

Multi-component dark matter from Minimal Flavor Violation

Shohei Okawa (APCTP)

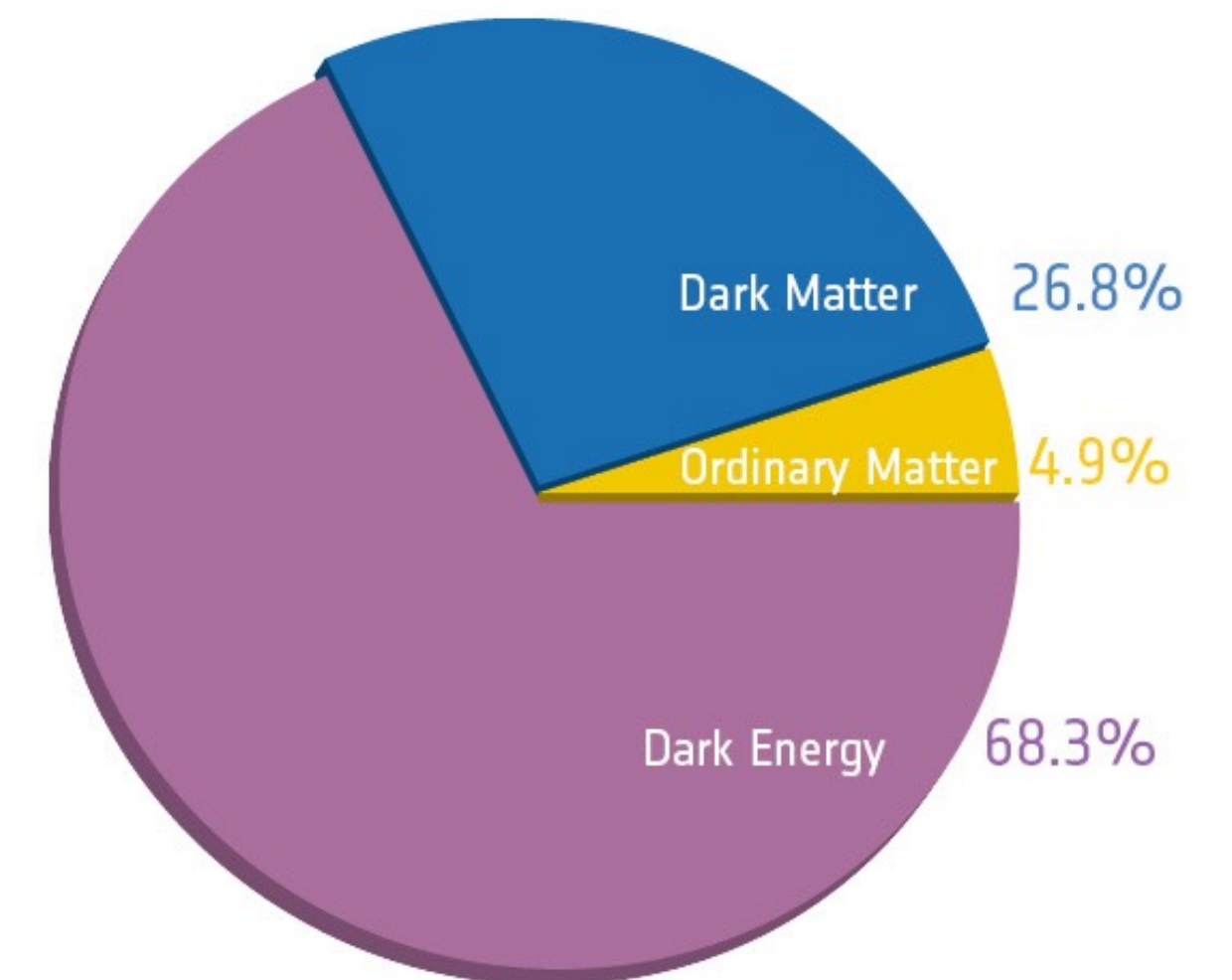


Based on JHEP 11 (2024) 114 with Federico Mescia (INFN LNF), Keyun Wu (ICCUB, Barcelona)
+ongoing works with Alessandro Dondarini (Florence), Giacomo Landini (IFIC Valencia)
Claudio Toni (LAPTH), Joel Swallow (INFN LNF)

The IBS Conference on Dark World
IBS Science and Culture Center, Daejeon / 28th October 2025

Dark Matter exists in the Universe

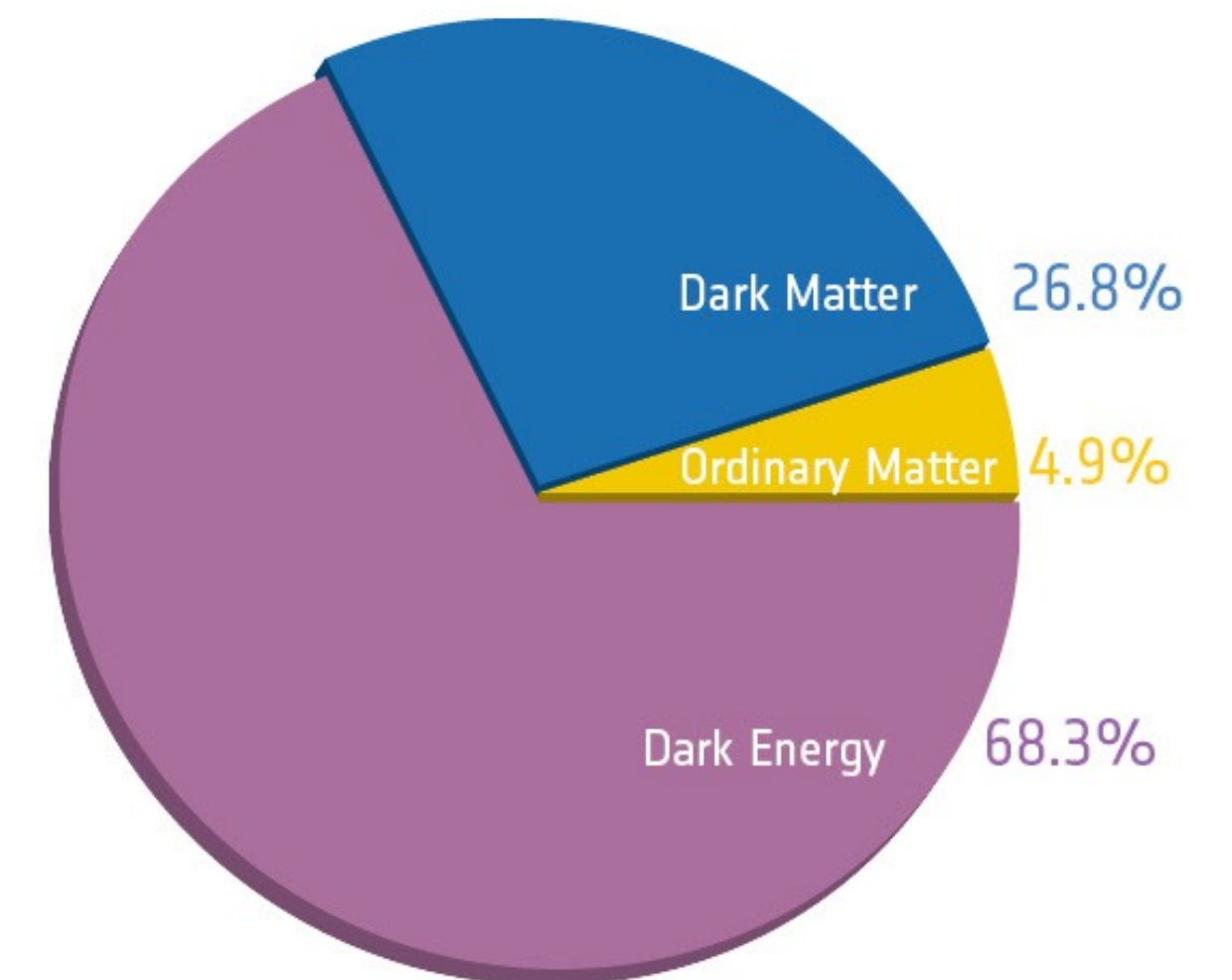
- ▶ constitutes **~25% of the Universe**
- ▶ behaves like matter ($p \sim 0$)
- ▶ plays a crucial role in cosmological structure formation
- ▶ **needs BSM!!!**



Dark Matter exists in the Universe

General properties

- ▶ stable (or lifetime longer than the age of the universe)
- ▶ collision-less (→ **no electric charge**)
- ▶ non-relativistic
- ▶ abundant in the universe ($\Omega h^2=0.12$)



We know almost **nothing** about dark matter

- ▶ **mass** scale?
- ▶ **elementary** or **composite**?
- ▶ non-gravitational interactions?
- ▶ single component or more?
- ▶ ...

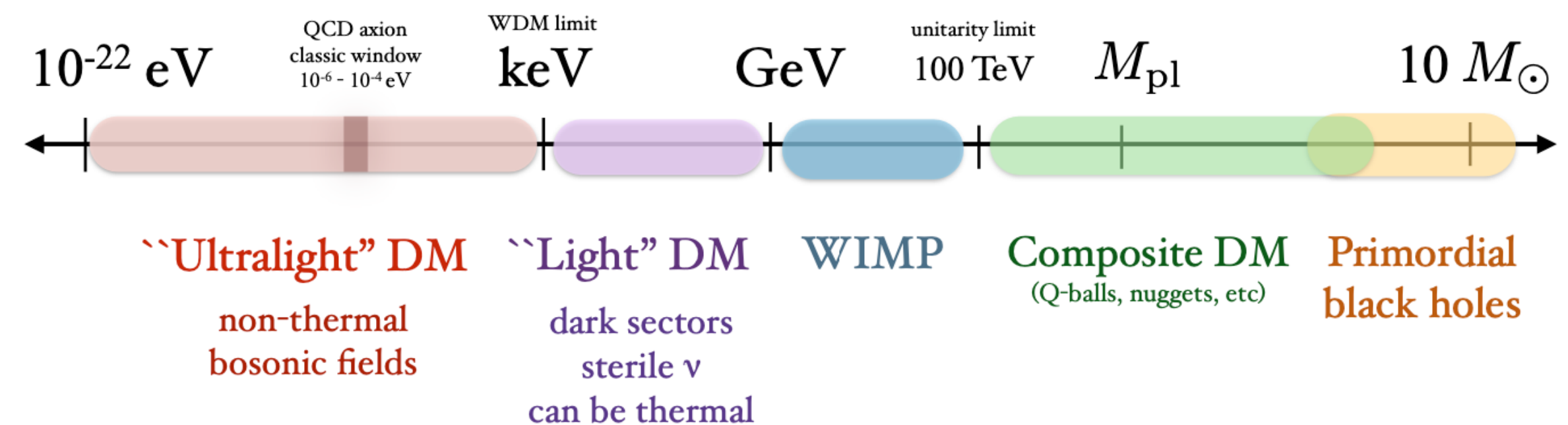


figure from 1904.07915

**Huge parameter space for viable candidates,
so what guiding principle is there?**

What if dark matter has flavor?

of matter fermions

I consider a framework where **DM fields are charged under SM flavor groups**

and show dark matter can naturally be

- stable

- multi-component

Also discuss some phenomenological implications of such multiple dark states

Flavor symmetry and dark matter

Flavor symmetry in the Standard Model

A large global flavor symmetry in the gauge sector

$$\mathcal{G} = \mathrm{U}(3)_{q_L} \times \mathrm{U}(3)_{u_R} \times \mathrm{U}(3)_{d_R} \times \mathrm{U}(3)_{\ell_L} \times \mathrm{U}(3)_{e_R}$$

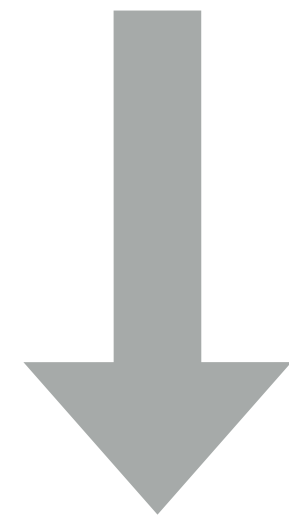
$$q_L^i, u_R^i, d_R^i, \ell_L^i, e_R^i \quad (i = 1, 2, 3)$$

- The matter fermions comprise **five** different gauge representations of Weyl fermions
- There exist **three** species, or flavors in each representation

Flavor symmetry in the Standard Model

A large global flavor symmetry in the gauge sector

$$\mathcal{G} = U(3)_{q_L} \times U(3)_{u_R} \times U(3)_{d_R} \times U(3)_{\ell_L} \times U(3)_{e_R}$$



broken by Yukawa interactions

$$\mathcal{L}_{\text{yuk}} = -\bar{q}_L Y_u \tilde{H} u_R - \bar{q}_L Y_d H d_R - \bar{\ell}_L Y_e H e_R + \text{h.c.}$$

$$U(1)_Y, U(1)_B, U(1)_L, U(1)_{L_e - L_\mu}, U(1)_{L_\mu - L_\tau}$$

broken spontaneously

broken to $U(1)_{B-L}$ by anomaly

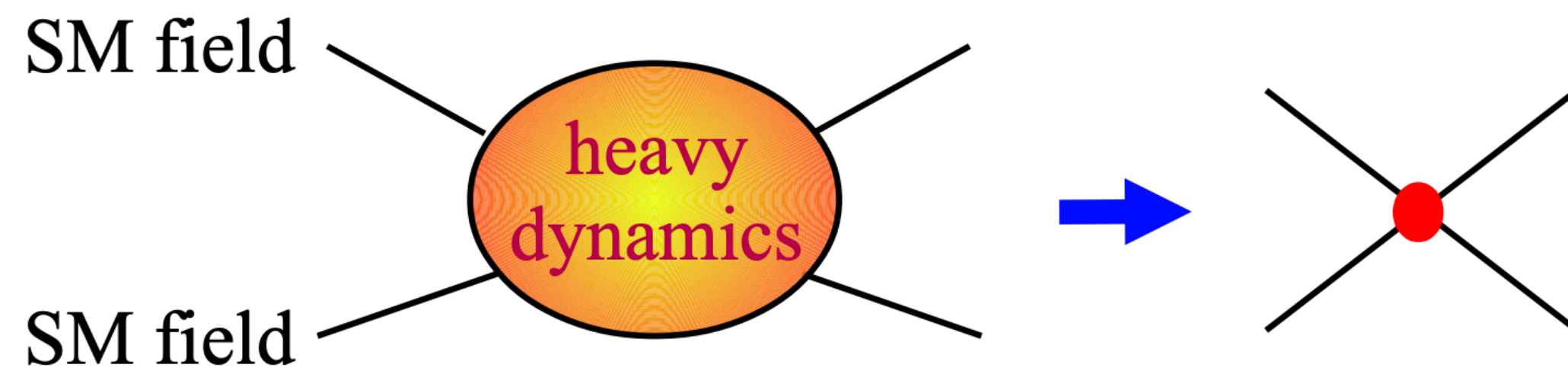
remnant symmetry

Flavor structure in physics **Beyond the SM**

$$\mathcal{L}_{\text{SM-EFT}} = \mathcal{L}_{\text{gauge}} + \mathcal{L}_{\text{yukawa}}$$

$$+ \sum_i \frac{1}{\Lambda_i^{d-4}} \mathcal{O}_i^{d \geq 5}$$

contact int. from heavy BSM dynamics



What is the flavor structure in BSM?

