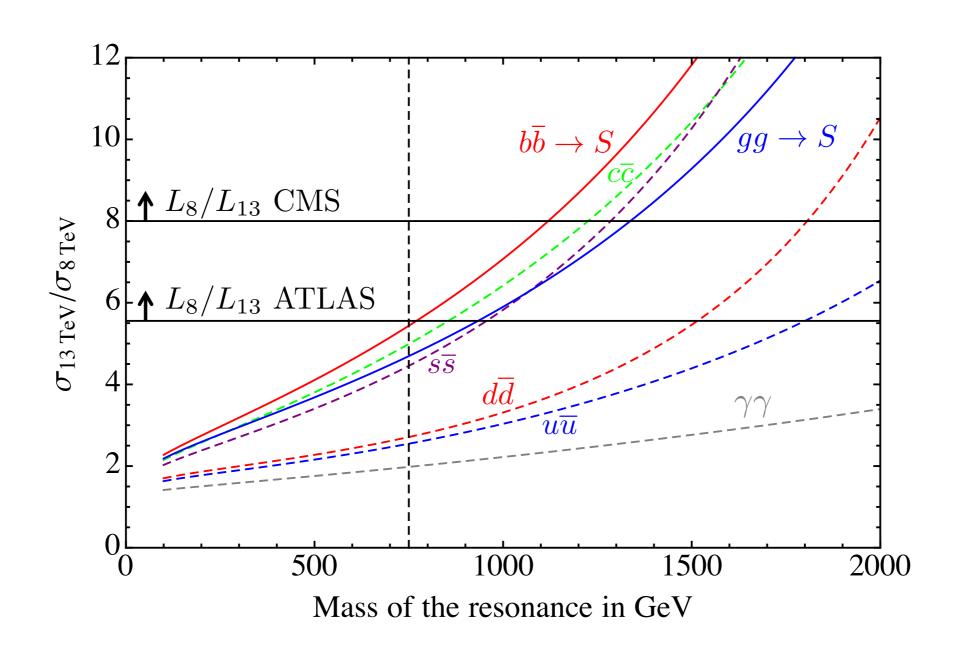
# Di-photon returns…

Mihoko Nojiri (KEK & IPMU )

based on a work with
Chengcheng Han, Koji Ichikawa,
Shigeki Matsumoto Michihisa Takeuchi
(1602.08100. will be published in JHEP soon)

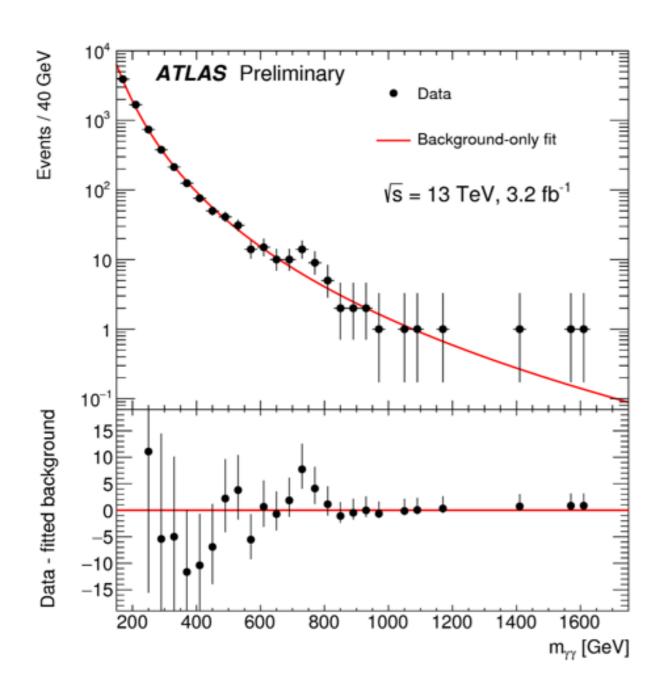
LHC 13TeV last year is still low in luminosity but powerful enough so that we may still learn someghing.



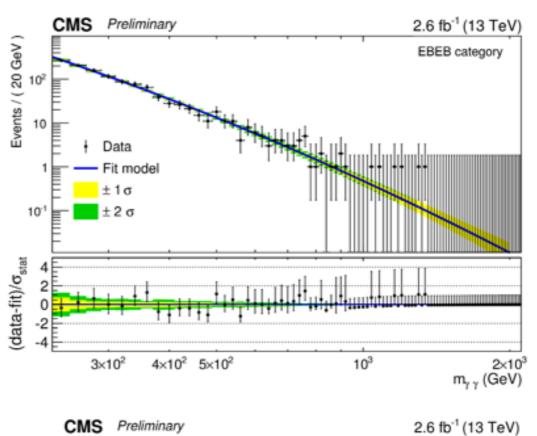
from 1512.04933 Franceschini et al

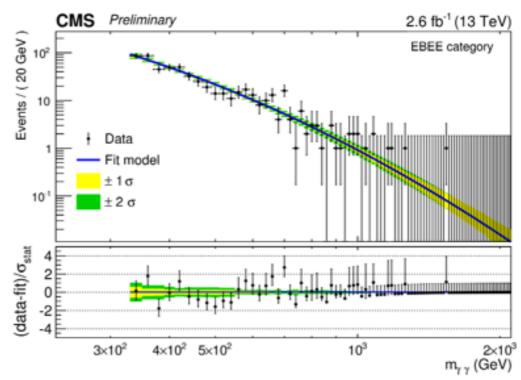
### of course there are nothing to worry about this...

