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Sphericity of peaks with non-Gaussianity

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Spherical collapse is assumed in most of the works on PBH formation from the primordial curvature perturbation. According to the peak theory[1], sufficiently high peaks of a Gaussian random scalar field statistically have spherical symmetric shapes in the homogeneous and isotropic universe, which guarantees the above assumption. However, the PBH formation is related to highly non-linear regime of the perturbation and the non-Gaussianity may play an important role. In the talk, I will explicitly show how this argument can be modified if the curvature perturbation obeys non-Gaussianity.

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