If flavor symmetry maximally broken in BSM,
we find

$$\frac{1}{\Lambda^2} (s_L \gamma^\mu d_L)^2 \Rightarrow \Lambda \gtrsim \text{PeV} - 100 \text{ PeV}$$

=> stringent bounds on heavy dynamics

=> BSM is not directly accessible?

Minimal Flavor Violation (MFV) hypothesis

[Chivukula, Georgi '87; Hall, Randall '90; D'Ambrosio et al. '02]

All flavor violation are caused solely by the Yukawa matrices

- Promote the Yukawa matrices to **spurious fields** transforming like

$$Y_u \sim (\mathbf{3}, \bar{\mathbf{3}}, 1, 1, 1), \quad Y_d \sim (\mathbf{3}, 1, \bar{\mathbf{3}}, 1, 1), \quad Y_e \sim (1, 1, 1, \mathbf{3}, \bar{\mathbf{3}}).$$

$$\text{under } U(3)_{q_L} \times U(3)_{u_R} \times U(3)_{d_R} \times U(3)_{\ell_L} \times U(3)_{e_R}$$

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- This makes Yukawa Lagrangian flavor singlet: $\mathcal{L}_{\text{yuk}} = -\bar{q}_L Y_u \tilde{H} u_R - \bar{q}_L Y_d H d_R - \bar{\ell}_L Y_e H e_R + \text{h.c.}$

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- All new physics interactions should also be flavor invariant

$$\mathcal{L}_{\text{NP}} = C_{ij} \left(\bar{u}_{Ri} \gamma^\mu u_{Rj} \right) \mathcal{O}_\mu \quad \rightarrow \quad C_{ij} = c_0 \delta_{ij} + \epsilon c_1 (Y_u^\dagger Y_u)_{ij} + \epsilon^2 \left[c_2 (Y_u^\dagger Y_u Y_u^\dagger Y_u)_{ij} + c'_2 (Y_u^\dagger Y_d Y_d^\dagger Y_u)_{ij} \right]$$

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$$\frac{1}{\Lambda^2} (s_L \gamma^\mu d_L)^2 : \frac{1}{\Lambda^2} \rightarrow \frac{(Y_u Y_u^\dagger)_{sd}^2}{\Lambda^2} = \frac{y_t^4 (V_{ts} V_{td}^*)^2}{\Lambda^2} \sim \frac{10^{-7}}{\Lambda^2} \quad \rightarrow \quad \text{orders of magnitude lower BSM scale is allowed!}$$

MFV stabilizes dark matter

[Batell, Pradler, Spannowsky '11]

- Consider a new QCD singlet in a **non-trivial representation of $SU(3)_{q_L} \times SU(3)_{u_R} \times SU(3)_{d_R}$**

$$\chi \sim (n_{q_L}, m_{q_L}) \times (n_{u_R}, m_{u_R}) \times (n_{d_R}, m_{d_R})$$

n's, m's are Dynkin coefficients; (1,0)=triplet, (1,1)=octet

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- General decay operators

$$\mathcal{O}_{\text{decay}} = \chi \underbrace{q_L \dots}_{A} \underbrace{\bar{q}_L \dots}_{\bar{A}} \underbrace{u_R \dots}_{B} \underbrace{\bar{u}_R \dots}_{\bar{B}} \underbrace{d_R \dots}_{C} \underbrace{\bar{d}_R \dots}_{\bar{C}} \times \underbrace{Y_u \dots}_{D} \underbrace{Y_u^\dagger \dots}_{\bar{D}} \underbrace{Y_d \dots}_{E} \underbrace{Y_d^\dagger \dots}_{\bar{E}} \mathcal{O}_{\text{weak}}$$

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 $SU(3)_c$ invariant

$$(A + B + C - \bar{A} - \bar{B} - \bar{C}) \bmod 3 = 0$$

 only $q\bar{q}$, qqq can be QCD singlet

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$SU(3)_c$ and
flavor invariant



$$(A + B + C - \bar{A} - \bar{B} - \bar{C}) \bmod 3 = 0$$

$$(n_{q_L} - m_{q_L} + A - \bar{A} + D - \bar{D} + E - \bar{E}) \bmod 3 = 0$$

$$(n_{u_R} - m_{u_R} + B - \bar{B} - D + \bar{D}) \bmod 3 = 0$$

$$(n_{d_R} - m_{d_R} + C - \bar{C} - E + \bar{E}) \bmod 3 = 0$$



$$(n_\chi - m_\chi) \bmod 3 = 0$$

$$n_\chi = n_{q_L} + n_{u_R} + n_{d_R}$$

$$m_\chi = m_{q_L} + m_{u_R} + m_{d_R}$$

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$$(A + B + C - \bar{A} - \bar{B} - \bar{C}) \bmod 3 = 0$$

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$$(n_{u_R} - m_{u_R} + B - \bar{B} - D + \bar{D}) \bmod 3 = 0$$

$$(n_{d_R} - m_{d_R} + C - \bar{C} - E + \bar{E}) \bmod 3 = 0$$



$$(n_\chi - m_\chi) \bmod 3 \neq 0$$

$$n_\chi = n_{q_L} + n_{u_R} + n_{d_R}$$

$$m_\chi = m_{q_L} + m_{u_R} + m_{d_R}$$

For χ to be stable, at least one of four equations should **NOT** be satisfied

Why flavored DM stabilized within MFV?

There is an unbroken **Z_3 symmetry** $\subset \text{SU}(3)_c \times \text{SU}(3)_{q_L} \times \text{SU}(3)_{u_R} \times \text{SU}(3)_{d_R}$ [Batell, Lin, Wang '13]

► Z_3 charge: $\psi \rightarrow U\psi$ with $U = (\omega^2)^{n_c - m_c} \cdot (\omega)^{n_q - m_q} \cdot (\omega)^{n_u - m_u} \cdot (\omega)^{n_d - m_d}$ where $\omega^3 = 1$

► All SM fields and Yukawa spurions are singlet

- quarks: $Q \rightarrow (\omega^2 \cdot \omega) Q = Q$
- other SM fields: $\phi \rightarrow \phi$

► Flavored DM: $\chi \rightarrow (\omega)^{n_\chi - m_\chi} \chi$

- χ is Z_3 non-singlet if $(n_\chi - m_\chi) \bmod 3 \neq 0 \rightarrow$ **stabilized**

Flavored DM candidates

(n, m)	$SU(3)_Q \times SU(3)_{u_R} \times SU(3)_{d_R}$	Stable?
(0, 0)	(1, 1, 1)	
(1, 0)	(3 , 1, 1), (1, 3 , 1), (1, 1, 3)	Yes
(0, 1)	($\bar{\mathbf{3}}$, 1, 1), (1, $\bar{\mathbf{3}}$, 1), (1, 1, $\bar{\mathbf{3}}$)	Yes
(2, 0)	(6 , 1, 1), (1, 6 , 1), (1, 1, 6) (3 , 3 , 1), (3 , 1, 3), (1, 3 , 3)	Yes
(0, 2)	($\bar{\mathbf{6}}$, 1, 1), (1, $\bar{\mathbf{6}}$, 1), (1, 1, $\bar{\mathbf{6}}$) ($\bar{\mathbf{3}}$, $\bar{\mathbf{3}}$, 1), ($\bar{\mathbf{3}}$, 1, $\bar{\mathbf{3}}$), (1, $\bar{\mathbf{3}}$, $\bar{\mathbf{3}}$)	Yes
(1, 1)	(8 , 1, 1), (1, 8 , 1), (1, 1, 8) (3 , $\bar{\mathbf{3}}$, 1), (3 , 1, $\bar{\mathbf{3}}$), (1, 3 , $\bar{\mathbf{3}}$) ($\bar{\mathbf{3}}$, 3 , 1), ($\bar{\mathbf{3}}$, 1, 3), (1, $\bar{\mathbf{3}}$, 3)	

- independent of **spin** and **EW representation** of χ
- Only the lightest flavored state is stabilized due to MFV
 - All heavy flavors quickly decay, and only the lightest flavor is DM (Batell+ '11; Lopez-Honorez+ '13)
 - Some heavy flavors are decaying but long-lived enough to serve as DM → **multi-component DM** scenario

[Mescia, **SO**, Wu, 2408.16812]

[Batell, Pradler, Spannowsky '11]

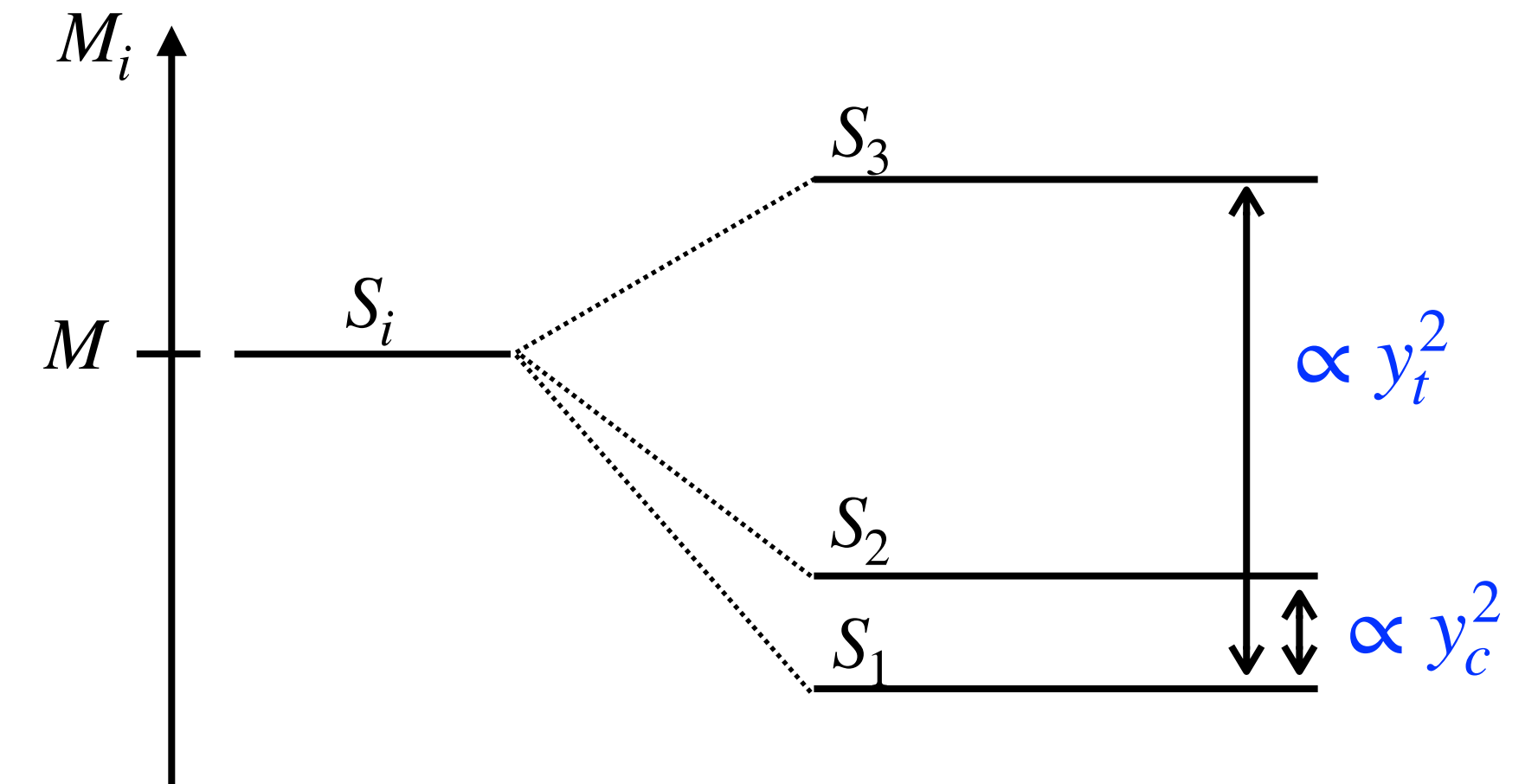
A benchmark study of a dark matter family

2408.16812 with Mescia, Wu

- A gauge singlet, $SU(3)_{u_R}$ triplet scalar $S \sim (\mathbf{1}, \mathbf{3}, 1)$
- Both mass and interactions controlled by the Yukawa couplings

