

High-Temperature Superconducting (HTS) Multiple-Cell Cavities for High-Mass Axion Search

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KAIST & IBS-DMAG

***** IBS-DMAG is the successor of IBS-CAPP

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Axion detection setup



Experimental Tests of the "Invisible" Axion

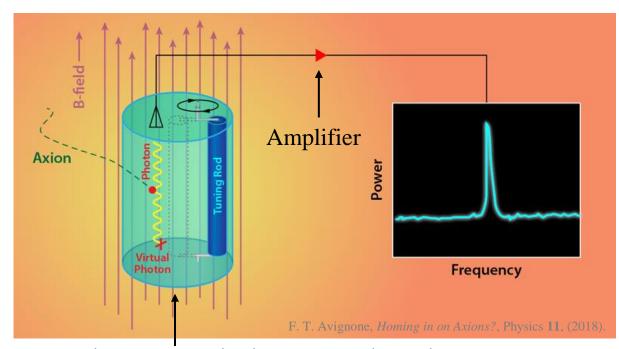
P. Sikivie

Physics Department, University of Florida, Gainesville, Florida 32611 (Received 13 July 1983)

Experiments are proposed which address the question of the existence of the "invisible" axion for the whole allowed range of the axion decay constant. These experiments exploit the coupling of the axion to the electromagnetic field, axion emission by the sun, and/or the cosmological abundance and presumed clustering of axions in the halo of our galaxy.

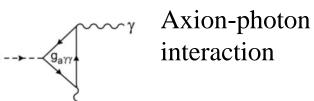
PACS numbers: 14.80.Gt, 11.30.Er, 95.30.Cq

- Cavity Axion Haloscope
- → The most sensitive method for detecting axion dark matter



Resonant microwave cavity in a cryogenic environment

Proposed by Pierre Sikivie in 1983



$$f_a = \frac{m_a c^2}{h}$$

When $f_a = f_{cav}$

→ resonant conversion

Unknown m_a

- → Resonant frequency tuning
- → Dielectric or metal rod

Motivation & Key Technologies

• Axion conversion power

$$P_{a\gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C \frac{\beta}{1+\beta} \frac{Q_l Q_a}{Q_l + Q_a}$$

• Scan rate

$$\frac{df}{dt} \propto \frac{B^4 V^2 C^2}{T_{SVS}^2} Q_l Q_a$$

 $Q_l \gg Q_a$ or $Q_l \ll Q_a$ Jinsu Kim, Patras (2023)



$$B^4, V^2, C^2, Q_l$$
 \rightarrow Enhancing $\frac{df}{dt}$

 $g_{a\gamma\gamma}$: Axion photon coupling ρ_a : Axion density in the local dark-matter halo

B: Externally applied B-field

V: Cavity volume

C: Cavity form factor

Q_l: Loaded Cavity Q-factor

 Q_a : Axion Q-factor

 β : antenna coupling coefficient

 T_{sys} : system noise temperature

1. Multiple-Cell Cavities



Maximize *V* for higher *f*



2. HTS Cavities

D. Kim et al., Journal of Cosmology and Astroparticle Physics 2020, (2020).

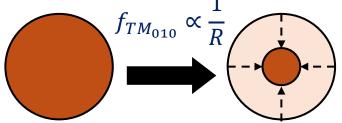
Multiple-Cell Cavities

• Goal B^4, V^2, C^2, Q_1 \rightarrow Enhancing $\frac{df}{dt}$

• Why are Multiple-Cell Cavities needed?