#### ATLAS and CMS



m=750 GeV  $3.6\sigma$  local ( $2.0\sigma$  global)





m=760 GeV  $2.6\sigma$  local (<1.2 $\sigma$  global)

 $\sigma_{\gamma\gamma} = 4.4 \pm 1.1 \text{fb} \text{ [arxiv:1512.04929]}$ 

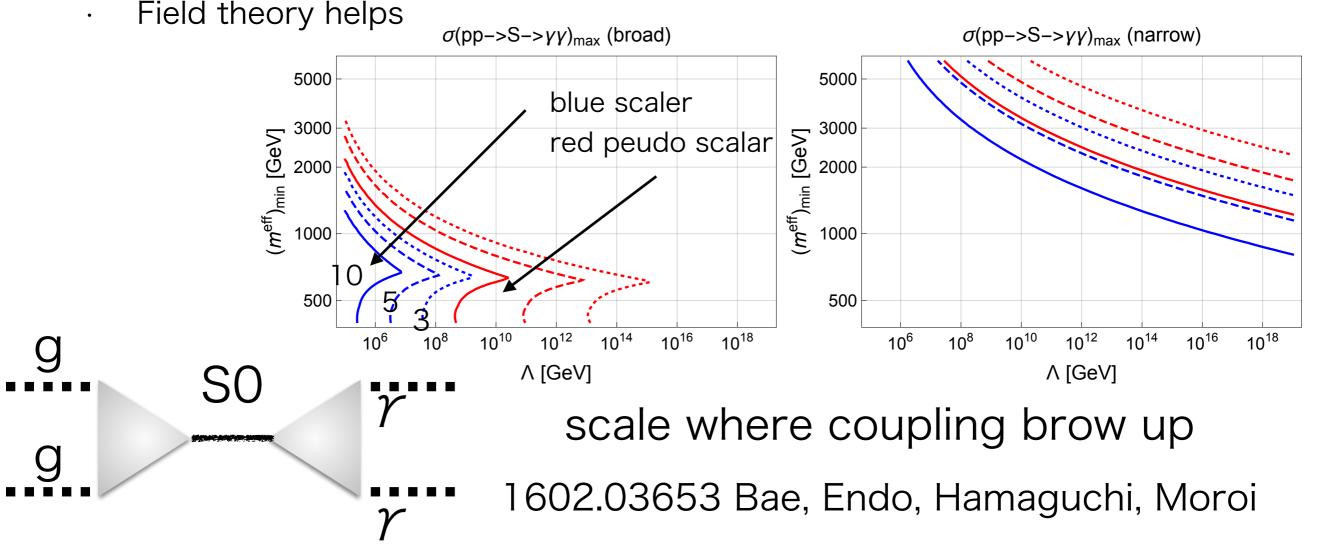
# effective theory approach vs UV completion (if any)

· Need New scalars couples to gg and  $\gamma$   $\gamma$  (fermion loops)

$$\mathscr{L}_{\mathrm{eff}} \supset \frac{C_{BB}}{m_{S_0}} S_0 B_{\mu\nu} \, \tilde{B}^{\mu\nu} + \frac{C_{gg}}{m_{S_0}} S_0 \, G^{a\mu\nu} \, \tilde{G}^a_{\mu\nu},$$

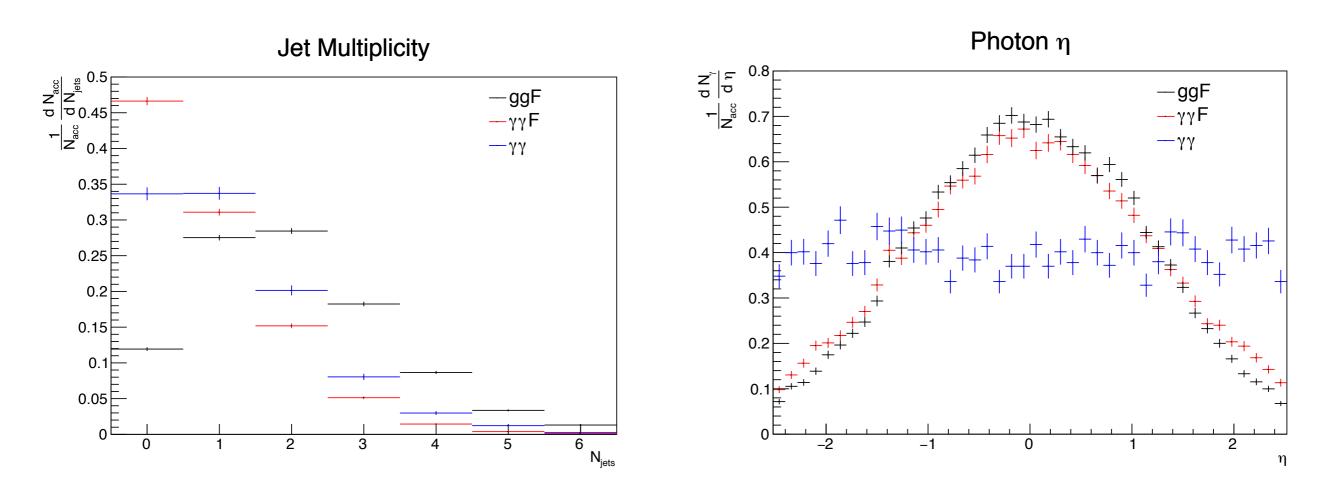
 To achieve the desired cross section, we need to many intermediate particles if it comes from fermion loop

(coupling is essentially  $\beta$  fun)



# with higher luminosity

 number of additional jets, photon eta distribution 1601.00638 Csaki, Hubisz, Lombardo, Terning