A benchmark study of a dark matter family

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● Scalar potential

$$V(H, S) = \{m_0^2 + \epsilon m_1^2 (y_u^i)^2\} S_i^* S_i \quad \rightarrow$$

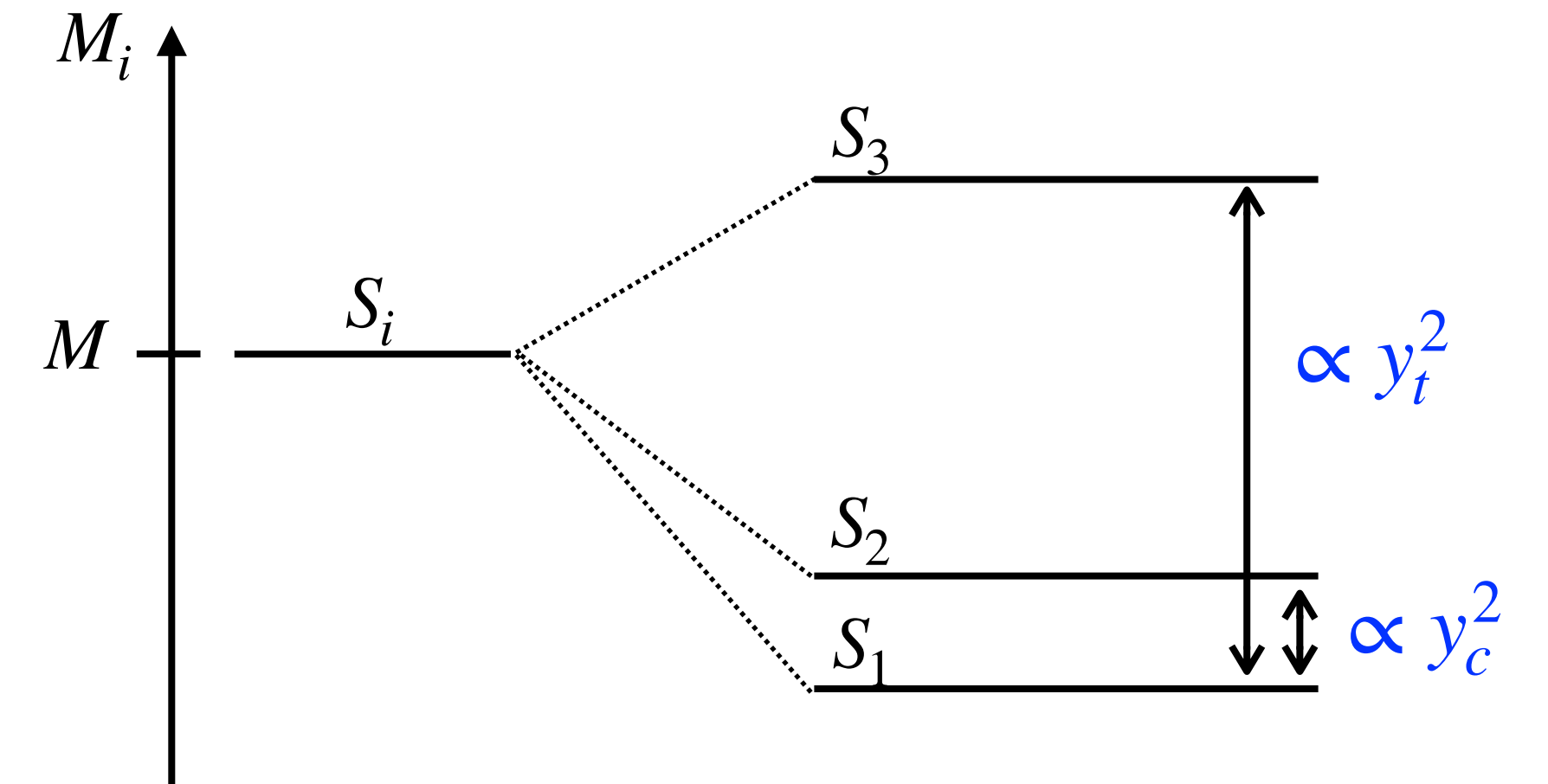
$$M_j^2 - M_i^2 = \epsilon m_1^2 [(y_u^j)^2 - (y_u^i)^2]$$

$$+ \frac{\lambda}{2} (b_0 + \epsilon b_1 (y_u^i)^2) (2vh + h^2) S_i^* S_i \quad \rightarrow$$

flavor diagonal, so doesn't lead heavy scalar decay

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Scalar potential

$$V(H, S) = \{m_0^2 + \epsilon m_1^2 (y_u^i)^2\} S_i^* S_i \quad \rightarrow \quad M_j^2 - M_i^2 = \epsilon m_1^2 [(y_u^j)^2 - (y_u^i)^2]$$

$$+ \frac{\lambda}{2} (b_0 + \epsilon b_1 (y_u^i)^2) (2vh + h^2) S_i^* S_i \quad \rightarrow \quad \text{flavor diagonal, so doesn't lead heavy scalar decay}$$

Dim-6 operators

$$\mathcal{L}_{d=6} \sim \frac{c_2^4}{\Lambda^2} \left(\bar{q}_{Li} (Y_u)_{ij} S_j \right) \tilde{H} (S_k^* \delta_{kl} u_{Rl}) + \text{h.c.}$$

$$\sim \frac{c_2^4}{\Lambda^2} \bar{u}_i (m_u^i P_R + m_u^j P_L) u_j (S_j^* S_i)$$



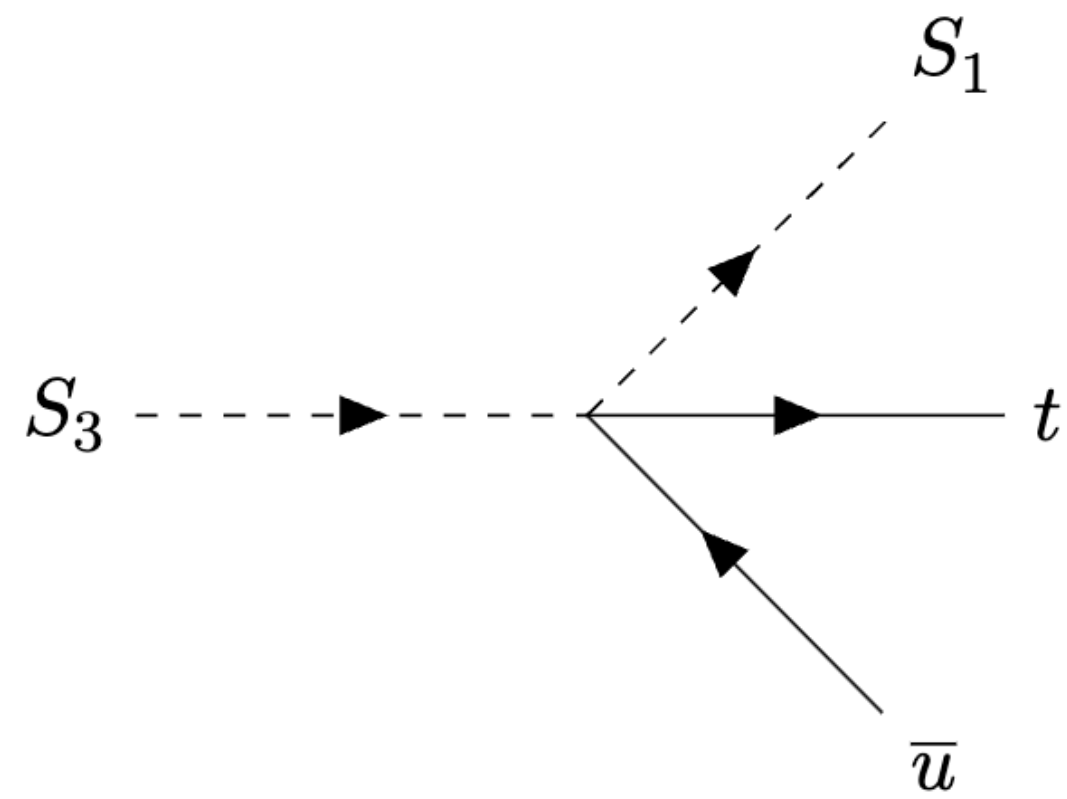
$$S_3 \rightarrow S_1 t \bar{u}, S_2 t \bar{c}$$

Heavy scalar decay is triggered!

Decay of heavy states

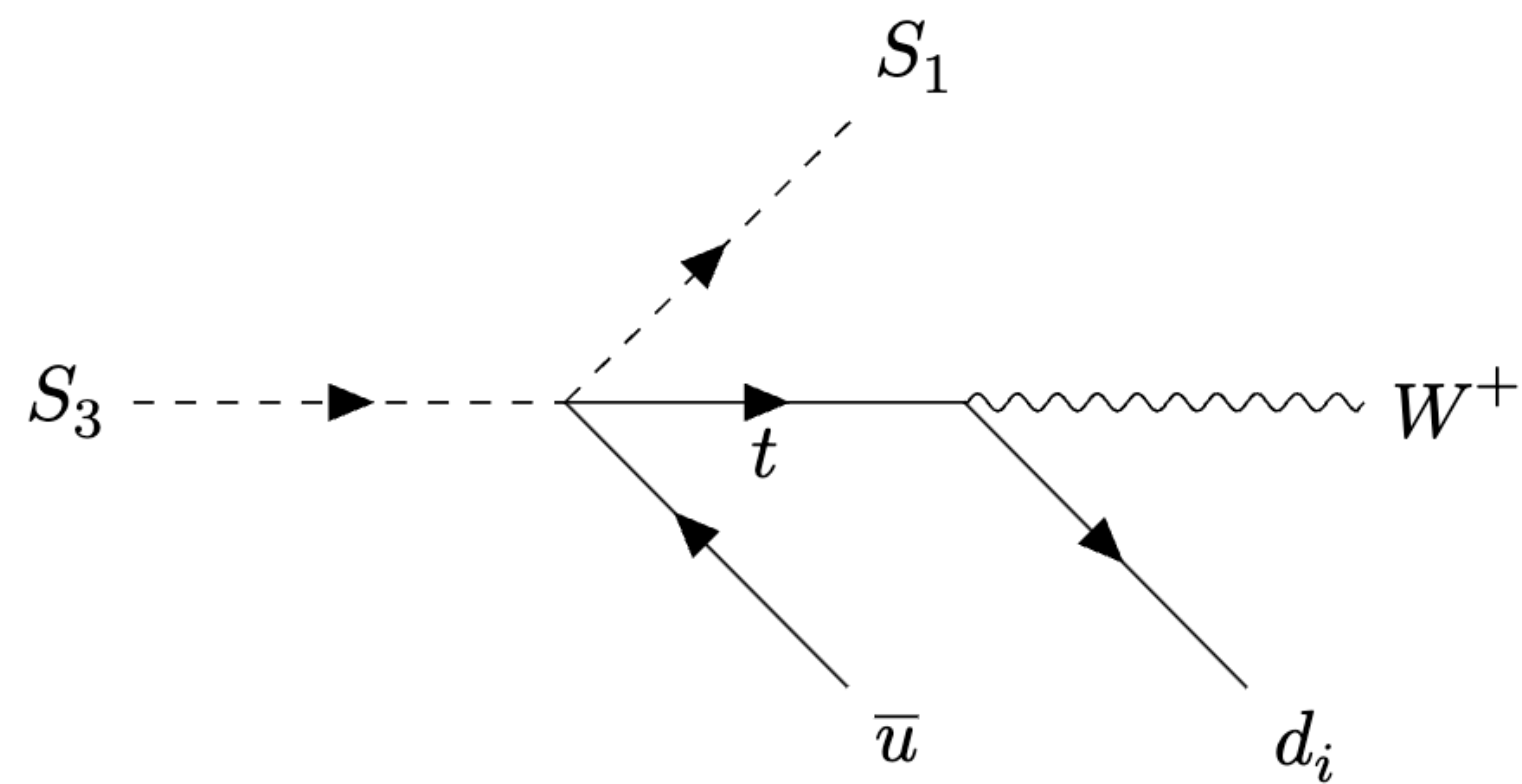
- Dominant mode depends on the mass splitting $\Delta M = M_3 - M_1$

$$\Delta M \gtrsim m_t$$



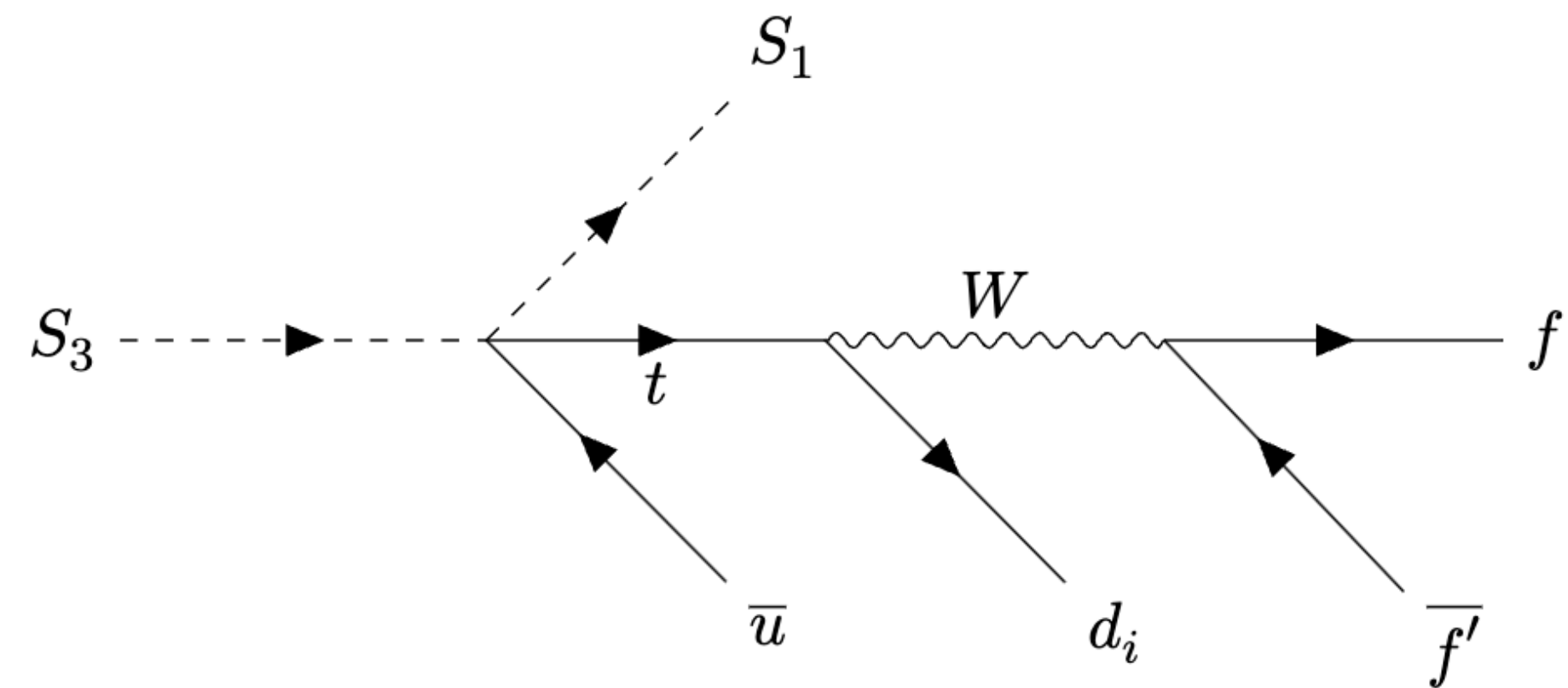
$$\Gamma \sim \frac{m_t^2 (\Delta M)^5}{480\pi^3 \Lambda^4 M_3^2}$$