* In the previous slide

For higher frequency searches

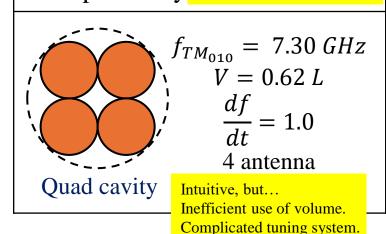


Unavoidable volume loss

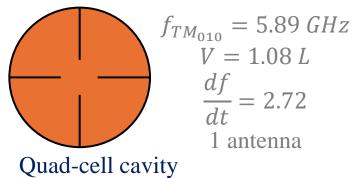
(Meanwhile, the fixed magnet bore radius)

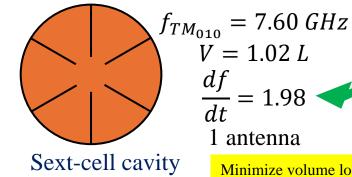
(Example: 100 mm diameter, 200 mm height, 5 mm wall thickness)

Multiple cavity Array of identical cavities



Multiple-cell cavity





V = 1.02 L1 antenna

> Minimize volume loss Use a single antenna → Simplified readout chain

Two times higher

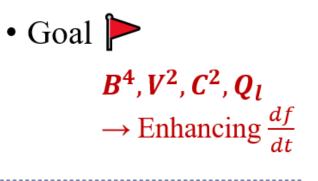
J. Jeong et al., Physics Letters B 777, 412 (2018).

 ω : resonant frequency

U: energy stored in the cavity

 P_d : power dissipated in the cavity

 R_s : surface resistance



• Why are HTS cavities needed?



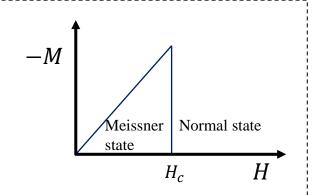
To improve the axion scanning speed,

Cavity quality factor:
$$Q = \omega \frac{U}{P_d} = \omega \frac{\frac{\mu_0}{2} \int |\mathbf{H}|^2 dV}{\frac{R_S}{2} \oint |\mathbf{H}|^2 dS} = \frac{G}{R_S} \propto \frac{1}{R_S}$$

→ Superconductor are ideal materials.

Magnetic fields on the order of 10 tesla are used at IBS-DMAG. However, Type-I superconductors are vulnerable to B-fields.

* In the previous slide



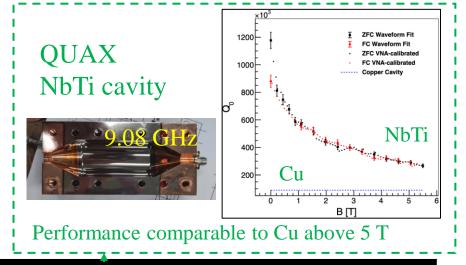
Material	T _c [K]	H _c [T]
Aluminum	1.2	0.01
Indium	3.4	0.03
Lead	7.2	0.08
60Sn-40Pb	7.8	0.2
97Sn-3Ag	3.7	0.02

https://www.mtm-inc.com/when-superconductivity-is-unanticipated.html

X Type-II superconductors

• Why are HTS cavities needed?

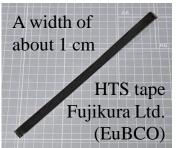




100 mK 8 GHz	R _S (B=0T) (Ohm)	$R_{S} (B=8T, \parallel_{c})$ (Ohm)	Critical Field (H _{c2})	Depinning Frequency
OFHC cu	$\sim 7 \times 10^{-3}$	$\sim 7 \times 10^{-3}$	None	None
NbTi (LTS) Gatti et al. PRD(2019)	$\sim 1 \times 10^{-6}$	$\sim 4 \times 10^{-3}$	√~ 13 T `\	~ 45 GHz
Nb ₃ Sn (LTS) Alimenti et al. SUST(2020)	$\sim 1 \times 10^{-6}$?	~ 25 T	O(10) GHz
Tl-1223 (HTS)	$\sim 1 \times 10^{-5}$	$\sim 1 \times 10^{-4}$ Calatroni et al. SUST(2017)	> 100 T (ab) Larbalestier et al. Nature(2001)	12 — 480 MHz Calatroni et al. SUST(2017)
ReBCO (HTS)	$\sim 1 \times 10^{-5}$ Ormeno et al. PRB(2001)	$\sim 1 \times 10^{-4}$ Romanov et al. Scientific Reports(2020)	> 100 T (ab) Larbalestier et al. Nature(2001)	10 – 100 GHz Romanov et al. Scientific Reports(2020)

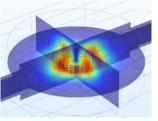
^{*} Depinning frequency: the frequency at which vortices move due to an external force.