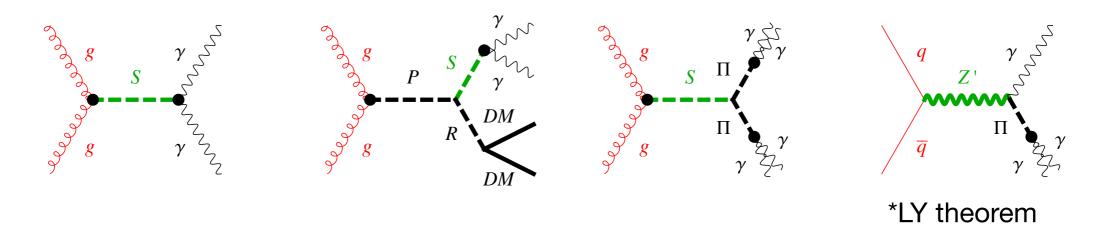


spin correlation of ISR is in principle sensitive to the CP of scalar eta distribution and  $\cos \theta^*$  of photon is sensitive to spin 0 or 2

#### Photon iet

#### A more complicated kinematics?

Compatibility between runs 1, 2 improved if S decays from a heavier particle.



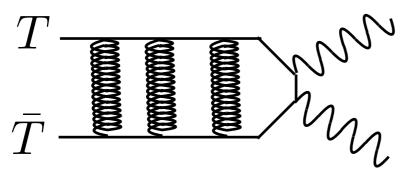
Tuning  $M_P \approx M_S + M_R$  needed to avoid  $p_T$ . S virtuality can fake S width.

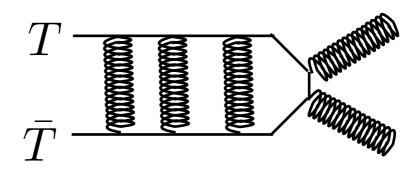
Or large  $S \to \Pi\Pi$  with  $\Pi \to \gamma \gamma$ , collimated and seen as a single  $\gamma$  if  $M_{\Pi} \ll M_{S}$ . Traveling in the detector material, 'photon jets' give more  $\gamma \to e^{+}e^{-}$ .

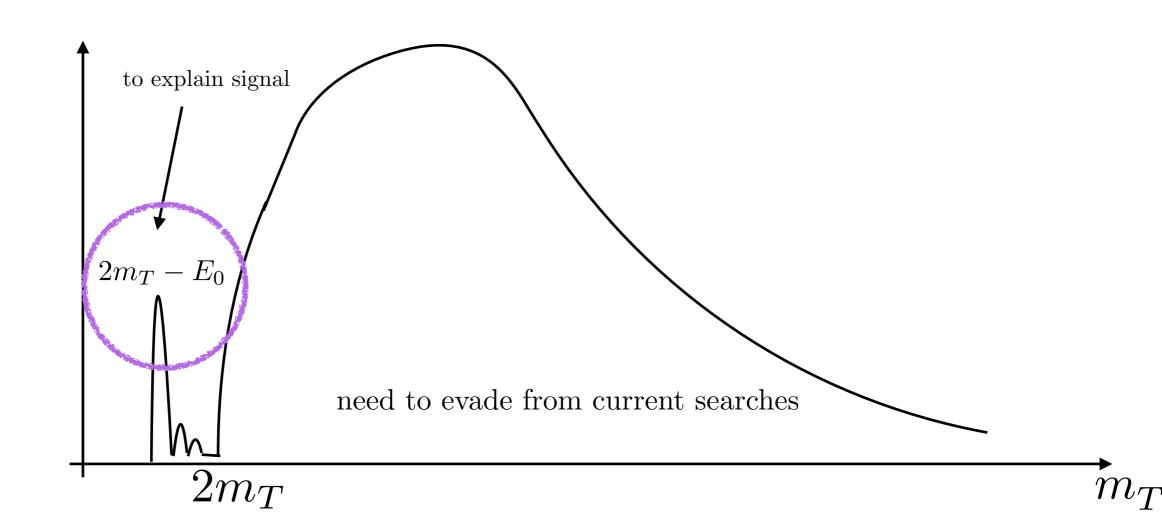
Or many collimated  $\gamma$ . Or not a peak. Or two nearby narrow resonances.

# Another possibility: bound state of heavy fermion

$$1S(J^{PC} = 0^{-+})$$
 at  $750GeV$   $m_T: 375 \sim 380GeV$ 







#### Proposed New Signal for Scalar Top-Squark Bound-State Production

#### Manuel Drees

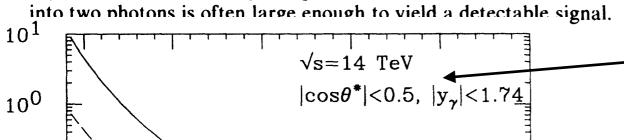
Physics Department, University of Wisconsin, Madison, Wisconsin 53706

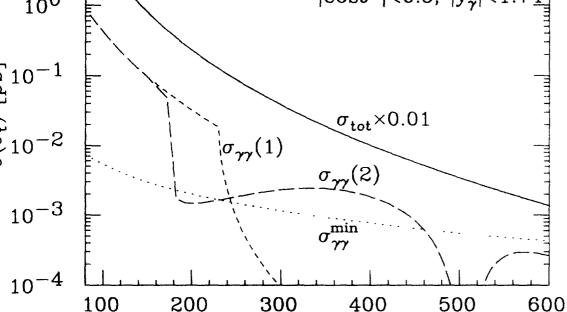
# SSC canceled Oct 31 of the year

Mihoko M. Nojiri\*

Theory Group, KEK, Oho 1-1, Tsukuba, Ibaraki 305, Japan (Received 4 October 1993)