$$m_t \gtrsim \Delta M \gtrsim m_W + m_d^i$$



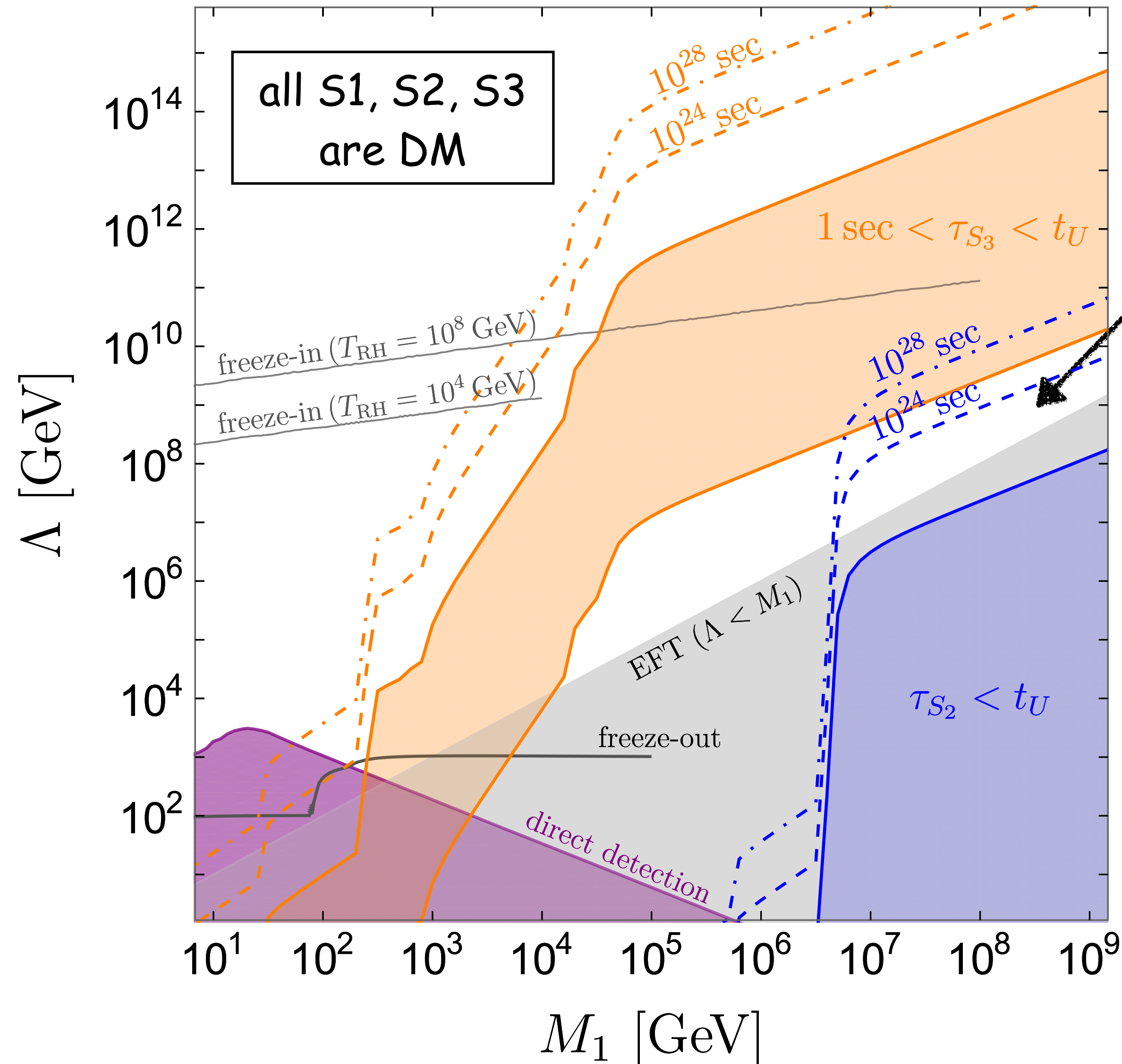
$$\Gamma \sim \frac{(\Delta M)^{11} |V_{ti}|^2}{41472\pi^5 \Lambda^4 M_3^2 m_t^2 v^2}$$

$$m_W + m_d^i \gtrsim \Delta M \gtrsim m_u + m_d^i + m_f + m_{f'}$$



$$\Gamma \sim \frac{(\Delta M)^{13} |V_{ti}|^2}{11612160\pi^7 \Lambda^4 M_3^2 m_t^2 v^4}$$

Parameter spaces for multi-component DM



$$\epsilon = 10^{-2} \simeq \frac{M_3 - M_1}{y_t^2 M_1} \simeq \frac{M_2 - M_1}{y_c^2 M_1}$$

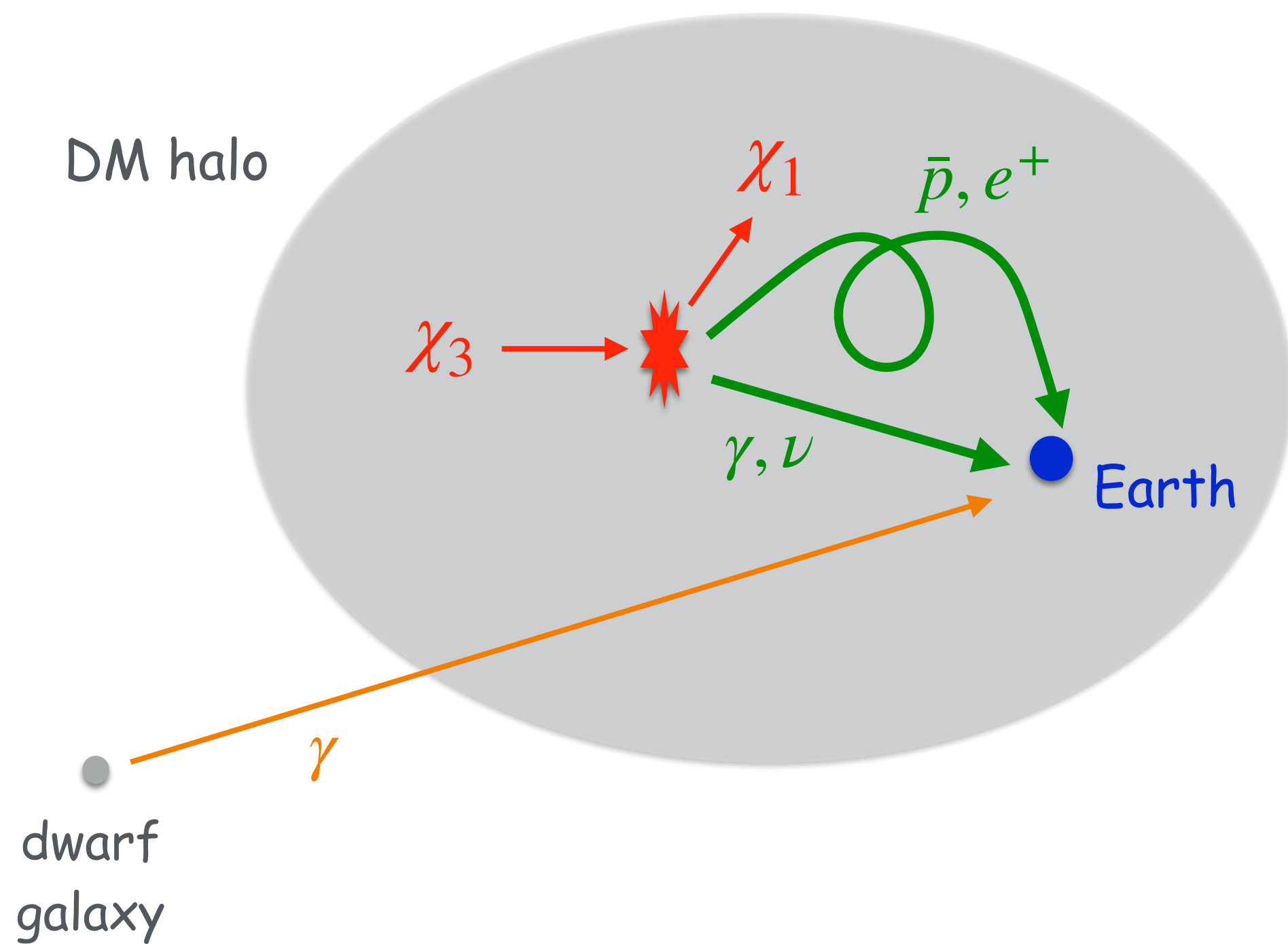
$$\lambda = 0 \quad (\text{no coupling to Higgs})$$

- $\tau_{S_i} > \tau_U \rightarrow \text{DM}$
- $\tau_{S_i} < \tau_U \rightarrow \text{not DM and has to decay prior to the BBN (we require } \tau_{S_i} < 1 \text{ sec)}$
- DM is composed of two or three components in the white region

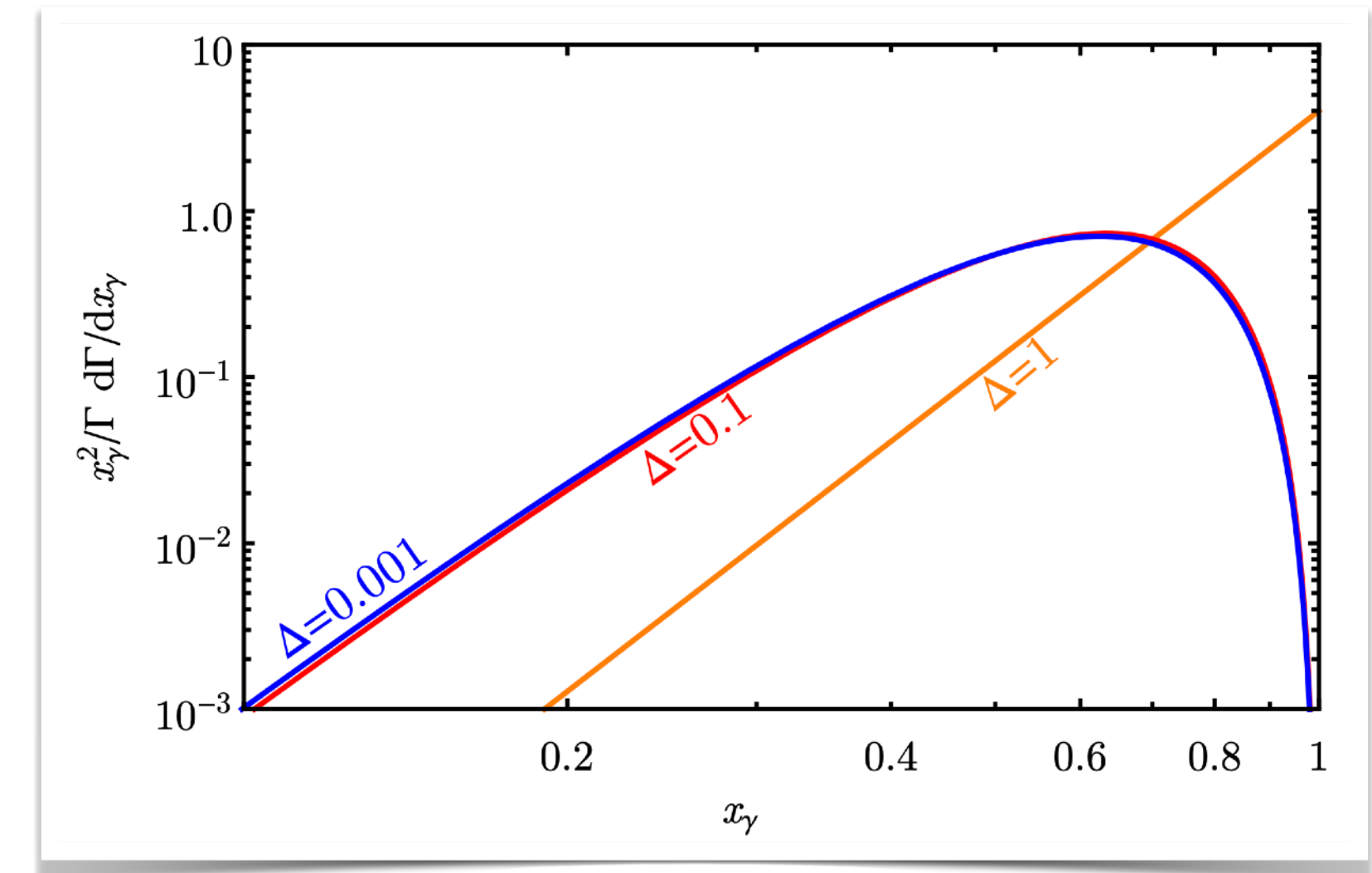
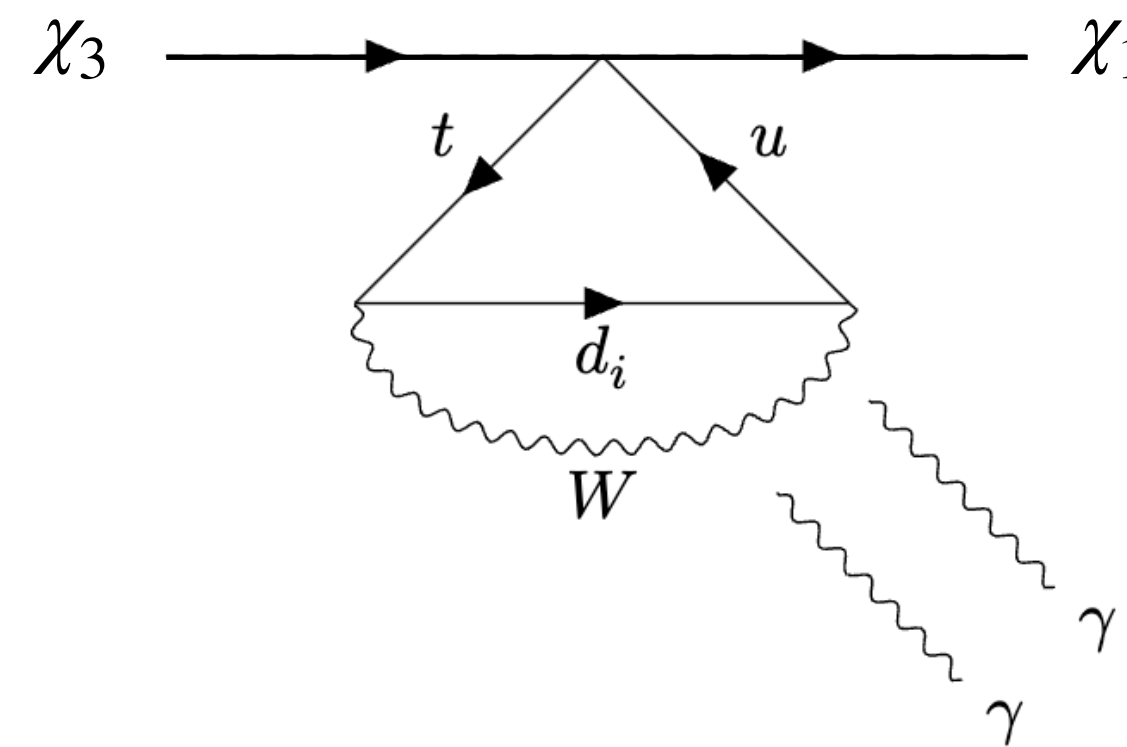
Signatures from this dark matter family?

