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## Relaxation of shell effect in giant dipole resonance width away from magicity

The isovector giant dipole resonance (IVGDR) serves as a crucial tool for probing a wide range of phenomena, from r-process nucleosynthesis to the determination of the strength of gravitational waves. Generally, the width of the IVGDR ( $\Gamma_G$ ) increases with temperature (T) in the range of 1 MeV lessimT

 $less sim 3\,\mathrm{MeV},$  with the possibility of saturation at higher temperatures [1]. However, in the low-temperature regime (T

less sim 1 MeV), studying  $\Gamma_G$  is particularly challenging due to the difficulty of achieving low excitation energies. Limited investigations in this regime suggest that the behavior of  $\Gamma_G$  is ambiguous, influenced by microscopic effects such as shell effects and pairing fluctuations, which hinder the expected thermal broadening of  $\Gamma_G$  [2, 3] .

Motivated by these challenges, we conducted a detailed study of  $\Gamma_G$  in the low-to-intermediate temperature range for nuclei near the N=Z=28 shell closure, where detailed analyses are currently lacking. Our work elucidates the relative importance of neutron-to-proton ratio (N/Z), shell closure, and thermal fluctuations in shaping the temperature dependence of  $\Gamma_G$  for nuclei near the doubly magic  $^{56}$ Ni. To isolate these effects, we studied  $^{62}$ Zn and  $^{68}$ Zn nuclei, populated via  $\alpha$ -induced fusion reactions. High-energy  $\gamma$ -rays ( $E_{\gamma}>4$  MeV) emitted from IVGDR decay were detected using the Large Area Modular BaF $_2$  Detector Array (LAMBDA) [4]. The measured spectra were analyzed using statistical model calculations implemented in TALYS.

A contrasting thermal behavior of  $\Gamma_G$  was observed for the two nuclei. For  $^{68}$ Zn, the width ( $\Gamma_G$ ) increases monotonically with temperature from its ground-state value. In contrast,  $^{62}$ Zn exhibited a suppressed width at low temperatures, consistent with the behavior of nearby nuclei with neutron and/or proton numbers close to 28. This suggests that the suppression of  $\Gamma_G$  at low temperatures is not a universal feature but is influenced by proximity to magic numbers, rather than N/Z asymmetry.

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