We study the production and decay of a scalar  $(\tilde{t}_1\tilde{t}_1^*)$  bound state  $\sigma_{\tilde{t}_1}$  at hadron supercolliders, where  $\tilde{t}_1$  is the lighter scalar top eigenstate. If  $\tilde{t}_1$  has no tree-level 2-body decays, the dominant decay modes of  $\sigma_{\tilde{t}_1}$  are gg or, if  $m_h < m_{\tilde{t}_2} < m_{\tilde{t}_2}$ , a pair of light scalar Higgs bosons h. Nevertheless, the branching ratio





 $m(\sigma_{\bar{t}})$  [GeV]

worse, I even forgot all about this...
SSC is... really really long time ago

# yes! we replaced the figure

LHC was 14TeV and 1year means 100fb-1 and it was supposed to start 2003 or even 2001 (I forgot) because it was competing with SSC

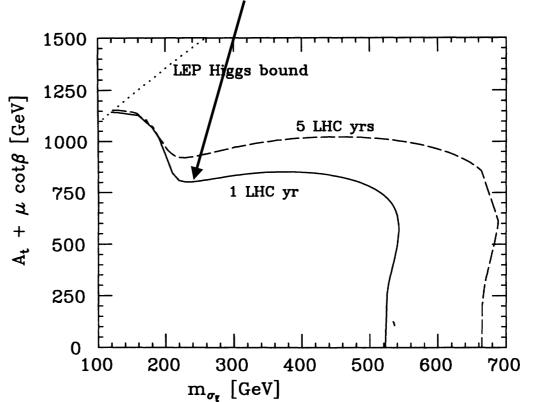


FIG. 10. Region in the pin  $m_{\sigma_{\tilde{t}}}$  and  $A_t + \mu \cot \beta$  that can lead 1 (solid curve) and 5 (dashed running the LHC at full lumfb<sup>-1</sup> per year). The region corner is excluded by LEP sea Higgs bosons. The curves have for  $m_t = 150$  GeV,  $\tan \beta = 3$   $m_{\tilde{t}_L} = m_{\tilde{t}_R}$ ,  $\mu = 750$  GeV, and depend little on these choices as discussed in the text.

# points of this PRL paper (at that time)

- · It was the time that we were not at all sure about jet signature at hadron colliders.
- We did not know how to use ISR jets for degenerate SUSY search.
   (CKKW is 2001, Alwall et al for BSM search using ISR is 2008) So this was a "dead zone" of SUSY search.
- Therefore predicting a new lepton/photon signature was very important. (It is still clean and easy channel.)
- Cosmology: We already knew about "coannihilation" in degenerate SUSY, where dark matter density constraint is non-trivial. (Griest Seckel is 1991, Drees Nojiri 1993 for SUSY)

### heavy scalar/fermion bound state as the origin of diphoton excess

- If the life time of a heavy colored particle X is longer than time scale forming XX bound state, there are a channel
- · pp  $\rightarrow$ X Xbar  $\rightarrow$  S  $\rightarrow$  gg,  $\gamma \gamma$ , ZZ we take X SU(2) singlet

$$\sigma(pp \to S_0 \to \gamma\gamma) = \frac{K}{s \, m_{S_0}} \frac{\Gamma_{\gamma\gamma} \, \Gamma_{gg}}{\Gamma_{\text{tot}}} \left[ \frac{\pi}{8} \int dx_1 dx_2 \, \delta(x_1 x_2 - m_{S_0}^2/s) f_g(x_1) f_g(x_2) \right]$$

$$\Gamma_{\text{tot}} = \Gamma_{\gamma\gamma}/c_W^4 + \Gamma_{gg} + 2\Gamma_X,$$

$$\Gamma_{\gamma\gamma} = 48\pi Y_X^4 \alpha^2 |\psi_0(0)|^2 / m_{S_0}^2,$$

$$\Gamma_{gg} = 32\pi \alpha_s^2 |\psi_0(0)|^2 / (3m_{S_0}^2),$$

$$\Gamma_{gg} = 32\pi\alpha_s^2 |\psi_0(0)|^2/(3m_{S_0}^2),$$

wave function of XX system at origin

4th power of hypercharge (···fit anything)

## bound state of X

· wave function at origin

QED part+ linear part

$$\left[-\frac{\nabla_{\boldsymbol{r}}^2}{m_X}+V(\boldsymbol{r})-E_0\right]\psi_0(\boldsymbol{r})=0,$$

$$V(\boldsymbol{r}) = -Y_X^2 \frac{\alpha}{|\boldsymbol{r}|} + V_{\text{QCD}}(|\boldsymbol{r}|).$$

numerically solve and fit

(consistent with Hagiwara Kato Martine, Ng 1990 for Y=0)

$$|\psi_0(0)| = \sum_{n=0}^4 a_n \left[ \ln(m_{S_0}/750 \,\text{GeV}) \right]^n, \qquad E_0 = \sum_{n=0}^4 b_n \left[ \ln(m_{S_0}/750 \,\text{GeV}) \right]^n,$$

	$Y_X$	$a_0$	$a_1$	$a_2$	$a_3$	$a_4$	$b_0$	$b_1$	$b_2$	$b_3$	$b_4$
							1				0.04078
	1/3	88.44	115.3	77.54	38.14	10.82	4.145	2.481	0.9416	0.2461	0.04143
	2/3	90.44	118.1	79.64	39.30	11.17	4.226	2.552	0.9726	0.2557	0.04341
1.65	1	93.82	122.9	83.20	41.25	11.76	4.363	2.672	1.026	0.2721	0.04682
	4/3	98.64	129.8	88.28	44.05	12.61	4.559	2.845	1.102	0.2960	0.05180
											0.05858
	2	112.6	150.7	104.5	51.83	14.40	5.162	3.366	1.322	0.3812	0.07999