Therefore, we looked at ReBCO.



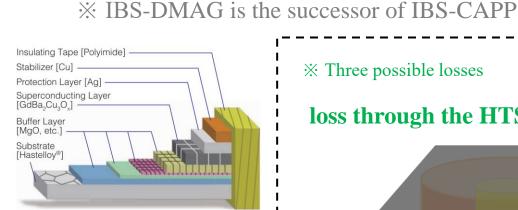
I assembled these resonators with Dr. Danho Ahn





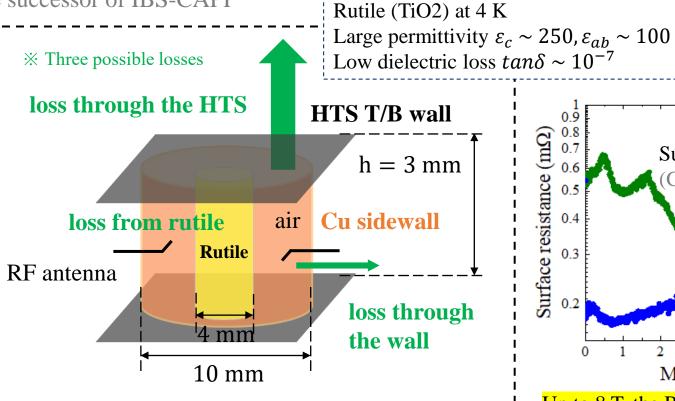
7.95 GHz $\text{TE}_{011} \text{ mode}$

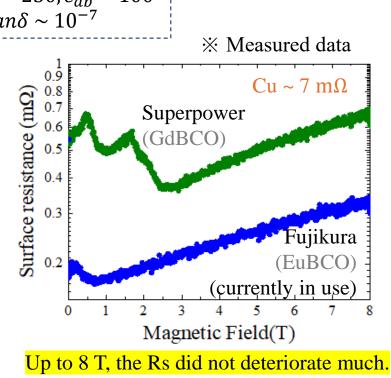
• Measurement of the Surface Resistance of HTS (ReBCO)

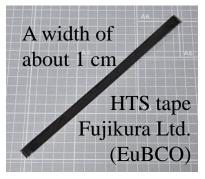


Various HTS tapes are commercially available from different global companies

e.g., Theva, Superpower, Fujikura, and others.





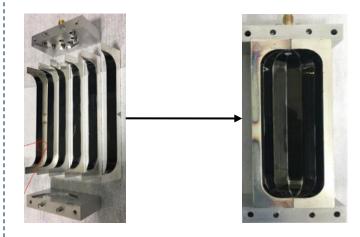


Commercially available 2D HTS tapes

→ Attach to the inside of the cavity

- HTS cavity developed by the CAPP-HTS team
 - **X** IBS-DMAG is the successor of IBS-CAPP

Method A – Gluing



Attach tapes to melon-shaped wedges and combine them

Method B — Soldering





The best-performing cavity

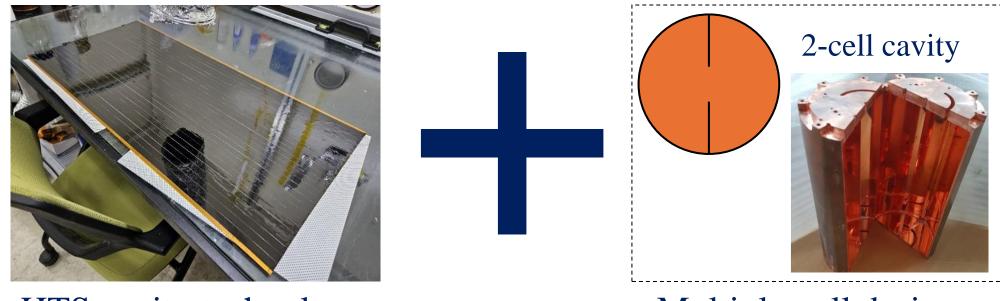
Frequency: 5.4 GHz

Q-factor: 13 M at 8 Tesla

 $(\rightarrow 100 \text{ times better than Cu cavity})$

Goal: Integration of the Two Technologies

We are moving toward integrating HTS cavity technology with multiple-cell design for the next-generation axion experiment.