I. Indirect searches for heavy decaying dark matter

with Dondarini, Landini, Mescia



● $\chi_{2,3} \rightarrow \chi_1 \gamma \gamma$ [Dienes+ 1406.4868; Ghosh+ 1909.13292; Herms, Ibarra 1912.09458]



We can also extend to

- ▶ other hadronic decay modes $\chi_{2,3} \rightarrow \chi_1 qq, \chi_1 qqqq, \dots$
- ▶ cosmology (CMB, BBN, Lyman- α , 21 cm line, etc.)

2. Flavored DM from rare meson decays

with Mescia, Swallow, Toni

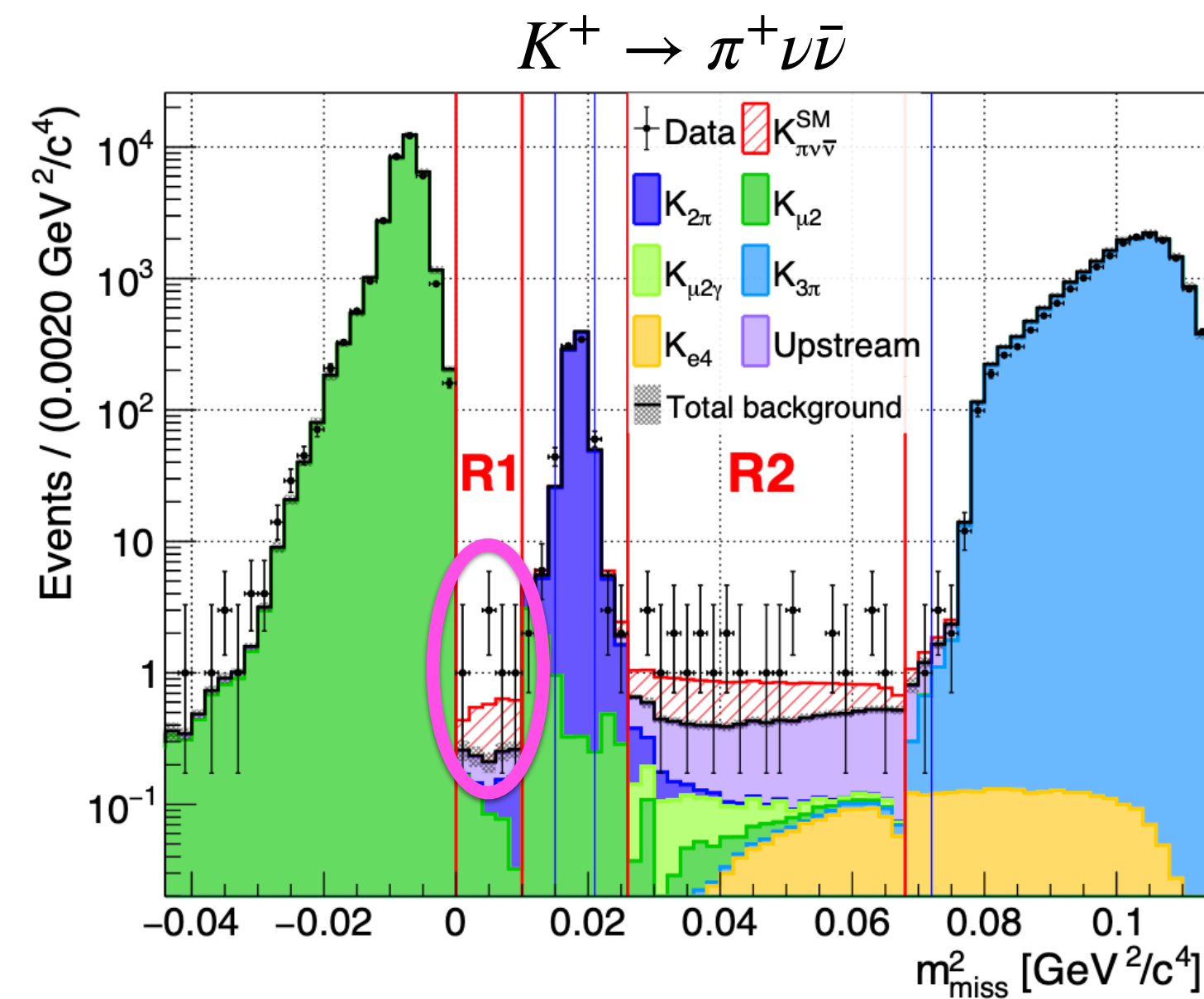
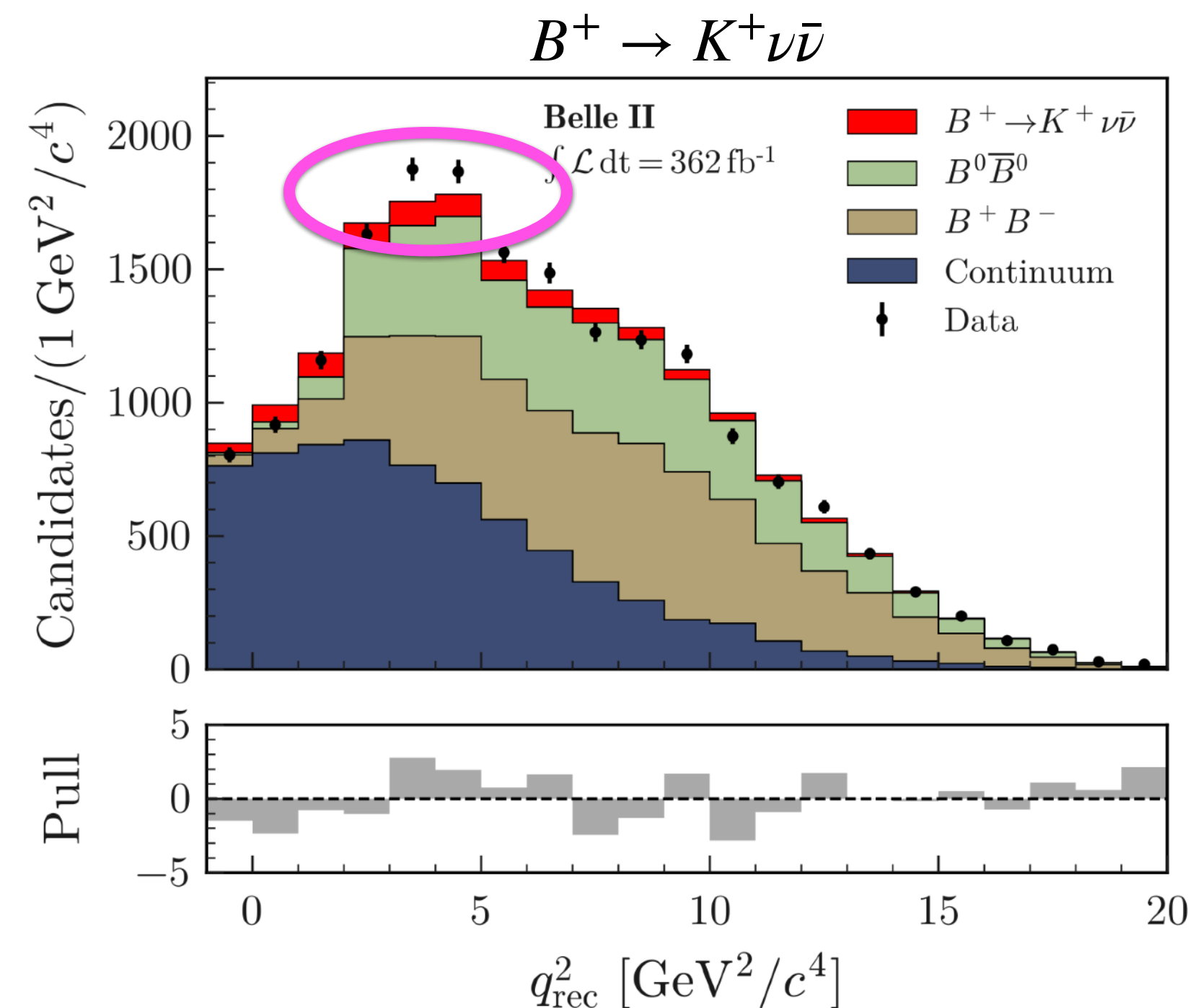
- Apparent flavor violating decays can naturally occur

e.g.) $b \rightarrow s\chi_3\bar{\chi}_2$ and $s \rightarrow d\chi_1\bar{\chi}_2 \Rightarrow$ mimic $B \rightarrow K\nu\nu$ and $K \rightarrow \pi\nu\nu$ processes!