### Abnormal candidate remains

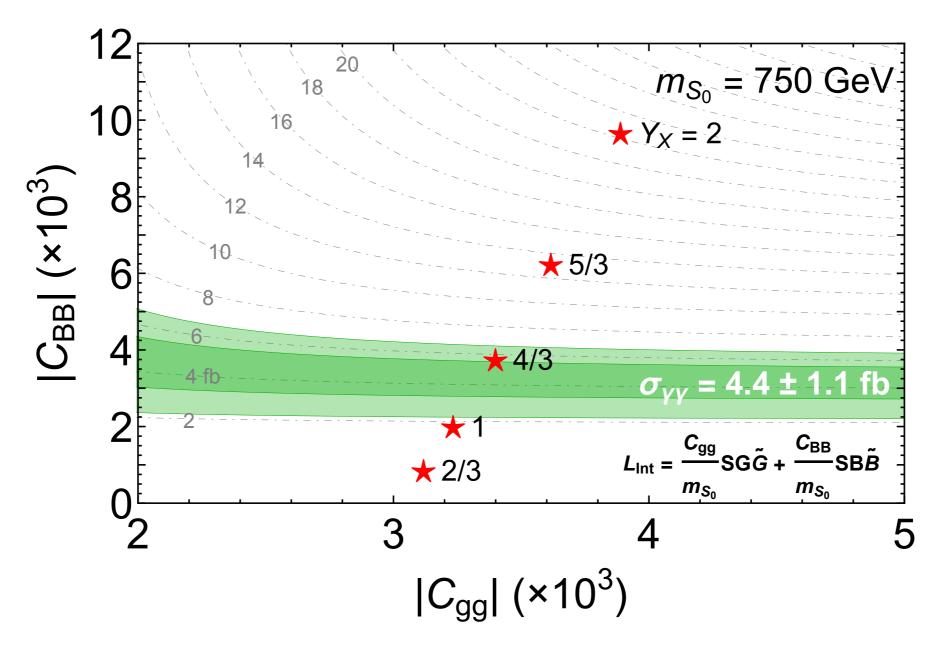


Figure 1: Red stars are predictions of our model on the  $(|C_{gg}|, |C_{BB}|)$ -plane with  $Y_X$  being 2/3, 1, 4/3, 5/3 and 2, respectively. Contours of the diphoton cross section as a function of  $|C_{gg}|$  and  $|C_{BB}|$  are also shown by gray-dashed lines. Darker (lighter) green-shaded region corresponds to the cross section experimentally favored by the diphoton excess at  $1\sigma$  ( $2\sigma$ ) level [3].

Numerically tough calculation, completely relay on Ishikawa-san

## This is nice

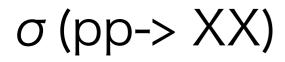
- There are no dim 4 operator involving X and SM particles
- The X decay is suppressed by some cutoff scale Λ
- Consistent with small width assumption to get signal

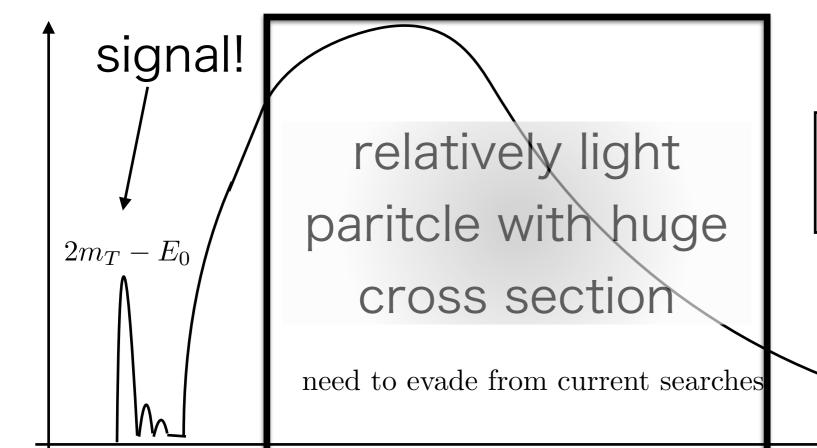
### Collider bound from open production

X decay pattern

1. X-> W j, Zj Hj etc is not allowed because of Q=4/3

2. X-> 2j, 3j (Kats and Strassler 1602.08819) weaker collider bound: large QCD background fat jet search at 13TeV?





3. This work

 $X \rightarrow \chi (ij, jjj) \chi = DM$ 

recast light squark search at LHC

 $m_{Tar{T}}$ 

# Quick consideration on the dark matter $\chi$

- $\cdot$  Z<sub>2</sub> odd
  - · X: strongly interaction particle (3,0,4/3)
  - $\cdot$   $\chi$ : weakly interacting partile
- · X and  $\chi$  interact through some higher dim operators such as

$$\mathcal{O}_F \sim (\bar{X}u^c)(\bar{\chi}u^c)/\Lambda^2$$
  $\mathcal{O}_S \sim (\bar{X}d^c)(\bar{u}^cd^c)\chi/\Lambda^3$ 

# (nearly model independent) $\Omega_{\chi}$

strongly interacting X work as the cleaner



- Early Universe (T~ mX)
- XX
   ⇒ gg large cross section: rapid annihilation density is small
- $\cdot \chi \chi \rightleftharpoons SM$  particles: maybe small and suppressed
- ·  $\chi$  SM  $\rightleftharpoons$  X SM rapid because so many SM particles in the thermal bath: leading small  $\Omega \chi$