HTS cavity technology

Multiple-cell design

The cavity design and copper components fabrication were carried out by Dr. Sungwoo and Dr. Sungjae.

0.64

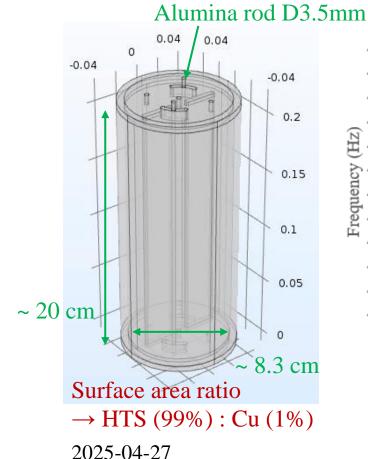
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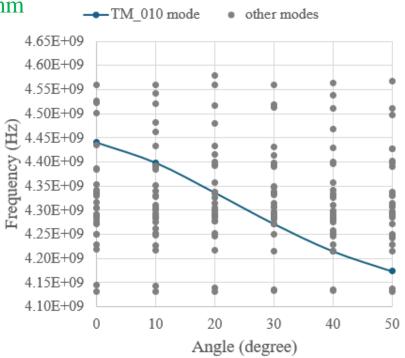
0.54

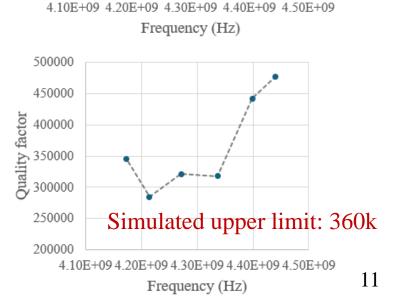
0.52

0.50

COMSOL simulation study





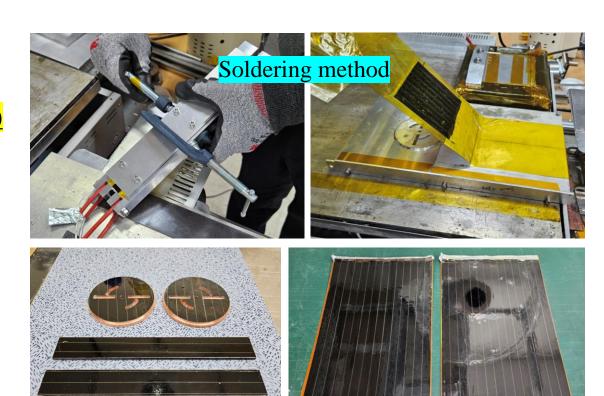


Frequency: 4.17 - 4.44 GHzMass: $17.27 - 18.37 \mu\text{eV}$

7th CUBES Workshop

Fabrication Process

- 1. HTS tapes are soldered onto copper components (endcap, sidewall, and partition)
- 2. Sidewalls are joined via laser welding & indium cold welding
- 3. 0.8 mm thick copper rings are tightly clamped around the sidewall and fixed in place using STYCAST 2850FT
- 4. The endcap is laser welded to both the sidewall and the partition.
- 5. Finally, they are electrically connected via indium cold welding.