@Belle-II
BaBar

@NA62
KOTO

- Mild excesses are reported by the Belle-II and NA62



2. Flavored DM from rare meson decays

with Mescia, Swallow, Toni

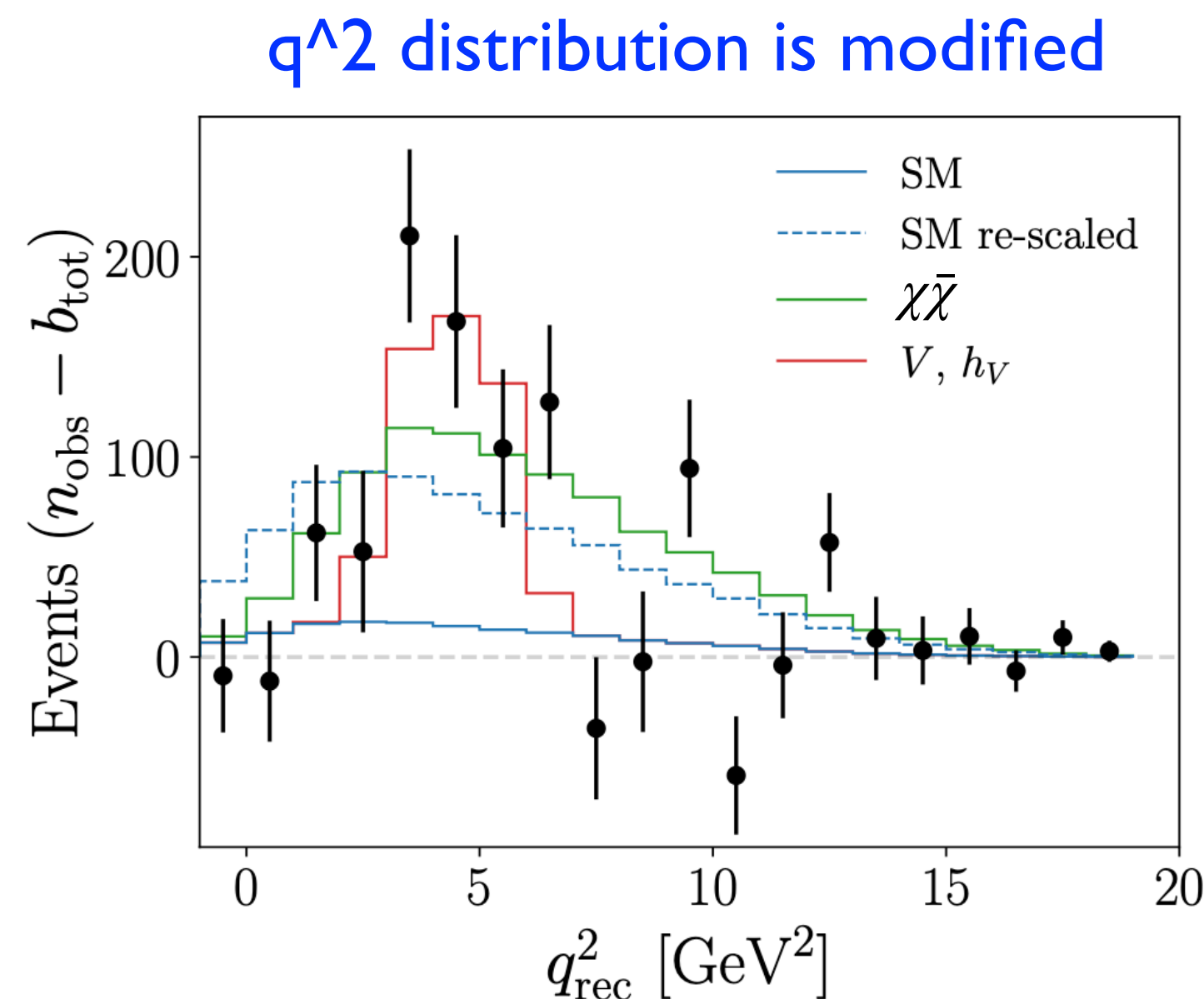
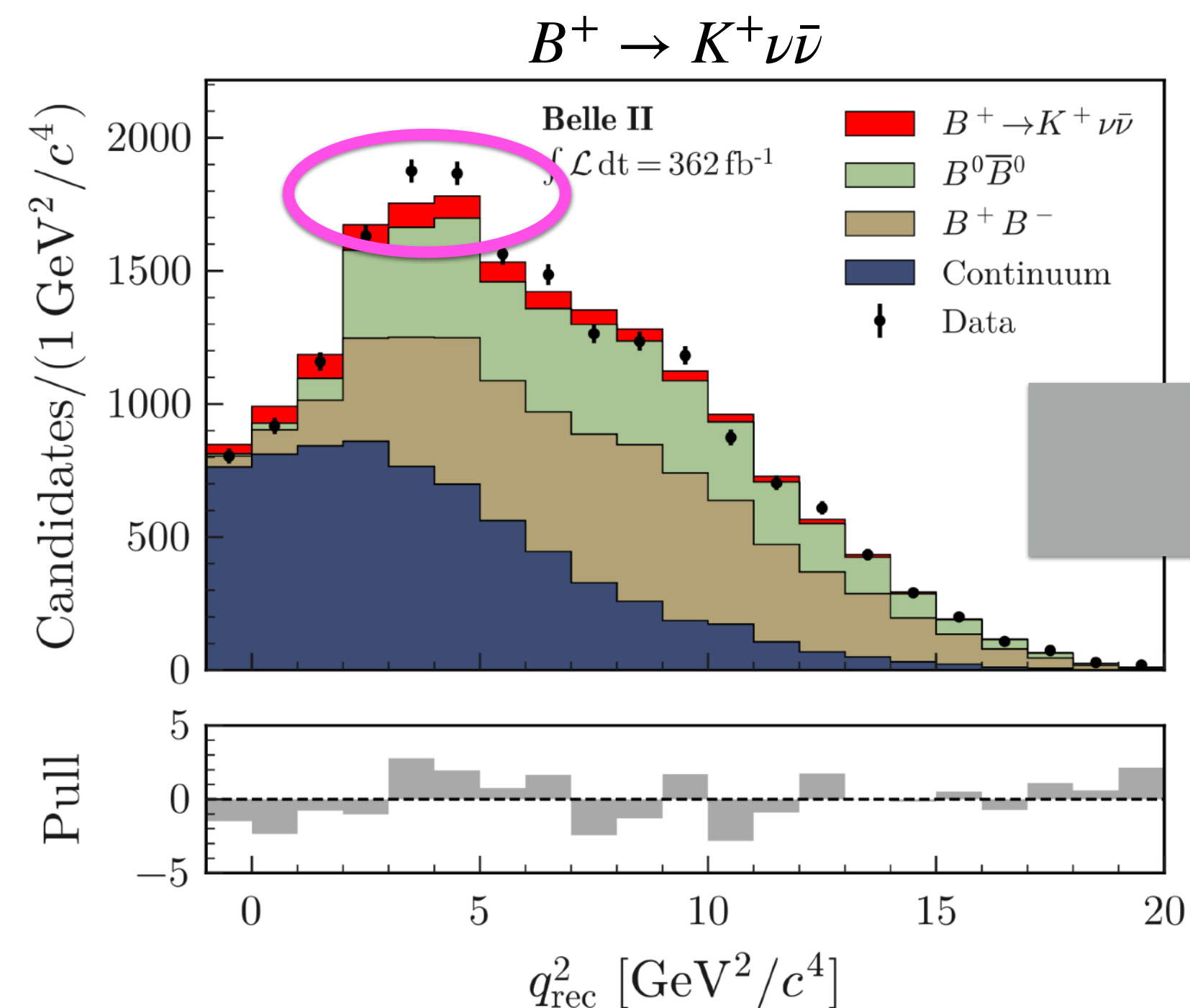
- Apparent flavor violating decays can naturally occur

e.g.) $b \rightarrow s\chi_3\bar{\chi}_2$ and $s \rightarrow d\chi_1\bar{\chi}_2 \Rightarrow$ mimic $B \rightarrow K\nu\nu$ and $K \rightarrow \pi\nu\nu$ processes!

@Belle-II
BaBar

@NA62
KOTO

- Mild excesses are reported by the Belle-II and NA62



Need more data, but future update might confirm some deviation in these decay modes

Summary

- The existence of DM requires BSM
- Non-trivial flavor structures in BSM may provide a new avenue for theories of dark matter
 - quark flavored DM + MFV $\rightarrow$ multicomponent dark matter
 - very rich phenomenology and cosmology
 - still many questions
 - relation to the flavor puzzles?
 - extension to lepton flavor?
 - $U(3)^5$, $U(2)^5$, or others?

Thanks for listening!

Back up

A benchmark study of a dark matter family

2408.16812 with Mescia, Wu

- A gauge singlet, $SU(3)_{u_R}$ triplet scalar $S \sim (\mathbf{1}, \mathbf{3}, 1)$
- Scalar potential within MFV ($\epsilon \ll 1$)

$$\begin{aligned} V(H, S) = & m_S^2 S_i^* \left(a_0 \delta_{ij} + \epsilon a_1 (Y_u^\dagger Y_u)_{ij} + \dots \right) S_j && \text{mass term} \\ & + \lambda S_i^* \left(b_0 \delta_{ij} + \epsilon b_1 (Y_u^\dagger Y_u)_{ij} + \dots \right) S_j (H^\dagger H) && \text{coupling to the Higgs doublet} \\ & + \left(\lambda_0 \delta_{ij} \delta_{kl} + \epsilon \lambda_1 \delta_{ij} (Y_u^\dagger Y_u)_{kl} + \dots \right) S_i^* S_j S_k^* S_l && \text{self-interaction} \end{aligned}$$

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 up to $O(\epsilon)$

$$\begin{aligned} V(H, S) = & \left\{ m_0^2 + \epsilon m_1^2 (y_u^i)^2 \right\} S_i^* S_i \\ & + \frac{\lambda}{2} \left(b_0 + \epsilon b_1 (y_u^i)^2 \right) (2vh + h^2) S_i^* S_i \\ & + \text{self-interaction} \end{aligned}$$

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 & + \left(\lambda_0 \delta_{ij} \delta_{kl} + \epsilon \lambda_1 \delta_{ij} (Y_u^\dagger Y_u)_{kl} + \dots \right) S_i^* S_j S_k^* S_l && \text{self-interaction}
 \end{aligned}$$



$$\begin{aligned}
 V(H, S) = & \left\{ \underbrace{m_0^2}_{\text{flavor independent}} + \epsilon \underbrace{m_1^2 (y_u^i)^2}_{\text{flavor dependent}} \right\} S_i^* S_i \\
 & + \frac{\lambda}{2} \left(\underbrace{b_0}_{\text{flavor independent}} + \epsilon \underbrace{b_1 (y_u^i)^2}_{\text{flavor dependent}} \right) (2vh + h^2) S_i^* S_i
 \end{aligned}$$

up to $O(\epsilon)$

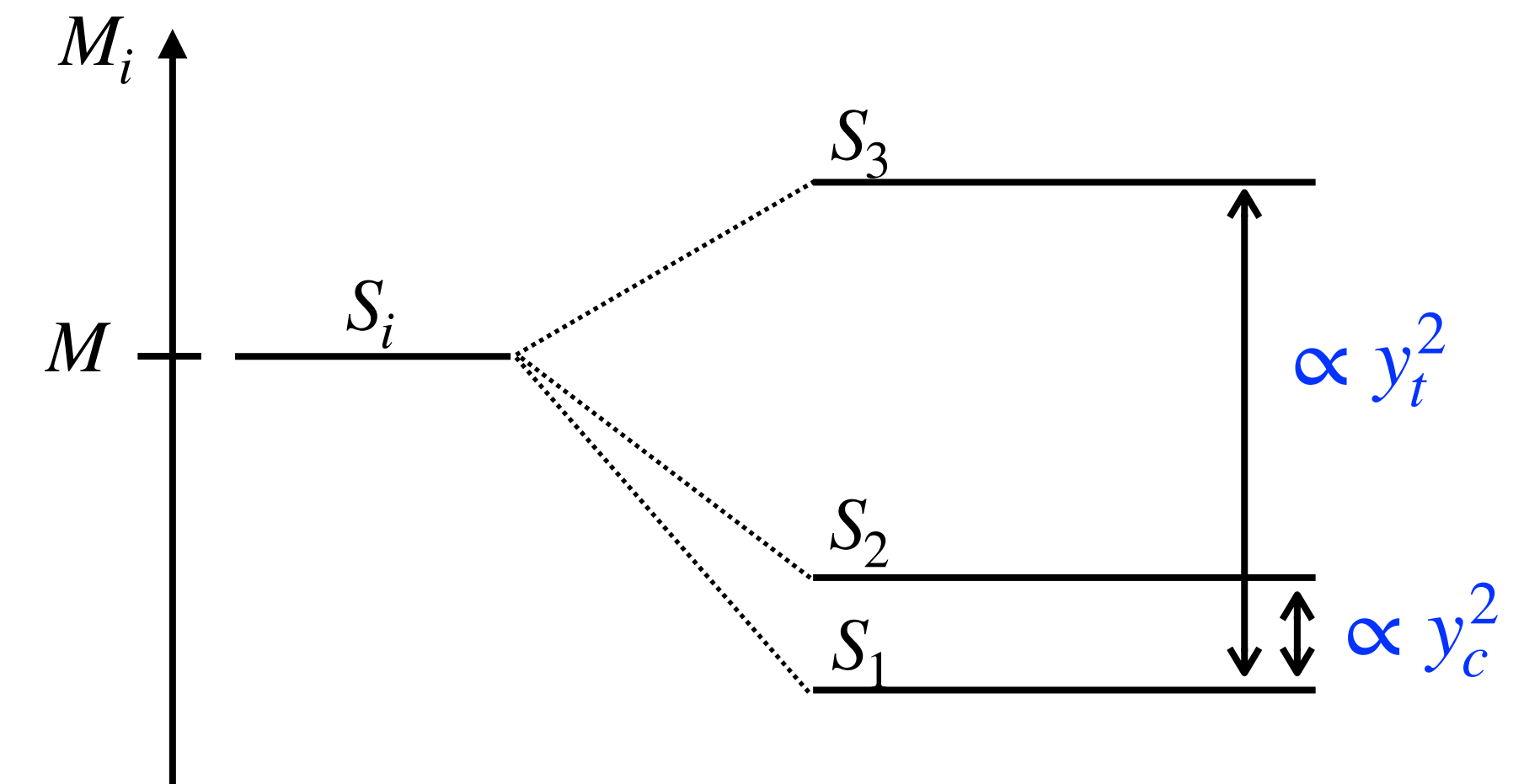
flavor independent flavor dependent

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 \end{aligned}$$



$\Rightarrow$ up to $O(\epsilon)$

$$V(H, S) = \{m_0^2 + \epsilon m_1^2 (y_u^i)^2\} S_i^* S_i + \frac{\lambda}{2} (b_0 + \epsilon b_1 (y_u^i)^2) (2vh + h^2) S_i^* S_i$$

$$M_j^2 - M_i^2 = \epsilon m_1^2 [(y_u^j)^2 - (y_u^i)^2]$$
 $\Rightarrow$ flavor diagonal, so doesn't lead heavy scalar decay

A benchmark study of a dark matter family

2408.16812 with Mescia, Wu

- We can also consider higher dimensional operators

● Dim-6 operators

$$\mathcal{L}_{d=6} = \frac{1}{\Lambda^2} \left(\sum_I c_{ijkl}^I \mathcal{O}_{ijkl}^I + c_{ij}^g \mathcal{O}_{ij}^g + c_{ij}^\gamma \mathcal{O}_{ij}^\gamma \right)$$

$$\mathcal{O}_{ijkl}^1 = (\bar{q}_{Li} \gamma^\mu q_{Lj}) (S_k^* i \overleftrightarrow{\partial}_\mu S_l),$$

$$\mathcal{O}_{ijkl}^3 = (\bar{d}_{Ri} \gamma^\mu d_{Rj}) (S_k^* i \overleftrightarrow{\partial}_\mu S_l),$$

$$\mathcal{O}_{ijkl}^5 = (\bar{q}_{Li} H d_{Rj}) (S_k^* S_l),$$

$$\mathcal{O}_{ij}^\gamma = (S_i^* S_j) F_{\mu\nu} F^{\mu\nu}.$$

$$\mathcal{O}_{ijkl}^2 = (\bar{u}_{Ri} \gamma^\mu u_{Rj}) (S_k^* i \overleftrightarrow{\partial}_\mu S_l),$$

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- Coefficients are determined by the Yukawa matrices

$$\begin{aligned} c_{ijkl}^4 &= c_1^4 (Y_u)_{ij} \delta_{kl} + c_2^4 (Y_u)_{il} \delta_{kj} \\ &+ \epsilon \left[c_3^4 (Y_u Y_u^\dagger Y_u)_{ij} \delta_{kl} + c_4^4 (Y_u Y_u^\dagger Y_u)_{il} \delta_{kj} + c_5^4 (Y_u)_{ij} (Y_u^\dagger Y_u)_{kl} + c_6^4 (Y_u)_{il} (Y_u^\dagger Y_u)_{jl} \right] \\ &+ \dots, \end{aligned}$$