### dark matter density $\Omega \propto 1/\langle \sigma v \rangle$

#### effective pair annhilation cross section

$$\begin{split} \langle \sigma v \rangle &= \sum_{ij} \langle \sigma_{ij} v \rangle \frac{g_i g_j}{g_{\rm eff}^2} (1 + \Delta_i)^{3/2} (1 + \Delta_j)^{3/2} \exp[-x \left(\Delta_i + \Delta_j\right)], \\ x &= m_\chi / T & \text{Griest and Seckel Phys. Rev. D43, 3191(1991)} \\ \Delta_i &= (m_i - m_\chi) / m_\chi & \text{mass difference} \end{split}$$

#### for our case

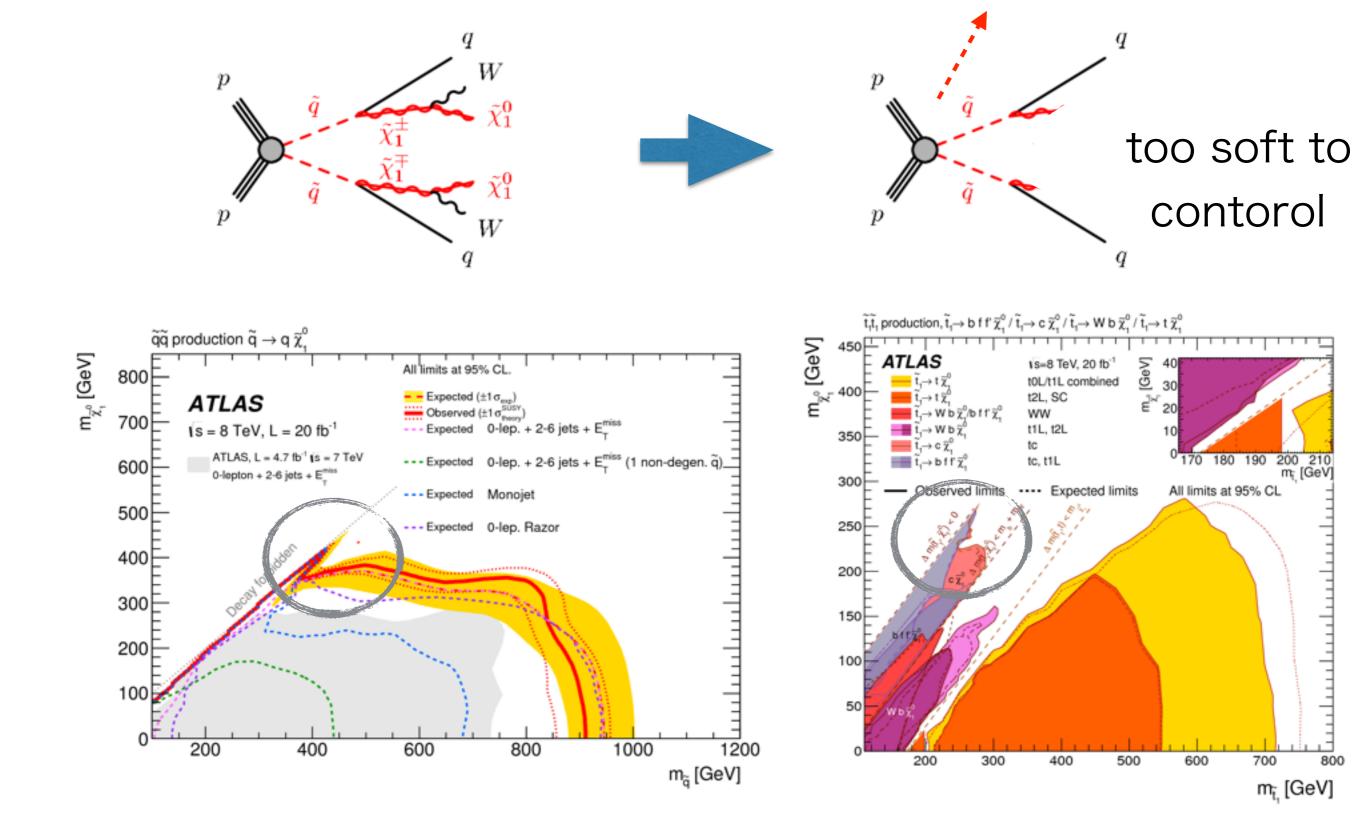
$$\langle \sigma v \rangle \simeq 2 \frac{43\pi\alpha_s^2}{27m_X^2} \frac{36(1+\Delta_X)^3 \exp(-2x\Delta_X)}{\left[g_\chi + 12(1+\Delta_X)^{3/2} \exp(-x\Delta_X)\right]^2}.$$

50 dark matter density is the **Scalar Dark Matter** function of X mass and **Fermionic Dark Matter** 40 Δm (GeV) 00  $\chi$  X mass difference dark matter density only XX and mass is small enough here 20 It is enough to exclude <del>10</del> this region 370 380 390 395 365 375 385  $m_X$  (GeV)

# Collider constraints

### degenerate BSM search

hard ISR (better prediction)



### degenerate heavy fermion search at LHC

- · generate p p  $\rightarrow$  X X upto 2jet matched sample.
- · force X to decay into X->  $\chi$  jj or  $\chi$  jjj to obtain efficiency.
- · signature is therefore jet (ISR, hard) + missing + soft activities
- · compare with ATLAS and CMS SUSY searches results using checkmate

#### 8TeV stop -LSP degenerate

			more ge	eneral	missing p	T search		
	$\not\!\!E_T$ [GeV]	$p_{T_{j_1}}[GeV]$	$\Delta \phi (j, p_T)$	$n_j$	$E_T/\sqrt{H_T}$	$m_{\rm eff}$ (incl.)	$\sigma_{ m obs}^{95\%}$	Ref.
M2 ×	340	340	0.4	<b>≤</b> 3	-	-	28.4 fb	[44]
SR5	350	$0.5 \rlap/\!E_T$	1.0	-	-	-	21 fb	[45]
SR6	400	$0.5 \rlap/\!E_T$	1.0	-	-	-	12 fb	[45]
SR2jm	200	300	0.4	$\geq 2$	$15\mathrm{GeV}^{1/2}$	1.2 TeV	21 fb	[46]

Table 3: Signal regions and upper bounds on the signal cross sections at 95% C.L. Here,  $p_{T,j_1}$ , H and  $m_{\text{eff}}$  (incl.) are the leading jet  $p_T$ , the scalar  $p_T$  sum of all jets and  $H_T + \not\!\!E_T$ , respectively.