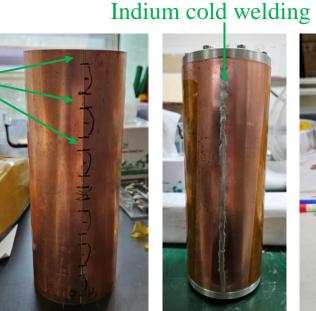


Laser welding

Fabrication Process

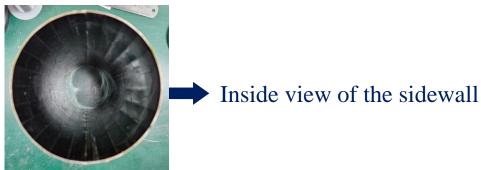
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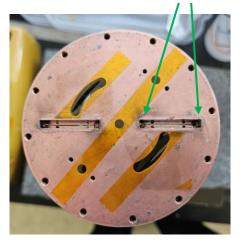
STYCAST



• Fabrication Process

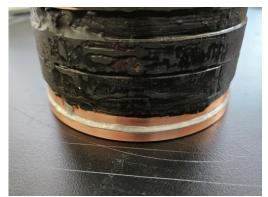
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Laser welding









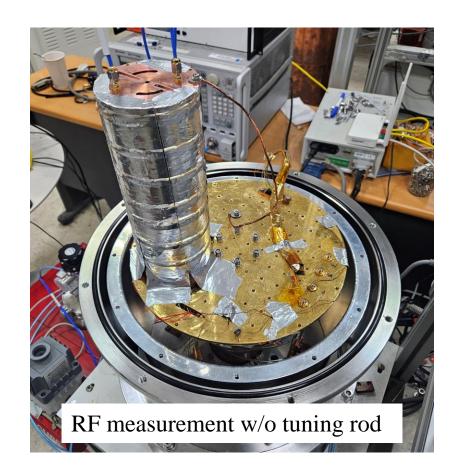
Indium cold welding

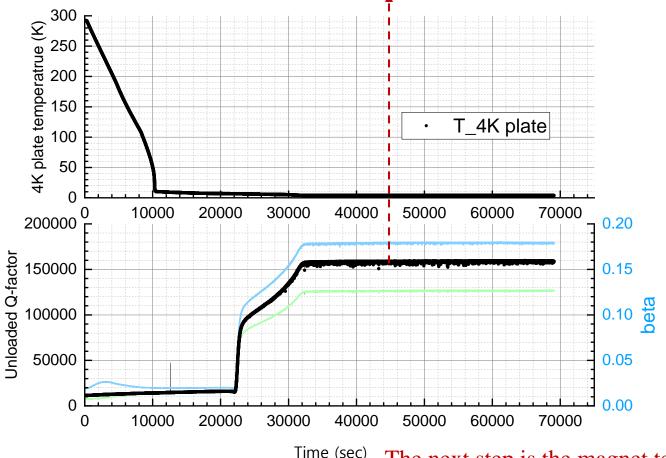
Performance Test



More than 3 times higher than that of a copper cavity

We achieved a Q-factor of $\sim 1.6 \times 10^5$ at a temperature near 10 K.





Time (sec) The next step is the magnet test.

Toward the Axion Search (Next Plans)



Alumina or Sapphire (high dielectric constant ε_r , low dissipation factor $tan\delta$)

- Design and assembly of the tuning system, including a dielectric tuning $rod \rightarrow May$
- Cryogenic performance tests with the tuning rod \rightarrow May
- To perform an axion search in the frequency range 4.17 – 4.44 GHz (mass range $17.27 - 18.37 \mu eV$), this cavity will be integrated into the **DMAG-12T** cryostat. → **Starting** in June



< DMAG-12T >

Refrigerator

- Model: LD400
- 10 mK (Base)
- 580 μW (100 mK)
- Dry type Magnet
- 12 Tesla
- 96mm Bore
- Nb₃Sn
- Vendor: AMI

Summary

- The axion haloscope continues to be the most sensitive and promising method for detecting axion dark matter.
- IBS-DMAG succeeded IBS-CAPP and continues to carry forward the axion search.
- To accelerate the search, IBS-DMAG is integrating independently developed multiple-cell design and HTS cavity technology.
- The HTS double-cell cavity, excluding the tuning system, has been fabricated and demonstrated a Q-factor more than 3 times that of a copper double-cell cavity at cryogenic temperatures.
- This cavity is scheduled to be installed in DMAG-12T in June for the next phase of the axion search experiment.