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2408.16812 with Mescia, Wu

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- Coefficients are determined by the Yukawa matrices

$$c_{ijkl}^4 = c_1^4 (Y_u)_{ij} \delta_{kl} + c_2^4 (Y_u)_{il} \delta_{kj}$$

$$\begin{aligned} \Rightarrow \mathcal{L}_{d=6} &\sim \frac{c_2^4}{\Lambda^2} \left(\bar{q}_{Li} (Y_u)_{ij} S_j \right) \tilde{H} (S_k^* \delta_{kl} u_{Rl}) + \text{h.c.} \\ &\sim \frac{c_2^4}{\Lambda^2} \bar{u}_i (m_u^i P_R + m_u^j P_L) u_j (S_j^* S_i) \end{aligned}$$

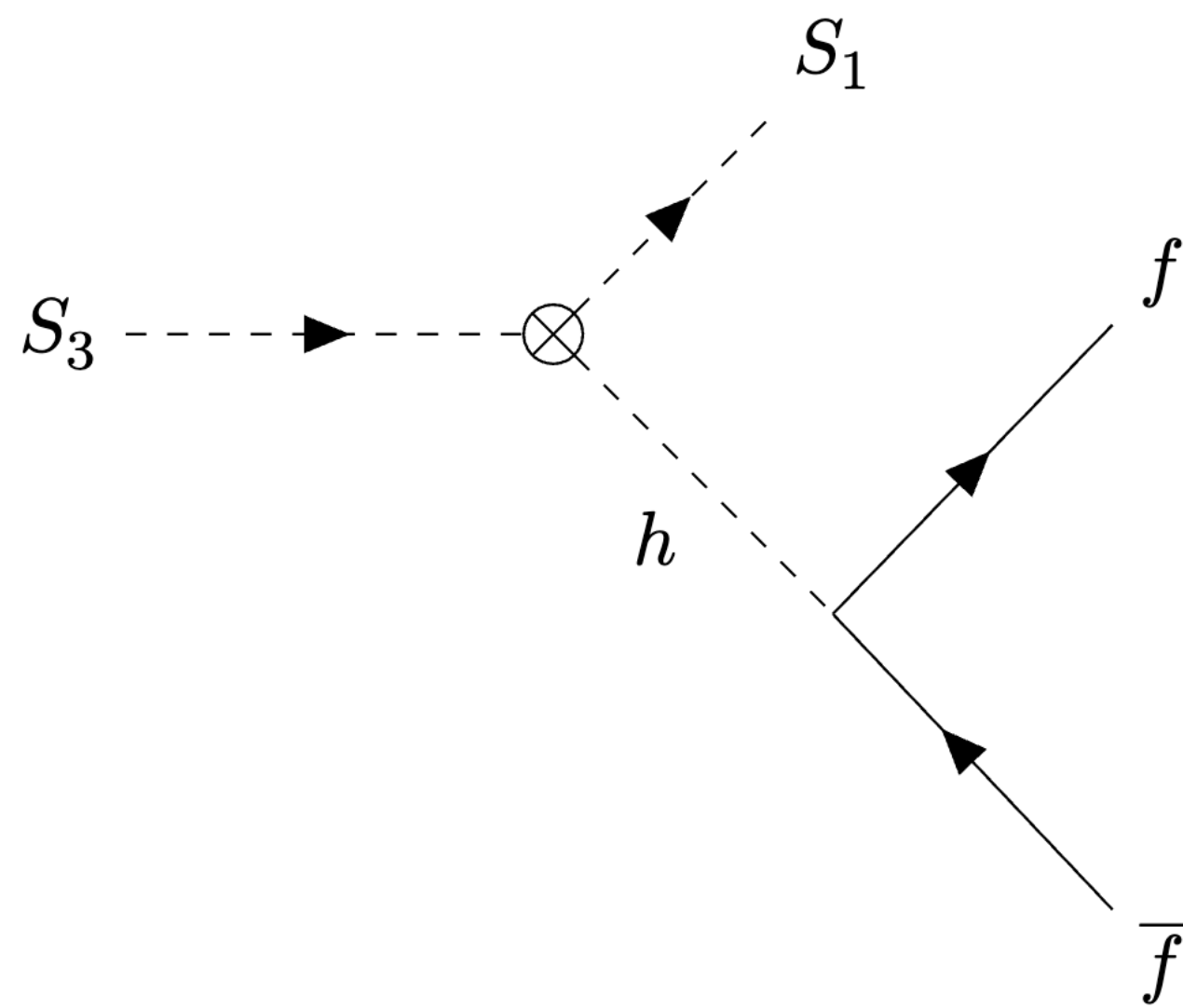
$$S_3 \rightarrow S_1 t \bar{u}, S_2 t \bar{c}$$

Heavy scalar decay is triggered!

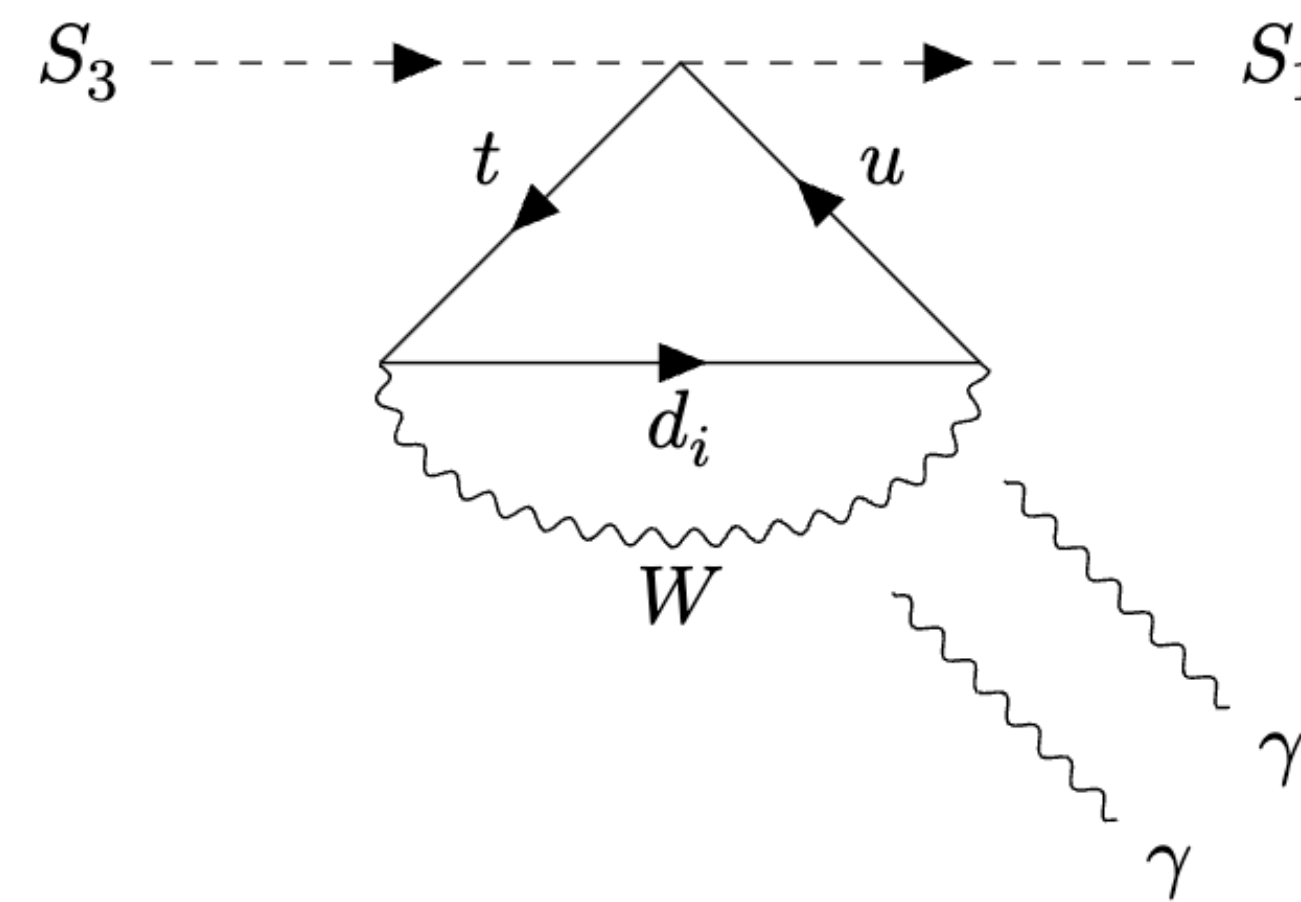
Decay at higher orders

Three-body decay into light particles is induced at higher orders or via loop diagrams

- appears at ϵ^2 order or two-loop level
- can surpass four or five-body ϵ^0 -order processes

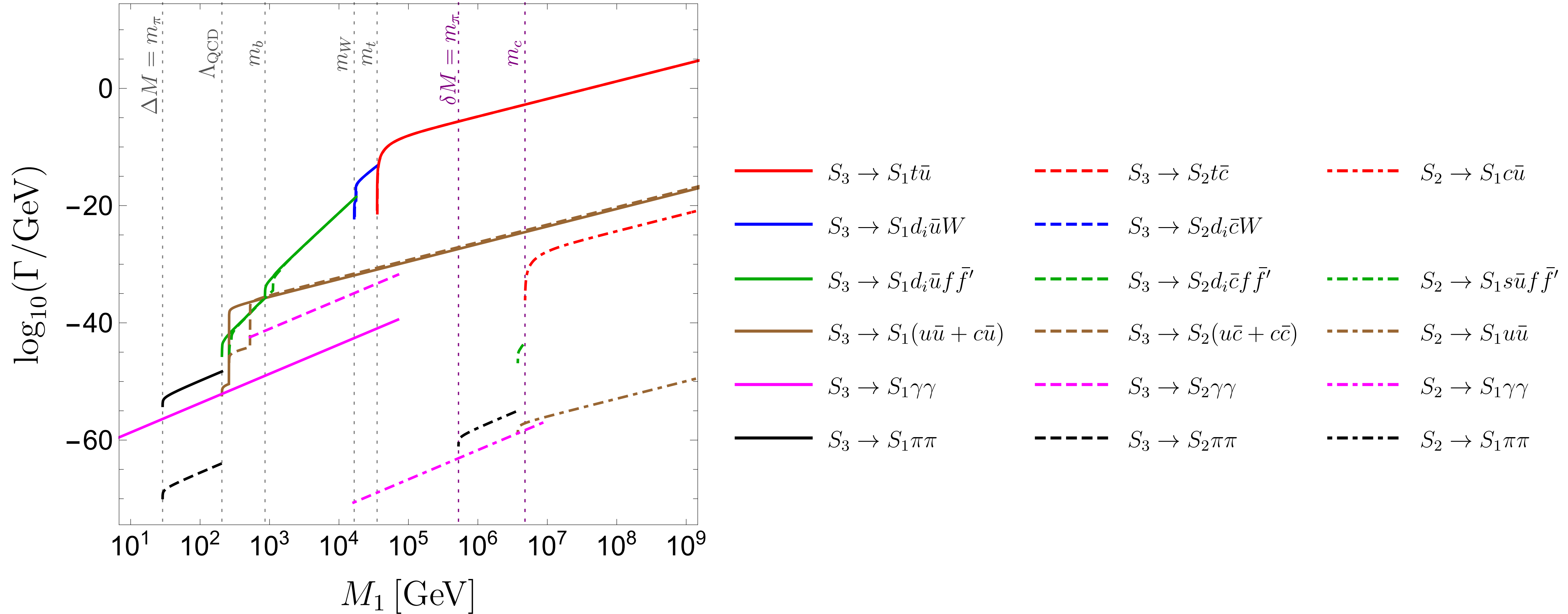


$$\sim \epsilon^2 Y_u^\dagger Y_d Y_d^\dagger Y_u$$

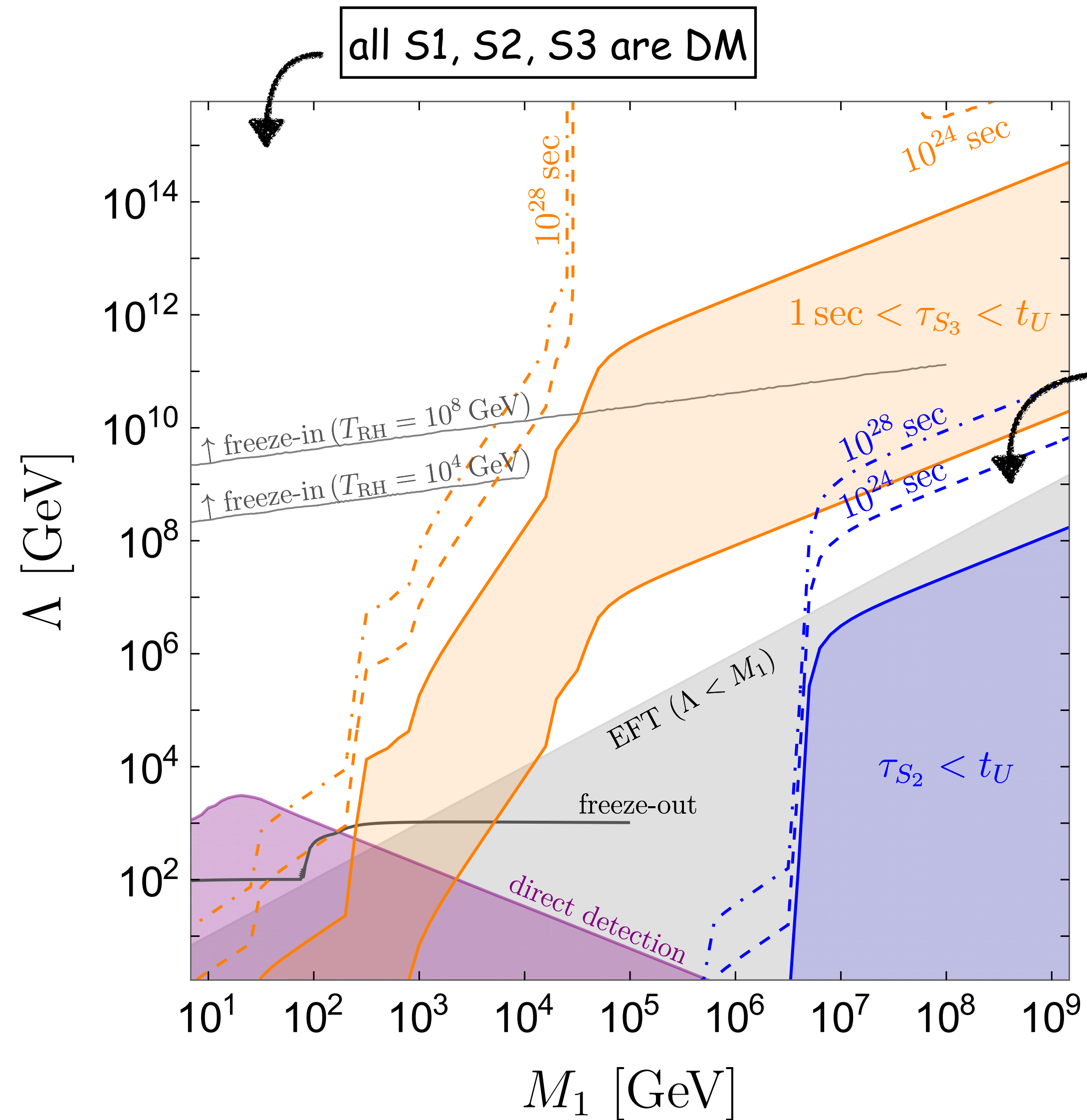


$$\sim \frac{Y_u^\dagger Y_d Y_d^\dagger Y_u}{(16\pi^2)^2}$$

Partial decay widths for heavy scalars



Impact of Higgs portal coupling (1/2)



$$\epsilon = 10^{-2} \simeq \frac{M_3 - M_1}{y_t^2 M_1} \simeq \frac{M_2 - M_1}{y_c^2 M_1}$$