#### 13TeV data

### background modeling at 13TeV

Table 1: The Standard Model background Monte Carlo simulation samples used in this paper. The generator the order in  $\alpha_s$  of cross-section calculations used for yield normalization (leading order (LO), next-to-leading order (NLO), next-to-next-to-leading order (NNLO), next-to-next-to-leading logarithm (NNLL)), PDF sets, parto showers, and tunes used for the underlying event are shown.

Physics process	Generator	Cross-section	PDF set	Parton shower	Tune
		normalisation			
$W(\to \ell \nu) + \text{jets}$	SHERPA 2.1.1	NNLO	CT10	Sherpa	Sherpa default
$Z/\gamma^*(\to \ell\bar{\ell})$ + jets	SHERPA 2.1.1	NNLO	CT10	SHERPA	Sherpa default
$\gamma$ + jets	SHERPA 2.1.1	LO	CT10	SHERPA	Sherpa default
$t\bar{t}$	Powheg-Box v2	NNLO+NNLL	CT10	Рутніа 6.428	Perugia2012
Single top ( <i>t</i> -channel)	Powheg-Box v1	NLO	CT10f4	Рутніа 6.428	Perugia2012
Single top (s- and Wt-channel)	Powheg-Box v2	NLO	CT10	Рутніа 6.428	Perugia2012
$t\bar{t} + W/Z/WW$	Madgraph 5.2.2.2	NLO	NNPDF2.3LO	Рутніа 8.186	A14
WW, WZ, ZZ	SHERPA 2.1.1	NLO	CT10	SHERPA	Sherpa default
Multi-jet	Рутніа 8.186	LO	NNPDF2.3LO	Рутніа 8.186	A14

# NNLO cross section and +NLO multijet

Table 3: Control regions used in the analysis. Also listed are the main targeted background in the SR in each case, the process used to model the background, and the main CR requirement(s) used to select this process. The transverse momenta of high-purity leptons (photons) used to select CR events must exceed 25 (130) GeV.

CR	SR background	CR process	CR selection
CRγ	$Z(\rightarrow \nu \bar{\nu})$ +jets	$\gamma$ +jets	Isolated photon
CRQ	Multi-jet	Multi-jet	SR with reversed requirements on (i) $\Delta\phi(\text{jet}, E_{\text{T}}^{\text{miss}})_{\text{min}}$
			and (ii) $E_{\rm T}^{\rm miss}/m_{\rm eff}(N_{\rm j})$ or $E_{\rm T}^{\rm miss}/\sqrt{H_{\rm T}}$
CRW	$W(\rightarrow \ell \nu)$ +jets	$W(\rightarrow \ell \nu)$ +jets	$30 \text{ GeV} < m_{\text{T}}(\ell, E_{\text{T}}^{\text{miss}}) < 100 \text{ GeV}, b\text{-veto}$
CRT	$t\bar{t}(+EW)$ and single top	$t\bar{t} \to b\bar{b}qq'\ell \nu$	$30 \text{ GeV} < m_{\text{T}}(\ell, E_{\text{T}}^{\text{miss}}) < 100 \text{ GeV}, b\text{-tag}$

CR

Channel 2il 2ir	~
Channel 2jl 2jr	11
Total bkg 237 16	53
Total bkg unc. $\pm 22 [9\%]$ $\pm 20 [12\%]$	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
MC statistics – ±1.8 [19	$\sqrt{6}$
$\Delta \mu_{Z+jets}$ ±6 [3%] ±5 [3%	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
$\Delta\mu_{W+\text{jets}}$ ±4 [2%] ±4 [2%]	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
$\Delta \mu_{\text{Top}}$ ±1.2 [1%] ±1.6 [1%]	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
$\Delta \mu_{\text{Multi-jet}}$ $\pm 0.05 [0\%]$ $\pm 0.09 [0\%]$	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
CR $\gamma$ corr. factor $\pm 8 [3\%]$ $\pm 6 [4\%]$	$% \left[ -\frac{1}{2}\left( -\frac{1}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2}\left( -\frac{1}{2$
Theory $W = \pm 1.4 [1\%] \pm 2.3 [1\%]$	<b>%</b> ]
Theory $Z \pm 6 [3\%] \pm 3.2 [2\%]$	<b>%</b> ]
Theory Top $\pm 2.7 [1\%]$ $\pm 2.1 [1\%]$	<b>%</b> ]
Theory Diboson $\pm 16 [7\%]$ $\pm 16 [10\%]$	$\mathscr{C}$
Jet/ $E_{\rm T}^{\rm miss}$ ±1.5 [1%] ±2.1 [1%]	<u>[6]</u>

# Our case: signal cross section is known in NNLO level(QCD)

$m_X$ [GeV]									
σ@8 TeV [pb]	4.35	4.01	3.70	3.41	3.15	2.91	2.69	2.49	2.31
$\sigma$ @13 TeV [pb]	20.34	18.86	17.51	16.26	15.12	14.07	13.10	12.21	11.39

Table 2: Pair production cross sections of the heavy fermion X at NNLO for various  $m_X$ .

#### cross section error 25%(NLO)→less than 10% (NNLO)

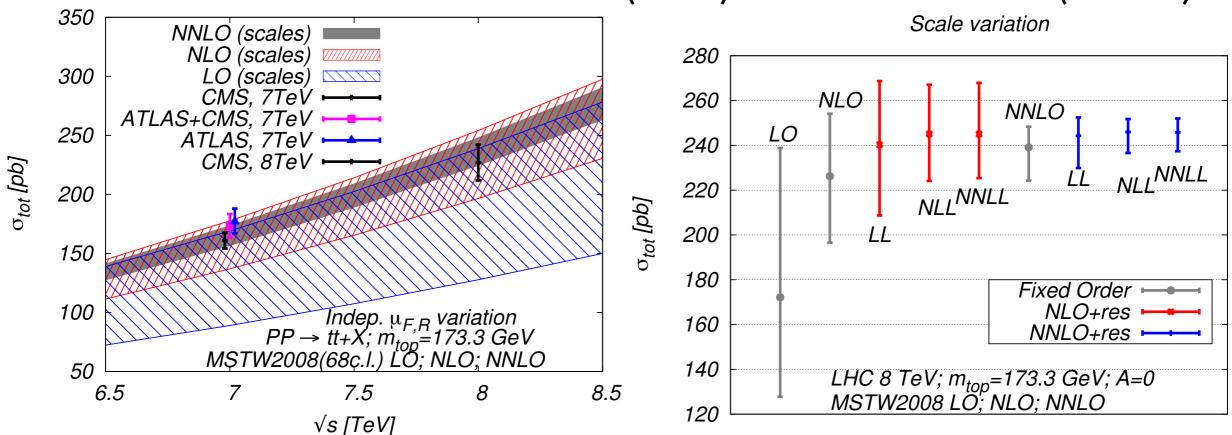


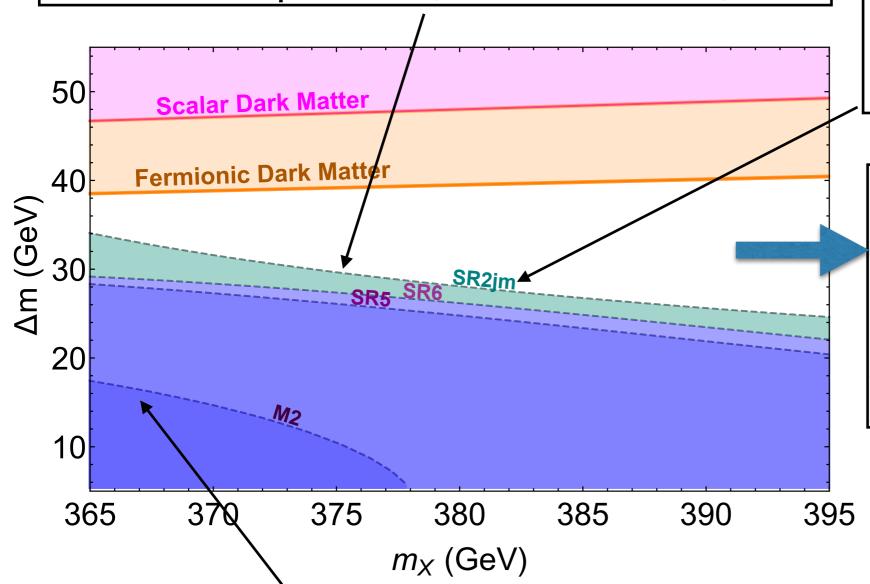
Fig. 2. – Scale dependence of the predicted cross-section at LO, NLO and NNLO at the LHC as a function of  $\sqrt{s}$  (left). On the right plot: detailed breakdown of scale uncertainty for LHC 8 TeV at LO, NLO and NNLO including also soft-gluon resummation at LL, NLL and NNLL.

### · Systematic error on efficiency

- ATLAS SUSY analysis (for SIGNAL)
  - NLO total cross section,
  - SUSY + ISR jets (matched tree) for signal efficiency: varying scale,
     factorization scale, and emission scale from 0.5 to 2. cross section times
     efficiency error is 16% for degenerate stop and we use this number
  - · CheckMate scheme (allow up to two sigma deviation-> 32%
- NLO cross section uncertainty is large for our case (25% for scale and pdf). NNLO available + NLO multi-jet calculation is desirable(faking Sherpa tt + multi-jet?) Regret we have not done this (too competitive dxxx new resonance)
- Tight cuts might improve the limit significantly provided controlled systematics with higher order generators

# results

The limit depends on signal error assumptions (we take 16%)



13TeV data gives best limit, and the limit is weaker than expected

limit is weaker here because mono-jet condition (pT1 vs ETmiss) two body decay case ( X-> jx) is excluded by multijets+ missing search

Tight mono-jet requirement leads worse sensitivity for heavy quark: (more ISR). Stop search is optimize for M2

# (personal) summary

- It is a responsibility of elderlies to tell ideas and disasters in the past to younger generations
- Heavy quark and dark matter for 750 GeV diphoton: We proposed a effective theory with rather few number of particles with a DM, Almost Model independent upper limit of the DM density.
- The LHC is not safe playground as it was in 1993, when we did not think about using ISR to search for the particle degenerate with dark matter. (Alwall et al PRD 79 (2009) 015005)
- Our case still survives (for narrow width). Checkmate scheme is very conservative. More stringent constraint this year.
- · LHC will exclude narrow width heavy fermion though  $\gamma \gamma$  channel.