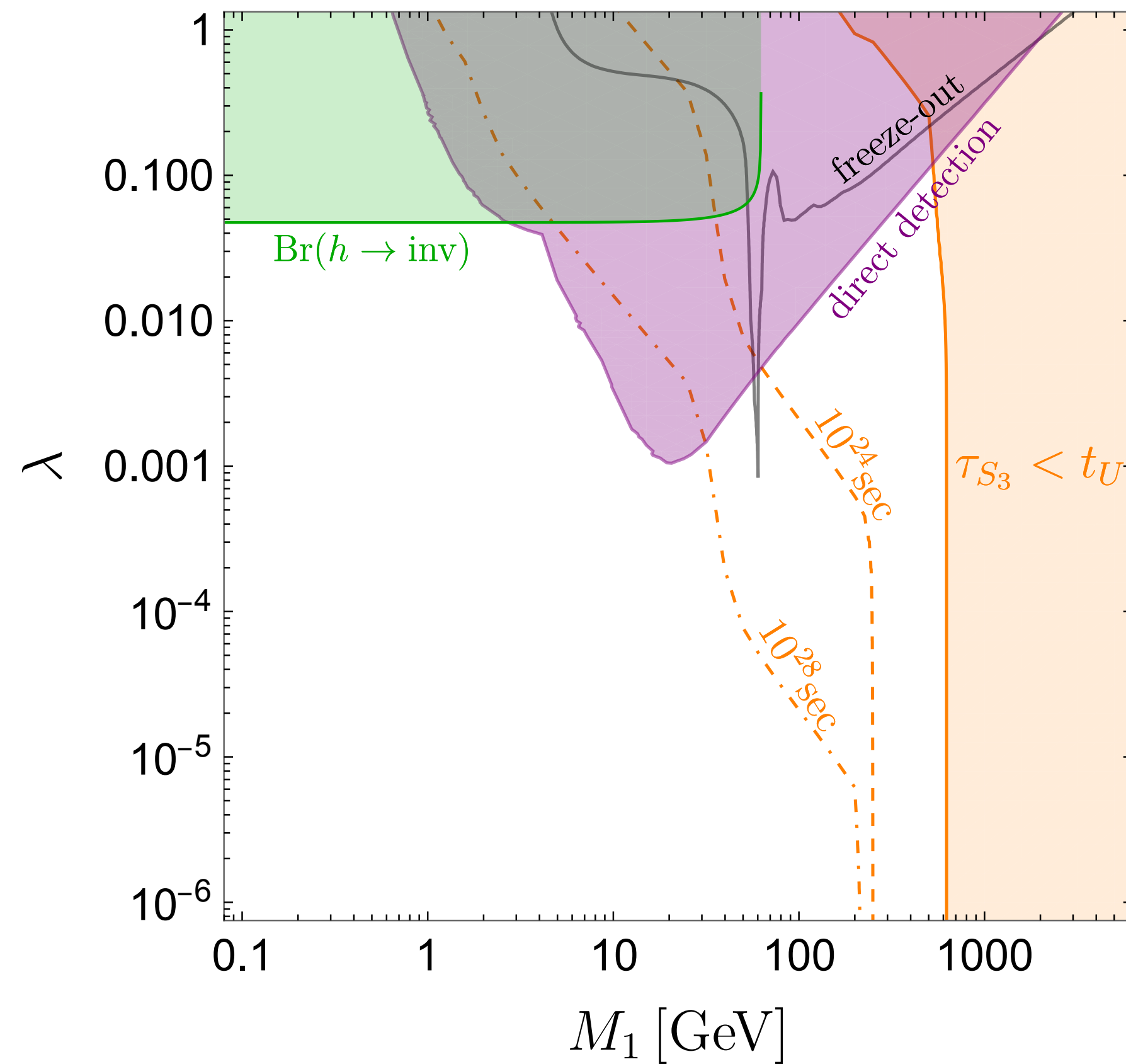
$$\lambda = 10^{-11}$$

- Heavy components are also DM if $\tau_{S_i} > \tau_U$
- White region is allowed
 - two-component between orange and blue regions
 - three-component above the orange region

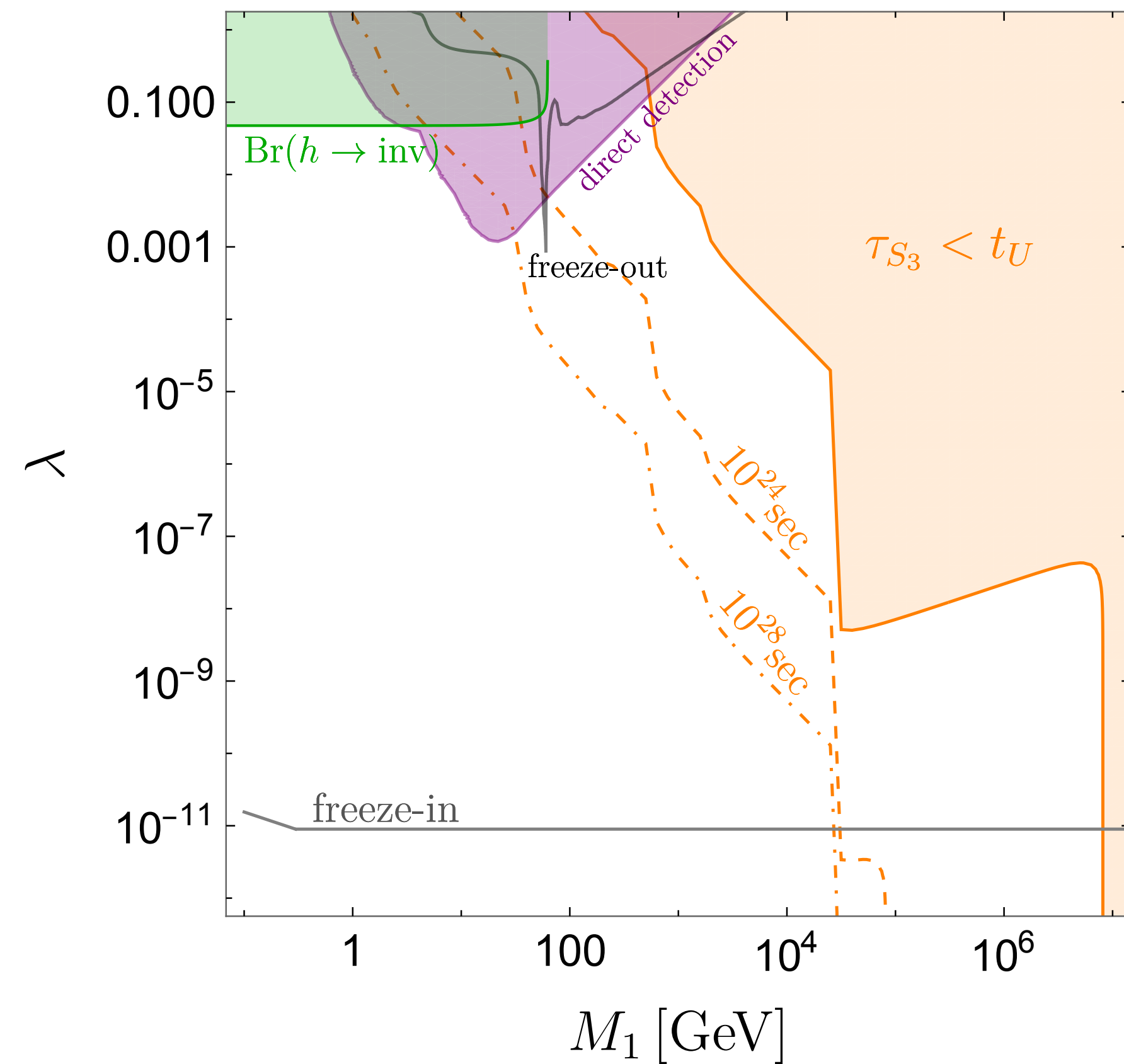
Impact of Higgs portal coupling (2/2)

$$\epsilon = 10^{-2} \simeq \frac{M_3 - M_1}{y_t^2 M_1}$$

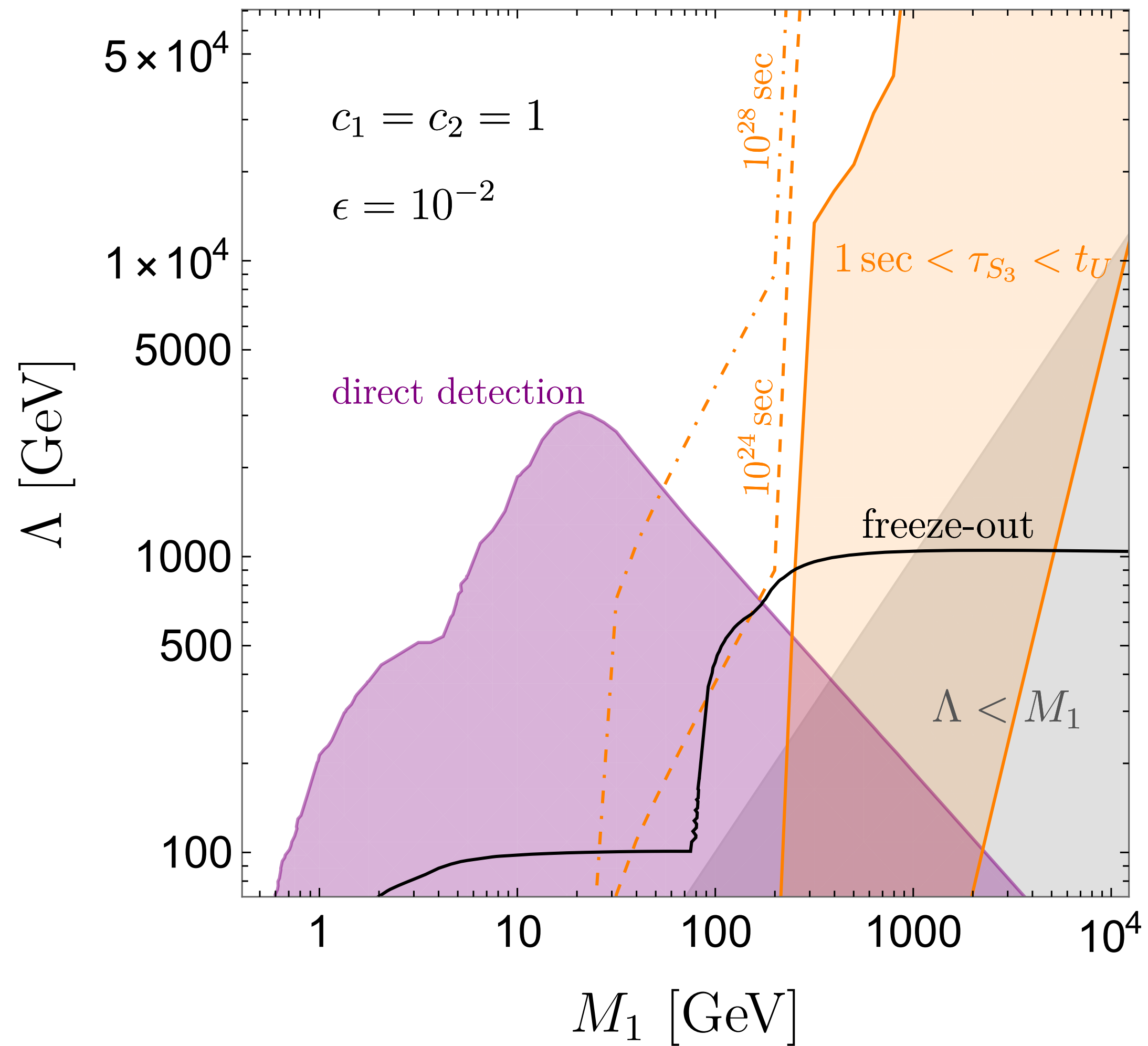
$\Lambda = 10^4 \text{ GeV}$



$\Lambda = 10^{13} \text{ GeV}$



Closer look at the WIMP region



$$\epsilon = 10^{-2} \simeq \frac{M_3 - M_1}{y_t^2 M_1} \simeq \frac{M_2 - M_1}{y_c^2 M_1}$$

$\lambda = 0$ (no coupling to Higgs)

- Only a limited mass range $M_1 \sim 180\text{-}210\text{ GeV}$ is allowed in the freeze-out scenario
- EFT is not justified in the region, $\Lambda < M_1$